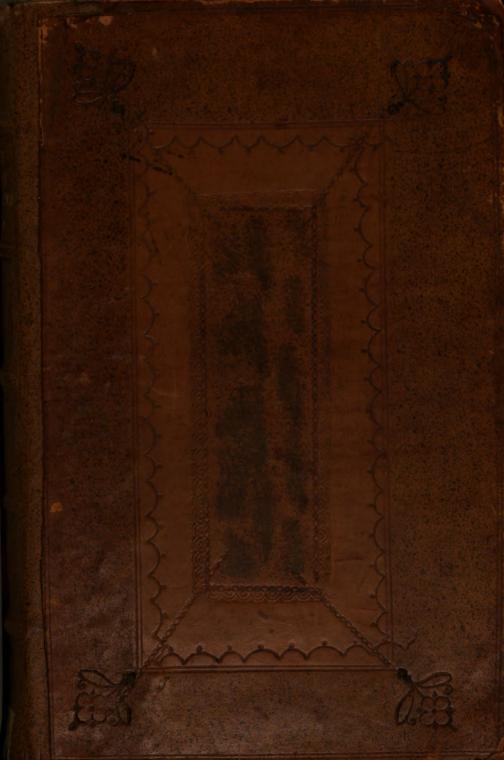
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Charles duBois.

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Mathematical Philosophy.

# LECTURE



FTER having difpatch'd the Matters of pure Aftronomy, we proceed unto the other Part of our Work, the Philosophy of the Famous Sir Ifaac Newton. For we are purpos'd to trace the Steps of that

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Great Man, and to fet forth his principal and most noble Philosophical Inventions in a more eafy Method; that fo we may bring that (as I may fay) Divine Philolophy within the Reach and · Comprehension of those, who are but indifferently perhaps exercis'd in the Mathematicks, and communicate the Knowledge thereof as far as may be. But fince that it is neceffary for any one that would undertake this Philosophy, that befides some Knowledge of Geometry, Arithmetic, and Altronomy, he should also be furnish'd with the Knowledge of the true Laws of Motions; and especially should understand something of the Nature and Properties of those Curve Lines, which are called the Conic Sections; the Nature of our Purpose, which, as was faid, was to direct our felves chiefly to Mathematicians of the lower B

lower Form, and who only understand the first Elements of Geometry and Aftronomy; therefore 'tis requir'd of us, that we should in the Beginning touch upon, and in fome measure explicate, as well the Conic Sections, as the of late demonstrated Laws of Motion; that no one through his being ignorant of these Things, may lofe his Labour in his Study of that Philosophy which we have now in hand. For indeed, as to the first Laws of Motions and Collisions, Des Cartes was fo milerably miftaken about them, when he went about to establish them, and hath so boldly impos'd upon the World falle Rules concerning Collision and Reflection of Bodies, that it is worth the while to endeavour to root out of the Minds of Men the Prejudices which have forung from thence.

We shall therefore begin with the Conic Sections: and before we go about any thing elfe, give fome Knowledge and Understanding of those Lines which are interwoven with all the Philofophy of the Famous Newton, who fhews that all the Paths, whether of the Planets or Comets of our System, are according to some or other of the faid Sections. But we shall not spend so much Time about this Matter, as to deliver the Conic Elements in any other than a fummary Way, or otherwife than by bringing into View out of the Writers of Conics, and effectially the Famous De La Hire, the Natures, and chief Properties, and Affections of these Curve Lines without their Demonstrations, assuming them for demonstrated. And forafinuch as although the faid Lines may be fet forth by mere Delineations in a Plane, as will be done afterwards; yet the Geometricians, as well the Moderns as those of old, have for the most part expounded them by the Sections of a Cone;

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Cone; and becaule also these Curves can in no other. Way be shew'd all together, and at once, and confequently, the mutual Relation and Cognation which is betwixt them, cannot in any other way of explicating them be so clearly made known. For these Reasons, I shall, in the first place, open the Natures of these Lines, and set them forth by the Sections of a Cone, proposing to explicate them by Delineations on a Plane afterwards

If you take any immoveable Point without a Plane in which a Circle is describ'd, and imagine a right Line drawn through that Point, and produc'd infinitely both ways, to be mov'd about the Circumference of the Circle, the Superficies which will arife from this Motion, is called a Conical Surface; and the Surfaces on both Sides the immoveable Point taken conjunctly, are termed Surfaces vertically opposite : The immoveable Point, common to both, is nam'd the Vertex: the Circle, the Bafe; and the Solid comprehended under the Conical Surface and the Bale, and which may be suppos'd to be infinitely produc'd, is called a Cone; to which Solid, that generated beyond the Vertex is both equal and like. The Right Line, which is drawn from the Vertex to the Center of the Circle which is the Bafe, is the Axis of the Cone: Which Right Line, if it be perpendicular to the Plane of the Bafe, the Cone is called a Right Cone; but if not, an Oblique or Scalene one. Further, a Plane howfoever pofited, fo that it paffeth not through the Vertex it felf, doth cut the Conic Superficies, and is called, A Secant or Cutting Plane; and another Plane which doth pass through the Vertex, and is every where Parallel to the Secant, goes by the Name of the Vertical Plane; and that Curve Line B 2 which

which the Conic Superficies describes in the Cutting Plane, is called a Conic Section; which Section varies according to the different Inclinations of the Cutting Plane to the Cone.

Hence will arife three Cafes: (1.) When the -Vertical Plane toucheth the Conical Surface or Surfaces, and then the Section in the Cutting Plane is called a Parabola. (2.) When the Vertical Plane neither touches nor cuts either of the Surfaces, then the Section is called an Ellipsi. (2.) When the Vertical Plane cuts one of the Surfaces, and confequently the other, then the Secant Plane also cuts both Surfaces, (fince it is Parallel to the Vertical one) and the Sections are called Hyperbole, Opposite, or opposite Sections. If therefore the Secant, and the Vertical Plane, be fo mov'd round in a Parallelism each to other, that the Vertical Plane doth fometimes cut the Bale, fometimes touch the Conic Superficies. fometimes is placed wholly without the Cone : It is manifest, that by those Conical Superficies, divers Species of Hyberbola's, divers Parabola's, and divers Species of Ellipsi, will be delineated in the Secant Plane. ' And moreover we plainly fee. what a near Affinity there is betwixt all these Lines. For if the Section be parallel to the Bafe, or even in a Scalene Cone, if it be fubcontrarily posited, it will be a Circle; which therefore is defervedly reckon'd amongst the Conic Sections, as being one of the Extremes of the Ellipsis; from which then, if you proceed by a gradual Change of the Inclination of the Cutting Plane, there will be produc'd infinite Species of Ellips; until at length the Inclination becoming Parallel to the Side of the Cone, the other Extreme of the Ellipsis passeth into a Parabola. But then the Inclination of the Cutting Plane being never fo little

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little changed further, there will arife an Hyperbola; of which there are infinite Species, according to the divers Inclination of the Vertical Plane within the Cone. So that the Ellipses do on this Side end in a Circle, and on that in a Parabola; the Parabola on this Side in an Ellipsi, and on that in an Hyperbola; the Hyperbolæ on one part in a Parabola, and on the other in a ftrait Line. But because the Conical Delineation of the Regular Curves may feem too difficult to many, I shall purfue it no farther, but proceed to that Exposition of these Lines, which is us'd by Cartes and others, and is perform'd by an eafy Delineation of them upon a Plane.

For a right Conception therefore of the Production and Nature of an Ellipsi, let ( Plate I. Fig. 1.) H and I be two Points, or two Nails or little Pegs, about which let there be put a Thread BHI; and then putting your Finger, or a Pin, to the Thread, and keeping the fame always in an equal Tenfion, move your Finger round from the Point B, until you return to the fame Point B again. By this Revolution of the Point B, is defcrib'd the Curve Line, called the Ellipsi, which differs from the Delineation of a Circle only in this, that a Circle hath only one Center, but the Ellipsis hath, as it were, two Centers; which indeed, if the faid Points H and I, their Diftance vanishing away, should come together into one, the Elliptic Curve would become perfectly Circular. But by how much the greater the Diffance is betwixt those Points, the fame Length of the Thread still remaining; so much the farther is this Figure remov'd from the Circular. So that according to the divers Proportion of the Diftance HI to the Thread BHI, or to the Line DK, which is equal to the fame Thread made lefs by the Diltance B 2

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Distance HI, divers Species of Ellipses will be described. But then, if the Length of the Thread shall be increas'd or diminish'd, in the same Proportion as the Diftance of the Points H and I fhall be increas'd or diminish'd, there will be defcrib'd indeed divers Ellipses, but which are all of the fame Species, or like to one another. From whence it appears, that Ellipses are not only innumerable in Magnitude but in Species alfo, and reach from a Circle to a Right Line : For like as when the Points H and I meet together, the Ellipfis becomes a Circle; fo when they are remov'd from each other half the Length of the Thread, it becomes a Right Line, both Sides meeting together. From whence also it is manifest, that every Species of Ellipfes is no lefs different from any other, than the Extremes of them are different on this Side from a Circle, and on that from a Right Line. It also appears from this Delineation. that if from a Point taken at Pleasure in the Elliptick Periphery, as the Point B, you draw two Lines to the two Central Points; thefe two Lines BH and BI taken together, will be equal to the greatest Diameter DK; and confequently that the Sum of them is always given: Which thing the Construction it felf shews. For that Part of the Thread, which is extended from I to B, and from thence back to H, is the fame with that which returneth from I to F, and from thence back to H; and again, that Part of the Thread which reaches from D to H, is the fame with that which reacheth from K to I, or DH is equal to IK; therefore IB + BH, which by the former is equal to ID + DH, is equal to ID + IK, that is, to KD.

And thus much for the Production of the Figure in a Plane; we shall now subjoin the Names of

of the chief Lines in it, and the most notable Properties thereof; fo as to give fome Sort of Knowledge at least of this most Noble Curve, for the more right understanding true Astronomy. and the Courfes of the Planets.

In Fig. 2. Plate 1. DFKR is an Ellipsi; C the Center; the Points H and I, the Foci thereof; DK the greater Axis, or the Transvers Axis, or the principal Diameter, or Latus transversum, the Transvers Side; FR is the leffer Axis: All the Right Lines paffing through the Center C are Diameters: All Right Lines terminated at the Periphery, and which are divided into two equal Parts by any Diameter whatever, are called Ordinates, or Lines orderly applied, to wit, with respect to that Diameter. Thus MG paffing through the Center, is a Diameter; and PK which is divided into two equal Parts by the fame, is an Ordinate thereof, or a Line orderly applied there-That Part of every Diameter, which is into. tercepted betwixt the Vertex thereof and the Ordinate, as Mu, is call'd the Absciffa or Abscifs thereof, (as being cut off from the fame Diameter:) A Line drawn from the Vertex of the Diameter, parallel to the Ordinates thereof, as no. is a Tangent to the Ellipsi in that Vertex. A Diameter parallel to the Ordinates of another Diameter, and which confequently hath its Ordinates parallel to the former Diameter, is term'd a Conjugate Diameter. Thus GM and VT are conjugate each to other, and the Ordinate PK is parallel to the Diameter VT, and the Ordinate KE to the Diameter GM. The Ordinate to the greater Axis, which paffeth through either of the Foci, as MA in the first Fig. is term'd the principal Latus rectum, or the Parameter of the greater Axis. 1. 1

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Now the most notable Properties of this Ellipfis are these: (1.) The Ordinates of every Diameter, which by the foregoing Definitions are bifected by the Diameter, are parallel each to other.

(2.) The Ordinates of the Axes are perpendicular to the Axes themfelves: But the Ordinates of the reft of the Diameters are oblique to their Diameters; and in Ellipses of divers Species, fo much the more oblique, at equal Diffances from the Axis, by how much the Proportion of the greater Axis to the leffer is the greater; but in the fame Ellipsi, fo much the more oblique, by how much the more remote the Diameters are from the Axes.

(2.) There be only two Conjugate Diameters, which are equal each to other; those, to wit, whofeVertices are at equal Diftances from theVertices of the Axes. Thus the Diameter VT is conjugate and equal to that other GM, where, to wit, VF is equal to MF, and VD equal to MK.

(4.) The obtufe Angle VCM of these two Diameters, which are conjugate and equal, is greater, and the acute Angle VCG is lefs than every other Angle contain'd by the reft of the Diameters that are conjugate to each other.

(5.) If the Lines  $\mu$  P and  $\nu$  B be Semi-ordinates to any Diameter, as MG; the Square of the Semi-ordinate  $\mu$  P is to the Square of the Semi-ordinate "B, as is the Rectangle  $M\mu \times \mu G$  to the Rectangle  $M_{\nu} \times \nu G$ ; that is,  $\mu P q$  is to the Restangle comprehended under the two Parts, into which the Diameter is divided by the Ordinate KP, as By is to the Rectangle under the Parts of the Diameter made by the Ordinate AB.

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(6.) The

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(6.) The Parameter, or Latus rectum of any Diameter is a third Proportional to that Diameter, and its Conjugate. That is (in Figure 1.) if the Diameter DK, is to its conjugate Diameter EF, as EF to Y, then Y is the Parameter or DK:EF: Latus rectum of the Diameter D K. Whence :: IF: AM A M, an Ordinate to the Axis thro' the Focus, is as above, equal to the principal Parameter, and is a third Proportional after the greater and leffer Axis. For the Axes are the principal Pair of conjugate Diameters.

(7.) The Square of every Semi-ordinate, (as y manne M I in the first Figure) is less than, the Rectangle 2017/10 made of any Abscifs whatever, (as I K) drawn into the Latus restum of its own Diameter, (or than 1997) I K X Y. And in the other Figure, " $P \mu q$  is less than the Rectangle made; of the Abscifs M  $\mu$ , and the Latus restum of M G. From which Defect, or ended this Section hath its Name.

(8.) If from any Point, as B in the first Figure, you draw the right Lines B H and B I to the Foci, the Sum of them will be equal to the greater Axis, as was shew'd above. And if the Angle I B H comprehended by those Lines be bisected by the right Line ba; the Line a is perpendicular to the Tangent V B in the Point B, that is, to the Curve in the Point of Contact.

(9.) The Curvature, with respect to the Center of the Ellips, is at divers Distances from that Center in the Quadruplicate Proportion of those Distances directly : As, if C K be double of C F, the Curvature in the greater distance K, shall be to the Curvature in the lesser distance F, as 16 is to 14 and if C K be Treble, C F, the Curvatute in K, will be to that in F as 81 to 1. And so of the rest.

10. The Curvature of the Elliptic Arches, with respect to the Focus, is in divers Distances from that Focus, in the simple Proportion of the Distance directly. Thus, if HD be half of HK, the Curvature at D, if you respect the Focus; XH will be half of that at K, respecting the same Focus, and so of the rest. And the Thing is the fame in a Parabola and Hyperbola.

(11.) The Diftance of a Body turn'd round in an Ellipsi, about the Focus H, from the fame Focus, is the greatest of all in the Point K, least of all in the Point D, and mean in the Points E and F; and that mean distance HF is equal to the greater Half-Axis D C or CK; as is manifest from the Production of the Ellipsi.

(12.) The vanishing Subtense of the Angle of Contact, parallel to the Distance from the Focus, at an equal perpendicular Interval from that distance, always remains given and unvaried in the fame Ellips, yea and in the fame Parabola and Hyperbola too. Thus if d Z be always given, g d also will always remain given in a distance infinitely small.

(12.) The Area of the Ellipsis is to the Area of the Circle circumscrib'd, as the lesser Axis is to the greater; and so are all correspondent Parts whatever amongst themselves, as MIK, mIK; and the Ordinates to the greater Axis, as MI are divided by the Elliptic Periphery always in the fame Proportion; so that MI is to mI always in the fame Proportion; to wit that of the lesser Axis to the greater. And we are to reason in the fame manner concerning a Circle inscrib'd in the Ellips.

(14.) All Parallelograms describ'd about the conjugate Diameters of the Ellips, and comprehending the Ellips, are equal. Thus the Parallelogram

logram  $\alpha\beta\gamma\delta$  is equal to that other  $\beta\gamma\delta$ ; and thus it is every where.

(15.) If a right Line always paffing through one of the Foci be fo mov'd, that the Elliptic Area defcrib'd by the fame is proportional to the time; the Angular Motion of a right Light drawn from the other Focus to the former Line, will be almost equable. Thus in the former Figure, if the Angular Motion of the Line H B be fo attempered, that the fame being according to the reciprocal Proportion of the Distance accelerated or retarded, doth defcribe the Area DHB proportional to the time, the Angular Motion K I B, about the other Focus I will be almost proportional to the time, and confequently without any notable Acceleration or Retardation, and nearly equable; that is to fay, where the Ellipses doth not differ much from a Circle.

Feb. 7. 1703.

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#### LECT. II.

O país now from the Ellipsi to the Parabola, let DI be an Infinite right Line, and IL another perpendicular to it. Then there being taken in the Line DI any Point F, let the Line FI be bifected in the Point T. And let there be taken Two Threads joined together in the Point T, one TI, the other TF. And let a Pin fixed to the Threads in the Point T be moved to the Right and Left, in such a manner, that when the Pin is in any other Position as in P, the Thread TI

TI which here becomes PL be always perpend dicular to IL, or, which is the fame Thing parallel to DI, but equal to the Thread TF, which in this Cafe becomes PF, ever passing thro' the Point F. And the Curve thus generated by the Pin, infinitely produced both ways, is called a Parabola. In which g P i T s R o is the Periphery ; ID the Axis or principal Diameter : F the Focus. The Point T the principal Vertex ; an Ordinate to the Axis through the Focus is equal to the principal Latus Rectum. All right Lines n i, or R Z Parallel to the Axis, are Diameters, as dividing the Lines i h and K T which are Parallel to the Tangents at their Verttices into Two equal Parts; and they are called Diameters belonging to the Vertices in which they terminate, as T, i.

Now the principal Properties of a Parabola are these.

(1.) Every Diameter or right Line parallel to the Axis, bifects all the Lines within the Figure which are parallel to the Tangent of the vertical Point. Which bifected Lines are as hath been faid called Ordinates.

(2.) The Ordinates of the Axis are perpendicular thereto: But the Ordinates of the reft of the Diameters are oblique to their Diameters; and fo much the more oblique, by how much the Vertex of the Diameter is further remov'd from the principal Vertex of the Parabola.

(3.) The Latus rectum, or Parameter to every Diameter, is a third Geometrical Proportional after any ableiffe and its femi-ordinate; that is the Latus rectum of the Diameter (in), (or that of the Vertex (i)) is y; if it be thus; as the Ableifs (iq) is to the Semi-ordinate (q k) fo is that Semiordinate (q k) to y,

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(4.) The

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(4.) The principal Latus rectum, or that belonging to the Axis, is equal to the Ordinate (h i) palfing through the Focus; and fourfold of F T, the leaft diffance of the Focus from the principal Vertex.

(5.) The Latus rectum belonging to any Vertex or Diameter, is also fourfold of the diffance of that Vertex from the Focus. Thus the Latus retum of the Vertex s is fourfold F s, and so it is every where.

(6.) The diffance of any Vertex or Point in the Parabola whatfoever from the Focus, is equal to the leaft diffance of the fame from the Line L L, which is perpendicular to the Axis, and is diffant from the principal Vertex by a Quarter of the principal Latus rectum. For by the Conftruction, the Line F P is equal to P L.

(7.) The Square of every Semi-ordinate, as (q k) is equal to a Rectangle made of the Latus rectum, of the fame Vertex as Y, and (i q) the Absciffe of the Diameter of the Vertex. And from the Equality of the magacoan, or Comparison in the Figure, berwixt the Rectangle and the Square of the Semi-ordinate, without any Excess for Defect, the Name of the Section is derived.

(8.) When therefore, the Latus rectum in any Diameter is given, the Absciffes are as the Squares, or in the duplicate Proportion of the Semi-ordinate. Thus TF is to TG as iFq is to gGq; and so likewise is iq to ir, as the Square of qT is to the Square of r1; and thus every where. From whence also, when the Abscifs of the Axis is equal to the principal Latus rectum, or fourfold of the diffance from the Vertex, it will be equal to its femi-ordinate.

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(9.) The Angle comprehended by any Tangent whatever, and a Line from the Focus, is equal to an Angle comprehended by the fame Tangent. and any Diameter or the Axis. Thus the Angles IiF and p in are equal. From whence indeed, (which thing is to be noted by the way) all the Rays of Light which fall upon the Concave part of the Surface produced by the Convolution of the Parabola about the Axis, which fall, I fay, upon the fame Parallel to the Axis, will be reflected from a Concave Paraboloid Figure to the Focus F, and will beget there a most vehement burning; from which Property indeed the Point F hath the Name of Focus; and hath communicated the fame Name to the like Points in an Hyperbola and Ellipfis.

(10.) A Parabola, like as an Hyperbola, doth not enclose a Space, but is stretched forth in infinitum.

(11.) A Parabolic Curve always tends more and more in infinitum to a Parallelism with its Diameters, but can never reach thereto.

(12.) If two Parabolz's be defcribed with the fame Axis and Vertex, the Ordinates to the common Axis will be cut off by the Parabolz in a given Proportion; and the Area's comprehended by the fame Axis and Ordinate, and the refpective Curves will be in the fame given Proportion to one another.

(13.) Every Parabolic Space, comprehended betwixt the Curve and the Ordinate, is to the Parallellogram made of the fame Bafe and Altitude in a Subfefquialteral Proprotion, that is; as 2 is to 3, and to the external Space in a double Proportion, or as 2 is to 1. So qiT is to qiI as 2 is to 3, and to i I T as 2 is to 1. From whence it becomes eafy to fquare the Parabola.

(14.) The

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(14.) The diffance betwixt the Vertex of the Axis, and the Point where any Tangent whatever interfects it, as I, is equal to the Abfcils of the Axis which belongs to the Ordinate apply'd from the Point of Contact. So T I is equal to T F; and thus it is every where.

(15.) All Parabolz's are like, or of the fame Species ; as are also all Circles.

(16.) If a Diameter be continued through the Point of meeting of Two Tangents, this Diameter will bifect the Line that joins the Contacts. Which Property of the Parabola is likewife to be applyed to the Ellipfis and Hyperbola.

And thus much for the Parabola. We now come to the Hyperbola (Fig. 4. Plate I.) Take a Staff or Rule of a fufficient Length as IB, let I and H be two Central Points answering to the Foei of an Ellips, in which let Nails be fastned; then there being tied to one end of the Stick a Rope or Thread, twofold shorter than the Stick. let the other end thereof be bor'd through, and fo fixed upon the Nail I; but as for the other end of the Rope let it be fixed by a Knot upon the other Nail H; which done, laying your Finger upon the Point B, where the Rope and Staff are tyed together, let your Finger descend to long that you have thereby now applyed and joyn'd the whole Rope to the Staff or Rule, the Staff having been in the mean while, as it needs must, wheel'd about the Gentre I. And thus you have defcrib'd by the Point B, the Vertex of the Angle H B L & Curve Line, X B D which is part of an Hyperbola; the whole confifting of that Curve wich will refult from the Curve X BD, which hath added to it the Curve Y D, the Product of the Rule and Work as turn'd to the other Side. 1.4

Side. Furthermore, if you transfer the Hole or Knot of your Rope to the Nail I, and fasten the end of the Staff upon the Nail H, you will defcribe another Hyperbola vertically opposite to the former, which is altogether like and equal thereto. But then, if without changing any thing in the Rule and Nails, you fhall only apply a longer Rope, you will defcribe an Hyperbola of a different Species from the former ; and if you shall still lengthen the Rope somewhat, you will have another Sort of Hyperbola ; until at length making the Rope double in length of the Rule, vou will have the Hyperbola chang'd into a right But if you alter the Diftance of the Nails Line. in the very fame Proportion, in which you change the difference betwixt the Length of the Rope and that of the Stick; in this Cafe you will have Hyperbolæ mark'd out, which are altogether of the fame Species, but have their fimilar Parts differing in Magnitude. And laftly, if the Length of the Rope and Rule be equally increas'di their Difference in the mean while, and the Interval of the Nails remaining the fame; not a different Hyperbola either as to Species or Magnitude will be defcrib'd, nor any other than a greater Part of the fame Hyperbola. And this for the Mechanical Construction of an Hyperbola in a Plane.

But it is to be acknowledg'd, that many Properties of an Hyperbola are better known from another manner of generating the Figure; which Way is this: (See Fig. 5. Plate 1.) Let LL and MM be infinite Right Lines interfeding each other in any Angle whatever in the Point C: From any Point whatever, as D or e, let Dc, Dd, be drawn parallel to the first Lines, or (ec, ed;) which with the Lines first drawn make the Parel-

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Parallelograms as DcCd, or ecCd : Now conceive two Sides of the Parallelogram as Dc, Dd, or ec, ed, to be fo mov'd this way and that way, that they always keep the fame Parallelism, and that at the same time the Area's always remain equal : That is to fay, that Dc and e c remain always Parallel to MM, and Dd or e d always Parallel to L1; and that the Area of every Parallelogram be equal to every other. one Side being increas'd in the fame Proportion wherein the other is diminish'd. By this means the Point D or e will describe a Curve-Line within the Angle comprehended by the first Lines: which is altogether the fame as was defcrib'd above, both by the Section of a Cone, and Cartes's Delineation. And in like manner, in the Angle vertically opposite will be describ'd a like and equal Hyperbola, if fo be the Parallelogram CcKd, equal to the former, be supposed to be mov'd in the same manner as before : Which Hvperbola's are, as was faid before, called oppofite Sections, or opposite Hyperbolz. Now in either of the two Figures, DK is the Transvers Axis, or Transvers Diameter of the Hyperbola. or the Opposite Sections: The Point C is the Center : The Points H and I the Foci. And in the 2d Figure, all the Lines paffing through the Center C, as i h are Diameters. But if Hyperbolæ be describ'd in the following Angles, as LCM, MCL, those Sections will be called the Following Sections; and if the Diftance of the primary Vertex of those Hyperbolz from the common Center C, as C B, or C 7, be equal to the Semi-tangent K, or K, at the primary Vertex of these, those Sections shall be called Conjugate Sections : And all the Figures together will be to be named the Hyperbolic System. Fur-

Furthermore, (i, h) the Ordinate to the Axis through the Focus, is equal to the principal Latus rectum, or the Parameter of the Axis; and an indeterminate Diameter, whether of the following Sections, or of the former, which is parallel to the Ordinates of any determinate Diameter, is called the Conjugate Diameter of the fame: and hath its Ordinates parallel to the former Diameter.

And now we come to the principal Properties of the Hyperbola, and the opposite Sections, which are as follows :

(1.) Any Diameter or right Line paffing thro' the Center, bifects all its Ordinates; that is, all the Right Lines which are terminated on both Sides by the Hyperbolic Periphery, and those parallelLines that are bifected by any Diameter whatever, are called the Ordinates of that Diameter.

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(2.) The Ordinates of the Axis are perpendicular to the fame: But the Ordinates of the reft of the Diameters are oblique to their Diameters; and fo much the more in divers Species, at equal Diftances from the Axis, by how much the Difference of the Angles including the Hyperbolæ is the greater; and in the fame Hyperbola, fo much the more oblique, by how much the Diameters are remov'd from the Axis.

(2.) If any Lines, as H h and Q s, be Semiordinates to any Diameter whatever, as K D; the Square of the Semi-ordinate H h is to the Square of the Semi-ordinate QS, as the Rectangle K H D H is to the Rectangle K Q, D Q • And fo the Square (b n) is to the Square (a K,) as the Rectangle (i b h b) is to the Rectangle (i a h a ); and thus every where.

(4.) The Latus rettum, or Parameter of every Diameter, is a third Geometrical Proportional after

after the Diameter, and the Conjugate thereof, (or its Tangent, which is equal to it:) That is, the Latus rectum of any Diameter, as DK is Y, if it be thus; as the Diameter DK is to its Conjugate  $\beta r$ , or its equall  $(\omega r)$ ; fo that Conjugate  $\beta \gamma$ , or that Tangent  $(\omega r)$  is to y. And as the Ordinate to the Axis through the Focus is the principal Latus restum, fo it is more than Quadruple of the leaft Distance of the Focus from the Vertex.

(5.) The Square of any Semi-ordinate whatever, as (Qr,) is greater than a Rectangle made of the Abfcils DQ, drawn into the Latus reflume of its own Diameter, as y: And in like manner, the Square of the Semi-ordinate (b n) is greater than the Rectangle of the Abfcils (i b) into the Latus reflume of the Diameter (h i.) From which waspCold, or Excels, this Section hath its Name.

(6.) If from any Point of the Hyperbola, as (B) in the former Figure, there be drawn Right Lines to both the Foci, as BH, BI, the Difference of these Lines will be equal to the Axis DK; as will eatily appear from the Delineation it felf.

(7.) If the Angle H B I, comprehended by Lines drawn to the Foci, be bifected by the Right Line E B, that Right Line will be a Tangent to the Hyperbola in the Point B.

(8.) The Right Lines L L, and M M, which enclose the Hyperbolæ, are Afymptots of the Hyperbolæ, *i. e.* they are fuch unto which on both Sides the Curve approacheth nearer and nearer, but is never able to touch or coincide with the fame.

(9.) The Species of Hyperbolæ are various, according to the different Magnitude of the Angle LCM comprehended by the Afymptots: But Ga that

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that Angle remaining the fame, the Species of the Hyperbola remains unchang'd; but according to the different Magnitudes of the Parallelograms, by which the Hyperbolæ are describ'd, Hyperbolæ of divers Magnitudes do arife: But if the Angle contain'd by the Afymptots be a right Angle, the Hyperbola is called Equilateral or Rectangular, and the Latus rectum of all the Diameters will (as it is in a Circle) be equal to their And laftly, if Hyperbolz be de-Diameters. fcrib'd about the fame Axis, in divers Angles of the Afymptots, the Right Lines perpendicular to the Axis will be cut off in a given Proportion by them all; and the Spaces likewife enclosed by the Right Lines or Ordinates, the produced Axis, and the Curves, will be in the fame given Proportion.

(10.) If the Diffances from the Center of the Hyperbola be taken in a Geometrical Proportion in one of the Afymptots, fo that CI, CII, CIII, CIV, CV, CVI, be continuedly proportional geometrically; and if from those Points there be drawn parallel to the other Afymptot, the Lines, II, II 2, III 3, IV 4, V 5, VI 6; the Spaces I 2, II 3, III 4, IV 5, V 6, will be equal amongst themselves. And confequently, if that Afymptot C M be fuppos'd to be divided, according to the Proportion of Numbers exceeding one another in a natural Series, those Spaces will be proportional to the Logarithms of all those Numbers.

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#### LECT. III.

AVING now expounded feverally the Curve Lines, called the Conic H Sections; let us now compare them to gether, and briefly confider, what Affinity there is betwixt them, what mutual Refpect they bear to one another, and what Difference there is among them.

Let the Point A therefore (fee Fig. 5. Plate 1.) be the Center of the Circle FXBY, and the common Focus of all the Sections: [ And it is indeed a certain Center, as it were, of all the Sections: And the Ordinate to the Axis through the Focus, or Latus rectum, doth in most of them more agree with the Diameter of the Circle paffing through the Center thereof, than the Axis it felf of the Section doth : ] Then let the Point F be the principal Vertex of all the Sections; and F X B Y a Circle, the Center of which Figure, as being only an Extream Ellipsi, falls in with the Foci, (where XY will be, if I may fo fpeak, the Latus rectum of the Circle passing through the common Focus or Center, and equal to the reft of the Diameters.) Let FGHI be an Ellipsi lefs Curve on the Vertex than the Circle is; the remoter Focus of which Ellipsis is the Point C; FH the principal Diameter or greater Axis; GI the lesser Axis; ef the principal Latus rectum, which is more than double to A F, the Diftance of the Vertex F from that Focus A, but lefs than Quadruple thereof. But it is to be noted, that C 2 another

another Ellipsi also may be drawn more Curve in F than the Circle; but then it is describ'd about the Point A, as the remoter of the Foci. But then after the greater Ellipsi, the Center thereof departing further in infinitum, there arifeth the Conic Section LdFcK, which we call a Parabola: which indeed is half of an Ellipsi infinitely long; the Axis whereof is the Infinite FH, and (cd) the Latus rectum : Which fame is Quadruple of the Diftance of the Vertex from the Focus AF. As for the Curvature of the Parabola in the Vertex F, it is less than that of the Ellipsi, as is easy to be seen. Then lastly, the Hyperbola MiFIN follows, whole Parameter, or principal Latus rectum (i1) is more than Quadruple to A F, the Diftance from the Focus : and the Curvature thereof in the Vertex F, is lefs than that of the Parabola, and will infinitely be diminish'd, the Angle TEV, contain'd by the Afymptots, being increas'd in infinitum, until at length the Afymptots falling into one Right Line, the Hyperbola it felf with its Afymptots, end in the Right Line OP perpendicular to the Axis. From whence it is to be noted, (1.) That the Conic Sections are in themselves a System of Regular Curves allied to each other; and that one is chang'd into another perpetually, when it is either increas'd or diminish'd in infinitum. Thus the Circle, the Curvature thereof being never fo little increas'd or diminish'd, passeth into an Ellips; and the Ellipsi, its Center going away infinitely, and the Curvature being by that means diminish'd, is turn'd into a Parabola : And when the Curvature of the Parabola is never fo little chang'd, there arifeth the first of the Hyperbolz; the Species whereof, which are innumerable, will all of them arife orderly by a gradual Diminution

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nution of the Curvature, until the Curvature vanishing away, the last Hyperbola ends in a Right Line perpendicular to the Axis. From whence it is manifest, that every Regular Curvature like to that of a Circle, from the Circle it felf unto a Right Line, is a Conical Curvature, and is diflinguish'd with its peculiar Name, according to the divers Degree of that Curvature. (2.) That the Latus rectum of a Circle is double to the Distance from the Vertex; that all the Latera recta of the Ellipses are in all Proportions to that Diftance betwixt the Double and Quadruple, according to their different Species : That the Latus rectum of the Parabola, is just Quadruple to that Distance; and lastly, that the Latera recta of Hyperbolæ are in all Proportions beyond the Quadruple, according to their various Kinds. (2.) That all the Diameters in a Circle and Ellipsi interfect one another in the Center of the Figure within the Section : That in the Parabola they are all parallel amongst themselves, and to the Axis; but that in the Hyperbola they interfect one another, but this without the Section, in the common Center of the opposite Sections. (4.) That the Curvature, with respect to the Focus in all these Figures, is increas'd or diminish'd proportionably to the Increase or Diminution of the Distance from the Focus. For although by reason of the Obliquity of the Tangents, the Curvature for the most part feems greater in a leffer Distance from the Focus, and less in a greater ; yet the true Curvature, which is to be defin'd by the Subtenfe of the Angle of Contact, is on the contrary greater in a greater Diftance, and leffer in a lefs, and greater or lefs in proportion to the Increase or Diminution of the Diftance ; as was above noted, and will be more fully open-C 4 eđ

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ed in the Sequel. And thus much for the Conic Sections.

And forafmuch as we shall make fome fort of Use of the Line called the Cycloid, we shall briefly describe it.

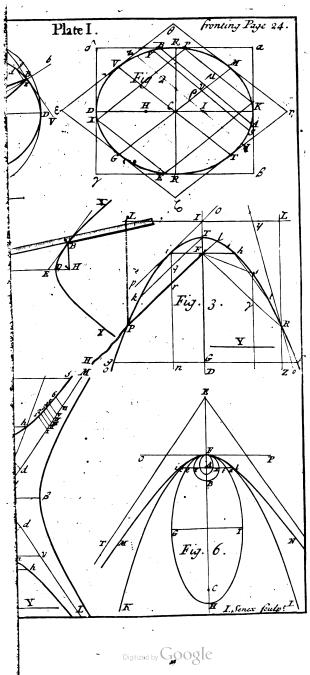
If upon the Right Line A E (fee Fig. 1. Plate 2.) the Wheel or Circle A B C D be roll'd along. until the Point A, in which it at first touched the faid Line, doth again after an entire Revolution meet and touch the fame in E; the generating Circle ABCD will describe the Line AE equal to its own Periphery, and the Point A by its Compound Mation will defcribe the Curve Line AFE, which is called a Trochoid or Cycloid : The Length of which Line is Quadruple to the Diameter of the generating Circle; and the Cycloidal Space comprehended by this Curve, and the Subtenfe A E, is triple the Area of the generating Circle. Moreover, any part whatever effimated from the Vertex, as FI is every where double the Chord of the Circle Fb, and the Tangent thereof GIH is perpetually parallel to the fame Chord F b. And thus much for the Cycloid.

Now after this preparatory fhort Explication of the Conic Sections, we come to our proper Work: Intending to proceed next to the true Laws of Motion, both those commonly known, and those which were lately found out and eftablish'd by the Famous Sir Isac Newton.

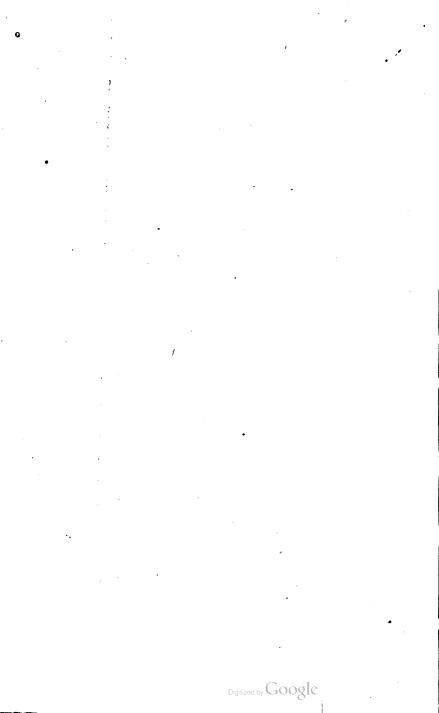
In the fetting forth of whose Noble Inventions, we shall generally make use of the very Words of that great Man; but yet so, that every where we shall endeavour to explicate, demonstrate, and to make clear and plain to all, what either Words or Things seem more obscure and difficult.

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#### DEFINITIONS.

(1.) **DODY** or Matter is an extended Subftance, Solid, or Impenetrable, of it felf merely Paffive, and indifferent to Motion or Reft; but capable of any fort of Motion whatever, and of all Figures and Forms. I call it a Substance extended, because that it possesses fome part of extended Space; but Solid and Impenetrable : not because it cannot be penetrated by Space, or perchance by other incorporeal Subflances, but because it is impenetrable by all other Matter; and upon that account it doth eminently claim the Name of Solid. I put in the Definition [its being indifferent to Motion and Reft] not that I reckon Motion, as well as Reft,a Thing plainly Negative or Privative, but because the Conception of a Body, as in Motion, is as easy and familiar as of a Body at Reft. I call it in it felf Paffive, becaufe we perceive nothing of Action or Energy, or of a Power of moving it felf, either in its Nature or Affections; but on the contrary from all the Phznomena of Motions we every where meet with its meer Inactivity. But I fay that it is capable of any fort of Motion and of all Figures and Forms; fince the daily Appearances in the World, and infinite Experiments, doe shew this to be the Nature of it. Time, Space, Place, and Motion, as being things fo well known to all, scarce need to be defined. But however, for the taking away fome Prejudices out of Mens Minds, it is very expedient, that with the famous Newton we should diftinguish these Quantities into Absolute and Relative, True and Apparent, Mathematical and Vulgar, and fo in a i i i i fore

fort describe them ; which for Order's fake shall

be done in the following Definitions. (2.) Time Abfolute, True, and Mathemati-cal, is an eternal and equable Duration, compounded of Parts, fucceeding each other in an immutable Order. For in it felf, and its own Nature it flows equably. Nor doth it depend on the Motion of Things, much lefs on their Reft, nor indeed upon their Existence. For whether any thing were mov'd or not, whether any Thing did exift, or nothing at all were in Being, it would be all one in this Cafe. Time flows equably, whatfoever relation any other Things have to one another.

(2.) Relative Time, or that which is Apparent and Vulgar, is fome fenfible and external Meafure of Duration (whether it be by Motion, or fome other way; whether it be Accurate, and even or uneven;) which is vulgarly used instead of the true time, as an Hour, Day, Month, Year, the Duration of the World, or any System from the beginning to the End, &c. In Aftronomy, Absolute Time is diftinguish'd from Relative, by the Equation of the Vulgar Time : For the Natural Days are unequal, which are nevertheless commonly taken for equal in the measuring of time. This Inequality Aftronomers correct, that they may measure the Heavenly Motions by a truer Time. It is poffible, that there may be not even Motion at all by which Time may be accurately measured. All Motions may be accelerated and retarded, but the flowing of absolute Time cannot be chang'd. The Duration or Perfeverance of the Existence of Things is the same, whether their Motions be fwift, flow, or none at all. Confequently this Duration is justly diftinguish'd from its sensible Measures, and collected from

from them by Aftronomic Equation. For this is that which Aftronomers have labour'd after; namely, that from the unequal Motions of the Heavenly Bodies, they might find an equable Motion about fome Center; from whence they may more eafily and accurately measure Duration, that flows equably.

(4.) True, Abfolute, Mathematical Space. is an Extension Penetrable, Indiscerpible, Immoveable, Infinite, Eternal, and every where like to it felf. Whether or no fuch an Extension doth really exift diffinct from Matter, is another Question. But that this is the common Notion of Space with all, must be allowed by every reasonable Man: and therefore, is to be taken as a Definition ; for so Geometricians do at first define a Circle. a Square, a Triangle, &c. not troubling themfelves with the Question, whether fuch Figures do really exift or no. We ought therefore to lay down a Description of Space should be laid down aforehand, least afterwards there should arise Strife about Words; as we may afterwards enquire whether it be the Idea of a Thing really exiftent.

(5.) Relative Space (which alfo as I fuppole is commonly called Place) is the Measure of Abfolute Space, or any moveable Dimension, which is defin'd and determin'd by our Senses, from its Pofition with respect to certain Bodies, and is commonly us'd by the Vulgar for immoveable Space. As the Dimension of an Aereal, Celessial, or Subterraneous Space, is defined by its Position in respect of the Earth. So Space, Absolute and Relative, are the same in Species and Magnitude, but do not always remain the same in Number : That is, if we consider the Space or Cavity contain'd in any Vessel, whithersoever the Vessel is mov'd,

mov'd, the Space or Cavity included within the Sides thereof will always be of the fame Nature. by reason that the nature of Space is every where fimilar to it felf; and will remain likewife of the same Magnitude, because of the given Magnitude of the containing Veffel. But it doth not remain the fame Space numerically, for that is changed perpetually by the Motion of the Veffel. In like manner, if the Earth be mov'd with an annual Motion about the Sun, the Space of our Air which relatively, and in refpect of the Earth remains still the fame, that is of the fame Nature and Quantity, will fometimes be one part of Abfolute Space, fometimes another, and fo will abfolutely and really be changed perpetually. For, indeed, as the Order of the Parts of Time is unchangeable, fo likewife is the Order of the Parts of Space; although the Things which are in them are continually mov'd and chang'd. For Times and Spaces are, as it were, the Places of themfelves, and of all other Things; which are placed in Time as to order of Succession, and in Space as to order of Situation. They are Places by their Effence, and it is abfurd to fay that the primary Places can be mov'd. These therefore are the Absolute Places; and the Translations which are from these Places, are the only Absolute Motions. But then, because the Parts of Space cannot be feen in themfelves, or diffinguish'd from each other by our Senses, instead of them therefore we use sensible Measures ; defining all Places from the Politions of Things, with refpect to fome Body which we look upon as unmov'd, and their Diftances from the fame : and eftimating all Motions with respect to the faid Places and fo far as we conceive Bodies to be tranfferr'd from them. And thus inftead of absolutePlaces

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and Motions, we make use of Relative ; and this indeed not unfitly in common Affairs : But in Philosophical Matters we ought to abstract from the Senses : for it is possible, that no Body is really quiescent, to which Places and Motions may not in this manner be referr'd to.

(6.) An absolute Place is that part of the absolute Space which the Body possesses.

(7.) A Relative Place is that part of Relative Space which a Body possesses. I fay that Place is a part of Space, not the Situation of a Body. or the ambient Surface as fome have defin'd it. For the Place of equal Solids are equal; and the fame quantity of Matter always possesses the fame Quantity of Space, of whatfoever Figure or Density it is. As for Example, The Places of a Sphere, and of a Cube of the fame abfolute Magnitude will be equal, or they will fill and be adequate to equal Places; although the ambient Surfaces, by reason of the diffimilitude of the Figures will be unequal; and fo in all other Figures. Further, the Motion of the whole is the fame with the Sum of the Motions of all the Parts, that is, the Translation of the whole from its Place is the fame with the Sum or Aggregate of the Translations of all the Parts from their Places; and confequently, the Place of the whole is the fame with the Sum of the Places of the Parts, and therefore is Internal, and in the whole Body. But Situations properly speaking have no Quantity, and cannot be faid to be greater or leffer, neither are fo much Places as Affections of Places.

(8.) Absolute Motion is a Translation of any Body or Substance from one absolute Place, or immoveable Space, into another absolute Place or immoveable Space.

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(9.) Relative Motion is a Translation of a Body from a Relative Place or fome moveable Space; into fome other Relative Place, or moveable Space; or a transferring of a Body from the Neighbourhood of fome ambient Bodies into the Neighbourhood of others; or lastly, a Translation of a Body from its Situation amongst fome certain Bodies into another Situation.

Thus in a Ship which is under Sail, the Relative Place of a Body is that part of the Ship in which it is; or that part of the whole Cavity which fuch a Body fills; and which confequently is mov'd with the Ship: And the Relative Reft of that Body, is the abiding thereof in the fame part of the Ship, or Cavity. But the true Reft thereof is its continuance in the fame part of the immoveable Space. From whence if the Earth did truly reft, the Body which relatively refts in the Ship, would be mov'd truly and abfolutely with the fame Velocity wherewith the Ship is mov'd on the Earth.

But if the Earth be also mov'd ; the true and absolute Motion of the Body will arife, partly from the Motion of the Earth in the unmov'd Space; partly from the Relative Motions, both of the the Earth and of the Body in the Ship; and from these Relative Motions will arise a Relative Motion of the Body on the Earth.So if that Part of the Earth in which the Ship is, be really mov'd towards the East with a Velocity of 10010 Parts, and the Ship be carried towards the Welt by the Wind with a Velocity of Ten Parts; and the Mariner walk in the Ship towards the East with one Part of Velocity : The Mariner will be mov'd truly and abfolutely in the unmov'd Space towards the East with 10001 Parts of Velocity. and Relatively on the Earth towards the West with Nine Parts of Velocity.

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#### LECT. IV.

E have already laid down fome Definitions by way of Preparation to the Newtonian Philofophy. We will now fuper-add a General Scholium appertaining to the two laft Definitions.

A General Scholium.] Reft and Motion Abfolute and Relative, are diffinguish'd one from another by their Properties, Causes, and Effects.

It is plain by what hath been faid, that altho' any two Bodies, each of which doth truly reft, do alfo reft betwixt themfelves; yet it doth in no wife follow from their refting betwixt themfelves, that they do truly reft: For there may be fome truly Quiefcent Body in or far beyond the Region of the Fixed Stars, with refpect to which both the faid Bodies do change their Position.

But from the Situation of Bodies in our Regions, in respect of one another, we cannot discover whether any of them keep a given Position in respect of that remote one; and so true Rest cannot be defined by their Situation between themfelves. The Property of Absolute Motion is, that those Parts which keep given Positions to the Wholes, participate of the Motions of those Wholes: For all the Parts of Revolving Bodies endeavour to recede from the Axis of Motion.

And the Impetus of moving Bodies arifes from the conjoint Impetus of each of their Parts: Therefore in Ambient Bodies, those move which are relatively at rest. And therefore true and abfolute

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folute Motion cannot be defined by a Translation from a Vicinity of Ambient Bodies, confider'd as at Reft. Those Ambient Bodies ought not only to be look'd upon as Quiescent, but also to be truly fo: But all included Bodies, besides their Translation from the Neighbourhood of Ambient Bodies, also participate of the true Motion of those Bodies; and that Translation being taken away, they are not truly, but only feem to be at Reft. For Ambient Bodies are to the included ones, as the outward Part of the Whole is to the inward one, or as the Shell to the Kernel. But if the Shell be mov'd, the Kernel or part of the Whole is also mov'd together, without a Translation from the Shell.

In like manner, if a Relative Place be moved, a Body therein plac'd is also mov'd; and a Body which is moved from a moved Place partakes of the Motion of its Place So the Motion of any one walking backwards and forwards in a Ship whilft it is under Sail, is greater or leffer in refpect of the Earth, or Shore, according as it tends towards the fame or contrary Part with the Ship. But if he stand still in any certain Part of the Ship, he partakes of its Motion, and moves with the fame Celerity : And if it tends towards the fame part of the Ship, in respect of the Earth it will be moved fwifter than the Ship, if to the contrary flower: And fo we ought to reason concerning the Motion of the Earth if it doth move. Therefore, all the Motions which are made from moved Places, are only Parts of Whole and Absolute Motions; and every entire Motion is compounded of the Motion of the Body from its first Place. and of the Motion of this Place from its Place, and fo on a till we come to an unmoved Place ; as appears in the above-mentioned Example. Whence Motions 2

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Motions entire and abfolute can be defined by unmoved Places only : And therefore abfolute Motions are referred to unmoved Places, and relative Motions to moveable Places. But Places are not unmoved, unlefs they all keep the fame Pofitions to one another from Infinity to Infinity; and therefore unmoved Places always abide and conflitute the Space which we call immoveable.

The Caufes by which true and relative Motions are diffinguished from one another, are the Forces impressed on Bodies to generate Motion. True Motion is neither generated nor changed. unless by a Force impressed on the Body it felf. For fince any Part of Matter whatfoever is inactive and merely paffive, it cannot be moved without fome Force impressed from some other Place, nor thrust from its State without some Force which may change its State. But relative Motions (fuch only as Cartes owns) may be generated and changed without Forces impressed on the Bodies themselves. For it is sufficient, if Forces be impressed on other Bodies, to which the Relation is made to alter that Relation in which the relative Reft or Motion of these confift, if those other Bodies give way. So indeed, according to Cartes, it is sufficient that the Earth only be revolved, in order to the relative Motion of the Fixed Stars; and that the Earth will be at Reft, while it is carried round the Sun in the Solar Vortex, if it is in the fame Ambient Parts of fubril Matter, altho' together with those Parts it annually performs a whole Revolution in the Ecliptick, and is absolutely moved about the Sun. Again, true Motion is always changed by Forces impreffed on the moved Body. But relative Motion is not neceffarily changed by these Forces: For if the fame Forces are fo impressed on other Bodies alfo

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alfo to which there is a Relation, that the relative Situation be preferved, that Relation will also be preferv'd, in which the relative Motion confifts. As if a System of Bodies be moved among themselves after what manner soever, and an equal Force act upon equal Parts of the Syftem, according to parallel Lines, altho' that Force really changes the true Motion of every Part, neverthelefs it will not change the relative one: For the Politions and relative Motions of the Parts acting equally and by parallel Lines, will remain among themfelves as they were before. Therefore every relative Motion may be changed, tho' the true be preferved, (viz. by the Mutation of the Motions of other Bodies, ) and preferved where the true is changed; as appears from the laft Example: And therefore true Motion doth not confift in Relations of any Kind.

The principal Effect whereby Absolute and Relative Motions are diffinguish'd one from the other, is a Force whereby a Body departs, or endeavours to depart from the Axis of the Circular Motion. For in a Circular Motion barely relative, this Force is none at all; but in a true and absolute one, it is greater or lefs according to the Quantity of the Motion. If a Bucket, which hangs upon a long Rope, be turn'd round perpetually, fo that the Bottom of it always remains parallel to the Horizon, and the Axis of the Motion perpendicular thereto, until the Rope by twifting is become very fliff: Then fill it with Water, and let both Bucket and Water be at reft ; then by a fudden Force will the Bucket turn about with a Motion contrary to the former; and by the Strings untwifting it felf, it will continue in this Motion. The Surface of the Water will at first be plain, and parallel to the

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the Horizon, as before the Motion of the Veffel: But after the Veffel by a Force impress'd on the Water by little and little, hath at length caus'd that the Water should begin to be turn'd about fenfibly in the Form of a Whirl-pool, as it were; it will depart by degrees from the Middle. and will ascend to the Sides of the Veffel, putting on a Concave Form; and will ascend with a swifter Motion, more and more, until at length performing its Revolutions in equal Times with the Veffel it felf, it comes to reft relatively in the fame. Now this Ascent shews an Endeavour of departing from the Axis of the Motion. For although the Receffion from the Axis of the Motion be in it felf perpendicular to the Axis, yet feeing the Veffel doth in that Place hinder the actual Receffion, the Force will be impress'd upon the next Particles, and become fenfible where it hath room ; and becaufe the true Circular Motion will be greater in Particles which are most remov'd from the Center, forasmuch as it is communicated to them from the Veffel first, and chiefly by reason of the greater Circles and greater Celerity which is towards the Circumference, the Parts more remote from the Center will recede the more from that Center: And thus that Afcent of the Water arifeth from its true Circular Motion, and is measured by its Endeavour of receding from the Center. And it is to be observ'd, that the true Circular Motion is in this Place altogether contrary to the relative Motion. For at first, when the relative Motion of the Water, with refpect to the Veffel, was the greateft of all, foras fmuch as the Veffel was whirl'd about, the Water remaining almost unmov'd; and confequently the Water it felf, which is contain'd, was most fwiftly mov'd to the contrary Part, in respect of D 2 rhe

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the Vessel, without any true Motion in it felf; then, I fay, that relative Motion excited no Endeavour of receding from the Axis; the Water as yet remain'd plain and Level: But after that the Relative Circular Motion of the Water decreas'd, and the True fenfibly begun, the Afcent to the Sides of the Veffel shew'd an Endeavour of receding from the Axis; which Endeavour shew'd the true Circular Motion, more and more increafing, until it became the greateft, which was when the Water came to reft in the Veffel relatively. It is plain therefore, that the faid Endeavour depends not upon a Translation of the Water, in respect of the Ambient Vessel. (Where the Veffel alone is moved, and from thence a relative Motion only is given to the Water.) And therefore true Circular Motion is not to be defined by fuch Translations. There is only one true Circular Motion of every revolving Body, to which one fingle Endeavour answers, as its proper and adequate Effect : But relative Motions, according to divers Relations to divers Bodies, and divers Situations, according as this or that Body is respected, are innumerable, and tend towards all Parts at once; and as it is with Relations in general, are destitute of all true Effects, any farther than they participate of true Motions. From whence also in their System, who would have the Heavens below the Sphere of the Fixed Stars to be turn'd round, and to carry the Planets along with them; the Planets which relatively reft in their Heavens, are notwithstanding truly mov'd, as well as the Heavens themfelves: For they change their Politions according to their different Periods of Revolutions, which is the Cafe of Bo- ' dies really moved. Accordingly the Stars themfelves, as Parts of the Revolving Spheres, partake of their Motion.

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Motion, and endeavour to recede from the Axis.

Therefore the Relative Quantities, which we have now diffinguish'd from the true, are not those very Quantities which they are reckon'd to, be, as the Space contain'd betwixt the Walls of a Chamber, the Diurnal Motion of the Stars, &c. but they are their fenfible Measures, (whether true or falle) which are vulgarly made use of instead of the true measured Ouantities themselves. Wherefore, if the Significations of Words are to be defin'd f om their Ule, by the Names of Time, Space, Place, and Motion, these Measures are properly to be underftood; and the Expression will be unufual and purely Mathematical, if the absolute Quantities themselves be understood. And therefore as they do Violence to the Holy Scripture, who there interpret these Words, as intending the absolute Quantities; fo also do those who from the Reft affign'd to the Earth, and Motion to the Sun, in the Words of the Scripture, are wont to dispute concerning the true Frame of the World, contrary to evident Reasons of Astronomy and Philosophy; as they do likewife, if fuch there be, who from the Words wherein it is predicted, That Time shall be no more, do from thence collect, that Eternal Duration, or Absolute Time, fhall be annihilated. Nor do those any whit less defile Mathematicks and Philosophy, who confound the true Quantities with their Relations and vulgar Measures.

Now to know the true Motions of Bodies, and actually to diffinguish them from the apparent, is indeed difficult; because the Parts of the unmoveable Space in which the Bodies are truly mov'd, do not encounter the Senses. Yet notwithstanding, the Case is not altogether desperate; for we D 2 have

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have certain Tokens and Arguments of the fame, partly from the apparent Motions which are the Differences of the true, partly from that Force which is the Caufe and Effect of the true Motions. As if two Globes, tied together by a Cord at an even Distance from each other, should be rolled-round about a Center of Gravity common to both, the Endeavour in the Globes of departing from the Axis of the Motion, would fhew forth it felf in the ftretching of the Cord; and from thence the Quantity of the Circular Motion might be computed. Then, if any equal Force whatever should be at the fame time impressed upon the Alternate, that is, the diametrically opposite Faces of the Globes, to increase or diminish the Circular Motion; that is, if one should be impress'd on one Part, and the other on the contrary Part at the fame time, the Increase or Decrease of the Circular Motion would be feen from the increas'd or diminish'd Tension of the Cord. And from thence, at length, would be found, Which are the Faces of the Globes on which the Force ought to be impress'd, for the augmenting the Motion most of all; to wit, the hinder Faces, or those which in the Circular Motion do follow. But the Faces which follow being known, and by confequence the opposite ones, or those which go before, the Determination of the Motion will be After this manner, both the Quantity known. and Determination of this Circular Motion might be found in any immense Vacuum, where there is nothing fenfible and external with which the Globes might be compar'd. If now there fhould be placed in that Space fome far diftant Bodies, keeping a given Polition one with respect to another, such as are the Fixed Stars in our Regions; it could not be known from the relative Tranfla-

Translation of the Globes amongst Bodies, whether the Motion were to be attributed to thefe or those; like as we upon the Earth cannot by any apparent Motion of the Fixed Stars, determine whether it be the Earth or they that is indeed mov'd: But if the Cord be minded, and it be found that the Tension thereof is the very fame which the Motion of the Globes requir'd, we might conclude that the Motion is in the Globes: and then at length from the Translation of the Globes amongst the other, collect the Determination of the Motion. For feeing that from the Tension of the Cord, it would be manifest, that the Motion is truly in the Globes, and not in the remote Bodies; the Motion of the Globes, as well in respect of Velocity as Direction, will eafily be determin'd by those Bodies, which are defervedly now to be look'd upon as unmov'd. And in this way we collect the annual Motion of the Earth, as being exactly proportional to the Centri-petal Force towards the Sun; and likewife eafily gather the Stability of the Fixed Stars from the annual Motion of the Earth. Then the Motion of the Earth, and Stability of the Fixed Stars being known, it is as eafy to deduce from thence the Velocity and Direction of the annual Motion. But in what manner true Motions are to be collected from their Causes, Effects, and different Appearances; and on the other hand, in what manner, from Motions either true or apparent, their Causes and Effects are to be gathered, will be taught more largely in the Process.

(10.) The Quantity of Matter is the Measure of the same, arising from the Density and the Magnitude thereof conjunctly.

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Air, which is as dense again as some other Air, and poffeffes double the Space, is Fourfold of the "other. And if a Cubic Veffel contain Air, which by compression is reduced into a leffer Cube. the Density in the leffer Cube will be to that in the greater, as the greater Cube is to the lefs: or in the Triplicate Proportion of the Sides reciprocally; and the Diffances of the Particles of Air which are like, and in like manner polited. will be in the Proportion of the Cubic Sides. The fame Thing is to be underftood of Snow and Powders condens'd by Compression or Liquefaand there is the like reason of all Bodies in whatfoever manner condens'd. We have no regard in this Place to a Medium pervading the Infterflices of the Parts, if there be any fuch. But we shall call this Quantity of Matter, which is to be reckon'd from the Denfity and Magnitude conjointly, every where hereafter Body or Mafs. And the fame is known by the Weight of every Body ; for an equal Quantity of Matter, of what fort soever it is, doth equally gravitate; as is manifest by Experiments of Pendulums which have been most accurately made, as will be taught in the Sequel. And from hence indeed, that we may note this by the way, It is certain, that either there is no Æthereal Medium pervading the Pores of Bodies; or if there be any, feeing it doth in no wife gravitate nor hinder the Motion of Bodies, it ought to be reckon'd Matter differing from that of all other Bodies; yea, in fpeaking properly, it deferves not the Name of Body or Matter at all. But we shall have occasion to fay more of this hereafter.

(11.) The Quantity of Motion is the Measure of the same, arising from the Velocity, and from the Quantity of the Matter conjunctly.

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The Motion of the Whole is the Sum of the Motions of all the Parts; and confequently in a Body double of fome other, and mov'd with equal Velocity, the Motion is Twofold of the Motion which is in the other Body; and in the Double Velocity of the greater Body Fourfold. The Quantity of Matter therefore is equal to the Rectangle of the Density drawn into the Magnitude : and the Quantity of the Motion equal to the Rectangle of the Velocity drawn into the Quantity of the Matter. From which Principle the Forces of Machines are eafily deduced. For wherefoever in the Equilibrium of Machines a Body is greater, there the Celerity of that Body will be fo much the leffer; and where the Body is the lefs, the Celerity will be fo much the greater; fo that the Quantity of Motion refulting from the Body, as drawn into its own Velocity. is equal on both Sides ; as will be more largely fer forth afterwards.

(12.) The innate Force of Matter is a Power of refifting, whereby every Body, as much as it can, perfeveres in its own State of Refling, or Moving uniformly strait forwards in a right Line.

This Force is proportional to the Body, and differs nothing from the Inactivity of the Body, but in the manner of conceiving it, by which it comes to pafs, that a Body is not without difficulty put out of its State whether of Reft or Motion. From whence, by a very fignificant Name it may be called the Force of Inactivity. But a Body exercifeth this Force only when it is acted upon by fome Force from without ; under which exercife of its Innate Force it is confidered in a Double Refpect ; to wit, as Refiftance and Impulfe. Refiftance, as far as it ftruggles with the imprefied Force,

Force, in order to preferve its own State; Impulfe, as the fame Body by not eafily giving way to the Force of the refifting Obstacle, endeavours to change its State. Indeed, it feems most proper to attribute Refistance to quiefcent, and Impulse to moving Bodies; and I should affign any Impetus whatfoever, where one of the Bodies is at Reft, to the positive Force of the moved Body, rather than to the Negative Force of the quiefcent one.

(13.) The impress'd Force is an Action exercis'd on a Body for the changing its State, whether of Reft or uniform direct Motion.

Thus this Force confifts in Action alone, and remains not in the Body at all after the Action. For the Body perfeveres in every new State by its fole Force of Inactivity. Now Force impress'd is from divers Caufes, as from a Blow, a Pressure, or Tendency to a Center.

(14.) The Centripetal Force is that, whereby a Body is drawn, impell'd, or in fome way or other tends to a Center.

Of this Sort is Gravity, whereby a Body tends to the Center of the Earth ; the Force Magnetic, or that whereby Iron tends to the Center of the Loadftone ; the Attraction or firetching of the Cord to retain the Stone that is whirl'd round in a Sling. • Hither alfo is to be referr'd that Force, whatfoever it is, whereby the Planets are continually held back from rectilineal Motions, and compell'd to revolve in Curvilinear Ones. The Quantity of this Centripetal Force is of Three Sorts, Abfolute, Accelerating, and Moving.

(15.) The Abfolute Quantity of Centripetal Force is the Measure of the same, greater or lesser, according to the Efficacy of the Central Cause, which propagates it from the Center all round about.

about. Thus the Strength of Magnets is different, and greater, Cateris Paribus, in the greater Magnet, and leffer in the lefs. The Attraction or Tension of the Cord greater in the Circumvolution of a greater Stone than in that of a lefs; and in the fwifter Circumrotation of the fame Scone than in a Slower. And thus we may conceive, that the Gravitation of Bodies to the Sun, which is a Body fo much greater than the Earth, is greater at an equal diffance than the fame is towards the Earth.

(16.) The Accelerating Quantity of this Centripetal Force is the Measure of it in divers Diftances from the same Center; which is proportional to the Velocity which it produceth in a given time.

As the Virtue of the one and the fame Magnet (in which confequently the Abfolute Quan. t ity remains the fame) is greater in a lefs diftance than in a greater; the gravitating Force in the Surface of the Earth is something greater about the Poles than about the Equator ; it is greater alfo near the Surface of the Earth than at a greater Distance from the Center. But this Accelerating Force, which is diffinctly to be noted, is at equal Diftances from the Center every where the fame, and this in all Bodies whether they be Heavy or Light, Great or Small, Solid or Fluid ; that is to fay, if you do here abstract from the relistance of the Air.' Which Thing is prov'd by the equally fwift Descent of all falling Bodies in Tubes emptied of Air; and from the Motion of all Pendulums, what Matter or Magnitude foever, vibrating together in like Circles or Cycloids.

(17.) The Moving Quantity of the Centripetal Force, is the measure of the same Proportional

onal to the Motion which it generates in a given time.

This Force is the propension of the whole Body towards the Center, which is effimated by the Quantity of the Force contrary thereto, which is requir'd to the hindring its Descent, which is called the Weight of a Body, and is greater in a Body which is greater; and greater in the fame Body by how much it is nearer to the Earth. The Absolute Quantity therefore of this Force we are treating of, is defin'd from Magnitude, or at least from the Strength and Efficacy of the Central Body. The Accelerating is that Force as perpetually decreasing in the Increase of the Diftance, and on the contrary. The Moving Force is the Weight it felf ; which arifeth from the Body or Mais drawn into the Accelerating Force. From whence the Absolute Force being given, the moving Force in a given Body will be as the Accelerating; and the Accelerating being given it will be as the Body. These three Forces therefore are referr'd to three Things, to Bodies, to the Places of Bodies, and to the Center of Force. The Motive Force respects the Body and the Endeavour and Propension thereof to the Center, as compounded of the Endeavours and Propensions of all the Parts. The Accelerating refers to the Place of the Body in the Medium as the Efficacy of the fame Absolute Force according to divers Diftances from the Center ; and the Absolute Force respects the Center or Central Body it felf, as endowed with fome Power. without which the moving Forces are not propagated round about; whether that Power or Caule be theCentral Body (as the Magnet in the Center of the Magnetick Force,) or the Earth in the Center of the gravitating Force, or be fome other Thing which

which doth not appear. At least, this is a Mathematical Conception, and fufficient for our prefent Purpofe; for we do not yet confider the Phylical Caufes of thoseForces. The Accelerating Force therefore is to the Motive as Swiftness is to Motion; for from the fame, as multiplied into the Quantity of the fame Matter, the Moving Force arifes, like as the Quantity of Motion arifeth from the Celerity multiplied into the Body. For the Sum of the Actions of the Accelerating Force upon each Particle of the Body, is the Moving Force of the Whole ; from whence, where the Accelerating Gravity is the fame, the Moving Gravity or Weight is as the Body. And in the fame Body where the Acceleration is diminish'd, as in the upper Regions, the Weight is likewife diminished. Thus where the Accelerative Force is Twofold lefs, the Weight of a Body Twofold or Threefold lefs, will become Fourfold or Sixfold lefs. Furthermore, we call Impulses and Attractions Accelerative and Motive in the fame Senfe. And we nfe the Words [Attraction, Impulse, and Propenfion to a Center] indifferently and promiscuously; We at prefent confidering these Forces not Phyfically but Mathematically, as was faid before, and now fay again, to Caution our Reader from understanding us, as Meaning and Defining Physical Caufes or Reasons of Motions, or attributing to Centres, which are Mathematical Points, true and proper Physical Force ; when at any time we fay that the Centers draw, or have Force in them. And fo far we have given you the Defini-tions requisite to be premis'd to Sir *lfaac Newton's* Philofophy.

Feb. 28. 1704.

LECT.

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LECT. V.

Axioms or Laws of Motions.

VERY Body perfeveres in its own prefent State, whether it be that of Reft, or uniform direct Motion; unlefs it be compelled by fome Force imprefs'd, to change that State.

Projectiles hold on their Motion, fo far as they are not hinder'd by the Reliftance of the Air, or their own Gravity. A Top, whole Parts by cohering continually draw themfelves from the Rectilineal Motion, ceafeth not to be whirl'd about, fo far as is not retarded by the Air, or the unevennefs of the Surface, on which it turns. But the greater Bodies of Planets and Comets maintain their Motions, whether Progreffive or Circular, much longer in Spaces less resisting. This Law of Motion is indeed the fundamental Law of all. and is most evident from the merely Paffive Nature of Matter; which makes it naturally as impoffible for'a Body of it felf to ftop its own Motion once begun, as it is for it to move it felf originally.

(2.) All Motion is of it felf Rectilinear.

For Motion cannot be conceived, but it must be directed and determin'd towards fome Place or other, and it will by the Law foregoing keep the fame Direction which it first had, until it be hinder'd or put out of its way by fome Extrinsic Cause.

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And confequently, whenever any Body is mov'd in a Curve, that Curvature muft needs proceed from External Force ; and therefore muft cease when that Force ceaseth. Which, when it doth, then by this and the foregoing Law, the Motion will be continued in a Right Line, which is the Tangent of the Curve, at the very Point of the faid ceafing Force, or in the laft Rectilineal Direction. Thus it is in a Stone wheel'd about in a Sling, which flipping out of the Sling is not now carried forward in its former Circle or any Circle at all, but in a Tangent of the former Circle; where indeed, by reafon of the Force of Gravity, compounded with the projectile Force, it describes a Parabolic Line : But of this afterwards.

(2.) All Bodies carried about, endeavour to recede from the Center of their Motion ; and by how much the Motion is the fwifter, this Endeavour is the greater.

For feeing Bodies do of themfelves tend unto a Rectilineal Motion, or that which is according to Tangents of Curves; and feeing all the Parts of the Tangents are further diffant from the Center of Motion, than the Parts of the Curves, unto which the Bodies are drawn by the Centripetal Force : It is manifelt, that that Endeavour of going off according to Tangents, doth as much draw back the Bodies from the Center, as the Centripetal Force draws them to it, and is exactly equal to the contrary Endeavour of the Centripetal Force.

(4.) The Mutation of Motion is proportional to the moving Force impress'd; and is according to the right Line in which that Force is impress'd. If any Force generates any Motion, a Double Force will generate a Double one, a Treble a Treble

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Treble one; and this whether the Force be imprefs'd all at once, or fucceffively.

And then this Motion imprefs'd, (for as much as it is always determin'd to the fame Part, with the generating Force) if the Body on which the Impreffion is, was before in Motion, either is added to the Motion thereof, as confpiring together with it; or fubducted therefrom, as being contrary thereto; and thus it either increafeth or diminifheth the former Velocity. But if the Impulfe be oblique, it is added obliquely, and compounded with the former Motion, according to the Directions of both. So that if it were at right Angles, the Velocity as confidered in the first Line will neither be increas'd nor diminifh'd.

(5.) Re-action is always contrary and equal to Action. That is, the Actions of Two Bodies acting upon each other, whether they be Impulfes or Attractions, are always directed each to the contrary Part, and are also equal.

Whatfoever preffeth or draweth another thing, is equally prefied or drawn thereby. If you prefs a Stone with your Finger, your Finger is equally prefs'd by the Stone. When a Horfe draws a Stone tied to a Rope, the Horfe will equally be drawn back to the Stone: For the Rope, which is diftended on both Sides, will, with the fame Endeavour of relaxing it felf, draw the Horfe to the Stone, as it doth the Stone to the Horfe; and will fo much hinder the Progress of one, as it forwards the Progress of the other. If one Body dashing upon another, shall by its Force in any fort change the other's Motion, it will also reciprocally undergo the like Change in its Motion, to the contrary Part, by the Force of the other, and this by reason of mutual Pressure. But then by these Actions are made equal Changes, not

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not of Velocities but Motions; to wit, in Bodies not otherwife impeded. For Mutations of Velocities made to the contrary Parts, fince the Motions are equally changed, are reciprocally proportional to the Bodies : We may thew the Matter briefly thus in Attractions. Let fome Obstacle be fuppos'd to be interpos'd betwixt the Bodies A and B, which attract each other, to keep them from meeting together. If either of the Bodies, as A, be more drawn towards B, than B is towards A; the Obstacle will be press'd more with the Action of A upon it, than with that of B, and confequently will not remain in an Æquilibrium. The ftronger Pressure therefore prevailing, will make that the System of the three Bodies will be mov'd directly unto that Part which is from A to B; and fo in a free Space will be mov'd in infinitum in that Direction, with a Motion continually acce-Ierated: Which is abfurd and contrary to the first Law of Motion. For by that, the System ought" to perfevere in its State, whether of Refting or moving right forwards; and confequently the Bodies will equally prefs the Obstacle, and fo will equally be attracted to each other. And the Thing is the fame if there be no Obstacle ; for the ftronger Motion will in the Meeting overcome the Weaker, and carry both Bodies to the fame Part, and according to its own Direction. Wherefore, either there is no Attraction in a System of Bodies where the first Law hath Place, such as the Solar one is, as we fhall hereafter clearly demonstrate it to be hereafter ; or the Attraction is mutual and equal. The Famous Sir Ifaac Newton hath made an Experiment of the Matter in the Magnet and Iron. If these be separately put in two proper Veffels, fwimming clofe to one another in a ftanding Water; neither will prevail over the other, E buc

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but with an Equality of Attraction on both fides. they will fustain their mutual Endeavours on one another, and at length being in Equilibrio will be at reft. So likewife, the Gravity betwixt the Earth and its Parts is mutual and equal. If the Globe of the Earth HEFGKI (fee Fig. 2. Plate 2.) be divided into Two unequal Parts by a Plane GE, the Gravity of the Part EGF, towards the rest of the Earth, will be equal to the Gravity of the reft of the Earth towards this Part : which is prov'd thus. Imagine the Earth to be divided by Parallel Planes into three Parts. EGF, HKI, EGKH; of which EGF and HKI, are equal to each other, and lye upon the middle Part EGKH. Here it will be manifest, that the middle Part EGKH, doth by its own Weight incline to neither of the Parts, but hangs, as we may fay, in Æquilibrio betwixt both, and fo refts. But the Extream Part HKI lies with its whole Weight upon the middle Part, and urgeth that towards the other Extream Part EGF. And therefore, the Force wherewith the Sum of the Parts HK I, and EGKH, tends towards the Third Part EGF, which is equal to the Weight of the same Parts, is equal to the Third Part EGF. And therefore if the Earth be divided by any Plane whatever, as EG into Two Parts EGF and EGI, the Force wherewith the greater Part EGI tends to the leffer EGF is equal to the Force wherewith the leffer Part tends to the greater; and unless those Weights were equal, the whole Earth would give place to the greater Weight, and in yielding to it would fly away, and there would no Place be found for it; which, as before, is abfurd, and contrary to the first Law.

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(6.) If

(6.) If of two equal Bodies, void of Elafficity, one of them which is in Motion meets the other at reft, upon the meeting they will both be carried forwards together, to the fame part, with half the Velocity of the Body which was moved. For the Body put in Motion, will in the Shock communicate its Motion fo long to the other quiefcent Body, till that moves with the fame Velocity with ir felf. For whilft the Velocity of the mov'd Body is greater than the Velocity of that which was before quiescent, the former Body will still force the other, and accelerate its Motion; but affoon. as the Body that was quiefcent hath gotten a Velocity equal to what the moving Body moves with it can force it no further, but follows it closely. And thus the Motion of the former being now equally divided betwixt them, it appears that they are both carried with half the Velocity which the former had before.

(7.) If two equal Bodies, void of Elasticity, do directly meet each other with the fame Velocity, they upon the Collision will both of them reft.

For so much as either of them tends to go forward it is repelled by the other : And fo thefe Two equal Forces, or Quantities of Motion, tending to the contrary Parts, will deftroy one another ; whereupon, there being no new Caule of Motion, they must needs both of them rest. In which Cafe, the Motion wholly perishethy contrary to the Opinion of Cartes, who would have the fame Quantity of Motion always to remain in the World.

(8.) If two unequal Bodies, deftitute of Elaflicity, meet one another with fuch Velocity, that by how much the greater exceeds the other in Magnitude, by fo much it is exceeded by the lefſeŗ

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fer in Swiftness, so that the Velocities are reciprocal to the Bodies; they will both rest after the meeting.

For the Motions which are directly opposite to each other, being as to the Quantities of them equal, they will definey one another, as before.

(9.) If a moving Body strike another at reft, (but both void of Elasticity) how unequal foever they be in Bulk and Quantity of Matter, they will both move after the shock with the same Velocity towards the same Parts, as in the Sixth Law: And the common Velocity will be so much less than the first, as both the Bodies together are greater than the Body first moved. For since all the Motion of the former Body is now divided between the two Bodies, the Velocity will be so much diminissed, as the Quantity of Matter to be moved is increased.

Corollary, Therefore, when the Bodies are given, there will be also given the Proportion of the Velocity of the moved Body before the Shock, to the common Velocity of the Bodies after the Shock. For as both Bodies together are to the moved Body, fo will the Velocity of the moved Body before the Shock, be to the common Velocity of both after the Shock.

(ro.) If two unequal Bodies, void of Elafticity, which are carried with equal Velocity to opposite Parts, hit against one another, the Quantity of Motion in both, taken together after the Collision will be the difference only of the former Motions; for the lesser Quantity of Motion on either Part will be equivalent to an equal Quantity of Motion on the other Part, and as above will destroy it; wherefore, there remains only the Excess of the Motion, as the fole Cause of it after the Shock. And the Case will be just the

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fame;

fame, as if the Body that had the greater Quantity of Motion, ftruck another at reft with that difference of Motions, and after the Shock ought to be calculated in the fame manner.

(11.) If two equal Bodies, woid of Elasticity, be mov'd with unequal Velocity towards the fame Part, upon their Collision there will remain the fame Quantity or Sum of their Motion, but the common Velocity will be the half of both the former Velocities put together.

For the Excels of Velocity will now be divided equally betwist both Bodies, and fo they will go away together with a mean Velocity.

(12.) If in two unequal Bodies, void of Elafticity, the Greater overtakes the Leffer, the common Velocity, after the Shock, will be greater, than half the Sum of the former Velocities. And on the contrary, it will be lefs when the leffer Body overtakes the greater. For if the Bodies were equal, it would, by the foregoing, be just half that Sum. Wherefore it will be more or lefs than half, in proportion to the Greatness or Smallness of the hindmost Body.

Corollary, Therefore the Velocities and Magnitudes of Bodies before the Shock being given, it will be eafy to compute the common Velocity of the Bodies after the Shock. For the Sum of the Motions, divided by the Sum of the Bodies, gives the common Velocity, when the Motions are made towards the fame Parts; or the difference of the Motions divided by the Sum of the Bodies gives the common Velocity, when the Motion is towards the contrary Parts.

Scholium, Thefe are the true Laws of Motion in Bodies, which yield fomewhat, but do not reftore themfelves, or are endued with no elastic Force; and the fame Laws may perhaps hold alfo  $E_2$  in

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in Bodies perfectly hard, fo that they be not Elaftic, But the Rules of Motion in Bodies perfectly Elaftic, or which reftore themfelves with the fame Force wherewith they are compress'd are altogether different from the former; and therefore tequire a diftind and separate Confideration. And forasmuch, as the Collifions of these Bodies do afford many both difficult and notable Phanomena ; and the famous Monfieur Hugens hath undertaken in a posthumous Work to explicate and demonstrate them ; but this indeed not without much going about, and a long Pomp of Arguments and Figures according to the manner of the old Geometricians, we fhall deliver the Laws of Motion of Elastic Bodies according to his Order, but in a briefer Method and one that is more natural; that fo Beginners may be able in some measure to comprehend the Certainty and Physical Origin of these Laws the first and fundamental one of which is this,

(12.) If a Body perfectly Elastic dalheth upon another Body of the fame fort which is Quiescent and Equal; after the Shock the Motion will be wholly transferr'd into that which was quiefcent before, and with the fame Celerity, but the Body which was mov'd before, will now reft. For the impelling Body, whether it were Blaflic or no, will by the Sixth Law commu-nicate half of its Motion to the other, and begin to go along with the other with the fame Pace; and by its Elasticity, the Force of which is equal to the Force of the direct Impulse, it will communicate the other half of its Motion ; from whence it comes, that the Motion of the Body before quiescent, will now be equal to that which the Impellent had before; and confequently, that feeing fo much of Motion as the Impellent. transfers to the other, fo much it lofeth of its own

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own Motion, the Motion upon the whole will be convey'd into the Quiescent, the Impellent having loft its Motion.

Corollary (1.) If a greater Body dasheth upon a leffer, the former will not reft, but only be mov'd more flowly; and the other which before refted will, in its being mov'd, gain a greater Velocity indeed than was in the Impellent, but a lefs Quantity of Motion.

Coroll. (2.) If a leffer Body dasheth upon a greater, is will not rest but go back; and the Quiescent will gain a less Velocity indeed, but a greater Quantity of Motion than was in the Impellent.

Coroll. (3.) If a Body, put into Motion, hit upon divers Bodies contiguous to one another, and quiescent, they will all rest but the last or furthest of them; and this will be mov'd with a Celerity, equal to, or greater or less, than that of the Impellent, according as the Impellent Body is equal to, or greater or less than the last Body. These *Carollaries* follow naturally from the present Law of Motion, and therefore seem to require no special Demonstration.

(14.) If two Bodies perfectly Elastic, which are equal, but mov'd with an unequal Celerity, dash one upon another, they, whether they were before carried to the fame part or to the contrary, will after the Contact be mov'd each with that Celerity which the other had before.

For if they tend towards the fame Parts, the common Velocity on both Sides being taken away, there will only remain the difference of the Velocities as the fole Caufe of the Change in the Shock; and fince by the foregoing Law that Velocity will be communicated to the flower Body, it follows, that the firiking Body fhall lofe E\_4 th a

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that Excels of Motion, and the flower Body get it : that is, in other Words but the fame Senfe. they will move with each others Velocities. And in like manner, we may demonstrate the same in the fecond Cafe, where the Bodies carried towards different Parts are supposed to strike one another. For the common Velocity being taken away on both Sides, the difference of Velocity. which after the Shock tends the contrary way, and will not at all change the former common Velocity, will remain as before, the fole Caufe of changing the Velocity; which by the foregoing Law will be transferred from the fwifter to the flower Body : Whence as before, it will follow, that elaftic equal Bodies after the Shock will move with each others Velocities.

(15.) Any Body how great foever, may be moved by any Body how small soever, with any Velocity what foever. This Law of Motion is indeed an Axiom, manifest in it felf, and wants no Demonstration.

(16.) When two Bodies, perfectly Elastical, are dash'd one upon the other, they depart from one another with the fame Celerity wherewith they approach'd one to the other ; that is, not with the fame abfolute Celerity perhaps, but with the fame relative Celerity. This Law indeed is the Foundation of all the following Laws of Motion. The Thing was before prov'd concerning two Elastical Bodies which are equal, when it was demonstrated, that in their departure from each other, there are the fame True and Absolute Celerities on both Sides, the Seats of them only being chang'd. And therefore it is necessary, that the Relative Velocity of departing from one another be the fame with that of coming towards one another. Now concerning Bodies unequal it is thus fhew'd. If a greater

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greater Body ftrikes a leffer, which is either quiescent, or at least moved flower, it will communicate some of its Motion to the quiescent or flower Body; and the Elasticity being laid alide. it will not reft; and by fuch Communication together with the quiefcent or flower Body it will continue to go on with a direct Impulse, and also by the Elastick Re-action accelerate that quiefcent or flow Body, until it recede from it felf with the fame Velocity, by which it had withftood its Motion, and compressed its Elasticity ; that is, by which it felf approached the other. Indeed, the greater Body must necessarily impress this Velocity on the leffer ; but it cannot imprefs a greater; (altho' the leffer Body of it felf is capable of a greater:) for as foon as the quiescent or slower moved Body hath gotten a Degree of Velocity equal to the Impulie or former Relative Velocity, it will fly thence, and will fustain no farther Impulse whatfoever. But if a leffer Body strike a greater, either quiescent or moved flower, it is impossible the leffer Body fhould impress the whole Excess of its Velocity on the quiescent or flower Body, (for that will be in that Cafe only where the Bodies were equal, as we just now faw in the 12th and 14th Cafes.) But in the Communication of the Motion, the Excels of the fwifter is loft, even when the Elaflicity is not confidered. And while the Bodies go on together in that manner, the hindmost will react on the foremost until they are separated with the fame Relative Velocity, with which at first they came together; for in this, and only in this Cafe can the Elaftic Force be equal to the Impulse : or rather fo far, and no farther can the leffer Body fuffer the Re-action. as in the former Cafe. But in those Bodies which mutually strike one ano-1 1 ther

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ther with unequal Velocities, we must take away the common Velocity from both, fo that it will generate the fame Velocity after the Shock but the Seats changed ; but then there will be left only the difference of the Velocities, as the fole Caufe to change the Velocity; which Caufe indeed will not cease, but in Acting and Re-acting, the Bodies will depart from one another with the fame Relative Velocity, with which they came together. And the Matter will every where depend on this, that the Elastick Forces, every where equal to the impressed, produce their Whole and Pure Effects only; which cannot be done any otherwife, than if the Relative Velocity of Receding, exactly answer to the Relative Velocity of approaching.

(17.) If two Bodies perfectly Elastical, do each return to the Impulse with the fame Celerity wherewith they rebounded from it; they will each of them, after the Second Impulse, require the fame Celerity as they had before the first Meeting. For by reason of the given Quantity of the Stroke in the Collision, there will be given therewithal a certain Rectangle, whofe two Factors are the Diftances from the Point of Concourfe, both the primary diffance, and that to which they return on both Sides after the 1st Conflict; if therefore we divide the Rectangle by the first Distance, there will come forth the fecond Diffance as the Quotient ; and if we divide it by the fecond Diftance we shall obtain the first Diffance for a Quotient; and fo perpetually. From whence it follows, that those Distances as describ'd in a given time, or the Velocities of coming to, and receding from one another, do answer to each other mutually, and follow one upon the other.

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(18.) In

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(18.) In two Bodies which meet one another. whether they be Elastic or not Elastic, there doth not always remain the fame Quantity of Motion as was before, but it may be greater or lefs. This Propolition, which directly contradicts Cartes, we deduced out of the Seventh Law, as to Bodies not Elaftic ; and it follows out of the last Law fave one, concerning Bodies Elaftic. For feeing the Quantity of Monion is estimated from the Celerity drawn into the Matter; and feeing in Bodies howfoever unequal, and unequally mov'd, the thing is indeed thus, that the Sum of their Velocities, or the Relative Velocity remains given, the Quantity of Motion will be very unequal, as the greater or leffer Body gains a greater or leffer Part of the entire refrective Velocity ; as will more clearly appear from that computation of Motions which will prefently follow.

(19.) If a Body perfectly Elastical, which is greater, meets a leffer one which is quiefcent, it will give a Velocity to it lefs than the double of its own. Bot feeing, after the Impulfe, the Bodies ought to be separated from each other with the fame refpective Velocity, with which they came to one another, that is, in the prefent Cafe with the Velocity with which the greater was mov'd before the Impulfe; if the Velocity of the quiefcent Body were double to the Velocity of the Body incutring, then after the Motion communicated to the Quiefcent, the Impellent ought to go forward with the fame Gelerity, which it had before, without any diminution of it : which is abfard.

(20.) If two Bodies perfectly Elastic, the Gelerities whereof are in reciprocal Proportion to their Magnitudes, meet one another directly and oppositely, they will both rebound with the fame Celerity Celerity with which they came to each other. For feeing the Force which arifeth from the mere Impulse of Bodies without any confideration of the Elasticity, is on both Sides equal, they by the Eighth Law will mutually fustain and destroy each other; fo that there will remain no Cause of Motion but the Elastic Force; which feeing it is on both Sides equal will beget equal Motions on both Sides; and confequently both the Bodies will rebound with the fame Celerity which , they had before.

Scholium, A Problem. There being given two unequal Bodies perfectly Elastic meeting one another directly, both of which are mov'd or one only. and the Celerity of both, or of the one, if only one be mov'd, being also given, to find the Celerities with which both are mov'd after the Meeting. Let it be made thus, as the Sum of the Bodies, is to the Double of the Second Body, fo is the given respective Celerity of Approach to the other Celerity. The Difference betwixt this last found Celerity, and the Celerity of the first Body before the Impulse (or in one Case the Sum of them, to wit, where the firft Body in the Motion goes before) will give the Celerity of the first Body after the Meeting; which Celerity being fubducted out of the whole respective Celerity given, the remainder will be the Celerity of the fecond Body after the Meeting. Which Rule is thus demonstrated. The Velocity of the first Body after the Meeting will be the Difference betwixt the Velocity of the first before the Meeting and the whole Velocity, where the Bodies are put to be equal, fo that the Sum of the Bodies is equal to the Double of the fecond Body, as appears from the 14th Law : It is therefore manifest, that all the Difference, that is, the Motion of

of the first Body after the Meeting, doth arife from the Difference of the Sum of the Bodies, and the double of the second Body; and confequently is proportional to the same. Which is the very Thing that the present Analogy supposeth.

For Example : Let the first Body be moved towards the right Hand with the Celerity of Six -Parts ; and the fecond to the contrary Part with the Celerity of Four Parts : Let the first Body alfo be quadruple of the fecond Body. The refpective Velocity therefore of approach will be of 10 Parts, 6+4=10. And the Sum of the Bodies will be of Five Parts. It will therefore be thus; as the Sum of the Bodies  $= \varsigma$  is to the Difference of them=2; fo is the whole respective Velocity=10 to  $\frac{2^{2}10}{5}$  = 4; the Difference of which Velocity, and of the Velocity of the first before the meeting=2, will give the Velocity of the first after the meeting. From whence the Velocity of the fecond after the Meeting will be found to be of Twelve Parts, QE I.

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But if the other Body doth reft, the Celerity thereof after the Meeting will eafily and immediately become known by the former Analogy. To wit, if the greater Body in the former Example be put to be unmov'd, the Motion thereof will immediately be found thus. For as the Sum of the Bodies =  $\varsigma$ , is to the Double of the fecond Body = 2; fo is the whole refpective Velocity = 4, to the Velocity of the Second after the Meeting =  $\frac{2^{N_4}}{5} = \frac{8}{3}$  or  $\frac{13}{5}$ . For the Difference betwixt the Celerity of the firft Body before the Meeting, which was none at all, and this Celerity, will be the very Celerity of the firft after the Meeting, and confequently the Velocity of the fecond will be Parts  $\frac{15}{5}$  or  $\frac{23}{5}$ .

(21.) The

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(21.) The Celerity which a greater Body perfectly Elastic, gives to a leffer perfectly quiefcent which is perfectly Elastic, hath that Proportion to that Velocity, which the leffer moved with the like Celerity gives to the greater which is quiescent, which the Magnitude of the greater hath to the Magnitude of the lefs. For by reason of the given respective Velocity in both, and the Sum of the Bodies also given, the Computation will be alike in both Cafes; to wit, as the given Sum of the Bodies is to the greater Body or the double of the leffer to the fought Velocity. The Velocities therefore are as the Bodies, QED.

Scholium, We shall in this Place, by way of Corollary, annex the three remaining Theorems of Hugens hitherto belonging, (albeit the Demonstration of them is longer than agrees to this Place) both because they are in themselves most noble Theorems,, and because, they may sufficiently appear manifest from a Calculation taught under the foregoing Problem.

(1.) Two Bodies perfectly Elastic meeting one another, that which is produc'd from drawing the Magnitudes of each into the Squares of their respective Velocities, being added together, both before and after the meeting of the Bodies, will be found equal on both Sides; if to wit, the Proportions both of the Magnitudes, and of the Velocities, be express'd in Numbers or Lines.

(2.) If any Body perfectly Elastic, meets another Body which is quiescent, whether greater or less; it will give a greater Celerity thereto, by an interpos'd Elastic Body of a mean Magnitude, which is likewise quiescent, than if it hit upon it without the Interposition of the other Body: And

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And it will confer upon it the greatest Celerity of all, if the interpos'd Body be a mean proportional betwixt the Extreams.

(3.) By how much the greater Number of Bodies perfectly Elastic, is interpos'd betwixt two unequal Bodies perfectly Elastic, whereof the one refts and the other is mov'd, by fo much the greater quantity of Motion will be produced in the quiefcent Body: But the greateft Motion of all will be convey'd through the Multitude of the interpos'd Bodies, if the interpos'd ones, together with the Extremes, do constitute one continued Series of Geometrical Proportionals.

And it is to be noted, that it appears from the two laft Theorems, according to the Author's Computation; that if there be given an 100 Bodies placed in a Line, which are in a double Proportion, and the Motion begins at the greatest, the Celerity of the least will be to the Celerity of the greatest, as 14,760,000,000 to 1 or thereabouts. But if the Motion begins with the least, the Quantity of the Motion will be increased in the End, in about that Proportion which 1 bears to 4,677,000,000,000. Where in the former Case is seen a most prodigious increase of Celerity; and in the latter, a more stupendious Augmentation of the Quantity of Motion.

But to conclude, what things Hugens afferted (that I may advertife this at length) concerning all Bodies, or at leaft concerning all Bodies perfectly hard, we have all along with our Famous Mathematicians Wallis and Newton, demonstrated of Bodies perfectly Elastic only. Nor certainly, ought they to be otherwife understood or affirm'd. For the Laws of Motion which agree to Bodies not Elastic, are for the most part altogether different from these, as is abundantly manifest from

from what hath been faid, and therefore ought in no wife to be mingled together with the Laws of Elastics. But as to what concerns Bodies imperfectly Elastic, we shall deliver their Laws out of the famous Newton in the following Lecture.

May 8. 1704.

#### LECT. VI.

Law. VERY Body will in the fame (22.) E Time defcribe the Diagonal of a Parallelogram with Forces conjunct, that it would do the Sides with those Forces separate.

(In Fig. 3. Plate 2.) Let the Body A be carried in a given time by the fingle Force A B, impress'd according to the Direction of the Line A B, from A to B; and by the fingle Force A C impress'd according to the Line A C let it be carried in the fame time from A to C; and let the Parallelogram ABDC be compleated; I fay, that by both the Forces impress'd together, it will be carried in the given time from A along the Diagonal unto D. For because these Forces impress'd together are not opposite one to the other, they can in no wife deftroy one another, but will beget a certain Motion which is in the middle betwixt both. For feeing the latter Force A C, acts according to the Line AC, which is Parallel and Equal to BD, this Force ought not at all to change the Velocity of coming to the Line BD, which is produc'd by the former Force. Therefore the

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the Body will come in the fame time to the Line BD, whether the latter Force be impressed or no; and confequently the Body, in the End of the given time will be found fomewhere in that Line BD. And by the fame Argument, fince the former Force A B, acts according to the Line AB, which is Parallel and Equal to CD, this Force ought not at all to change the Velocity of coming to the Line CD, which was generated by the latter Force. Therefore the Body would come in the fame time to the Line CD. whether the former Force were impress'd or no; and confequently in the End of the given Time will be found fomewhere in that Line C D. And therefore it is necessary, that in the End of that time the Body should be found in D, the Concourse of the two Lines BD and CD. Furthermore, feeing the fame Thing may in altogether the fame manner be demonstrated of innumerable Points d, d, d, &c. in the same Diagonal Line; it is manifest, that the Body with these Forces conjoin'd, ought always to defcribe this Diagonal Right Line. Q. E. D.

Coroll. (1.) The Forces being given, the Velocity arifing from the Conjunction of them will be fo much the greater, by how much the Directions of the first Forces do the more conspire together, or by how much the Angle BAC is the lefs; and fo much the lefs as the Directions of those Forces are the more opposite to one another, or the Angle BAC is the greater: And the Velocity of both Directions which tend to go according to the Parallel Lines AC, BD, and AB, C D, parallel to the Lines B D and C D, or any others whatfoever, is in no wife chang'd by the conjunction of these Forces, but always remains F the

the same, as is manifest from what has been demonstrated in this Proposition.

Coroll. (2) The fame Diagonal Line A D, may be defcrib'd from the conjunction of innumerable Pairs of Forces. Thus, if inftead of the former Force A B you fuppole another, to wit A E; and for A C put A F, and then compleat the Parallelogram A E D F, the Line A D being the common Diagonal: The Body A from the conjunction of these Forces will defcribe the fame Line A D which it did before, as is manifeft from the Proposition. And there is the fame Reason for any other Pair of Forces whatever, by which the Sides of a Parallelogram, whole Diagonal is A D, ought to be defcrib'd.

Coroll. (2.) Therefore the Forces being given both in Magnitude and Direction, there is alfo given one right Line to be describ'd, to wit the Diagonal of a Parallelogram : but the describ'd Line or Diagonal being given, the Forces are not thence given, nor the Directions by which that Parallelogram was describd. For the Sides of a Parallelogram being given, and the Angle included, there is given therewithal the Parallelogram it self, and confequently the Diagonal of that Parallelogram; but a Line being given in Length and Direction for a certain Diagonal, the Parallelogram it felf is not from thence given ; to wit, because the Line may be the Diagonal of innumerable Parallelograms. For as the Sides of the Parallelogram, without the included Angle, do determine no certain Diagonal; fo neither doth the Diagonal without the adjacent Angles determine any certain Sides.

Coroll. (4.) Where the primary Forces BA, B,D (fee Fig. 4. Plate 2.) are equal, and comprehend the Angle A BD of 120 Degrees, the Velocity refulting from the conjunct Forces, will be the

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Tame as that of either of the feparate Forces; and the Directions of the Force only will be chang'd; for in this Cafe the Triangles A B C and B C D will be Equilateral, and compose a Rhombus; and the Diagonal therefore B C will be equal to either of the Sides.

Coroll. (c.) Where the primary Forces are equal, and the Angle included by the Sides is a right Angle, the Velocity arising from the Forces conjoin'd will be incommensurable to either of them separate; to wir, because the Diagonal of a Square is incommensurable to the Side.

Scholium. What has been spoken in this Propolition, and its Corollaries, concerning real Motions and Velocities, is to be applied to any Endeavours or Tendencies to Motion whatever. Thus, if the Body A in the former Figure be impelled by two Forces, which have that Proportion amongst themselves, which the Lines A C and A B have, and also are impell'd according to the Directions of the fame Lines, or be press'd, or drawn in those Directions, or any other way read according to the fame, although actual Mionion should not prefently follow by rea-Ion of some Obstacles, yet notwithstanding the Impulse or Force ariling from the conjoin'd Forces, tends according to the Direction of the Diagonal AD; and the Velocity to be produced is to be express'd or represented by, the Line AD 1 as will more eafily be understood from what follows

(a2.) All Forces and Motions whatever may be refoly'd into innumerable Forces and Motions; and on the contrary, direct Forces, and rectilinear Motions, may be compounded of innumerable oblique Motions and Forces,

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Thus in the former Figure, the Line and Dire-Ation of the Motion is the fame, whether it be compounded of the Forces A B, A C, or of the. Forces AE, AF, or arife from one fingle Motion impress'd according to the Line AD. And on the other Hand, any Motion whatever along the Line A D, although it may arile perhaps from one fingle Force impelling right forward, yet may be confidered as compounded of A B, A C. or AE, AF, and innumerable other the like; forafmuch, as the very fame Motion would arife from all those combin'd Forces. And in the fame manner are we to reason of Motions more compounded. For in the first Place, having confidered a Diagonal Line as refulting from two Forces combin'd ; we may then reduce those two unto one fingle Force, and conceive a third Force as superadded ; which being join'd to the former, will produce a Motion along another Diagonal of fome fecond Parallelogram, and then may we in our Conception superinduce a fourth Force, and after that a fifth, and fo on infinitely. Nor can indeed any direct Force, where there is occasion to resolve it into more, be otherwise resolv'd than thus. Now this Composition and Resolution of Forces occurs very frequently, and is abundantly confirm'd from Mechanics, as we fhall now fhew with our Author.

If unequal Rays OM, ON (fee Fig. 5. Plate 2.) going forth from O, the Center of Tome Wheel, do by the Threads MA, NP, fuftain Weights in Equilibrium, and the Forces of the Weights unto the moving of the Wheel be requir'd : Through the Center O let the Line KOL be drawn, meeting the Threads which fuftain the Weights perpendicularly; and from the Center O with OL, the greater Interval of the two οκ

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OK, OL, let there be describ'd a Circle meeting the Thread MA in D; then through O and D let there be drawn the right Line OD; to which let DC be perpendicular, and AC Parallel; and let the Parallelogram DCA be compleated. Now because it nothing matters, whether the Points of the Threads K L D be fastned or not fastned to the Plane of the Wheel; the Weights will be of the fame Force, whether they be hanged on the Points K and L, or those D and L; for (fetting alide the Weight of the Thread) the Gravity of the fame Body is the fame wherefoever the Thread is fixed, fo it be in a Line perpendicular to the Horizon : Let therefore the whole gravitating Force at A be reprefented by the Line AD as the Diagonal of a Parallelogram, that from the Proportion of the Diagonal to the Side of the Parallelogram, we may come to know where it is, that one of the Forces is none at all. Now that whole Force which A D defigns may be refolv'd into innumerable Pairs of Forces, but feeing others are foreign to our Purpole, let it be refolv'd into Dc (or AC) and DC; the one to wit according to the direction of the protracted Radius DO, the other Perpendicular to the fame Radius. One of these Forces AC or cD, by reason that it draws the Radius O D directly from the Center (for it tends from D to c in the Protracted Radius) is of no force at all for the moving the Wheel; but the other Force DC which draws the Radius DO perpendicularly, is of the fame Force as if it drew the Radius OL, equal to OD perpendicularly: But feeing the Wheel doth by the Hypothesis reft in Equilibrio, the Weight P will be to the Weight A, as the Force DC is to the Force DA. For the whole Force of the Weight P F 2 draw

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draws the Radius OL perpendicularly, and for confers its whole Force to the moving of the Wheel; but only that Part of the entire Weight A represented by the Line A D, I fay, only that Part of this Weight which is expounded by D C, draws the Radius OD, which is equal to OL perpendicularly, the Force of the other Part which tends according to the Radius O being wholly loft; that Part therefore DC only avails to the moving of the Wheel. Since therefore. because of the Equilibrium suppos'd on both Sides, the entire Force of the Weight P is equivalent only to a certain Part of the Weight A, to wit, to DĆ; it is manifest, that the Weight A ought to be fo much greater than the Weight P, by how much the Diagonal DA is greater than the Side DC; and that because of the Declination of the Body A from the perpendicular DC. As therefore the Weight A is to the Weight P, fo' is DA to DC; that is, because of like Triangles ADC, DOK, as OD or OL is to OK, therefore the Weights O and P, which are reciprocally as the Rays OL and OK which are placed in a ftraight Line, will be of the fame Force on both Sides, and confequently fland in an Equilibrium. And this indeed is the moft known, and the fundamental Property of the Balance, Leaver, and Axis in Peritrochio, which from the refolution of Forces is eafily demonstrated. But if either of the Weights be greater than in this Proportion, its ftronger Force will prevail, and fuffice to move the Wheel. But if the Weight  $\pi$ , equal to the Weight P, be partly hang'd upon the Thread N *π*, and doth partly lye upon the Oblique Plane  $\pi G$ ; let N H and  $\pi H$ be drawn, the former Perpendicular to the Horizon, the latter to the Plane  $\pi G$ ; and let the Pa-2 . rallelogram

rallelogram **T**NRH be compleated. And if the entire Force of the Weight . tending downwards, be represented by the Line NH, it may be refoly'd into the Forces  $\pi$  N. R N. Now if to the Thread  $\pi$  N, fome Plane as  $\pi Q$  were perpendicular, cutting the other Plane  $\pi G$  in a Line parallel to the Horizon, and the Weight  $\pi$  lay wholly upon these two  $\pi Q$ ,  $\pi G$ ; the Weight  $\pi$  would press these Planes perpendicularly, to wit  $\pi Q$ by the Force  $\pi$  N, and the Plane  $\pi$  G by the Force **R** N. And therefore if the Plane  $\pi O$  should be taken away, that the Weight fhould firetch the Thread, because the Thread doth now by fustaining the Weight fupply the Place of the Plane which is taken away, it will be ftretch'd with the fame Force **<b>w** N wherewith the Plane was before prefs'd. From whence the Tension of this Oblique Thread will be to the Tension of the other perpendicular Thread PN, as # N is to NH; and therefore if the Weight a be increas'd in the **Proportion of NH to N** $\pi$ , it will fuftain the Weight A, and the Wheel will not be mov'd. From whence, if the Weight  $\pi$  be to the Weight A in the reciprocal Proportion of the leaft Diftances of their Threads AM PN from the Center of the Wheel, or as KO to OL, and also in the direct Proportion of NH to  $\pi N$ ; that is, joining both Proportions together, as the Rectangle KO x NH to the Rectangle OL  $\times \pi N$ , the Weights will be of equal Force to the moving of the Wheel; and confequently they will futtain each other in an Equilibrium, as any one may easily find upon Tryal.

Corollary (1.) From hence we may difcover a new way for measuring all leffer Weights from one given Weight. For if the Plane  $\pi$  G, perfectly polified, be placed gradually at divers Degrees of F 4 Inclination

Inclination, the fame Weight a or P will be equivalent to divers Weights whatfoever lefs than it felf: to wit, in the Proportion of the Line  $\pi N$ to HN. And confequently, if a Table should be made, of the Proportions of the Lines \* N and HN in the divers Degrees of Inclination; it will be easy from the Inclination of the Plane \* G. and one only given Weight \* or P to examine and determine the Weights of all Bodies lefs than T or P.

Coroll. (2.) Hence likewife we may effimate the Velocities or Weights of Bodies, descending or declining in any Plane whatfoever : Let AB be the inclining Plane, and (fee Fig. 6. Plate 2.) of the Body defcending along that Plane, or leaning upon it; let the entire Force of its Gravity be represented by the Line df perpendicular to the Horizon; and let that whole Force be refolv'd into two Forces f c, and f g, of which let the one be perpendicular to the inclined Plane, for the bearing of which therefore that Plane adequately fufficeth ; and let the other be put Parallelwife with respect to the inclined Plane, which therefore is fitted for exciting a Motion, or at least for procuring an Endeavour towards Motion without any Impediment. The Motion therefore or Weight in the inclined Plane, is to the Motion or Weight in the Plane perpendicular to the Horizon, as the Side fg is to the Diagonal Line fd: that is, because of the likeness of the Triangles fg d and ABC as AC is to AB, or as the Radius is to the Secant of the Angle BAC; which is a Proposition very well known in Mechanics.

Caroll. (2.) From hence also the Force of the Wedge appears. Let (Fig. 7. Plate 2.) CCA be a Wedge ftruck by a Maller with a direct Blow; let the

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the whole Force of the Stroke be expressed by the Line DA; and let it be refolv'd into two Forces DQ and DR; the one of which let be perpendicular to the Face of the Wood CA, and confequently directly fet to remove the fame Face, and the other DR parallel to the fame Face, and confequently polited to go forward directly; and let the fame be underftood of the other half of the Wedge DAC. The removal then of the lateral Obstacle, according to the Line DQ, is to the Progress of the Force downwards, according to the Line DR, as DQ is to DR; that is, because of the Triangles DQA, DCA, which are like, as DC is to DA; or the force of the other Part being computed, as CC is to DA; which alfo is a most known Property of a Wedge, and univerfally receiv'd in Mechanics. Or, if we be minded to difpatch the Matter with Sir Isaac Newton out of what has been before demonstrated : The Weight \*, in the last Figure fave one, lying upon the two oblique Planes  $\pi Q$ ,  $\pi G$ , will have the Nature of a Wedge betwixt the interval Faces of the cloven Body; and from thence the Force of the Wedge and Mallet will be known. for the Force wherewith the Weight # preffeth upon the Plane  $\tau Q$ , is to the Force wherewith the fame is impell'd, either by its own Gravity, or by the Stroke of the Mallet, according to the Line perpendicular to the Horizon, as  $\pi \dot{N}$  is to NH; and is to the Force wherewith it preffeth upon the other Plane  $\pi G$ , as  $\pi N$  is to NR. Nay the Force of the Screw likewife may be collected by the like division of Forces, forasmuch as the fame, in our Author's Opinion is a Wedge forced by a Leaver.

Scholium.

Scholium, The use therefore of this Composition and Resolution of Motion, appears of very wide Extent, and from its Clearness demonstrates its own Truth, fince all the Mechanic Science, which is diversly demonstrated by Authors, depends upon the Things which have been now faid: For from these are derived the Forces of the Machines which are wont to be composed of Wheels, Screws, Leavers,

and Weights, alcending directly or obliquely, and the reft of the Mechanic Powers; as also the Force of Muscles for moving the Bones of living Creatures.

Octob. 22, 1704.

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### LECT. VII.

T Control HE Quantity of Motion which is collected, by taking the Sum of Motions made to the fame Part, and the difference of those made to the contrary Parts, is

made to the contrary Parts, is not chang'd by the Actions of Bodies one upon another.

For Action, and the contrary to it Re-action, are equal by the Fifth Law; and confequently by the Fourth, they make equal Mutations of Bodies towards the contrary Parts. Therefore, if the Motions be made to the fame Part, whatfoever is added to the Motion of the Body, which flies away, will be fubducted from the Motion of that which follows, fo that the Sum fhall remain

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main as before. But if Bodies meet one another in the fame Line, there will be an equal fubduction from the Motion of both, and confequently the difference of the Motions made to contrary Parts will remain the fame. As if the Spherical Body A be Threefold greater than the Spherical Body B, and have two Parts of Velocity; and B follows in the fame right Line with 10 Parts of Velocity; and confequently the Motion of the BodyA, refulting from its Velocity and Magnitude together, is to the Motion of the Body B, effimated in the fame manner as 6 is to 10; therefore the Sum of the Motions to the same Part is of 16 Parts. In the meeting together therefore of the Bodies A and B, if the Body A, according to the Quantity of its Elasticity, doth gain three Parts of Motion, or four or five, the Body B shall lose fo many; and confequently the Body will go forward after Reflection, with Nine Parts, or Ten, or Eleven, and B with Seven, or Six, or Five, the Sum of Sixteen Parts, remaining always as before; as the thing will always happen in Bodies not at all, or at most in a less degree Elastic. But if the Body shall gain 9, 10, or 11, or 12 Parts, and so go forwards, after the meeting with Fifteen Parts in all, or Sixteen, or Seventeen or Eighteen ; the Body B, whilst it loseth fo many Parts as A gains, either will go forward with one Part having loft Nine; or it will reft, its progreffive Motion of Ten Parts being wholly loft; or it will go back with one Part, having loft its Motion, and (as I may fay) one Part more, or it will retrocede with two Parts, because of the progressive Motion of twelve Parts which was taken away. And fo the Sums of the confpiring Motions 15+1, or 16+0; and also the Differences of the contrary Motions 17---1, or 18---2, will always be of Sixteen Parts, as it was before the Meeting or Reflection ; which will

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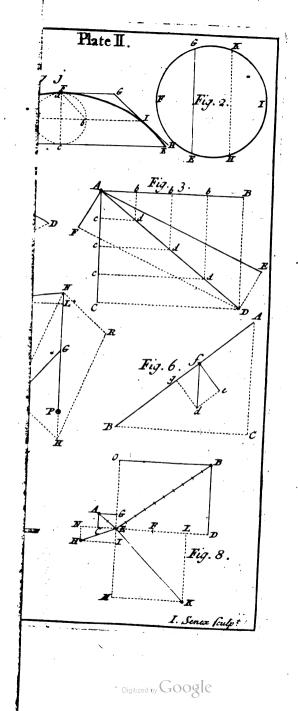
will happen also in Bodies imperfectly Elastic, as may sufficiently be understood from the Laws of Motion before delivered, and is afterwards to be faid concerning Bodies imperfect-ly Elastic. But the Motion wherewith Bodies go forwards after Reflection being known, there will be found the Velocity of the fame after the Reflection, by faying it is to the Velocity which was before the Reflection, as the Motion after is to that which was before. As in the last Case, where the Motion of the Body A was of Six Parts before the Reflection, and of Eighteen afterwards, and its Velocity was of two Parts before the Reflexion; its Velocity will be found to be of Six Parts after the Reflexion, by faying according to the Golden Rule; as Six Parts of Motion before the Reflexion is to Eighteen Parts afterwards, fo is the Velocity of two Parts before the Reflexion to Six Parts of Velocity after. For feeing the Quantity of Motion doth arife from the Velocity and Magnitude conjunctly, in a given Body the Quantity of Motion will be effimated from the Velocity alone, and confequently the Quantity of Motion and Velocity will be directly proportional to each other. But if Bodies not Spherical, or which move in divers right Lines, fall one upon another obliquely, and their Motions after the Reflexion be required ; the fituation of the Plane, by which the concurring Bodies are touched in the Point of Concourfe, is to be confidered and known. Thus the Motion of both Bodies is to be diftinguish'd into two, one perpendicular to the Plane, the other parallel to the fame ; but the parallel Motions, by reason that they are in no wife opposite to each other, the Bodies acting upon one another, according to a Line perpendicu-lar to this Plane, the fame are to be retain'd after

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after the Reflexion as well as before; and equal Mutations to the contrary Parts are to be attributed to the perpendicular Motions, in fuch fort, that the Sum of the confpiring Motions, and the difference of the contrary ones remain always as before. As for Example : Let the Body A (fee Fig. 8. Plate 2.) which is Spherical, and perfectly Elastic, be threefold of the Body B, which is also Spherical and perfectly Elaftic, and let A have two Parts of Velocity, reprefented by the Line A E. divided into two equal Parts; and let the Body B meet it obliquely according to the right Line BE, in the Angle A E B, with ten Parts of Velocity, reprefented by the Division of the Line BE into Ten Parts equal amongst themselves, and to the former; let the Angle A E B be bifected by the right Line O E M; Let A G and BO be let down perpendicular to the Line EO; and the Parallelograms ACEG, BOED be compleated. The Plane then which paffeth through O M, will be that by which the Spherical Bodies will be touched in the Point of Concourse : and the oblique Motions along the Diagonals AE and BE will be diftinguish'd on both Sides into two, to wit AE into AG and AC, and BE into BO and BD; one of which Motions AG and BO, or CF and ED are perpendicular to the Plane of Concourfe: to which alone therefore, as being directly opposite to each other, and tending to the contrary Parts EC and ED, all the change of the Motions in the Concourse is to be referred, in the mean while that the other AC and BD, or GE and OE which are parallel each to the other, and in the Point of Concourse tend wholly unto the fame Part, are fo far from being contrary to one another, that they are rather to be reckon'd to confpire together direaly,

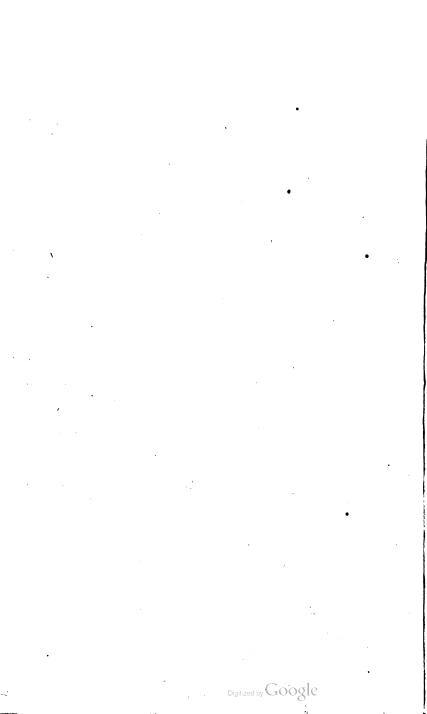
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rectly, and confequently are to be retain'd after the Reflexion as well as before. Wherefore let EI be equal to EG, and EM equal to EO; and that we may estimate the Mutations of Motions made to the contrary Parts, and to be direand according to the Line C D; let us make the computation according to the Twentieth Law of Motion, borrowed from Hugens. Let it be made then, as the Sum of the Bodies A and B = 4 is to the double of the Body B = 2; fo is C D, the respective Celerity of the approach, which is of Twelve Parts, (for becaufe the Triangles A GE, BOE are like. A G or CE is to BO or ED as A B = 2 is to E B = 10 ; and confequently, AE + EB = 12) to the half of GD = CF =6. And the Difference, betwixt the Celerity of Six Parts, and the Celerity of the Body A before the Impulse which was of Two Parts, which is, equal to 4, will give the Celerity wherewith the Body A will be mov'd after the Concourse: which Celerity being taken away put of the whole respective Celerity which was before the Impulse, to wit, 12 - 4 = 8, there remains the Celerity of the Body B after the meeting, i Let therefore BN be of Four Parts, and EL of 8, and the Parallelograms ENHI, and ELK'M being compleated , and the Diagonals E-H and E.K. being drawn, the Bodies A and B in the fame time in which they haftened to the Meeting before, according to the Diagonals A E and B E will come after the Meeting, to the Points H and K, being reflected into the Diagonals E H, and EK; and the Motion of the Body A will be of 4x 3 == 12 Parts ; and the Motion of the Body  $B = 8 \times 1 = 8$  Parts, the difference of which Motions is Four Parts, which was alforthe difference of the Motions before the Meetings Wherefore dine. 2



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Wherefore in this Cafe the Quantity of Motion which is collected by taking the difference of the Motions made to the contrary Parts, is not changed by the Action of the two Bodies upon one another; and confequently in Bodies dafhing one upon another obliquely, this Rule holds good, as well as in Bodies which directly meet one another. Now from these Reflexions, there are wont to arise circular Motions of Bodies about their own Centres: But we shall have no occasion to confider these Cafes in what follows; and it would be too long to demonstrate all the Things hereto appertaining.

### A Lemma to the 25th Law.

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If two right Lines given in Polition, A C and B D, (Fig. 1. Plate 3.) be terminated at the given Points A and B, and have a given Proportion to one another; and the right Line C D, wherewith the indeterminate Points C and D are join'd, be divided in the given Proportion in K/; I fay, that the Point K will be placed in a right Line given in Polition

For let the right Lines A C and B D (if they be not Parallel) meet together in the Point E; and in BE let BG be taken in the fame Proportion to A E as is B D to A C: And let F D be equal to E G. Here E C will be to G D, that is, to E F, equal by Hypothefis to G D, as A C is to B D, and confequently will be in the Proportion given; therefore the Triangle E F C will be given in Species (to wit as having the Angle C E F, and the Proportion of the Sides about the fame Angle given) let C F be cut in L in that given Proportion, and fo there will be given in Species the Triangle E F L (by reafon

of the given Proportion of the Sides about the given Angle EFC ;) and therefore the Point L will always be plac'd in the Line EL given in Polition. Join LK; and becaule of FD which is given, as being equal to E G given, and the proportion of LK to FD which is given, the fame to wit as that of CK to CD, LK will be given. Let EH be taken equal to LK, and EL K H will be a Parallelogram, for LK is parallel to FD, and confequently to EH the protracted part of the fame Line, to which it is by the Hypothefis equal. Therefore the Point K is placed in HK a fide of a Parallelogram which is given in Polition. Q. E. D. But then, if the right Lines A C, BD, be parallel each to other, the Point of Concourse will be infinitely distant, that is none at all; and all the Lines EC, EL, HK, ED, will be parallel to one another. (fee Fig. 2. Flate 2.) In which Cafe the Lemma is thus demonstrated. Let the Points terminating the Lines A C, B D, which have a given Proportion, be join'd by the Lines A B, C D, and let these joining Lines be protracted to meet together in Q; through the Point K which divides the Line CD in the given Proportion, let HK be drawn parallel to AC and BD: I fay, that the Point K is placed in the right Line HK given in Polition. For wherefoever the Points C and D are taken in the right Lines A C, and B D, the Line joining these Points will tend to the Point Q, as in the Points c and d, and the joining Line c d will be divided in that given Proportion by the Line HK: For according to the Hypothelis, and in this Figure A c is to B d as A C is to B D; as also according to the Hypothesis and in this Figure ck is to cd as CK is to CD. It is manifest therefore

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fore in this Cafe, that the Point K is always placed in the right Line given in Polition.

Coroll. (1.) If two Points go forwards together with an uniform Motion in right Lines, and their diftance be divided in a given Proportion, the dividing Point will be placed in a right Line given in Polition ; and that Point as K will be mov'd uniformly in that right Line. For, because of the Uniform and Even Motion of both Points, the Lines of Motion which they defcribe at the fame time will always be in a given Proportion, to wit, in the Proportion of the Celerities which are on both Sides equable : From whence it is manifeft by the things already demonstrated, that the Point K will always be carried in the right Line HK. But that it is carried uniformly, and with an equable Motion, will be thus demonftrated : HK is always equal to EL, and EL increaseth in the same Proportion as the Lines EC and EF proportional to it, which Lines are also proportional according to what hath been already faid, to AC and BD, along which the Bodies are mov'd at the fame time. EC therefore is to EF, as AC to BD; from whence, fince those Lines by the Uniformity of the Motion do increase equally; E L also, and H K, which is proportional to the same, will alfo increase equally; or, which is the same thing, the Point K will be carried with an uniform and equable Motion along the Line HK. Q.E.D. And we may argue in like manner in the fecond Cafe where the Lines were suppos'd parallel. Nor is there need of more Words in fo plain a Matter. The truth of the Lemma will likewife be concluded concerning a folid Place by almost the like Demonstration, viz. if you conceive a Plane cutting the leaft diftance of the Lines G

Lines in the fame Proportion and perpendicular to the fame diffance; and if you imagine Lines let fall perpendicular to the faid Plane, the Demonstration will be as in this Proposition, if inftead of the Lines of Motion, supposed in different Planes, we use the Lines joining the Perpendiculars, which will be in the fame Plane.

Coroll. (2.) If both Points go forwards unto the fame Part, the dividing Point will alfo go forwards unto that Part: From whence in every Cafe, that dividing Point K will either reft or be mov'd uniformly in a right Line. If one of the Points be mov'd unto this Part, another to the contrary, the Point K will be mov'd more flowly unto one Part or the other, according as the Proportions of the greater Celerity, or of the diftance from the Point K fhall require : Or laftly, if those fhall be Proportions of Equality, and prevail on neither fide, the dividing Point will be mov'd to neither Side, but wholly reft.

(25.) The common Center of Gravity of a Syftem of Bodies doth not change its State either of Motion or Reft. from the Actions of the Bodies amongst themselves, (whether they be Attractions or Impulses;) and therefore the common Center of Gravity of all Bodies acting upon one another (Actions and Impediments, whether External or otherwise gotten being excluded) doth either reft. or is mov'd uniformly straight forwards, For if two Bodies or Points, as C D go forwards with an uniform Motion in the right Lines A C, B D, and their (see Fig. 1, 2. Plate 2.) distance CD be divided in a given Proportion; (as the Line always passing through the Centers of the mov'd Bodies is divided by K the common Center of Gravity of both in a given Proportion, to wit, that which is reciprocal to the Bodies) their common

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mon Center of Gravity K, will either reft or be mov'd uniformly in theLine KH. Therefore if any Number of Bodies be mov'd uniformly in right Lines, the common Center of any two of them either refteth or goeth forwards uniformly in a right Line ; because that the Line connecting the Centers of those Bodies which go forward uniformly in right Lines, is divided by the common Center of Gravity in a given Proportion; in like manner, the common Center of these two, and any third Body whatever, either refts or goes forwards uniformly in a right Line; because that the distance of the Center of Gravity of these two, and of the Center of the third, is divided thereby in a given Proportion, to wit, a reciprocal Proportion to the third Body, and the System of the two : For the common Center of Gravity of the two goes forwards uniformly in a right Line. and confequently is to be reckon'd as if it were the Center of a fingle Body. In the fame manner, the common Center of Gravity of these Three and any Fourth, either refts or goes forwards uniformly in a right Line; because the distance betwixt the Center of Gravity of the Three, and of a Fourth, is divided in a given Proportion; and lo on in infinitum. Therefore in a System of Bodies which are altogether free from Actions upon one another, and others also impress'd from without; and which confequently do either reft or are mov'd uniformly in feveral right Lines, the common Center of Gravity of all either refts or is mov'd uniformly straight forwards. Moreover, in a System of two Bodies acting upon one another, fince the Diftances of the Centers of both from the common Center of Gravity is reciprocally as the Bodies, the Relative Motions of the fame Bodies, of coming to that Center or de-G 2 parting

parting from the fame, (whether the one be from Attraction, or a Centripetal Force, or the other be from an Impulse or a Centrifugal Force) are equal betwixt themfelves, and the Velocities of Accels or Receis are reciprocally proportional to the Bodies, that is, directly proportional to the Diftances from the Center of Gravity of both. From whence, by those Actions the distance from the Center will be proportionally increas'd or diminish'd; and therefore that Center is neither advanced forwards nor retarded, nor fuffers any change in its own State as to Motion or Reft, from the equal Changes made to the contrary Parts, and confequently the Actions of these Bodies amongst themselves, whether they repel one another or attract, do not any ways alter the State of the common Center of Gravity. Now in a System of more Bodies, because the common Center of Gravity of any two acting mutually upon each other, doth not in any wife by reafon of that Action change its State; and the common Center of Gravity of the reft fuffers nothing therefrom; but the diftance of the Centers of these two is divided by the common Center of all the Bodies into Parts reciprocally proportional to the Sums total of the Bodies whole Center of Gravity they are, and confequently those two Centers keeping their State of Motion or Reft, the common Center of Gravity of all will also keep its State; from hence it is manifest, that that common Center of all never changeth its State as to Motion or Reft, becaufe of the Actions of two Bodies betwixt themselves. But in such a System of all, the Actions of all amongst themselves are either of two Bodies; in which Cafe the common Center of Gravity of the whole System is nothing

thing chang'd, as we have already fhewn; or compounded of the Actions which are betwixt Couples of Bodies; and therefore they will fuperinduce no change in the State of Motion or Reft to the common Center of Gravity. For if by the Action of A upon B, the State of the Center of Gravity is nothing chang'd, nor by the Action of C upon B; neither will the fame be diffurb'd by the conjunct Forces of C and A upon B. Wherefore, feeing that the common Center of Gravity, when Bodies do not act upon one another, either refts or goes forwards uniformly in fome ftraight Line; the fame will continue notwithstanding the Actions of Bodies one upon another, either always to reft or to go forwards in a right Line uniformly, unless it be moved. from this State by fome extrinsic Force impress'd upon the System. There is the fame Law therefore of a System of Bodies as to perfeverance in the State of Motion or Reft, that there is of one fingle Body. For the progressive Motion, whether of a folitary Body or of a Syftem of Bodies, ought always to be effimated from the Motion of the Center of Gravity.

Octob. 30, 1704.



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### LECT. VIII.

(26.) HE Motions of two Bodies included in a given Space, and partaking of the Motion thereof, are the fame amongft themfelves, whether that Space reft-

eth, or that the fame is mov'd uniformly ftraight forward without a Circular Motion.

For the difference of the Motions tending to the fame Part, and the Sums of those which tend to the contrary Parts, are the fame at the beginning in both Cafes (by the Hypothesis); and from these Sums or Differences arise the Congreffes or Shocks whereby the Bodies encounter one another, [to wit, of the Sums of the Motions tending to the contrary Parts, and the Differences of the fame when tending to the fame Part.] Therefore by Law 4th, the Effects of the Congresses will be equal in both Cafes, and therefore the Motions amongst themselves in one Cafe will remain equal to the Motions amongst themfelves in another. For the common and uniform Motion of the Space, and included Bodies which tends to the fame Part, will either by equally accelerating all, as in cafe they all tend to the fame Part with the Space it felf, or by adding fo much to one as it takes away from another, as in those which tend to the contrary Parts, make that the Forces at the Meetings of the Bodies will be in no wife changed. This fame thing is prov'd by a manifest Experiment; for all Mo-- tions 

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tions are in the fame manner in a Ship, whether it refteth, or be carried uniformly ftraight forwards.

(27.) If Bodies be mov'd in any wife amongft themfelves, and be preffed with equal accelerative Forces according to parallel Lines, they will all continue to be mov'd in the fame manner amongft themfelves, as if they were not preffed with thofe Forces. For thofe Forces will move all the Bodies equally as to Velocity, whilft they act according to the Quantities of the Bodies to be mov'd, and in parallel Lines; and therefore they will not change their Positions and Motions in respect of one another.

#### A Lemma to the Experiments following :

The Velocity of a pendulous Body in the loweft Point of any Circle, is always as the Chord of the Arch which is defcrib'd in the falling. (See Fig. 2. Plate 2.)

Let CAB be a right Angle, C or H a moveable Body hanging by the Thread CA or HA upon the Center A, which will fall down in the Arch CB or HB. I fay that the Velocity of the Body C, in the lowest Point B, is to the Velocity of the Body H in the fame Point, or rather the Velocity of the fame Body falling firft along the Arch C B, and afterwards the Arch HB, is as the Chord CB is to the Chord HB. For the Velocity of the Body, falling through the Arch C B, as we shall demonstrate by and by (Coroll. c. Prop. 6. following) is in the loweft Point B (that Velocity, to wit, wherewith the Body would go on to be mov'd in a right Line, which toucheth the Circle in B, if it should leave the Thread in B) is, I fay, the fame, as that which G 4

which the Body would have in the Point F, if it had fallen perpendicularly along CF. And by the fame, the Velocity of the Body which falls from H according to the Arch H B, is the fame as that which it would have in the Point K, if it fhould fall perpendicularly along H K : [the fame Celerity, to wit, being impress'd in Spaces betwixt parallel Planes, whether the Transit through those Planes be perpendicular, as in Bodies falling perpendicular to the Horizon; or whether it be oblique, as in pendulous Bodies describing circular Arches, as will be more fully made appear afterwards.]Therefore the Velocity of a Body descending along the Arch HCB, is to the Velocity of the fame defcending along the Arch H B, as the Velocity of a Body falling along CF, is to the Velocity of that which falls down HK. But (by Coroll. Prop. 4. beneath) The Velocity of a Body falling down in CF, is to the Velocity of a Body falling down in HK, in the fubduplicate Proportion of the Line CF to the Line HK; and (as will appear from Prop. 2. following) the Chord C B is to the Chord H B in the fame fubduplicate Proportion of CF to HK. From whence it follows, that the Velocity of the Body descending along the Arch CB, is to the Velocity of a Body falling down the Arch HB in the loweft Point B, as the Chord C B is to the Chord H B. O. E. D.

Corollary. From hence is to be corrected the Error of Hugens, or rather of his \* De vi Centrifuga. p. 426,427. portion of Velocity in the Point B, is the fame as that of the Lines themfelves C F, H K, when it is only the fubduplicate Proportion of the fame, as we have already

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ready demonstrated from the Principles of Hugens himself.

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A general Scholium. The truth of these Laws hath heretofore been prov'd by Sir Christopher Wren before the Royal Society by the Experiment of Pendulums; which thing the famous Mariotte hath taken upon him to fet forth and declare in an entire Book. But that this kind of Experiments may agree with the Theories, regard is to be had not only to the Elasticity of the Pendulous Bodies, but also to the Resistance of the Air. Let the Bodies A and B (fee Fig. 4. Plate 2.) hang by the parallel Threads A C and B D upon the Centers C and D : From these Centers, and at equal Intervals, let there be describ'd the Semi-Circles EAF, GBH, bisected respectively by the Radii CA and DB. Let the Body A be drawn unto any Point of the Arch E A•F as R, and the Body B being withdrawn, let it be let down from thence, and return after one Vibration, compounded of going and returning, to the Point V. Here R.V is the Retardation from the reliftance of the Air. Of this R V let ST be made a fourth part, placed in the middle; and let R Q be equal to QV; and thus ST will reprefent the retardatio nvery near which is in the Descent from S to A. For if in the Double both Afcent and Descent, the Retardation be R V, the Retardation in one Ascent or one Descent will be one fourth Part thereof. And feeing two Arches be greater, and two be lefs than the Arch QA, the resistance of the Air is to be taken neither in the greatest Arches, nor in the least, but in a mean betwixt them. From whence the fourth Part \$ T is neither to be placed at the higheft Part R, nor at the loweft V, but in a middle one which is betwixt both. Now let the Body B be reftor'd. into

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into its Place : Let the Body A fall from the Point S; and the Velocity thereof in the Point of Reflection A, will, without fensible Error, be the fame, as if it had fallen in a Vacuum from the Point T; the Body A by falling fomething higher, compensating the Resistance of the Air; and therefore, according to the just now demonftrated Lemma, let this Velocity of the Body in the Point A be represented by the Chord of the Arch T A. After the Reflexion, let the Body A come unto the Place s, and the Body B to the Place k, whether the Bodies be Elastic or not. Let the Body B be taken away, and the Place (u) be found, from which if the Body A was let down, and should after one entire Vibration return to the Point (r); (s t) may be a fourth Part of (r u) which (s t) is fituate also as before in the middle : And by the Chord of the Arch (t A) let the Velocity which the Body A hath after its Reflexion in the Point A be represented ; for (t) will be the true and correct Place unto which the Body A, the refiftance of the Air being taken away, ought to have ascended; and by the like Method will the Place (k) be to be corrected, that to wit, to which the Body B afcends, and the Place (1) to be found, that to wit, unto which the Body ought to have ascended in a Vacuum. By this means we may try all this fort of Experiments in like manner as if we were placed in a Vacuum. Then at last the Body A is to be drawn into the Chord T A, which reprefents its Velocity, that the Motion of it in the Point A just before the Reflexion may be had; and afterwards into the Chord (t A), that the Motion of the fame in the Point A prefently after the Reflexion may also be had : As in like manner the Body B is to be drawn into the Chord (B1) that the Motion of it prefently after

**G1** 

after Reflection may be had ; and by the like Method where two Bodies are let down together from divers Places, there are to be found the Motions of both, as well before as after the Reflection, and then those Motions are to be compar'd at length betwixt themfelves, and the Effects of the Reflexion to be collected. By experiencing the thing in this manner in Pendulums of ten Feet, and this in Bodies both equal and unequal, and by making that the Bodies should meet together from very great Intervals, as of Eight, Twelve, or Sixteen Feet, the Famous Sir Ifaac Newton always found, without an Error of Three Inches in the meafures, when the Bodies did directly meet each other, that the change of the Motion to the contrary Parts was equally in both Bodies, and confequently that Action and Re-action, according to the fifth Law were always equal. As if the Body A should fall upon the Body B, which is at reft, with Nine Parts of Motion, and feven Parts. being loft in the Conflict, fhould go forward after the Reflexion with two Parts only, the Body B would rebound with those Seven which A loft. If Bodies meet one another, A with Twelve Parts of Motion, and B with Six; and A returns with Two, B would return with Eight; there being made to wit, a fubduction of Fourteen Parts on both Sides. Let Twelve Parts be fubducted from the Motion of A, and there will remain nothing; let there be fubftracted other two Parts, and then there will be a Motion of Two Parts unto the contrary Part. And fo as to the Motion of the Body B of fix Parts, by fubducting 14 Parts, there will arise Eight Parts of Motion the contrary way. But if the Bodies should be carried to the fame Part, A more swiftly with fourteen Parts, and B more flowly with five

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five Parts, and after the Reflexion A fhould go forward with five Parts, B would go forwards with fourteen, there being made a Translation of Nine Parts from the Body A to the Body B, and fo in all other Cafes. The Quantity of Motion which was collected from the Sums of the confpiring Motions, and the difference of the contrary ones, was never found in the Tryals of the forefaid great Man to be changed by the Congress or Collision of Bodies. For the Error of one Inch or two in the Measures is to be attributed to the difficulty of performing every thing with due exactness. For it was not an eafy thing to let down the Pendulums just at one time, fo that the Bodies should dash one upon another in the lowest Place AB: then to note the Places s and k, to which the Bodies ascended after the Congress, was difficult : and in the Balls themfelves which were to be us'd, the unequal Denfity of the Parts, and the irregular Texture arising from other Causes must needs bring in fome fort of Errors. But further, left any fhould object, that the Rule, for the proving of which this Experiment was invented, doth prefuppose either that the Bodies are absolutely hard, or at least perfectly Elastic, of which fort there are none perhaps to be found in natural Compositions; we add, that the Experiments now defcrib'd do fucceed as well in foft Bodies as in Bodies Hard and Elaftic; these Experiments not depending at all upon the Condition of the Hardness or Elasticity of Bodies. For if the thing were to be tryed in Bodies not perfectly Hard or Elastic, the Re-flexion ought only to be diminish'd in a certain Proportion according to the Quantity of the Elaflic Force which is diminish'd. In the Theory of Wren and Hugens, Bodies abfolutely hard return from one another with the Relative Velocity of the

the Congress; but with the famous Wallis it is altogether to be faid, that this holds in Bodies per-. perfectly Elastic only; and it is to be afferted, that other Laws have place in Bodies not Elastic, whether foft or hard, that obtain not in Elaftic ones ; as is abundantly manifest from what has been above opened. And particularly those Bodies only which are perfectly Elaffic, do after mutual Collisions return from one another with the Velocity of the Congress, according to the 16th Law of Motion thereto belonging. In those which are imperfectly Elaffic, the Velocity of the Return is to be diminish'd together with the Elastic Force, and in the proportion of the diminution of the fame : becaufe that that Elaftic Force (unlefs where the Parts of the Bodies are hurt by the Meeting, or fuffer fome kind of Extenfion as under a Mallet) feems to be in it felf certain and determinate, and makes Bodies to return from one another with that Relative Velocity, which is in a given Proportion to the Relative Velocity before the Concourfe. Which thing Sir Ilaac Newton did thus experiment in Balls made of Wool, most straitly wound up and preffed together; first by letting down Pendulums, and measuring the Reflexion, he found the Quantity of the Elastic Force; then by this Force he computed the Reflexions which were to be expected in other Cafes of Congress, and the Experiments answer'd. The Balls always return'd from one another with a relative Velocity, which was to the Relative Velocity of the Concourfe, as the Number Five is to Nine. Balls of Steel were almost perfectly Elastic, for they return'd almost with the very Velocity of the Concourse; those of Cork with fomething less; but in Glass Balls the Proportion was of about Fifteen to Sixteen. And

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And by this means, the Fifth Law of Motion as to Stroaks and Reflexions was prov'd by Dr. Wallis's Theory, which plainly agrees with Experience. Sir Ilage Newton doth also fhew in this Place briefly that this Rule holds also in Attractions; to wit, that the Quantity of Motion, collected by gathering the Sum of the Motions made to the fame Part, and the difference of those which are made to the contrary Parts, is not changed by the Actions of Bodies amongst themselves ; whole reasoning in this Matter we confidered above under the Fifth Law, and confequently shall at prefent omit it, and come to the reft of his Obfervations belonging to the prefent Place. As therefore Bodies in Concourse and Reflexion are of the fame Force, whole Velocities are reciprocally as the implanted Force or the Bodies themfelves. as may be underftood from the Eight and Seventeenth Laws, and Hugens's Eighth Proposition ; fo in moving Mechanic Instruments, Agents are of the fame Force, and fustain each other by contrary Endeavours, whole Velocities, effimated according to the determination of their Forces, be reciprocally as their Forces. Thus Weights are of equal Force to the moving the Beam of a Pair of Scales, which in the Vibration of the Balance are reciprocally as their Velocities up and down; that is, the Weights if they alcend and descend perpendicularly, are of equal Force betwixt themfelves when they are reciprocally, as the Diftances of the Points on which they are hanged from the Axis of the Balance. But if being hinder'd by oblique Planes, or fome other Obstacles, they afcend and defcend obliquely, those Weights are Equipollent, which are as the perpendicular Afcent and Descent, and this because of the determination of Gravity downwards. Thus in a Wheel

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Wheel or Pulley, the Force of the Hand which directly draws the Rope, is to the Weight ascending either directly or obliquely, as is the Velocity of the perpendicular Afcent to the Velocity of the Hand drawing the Rope, the Hand will fuffain the Weight in Equilibrio. In Watches, and the like Instruments which are fram'd of little Wheels put together, if the contrary Forces for the furthering and hindring the Motion of the Wheels be reciprocally, as the Velocities of the Parts of the Wheels on which they are impress'd, they will fustain one another. The Force of a Screw to prefs a Body, is to the Force of the Hand turning round the Handle, as the circular Velocity of the Handle in that Part where it is prefs'd by the Hand to the progreffive Velocity of the Screw towards the prefied Body. The Force with which the Wedge lies upon the Two Parts of the cloven Wood, is to the Force of the Mallet upon the Wedge, as the Progress of the Wedge according to the determination of the Force impress'd upon it by the Mallet, is to the Velocity wherewith the Parts of the Wood give Place to the Wedge according to Lines perpendicular to the Forces of the Wedge. And the Reason is the same in all Machines. Their Efficacy and Ule confifting only in this, that by diminishing the Velocity we may increase the Horce, and on the contrary. From whence in all Kinds of fit Inftruments that fo much talk'd of Problem is folv'd, of moving a given Weight by any given force, or of overcoming any other given Refiftance by any given Force how fmall foever. For if Ma-chines be fo form'd, that the Velocities of the Agent and Reliftents are reciprocally as the Force, the Agent will fustain the Refistent, and overcome the fame with a disparity of Velocities :

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Velocities : Certainly, if the difparity of Velocities be fo great, that it will overcome even all Reliftance, which is wont to arife as well from the Attrition of Bodies contiguous, fliding one upon another, as from the Cohelion of Bodies continuous, and which are to be separated one from another, and the Weights of Bodies to be lifted up; all that Refiftance being overcome. the Force redounding will produce an Acceleration of Motion proportional to it felf, partly in the Parts of the Machine, partly in the refiding Body. But to treat of, and handle Mechanics a's it ought to 'be done, belongs not to our Purpole ; we by thele Things only fhew how far and wide this Rule extends, and how certain the Fifth Law of Motion above delivered is. For if the Action of the Agent be eftimated from its Force and Velocity conjunctly, and the Re-action of the Reliftent from the Velocities of each of its Parts, and their Force in relifting, ariling from their Attrition, Cohefion, Weight and Acceleration ; Action and Re-action will be in all Kinds of Instruments equal one to the other; and fo far forth as the Action is propagated by the Inftrmment, and at length impress'd upon every resisting Body, its last Determination will always be contrary to the determination of the Relaction.

Corollary, From these two Laws of Motion now fufficiently explain'd and prov'd, the gross Errors of Des Cartes about the same do manifestly appear; whose Laws of Motions, are to far from agreeing every where with the true Laws of the same, that they are rather found every where to disagree with them. And consequently it is no wonder, if he in like manner err'd in the rest of the Phenomena of Nature. The Laws of Motion being now dispatch'd, we come unto the Propositions. November 6th, 1704. LECT. \*\*\* \*\*\* \*

#### LECT. IX.

#### **PROPOSITIONS.**

I. THE laft Proportion of the Tangent Subtenfe or Chord to the Curve Arch which belongeth to them, is the proportion of Equality; that is, the Tangent, Arch, and Chord, where the leaft Arch of all, or that which vanifh-

eth away is taken, do at length end in one and the fame Line. And the fame thing is to be underftood of the Sine. Let A b be the Arch of a Circle or fome other Curve, which is as little as may be ; let A f be the Tangent thereof, and A b the Subtenfe : I would know what is the Proportion of these Lines one to another if they be taken as near as may be to the Point A, or if the Point b doth as it were coincide with the Point A, and I fay, that the Proportion of the Arch, whether to the Tangent above, or to the Subtenfe below, is the Proportion of Equality. For from the Nature of Curves it is manifest, that all the difference betwixt the Tangent and Subtenfe of any Arch whatever doth arife from the length of the intermediate Arch, and that the difference is always fo much the greater, as the Arch which is taken is the greater, and fo much the less by how much the Arch which is taken is the lefs; from whence it follows, that in an Arch the least that may be the difference will be the least that may be, and in an Arch infinitely fmall, fuch 25

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as we now intend, the Difference will be infinite-Iy fmall, or none at all. And if the Difference betwixt the Tangent and the Subtenfe be none at all, much more shall the Difference betwixt the Tangent and the intermediate Arch be none at all, or the Difference betwixt the Subtenfe and that Arch; fince that Arch is every where of an intermediate Length Betwikt the Tangent and the Subtenfe. And this Equality of the leaft Tangents, Arches and Subtenfes, and of Sines alfo, is a Thing which Geometricians have always fuppofed and acknowledged ; whilft they have confidered the Perimeters of Curves, as being innumerable Sides of Polygons, and as arising from the Coalescence of the inscrib'd and circumscrib'd Figures, the Difference betwixt them vanishing awav.

Corollary, If therefore it can be demonstrated, that d b D b (Fig. 5. Plate 2.) the Subtenses of the Angles of Contact, are always berwixt themselves in the duplicate Proportion of the Subtenses A b, A B, as will prefently be done, it will from thence follow, that the same vanishing Subtenses are also in the duplicate Proportion of their contentinous Arches A b, A B, or of the Sines c b, C B; because the Subtense A b doth in the foregoing Case fall in, and become the same with the Arch A b or its Sine c b, and so doth at length the fubtense A B with the Arch A B, or its Sine C B-II. The Subtenses of the Angles of Contact

in Circles, are always in the duplicate Proportion of the Subtenfes of the conterminous Arches.

Let there be in the lame Figure any two Arches as A B and A b; and D B, d b (equal to A C and A c, the verfed Sines of the fame Arches) the Subtenfes of the Angle of Contact. To these Subtenfes (by Book the 3d, Prop. 31. of the Elements)

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Elements) G B and G b drawn from the Point G will be perpendicular; let the Rectangles ADBC, and A d b c be compleated. This done, the Square of A B (by VI. 8. of the Elements with VI. 17. of the fame) is equal to the Rectangle of A G into A C or D B; and in like manner, the Square of A b is equal to the Rectangle under A G and A c or d b. And confequently, the Proportion of the Square of A B to the Square of A b, is the fame as that of the Rectangle under A G and D B, fo that under A G and d b; that is, (by VI. 1. Elements) the fame as of the Line D B to the Line d b. Q. E. D.

Corollary: Therefore any Subtenfe whatever of the Angle of Contact, as D B or d b is equal to the Square of the Chord applied to the Diameter of the Circle. For as A G is to A B, to is A B to A C or D B: From whence by the Golden Rule, B D =  $\frac{A B \times A B}{A G}$  or  $\frac{A B q}{A G}$ . And in like manner, A G is to A b as A b to A c or d b; from whence d b =  $\frac{A b q}{A G}$  Q.E.D.

Coroll. (2.) In the leaft Segments of perspective Glaffes, the Altitudes or Axes of the Segments AC, and Ac, are to be reckon'd to have the fame Proportion betwixt, themfelves, which the Squares of the Bases or Apertures E b and R B Or. have. (see Fig. 5. Plate 3.) For we have fhew'd that A C and A c have the fame, Proportion as the Squares of the Subtenfes ;, and feeing in very fmall Arches the Subtenfes or Sines, or their Doubles R B and E B are almost in the fame Proportion amongst themselves; it follows, that the Altitudes also A C and A c have almost the fame Proportion as the Squares of the double Sines R B and E b, that is, of the Apertures. Q. E. D. Coroll H 2

Coroll (2.) In very small Angles, the Excels of the Secants above the Radius, are also very near as the Squares of the Subtenses, or Sines, or Tangents, or even of the Arches. For those Excesses (fee Fig. 5. Plate 2) b f and BF, do in that Cafe coincide with the Subtenfes of the Angle of Contact b d and B D, and confequently have the fame proportion amongst themselves as they. Thus you may fee in the Tables of Secants, that the Radius of the Circle being put to be of . 10000000 of equal Parts, the Excels of the Secant of two first Minutes is of two Parts, and that of the Secant of Four first Minutes is of Eight Parts ; from whence the difference of the former Secant, and the Radius, is fourfold of the difference of the latter Secant to the double Arch, and the Radius ; that is, those Differences are betwixt themselves as the Squares of the Arches; and fo of the reft.

Coroll. (4.) The Nascent or Evanescent Subtenfes of the Angle of Contact, are in the duplicate proportion of the conterminous Arches. For they are every where, by what hath been demonstra-ted before, in the duplicate Proportion of the Chords. But feeing the Chords do at last end in the Arches, that is, in Diftances 'infinitely fmall do coincide with them, and are equal to them, as we demonstrated above, those Subtenses in like manner will in the prefent Cafe be in the duplicate Proportion of the Arches.

Coroll. (s.) From whence also in the fame Cafe, according to the first Corollary of this Proposition, the vanishing Subtense of the Angle of Contact, will be equal to the Square of the Arch it felf, applied to the Diameter of the Circle.

Coroll. (6.) Hence is gathered that Noble and Fundamental Theorem of Sir Ilaac Newton, and

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of Mr. Hugens also; to wit, that in the circular Motion of a Body, the Centripetal Forces, or the Gravities toward the Center are every where, as the Squares of the Velocities of the Arches, defcrib'd at the fame time, applied to the Diameters or Radii of the Circles. For (in Fig. 6. Plate 2.) let the revolving Bodies B and b defcribe the Circumference of the Circles BD and b d, and in the fame given Time let them describe the infinitely fmall Arches B D and b d; because by their Innate Force alone they would describe the Tangents B C, b c equal to these Arches; by the first Law of Motion it is manifest, that there is some centripetal Force which perpetually draws back the Bodies from the Tangents to the Circumferences of the Circles; and confequently, these are one to another in the first Proportion of the Nascent Lines C D and c d; that is, as  $\frac{B D q}{BG}$  is to  $\frac{b d q}{bg}$ ; or, by taking half the Divisors, as  $\frac{BDq}{BS}$ 

to  $\frac{b d q}{b s}$  and because the Velocities are in the Proportion of the Arches describ'd at the fame time, those Forces will be as the Squares of those Velocities applyed to the Radii of the Circles. But if the Circles be equal betwixt themfelves, then by reason of the given Diameters those Forces will be as the Squares themselves of the Arches defcrib'd at the fame time, or of the Velocities; as we will fhew more fully afterwards.

Coroll. (7.) By means of the foregoing Corollary, we gather the Proportion of the Centripetal Force to any known Force, as that of Gravity. For fince that Force in the time that the Body describes the Line BC, or an Arch equal to it, H 2 impells

impells the fame along the Line CD; which in the beginning of Motion is equal to the Square of that Arch B D, applied to the Diameter of the Circle. And fince every Body by an uniformly accelerated Motion, or the fame Force continued towards the fame Part always, doth defcribe Spaces in the duplicate Proportion of the Times, as will prefently be shewed (under Prop. 4.) that Force in which time the revolving Body doth defcribe any given Arch, will make the fame Body, going right forwards, describe a Space equal to the Square of that Arch applied to the Diameter ; and confequently, is to the Force of Gravity as that Space is to the Space which an heavy Body in falling doth defcribe in the fame time. As for Example, from the Experiments of Pendulums, and other ways it is manifeft, that all Bodies whatever describe in a Vacuum 16,14 English Feet in one Second of Time whilft they are falling by the Force of Gravity : I therefore would know what Proportion the Centripetal Force, whereby the Moon is held in its Orbit bears to the Force of Gravity with us ? For this purpose the Square of the Arch of the Moon's Orbit, which is defcribed in one Second of Time, is to be divided by the Diameter of the Orbit; and fo we shall find the Line which the Moon (if the circular Motion thereof were destroyed, and it descended as an heavy Body towards the Earth) would describe in the fame time. The mean Distance of the Moon from the Center of the Earth is about Sixty-Times the Earth's femi-Diameter. or of about English Feet in Number 1257696000. The Circumference therefore of its Orbit, if reduced to a Circular, will be of about 7897834380 Feet; which Periphery, fince the Moon describes it in the Space of a periodical Month, or 27 Days, 7 Hours, 43 Minutes, that is, in 2360580 Seconds ; let

let the Circumference 7897824280 be divided by 2260580 the Seconds which belong to the fame, and the Quotient 2346 will give the Length of the Arch described by the Moon in a second of Time in English Feet, the Square whereof 11128976 being divided by the Diameter 2515392000 will give 100442 Parts of an English Foot to be defcribed in one Second by the Moon falling, and in one Minute 16[14 Feet, or thereabouts; therefore the Centripetal Force, or Gravity of the Moon, is to the Centripetal Force of Bodies with us upon the Surface of the Earth, as 443 Hundred Thousandth Parts of one Foot is to 16|14 Feet, that is almost as 1 to 2600. And confequently, the Force of Gravitation towards the Earth at the diffance of the Moon is only the 2600th Part of the Force of Gravity with us.

III. The Velocities of a Body accelerated by any uniform urging Force whatever, are betwixt themfelves, as the Times are wherein that uniform Force is impress'd; that is, in Double the Time Double, in Treble the Time Treble of it felf, and in Four times the Time a Quadruple Velocity. For if the Accelerating Force be Equable and Uniform, which is here fuppos'd; and the Body confequently, whether it rested at the First, or was mov'd with any Celerity, doth receive equal Degrees of Velocity, and an equal Increase in equal Time; it is manifest, that the Velocity of the Body is exactly proportional to the Time. For if in every given Particle of Time that Force doth generate a certain Velocity, it will be able in the next equal small Portion of Time to generate a Velocity like and equal to the former; and fo likewife in the Third, and Fourth, and Fifth, Oc. Particle of Time, and so in infinitum. From whence the entire Velocity will every where be às

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as the Space of Time, in which that Accelerating Force is impressid on the Body. Q. E. D.

Corollary. Seeing therefore it is manifeft by Experiments, that all Bodies whatever being accelerated by the Force of Gravity, do receive increafes of Velocity every where proportional to the Times; it is manifest, that the Force of Gravity doth act uniformly, and doth affect Bodies most fwiftly defcending, as well as quiefcent. From whence it appears, that the Gravity of Bodies is to be afcrib'd to no Preffure of the Air, or Impulse of either, nor to the Mechanical Endeavour of any Mater towards Motion. For all these Impulses or Endeavours would most of all affect a Body at Reft, and by how much the more fwiftly the Body should be mov'd, they would so much the lefs continually urge it, until at length, the Celerity produc'd becoming equal to the Force of the Caufe which generates it, all the Impulse would ceafe, and no Acceleration of Motion follow thence forward.

#### Some Lemmata to the 4th Prop.

(1.) Odd Numbers being added to themfelves continually do make all the Square Numbers. Thus One is the first of odd Numbers, and the first alfo of Square Numbers. But if the Number 2, which is the Second odd Number be added to One, there is made the Number Four, the Second of Square Numbers; and if the Number Five be added to Four there is made the Number Nine, the Third Square Number, and fo on *in infinitum*. We shall bring Two Demonstrations of this Lem-

Pract. Arithm. Book V. Chap. 1. Theor. 7. ma, one out of Tacquet, the other out of our own Store. Tacquet shews the Thing thus. In the natural

tural Progression, saith he of odd Numbers, r, 2, 5, 7, &c. the total Sum is equal to the Square of the Number of the Terms. For according to the Nature of Arithmetical Progression, the Sum of all the Terms is equal to the Product of Half the Sum of the Extremes, drawn into the Number of the Terms : But half the Sum of the Extremes of an Arithmetical Progression of odd Numbers. beginning with Unity, is equal to the Number of the Terms, (for it goeth on from Unity by Two's fuperadded, whilst the Number of the Terms increafeth only by Ones) and confequently, that . Product is equal to the Square of the Number of the Terms, and therefore the Sum Total of odd Numbers, beginning with Unity, is equal to the Square of the Number of the Terms, Q. E. D. We demonstrate it thus: Let (a c) or (a b) (fee Fig. 7. Plate 2.) be Unity, and (a d) the Square of Unity; I fay, that the Addition of the odd Numbers, 3, 5, 7, 9, Oc. is necessary to make the Squares (a h, a n, a s, a z, a A) of all Num-bers proceeding from Unity; for to the making the Second Square, or the Square of the Number Two, there are Three other Squares of Unity to be added to (a d) to wit, the Two lateral Squares (c d and d f) and the Diagonal Vertical Square (i g.)And then through all the reft of the Terms, the Number of the Squares to be added is always to be increas'd by Two for the making up of the reft of the Squares; to wit, three Squares (ki, 1 h, and g g) corresponding to the Three which were added before are first to be added, then another Square (h p) because that the Square, added by the Side of the Diagonal, doth always require a Pair of corresponding Squares to be fuper-added, to which at length is to be added another Diagcnal Square (m o). And thus it is every where (the Number

Number of the Squares to be added always, exceeding the former by Two,) that the Squares (a d, a h, a n, a s,  $\mathcal{O}$  c.) from Unity may be compleated. From whence it plainly appears, that the continual Addition of uneven Nambers begets all Square Numbers, Q. E. D. But he that will be content with an Induction, carried on without End, may fafely enough pafs By this way of demonstrating the Thing; howbeit it is indeed fo eafy, that it will not require much Attention of Mind to comprehend it.

Lemma (2.) If a Body doth in a given time depart gradually and uniformly from Reft, and by that means describe a certain Line ; the fame Body in the fame given time will, from the laft Celerity acquired, if it be uniformly continued, describe a Line double to the former. For fince the Body in departing from Reft acquired a certain Degree of Velocity by equal increases, the Line describ'd by the fame, will be to be diftinguish'd into innumerable Lines greater each to the former gradually : And if those little Lines gradually in-creasing, were disposed not Length-wife but orderly at the Sides, they would compole a certain Triangle, (a b c) or at least, according to Cavallier's Method of Indivisibles, are to be reckon'd to compose one ; where the Vertical Point of the Triangle, to wit (a) is the Point of Reft, and the Bafe (c b) defigns the laft (fee Fig. 8. Plate 2.) Line of Motion, and the reft of the Parallel little Lines, the Lines of the diverse Velocity which the Body had paffed through. Now if we had put the greatest Line (1 b) to have been measured in the same full time, or had dispos'd from the Point (a) to the Base (1 b) so many Times equal to the greateft, as we had before dispos'd Lines gradually increasing, we had compos'd

pos'd a Parallelogram, double to the former Triangle, (I. 41. of the Elements.) And confequently, the uniform Motion, gradually increas'd from the Point of Reft, is in the given time double to the former, Q. E. D.

IV. The Lines which Bodies by any urging uniform Force do describe, are in the duplicate proportion of the Times, i. e. if the Times be Seconds, One, Two, Three, Four, Five, &c. the whole Lines describ'd will be amongst themselves, as One, Four, Nine, Sixteen, Twenty-five, Oc. which are the Squares of the former. For if any Body whatever, by any urging uniform Force whatever, shall as it falls describe in the least Portion of Time, as one Second, fome Line; in the fecond equal Portion of Time it will defcribe a Line equal to the former, by reason of the continuation of a Force equal to the former ; and because of the gradual Acquisition, and Increase of the Velocity of Motion, it shall by Lemma the Second describe a Line double to the former; therefore from both Caufes conjoin'd it will now describe a Line treble to the former. But in the Third Particle of Time, by reason of the Force, of Gravity still acting, a Line equal to the former will be describ'd; and because of the Velocity of the former, which was double to B, continued in the equal Time a Line will be describ'd double to the former, that is, guadruple of the first, and so from the Forces conjoin'd a Line will be defcrib'd five-fold of the First ; and fo forwards from the continual Impression of Gravity a Line equal to the first will always be to be added, and then another Line equal to the first by reason that the Velocity is continually increas'd by one Part; and confequently, two Parts or Lines equal to the first are every time to be added; and confeqently, the

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the entire Lines defcrib'd in every fucceffive Particle of Time will be to be defign'd perpetually by odd Numbers. Seeing therefore (by Lemma 1.) the odd Numbers added one to another do orderly make all the Square Numbers, the Lines of thefe Moments added together will neceffarily make the entire Lines of the Moments to exceed the latter the foregoing in a duplicate Proportion, or in the Proportion of a Square Number to a Square Number. Thus if in one Second of Time Bodies be carried downwards by the Force of Gravity about Sixteen English Feet, as is manifeft from Experience; they will in two Seconds be carried Sixty-four Feet, and in three one Hundred Forty-four Feet, or thereabouts.

Or according to Galilaus in his Systema Cosmicum, we may demonstrate the Proposition thus. (Fig. 8. Plate 2.) Let equal Times be reprefented by equal Lines a b, b c, c d, d e, and the Velocity in the end of the first time bg, by the Line (b 1). Seeing then, that Velocity which the falling Body hath in that Place, was acquir'd not together and at once, but gradually, and in a certain Space of time, reprefented by the whole Line(a b) from the continual and uniform accelerating Force; therefore it is neceffary, that it fhould have had all the leffer Degrees of Velocity before it got that Velocity; from whence those former Degrees of Velocity will be reprefented by leffer Lines drawn from the Parts of the Time ab parallel to the Line; and feeing the

**Prop. 3.** foregoing. Velocity doth increase uniformly with the Times, those Lines according to the Method of indivisibles will constitute and compose the Triangle (a b 1). Therefore the whole Line which shall be described from all those Velocities join'd together,

ther, will be proportional to the Aggregate of all those Lines, that is, to the Triangle (a b 1); and will rightly be represented by that Triangle. But in the fecond time, when the Body shall now have acquir'd a Velocity proportional to the Line (b 1) and reprefented by the fame; with that only Velocity continued, it will describe a Line double to the former, and confequently to be reprefented by the Parallelogram (a b 1 h) or (b 1 kc) double to the Triangle (a b 1); and over and above, by the new Velocity, arifing as before. from the perpetual and uniform Incitation of the fame Force, a Line will be defcribed equal to the first Line ; therefore if you add both Forces together, in the fecond Time, the Line describ'd will be treble to the former, and to be reprefented by the Trapezium; and the Sum of the Lines describ'd in the first and second Time, will be to the Line described in the first time alone as the Triangle (a c 2) is to the Triangle (a b 1); that is, in the duplicate Proportion of the homologous Sides (a c) and (a b) which reprefent the Times; or as the Squares of the Times themfelves. In like manner, in the third Time the Body with the Celerity hitherto acquir'd, or the mere permanence of the Motion now got, will defcribe a Line to be represented by the Parallelogram (c 2 n d); and by the new added Force arifing from Gravity still, and continually uniformly inciting, will describe a Line to be reprefented by the Triangle (2 n 2.) From whence the Line describ'd in the third Time will be fivefold of the First, and to be represented by the Trapezium (2 c d 2); and the Sum of the Lines in the first, second, and third Times, will be to the Line described in the first time only, as the Triangle (a d 2) is to the Triangle (a b 1) or as the

the Squares of the Times (a d); and (a b) and fo in infinitum.

Corollary, Since according to what hath been before demonstrated, Celerity is every where proportional to the time, and feeing the Lines de-Icribed by Bodies falling down be in the duplicate Proportion of the Times, the fame Lines will alfo be in the duplicate Proportion of the Celerities, or as the Squares of the Velocities. As for Example, if the Velocities laft acquird of two Bodies falling to the Earth, be one to the other as the Number Two is to One, the Heights of the Fall shall be betwixt themselves as Four is to One. If the Velocity of one Body be treble to the Velocity of the other, the Height of the Descent of the same shall be Ninefold of the Height of the Descent of the other. And lo on ad infinitum.

Nov. 12, 1704.

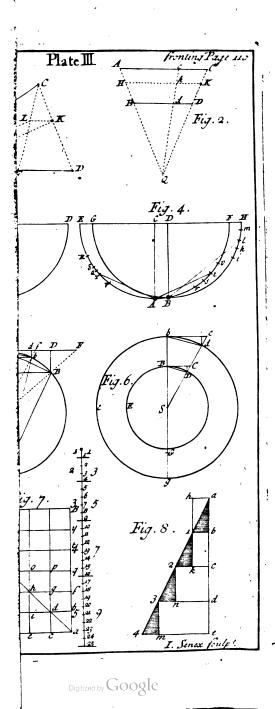
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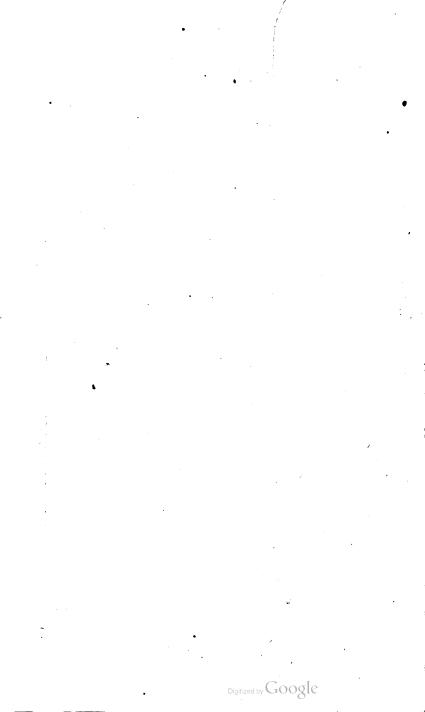
### L е с т. X.

F a Body fhall begin to tend upwards with that fame Velocity which it had acquir'd in the End of its Defcent, it will afcend to the fame Altitude in the fame Time, from

whence it before defcended, and shall equally lofe its Velocity in equal Times.

For by Force of what was demonstrated in the last Proposition it appears, that the Velocity once acquird, as (3.d) will always defcribe





fcribe an equal Parallelogram, whether the Body afcend or descend : But seeing the new Force of Gravity in the Descent increases by the Triangle (2 m 4 3) and in the Ascent diminisheth the same by an equal Triangle; it is manifest that the Trapezium now to be described in the Ascent will be equal to the Parallelogram before describ'd in the Descent, to wit the Trapezium 3 2 cd; and fo on From whence the Lines defcrib'd which rare proportional to these Trapeziums, and the Velocities proportional to the Bales of the Trapeziums, will every where be the fame in the Afcent that they were in the Descent ; until at length the Body reacheth to the laft Point of the Afcent in the fame time that it had descended from it.

VI. The Celerities of heavy Bodies acquir'd by -defcending upon divers Inclinations of Planes will -be equal, if the Elevations of the Planes or their perpendicular Altitudes be equal.

Let E G be a Line perpendicular to the Horizon, and E F a Line inclin'd to the Horizon (Fig. 1. Plane 4.) in any Angle whatever, and let G A be perpendicular to E F. I fay, that any heavy Body whatever will acquire the fame Velocity in defcending along the inclined Line EF, which it would acquire in the Line EG by the perpendicular Fall. For from what

was demonstrated before, the Force of Gravity in the oblique Plane

E F, is to the Force of Gravity in the perpendicular E G, as A B is to A C, or, on account of the likeness of the Triangles A C B, E F G, as E G to E F; or because the Triangle E G A is like to them, as E A is to E G. From whence, by reason of the divers Forces, the Motion and Velocity of the Body descending along E A

E A in the inclined Plane, will be to the Motion and Velocity of the Body defcending along E G, the Time of the perpendicular Defcent being given, as E A is to E G, or as E G is to E F; and the Velocity of the Body defcending along E A, unto the Velocity of the fame defcending along E F, will be in the fubduplicate Proportion

VI. 8. Elements.

of EA to EF, that is, in the Proportion of EA unto EG. Therefore the Velocity of the Body in

the Point A, is to the Velocity of the Body in the Point F, and to the Velocity of the fame defcending perpendicularly in the Point G, in the fame Proportion, to wit, that of the Line E A to the Line E G, or that of the Line E G to the Line E F. From whence it appears, that those Velocities are equal one to the other. Q. E. D.

Coroll. (1.) Whilft a Body falling perpendicularly defcribes the Line E G, another falling obliquely will defcribe the Line E A, determined by the Perpendicular G A.

Coroll. (2.) The Time of the perpendicular Fall is to the time of the oblique Fall, in the fubduplicate Proportion of the Line E A to the Line E F; or as the Line E A is to the Line E G, that is, in the proportion of the perpendicular Altitude E G to the oblique Line E F. From whence, by how much the Velocity is diminish'd on account of the diminution of the Force, it is increas'd, by reason of the increase of the time; fo that in the same perpendicular Aktitude there always remain the same Velocity, whatfoever may be the obliquity of the inclined Plane to the Horizon.

Coroll. (2.) The Times of Defcents upon Planes diverfly inclin'd to the Horizon, but whole Elevation or perpendicular Altitude is the fame,

fame, and betwixt themfelves as the Lengths of the Planes. For the time of Defcent by EF is to the cline of Defcent by EG, (Fig. 1. Plate 4.) according to what hath been already demonstrated, as EF is to EG; and the time of Defcent thro? EG is to the time of Defcent thro? EH as EG to EH; from whence by equality of Proportion, the time of Defcent through EF will be to the Time of Defcent thro? EG; as EF is to EH, Q.E.D.

Coroll. (4.) If a Body defcend from the fame perpendicular Altitude, with a continued Motion through how many foever and whatfoever contiguous Planes, as EI, IK, KL, howfoever inclin'd, it will always acquire the fame Velocity in the Endy that, to wit, which it would have acquired by falling perpendicularly from the fame Heighth. For by the Dermination of Mr. Hugens, the fame will be the Velocity, according to what hath been already demonstrated, of the Bodv (Fig. 1. Plate 4.) falling to the Point I, whether it fall along EI or MI; from whence the Velocity will also be the fame in its going along IK, as in falling along NK; from whence the Velocity will be the fame at the Point K, whether the Fail were through EI or IK or along MK, or even NK; from whence will follow that there will be the fame Velocity in falling along K L at the Point L, which would be the Delcent were abcording to one fingle Plane NL, for two Planes MK and KL, or even three Planes, EI, IK, KL ; the fame to wit, according to what hath been already demonstrated, which the Body falling perpendicularly could acquire at the Point G. Q. E. D.

Coroll. (5.): Hence it is allo manifelt, according to the Determination of the fame Hugens, that a Body descending along the Circumference of a Circle, or a Cycloid, or any Curve Line whatever, the fame

fame Velocity will always be acquir'd, if it defcend from an equal Altitude; and that that Velocity will be fo great as the Body ought to acquire by a perpendicular Fall from the fame Altitude. For Curve Lines are as it were, compounded of many innumerable right ones; and fince the Proposition is true in any rectilinear Perimeters whatever, and how many foever, it will be true likewife where they are in Number infinite, that is, where they end in Curve Lines. Q. E. D.

Coroll. (6.) Hence it also appears, that if an heavy Body's Motion be turned from a Descent upwards, it will afcend unto the fame Heighth from whence it came, along whatfoever Plane contiguous Surfaces, and in what fort foever inclin'd it shall be carried. For as before, (Prop. g. Fig. 1, Plate 4.) the fame will be the Velocity in any Point K and I. whether the heavy Body defcend or whether it alcend ; from whence certainly the fame will be the Limit of the afcending or defcending Velocity, the fame the Term of it at the Point E. From whence alfo, if there be an infinice Multitude of Planes ; that is, if the Surface be Curve, the Body will arife along the Curve Line also unto the Height from whence it came, and no higher.

Coroll. (7.) If a Body falls perpendicularly, or defcends along any Surface whatever; and then fhall, from the Force acquir'd by the Defcent be carried upwards along any other Surface; it will have in Afcending and Defcending the fame Velocity in Points of equal Heighth. And if the Plane or Surface of the Afcent be like and equal to the Surface of the Defcent, it will afcend in equal time in which it defcended. These Things do fo clearly follow from the Things already demonftrated, that there needs no more Words about them. Lemma's

#### Lemmata to the 7th Proposition.

(1.) If a Curve Line be of that Nature that it will every where fustain the Force of Gravity in proportion to its length; fo that by how much the Part of the Line to be describ'd is the greater, by fo much will the accelerating Force be the greater, and altogether in the fame Proportion; and by how much the part of the Line which is to be defcrib'd is the lefs, by fo much the lefs will the accelerating Forces be, and this in the fame Proportion ; the Times of Descent along such a Curve, whether the Arches defcrib'd be greater or lefs, will always be equal to one another. For the Velocity in the given Time is as the moving Force; if therefore the Line to be describ'd be alfo as the fame moving Force, it will be likewife as the Velocity; but if the Velocity of the Motion be every where as the Line to be described, it is manifest, that any Line whatever, whether great or fmall, ought to be defcrib'd in the fame time. But that this is the Nature and Property of the Cycloid is what comes to be demonstrated in what follows.

Lemma (2.) Let DAC (Fig. 2. Plate 4.) be a Semi-cycloid, DFA half of that Circle which produced the Cycloid; and from any Point in the Cycloid as B, let there be drawn the Line BE parallel to the Bafe DC, meeting the Semi-Circle DFA in E; then let the Chord AE be drawn, and from the Point B in the Cycloid the Line BL, parallel to the Chord AE; the Line BL will be a Tangent to the Cycloid in the Point B.

Lemma (2.) And the Arch of the Cycloid A B will be equal to the double Chord A E. These

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two

two Lemmata are manifest from the Elements of

See Wallis's Works Vol. I. P. 533, Grc. VII. In a Cycloid inverted, whole Axis is ere-Ated perpendicular, the Times of the Defcent wherein a Body let down from any Point whatever in it, comes to the loweft Point A, always equal betwixt themfelves.

Let BA and OA be any Arches whatever in the Cycloid, and BL and ON the Tangents in the Points B and O; and let E A and FA be Chords of the Semicircle D F A or parallel to the faid Tangents, by Lemma the fecond : let. A F be produced to the Point K. There are therefore by the third Lemma, Lines to be defcribed by the Body placed in one Cafe at B, in the other at O. which are betwixt themselves as the Chord E A is to the Chord FA; but the Force in the direction of the Tangent BL, or the Chord FA parallel thereto, is in the fame Proportion to the Force in the direction of the Tangent ON, or AF which is parallel thereto. For (Fig. 2. Plate 4.) as the Square of EA is to the Square of FA, fo is the verfed Sine EP to the verfed Sine FP; or fo is KM to FP; or fo is KA to FA. Therefore the Chord AE is a Geometrical Mean proportional betwixt the Chord AF and the Line AK; and confequently AF, AE, AK .... But the Force of Gravity, according to what hath been demon-Corol. 2. to ftrated, which is in the Plane A'E, is Lan 23.

A F, as A K is to A E; that is, as the Chord A E is to the Chord A F; and thus every where. But the Line to be defcribed is as the fame A E is to the fame A F; and confequently the accelerating

lerating Force is every where in the fame Proportion as the Line which is to be defcribed, and therefore the Times of the Defcent are every where equal. Q. E. D.

Coroll. (1.) If therefore we form other entire Semi-Cycloids QRT, QSC like and equal to the former AT and AC, the Vertices whereof touch the Bafe of the other at the Points T and C; and the heavy Body V hangs from the Center Q upon the Thread QRV, which is equal to QDA, or the double of DA; and be mov'd betwixt those Semi-Cycloids QRT, QSC, the pendulous heavy Body will, from the Evolution of the Thread defcribe the entire primary Cycloid, as is manifest from the Properties of the Cycloid, and will perform the Vibrations of what Amplitude foever, even to the greatest of them TAC, in the fame times exactly, and fo that the Center of the Ofcillation shall always be in the Curve Line it felf TAC.

Coroll. (2.) Seeing all Vibrations whatever in a Cycloid are always in equal Times, and feeing the leaft Vibrations in the leaft Arch of the Circle, the Radius whereof is Q A, and in the leaft Arch of the Cycloid T A C, by reason of the manifest Coincidence in this Case of the Arch of the Circle and of the Cycloid in the lowest Point are the fame; it is manifest that the time of every Vibration in the Cycloid is equal to the time of the leaft Vibration in a Circle, the Radius whereof is double to the Diameter of the Circle which produced the Cycloid.

Coroll. (3.) By reason also of the same Coincidence of the least Arches of the Circle and Cycloid in the lowest Point, the Vibrations in the Circle will be so much the more in equal Times, by how much the describ'd Arches are the less; I 3

fo that in very fmall Arches they may very well be reckon'd to be in equal times.

Coroll.(4.) Therefore in Pendulum Clocks which have longer Strings or Wires for the pendulous Bodies to fwing by ; the Times of the Vibrations of the leffer Arches which are defcribed, come nearer to an Equality, than they do where the Strings or Wires are fhorter; and confequently, the former Clocks are to be preferr'd far before the latter.

Coroll (s.) The Times of Vibrations in divers Cycloids are in the fubduplicate Proportion of the Cycloids or Radii QA; or the Lengths of the Pendulums are in the duplicate Proportion of the Times; this will eafily appear from what was demonstrated (Prop. 4.) before, as applied to the prefent Case. But it is to be noted, that the same Thing is also to be understood of the Times of Vibrations in Circles as well as in Cycloids. Thus, becaufe a Pendulum of 29L25 Inches perform any Vibrations whatever in a Cycloid, and also the least Vibrations in a Circle in the time of one Second, a Pendulum of 157 Inches would make the like Vibrations in the time of two Seconds, and one of 252125 Inches in the Space of 2 Seconds.

Coroll. (6.) Since the Times of all Vibrations whatever are equal in a Cycloid only; and upon this Account only are to be reckon'd equal in the leaft circular Arches; to wit, becaufe the Arches of the Cycloid and Circle coincide no where elfe but in the loweft Points, they being in every other Place fufficiently different from one another: It is manifeft, that the Vibrations in Arches of a Circle are fo much the lefs Ifochronal, by how much the greater they are; and confequent-

Horolog. O-Scillat. P. 9.

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ly, that in larger Arches they are fat enough removed from Isochronism. And according to Hugens, the time

time of Descent in a Quadrant of a Circle, is to that which is in the least Arch, as 24 to 29, supposing that the Vibration is made in a Vacuum. Which therefore will arise to a very fensible Difference when we compare the greatest Vibrations and the least together.

Coroll. (7.) Becaufe it appears from Experiments, and the Computation made thereupon, that each fingle Vibration to and fro, where the Pendulum is 96185 Inches long (each one I mean in a Cycloid, and the leaft in a Circle) is perform'd in 94125 Thirds of Time, or 1".  $24''_{125}$ . And becaufe, by what Hugens hath demonstrated, the Time of Vibration is to the time of the perpendicular Fail along the Quadruple of the Diameter of the generating Circle; or along double the Length of the Pendulum = 193176 Inches, or  $16_{11}$  English Feet, as the Circumference of the Circle is to the Diameter doubled;

or as  $94\lfloor 25$  Thirds of Time to 60 or one Second, [for  $355: 226: 94'' \lfloor 25: 60''' = 1'';$ ] thence it follows, that an heavy Body will defcend by the

Horol. Ofc. P. 57, 58. & de vi Centrifug. Prop. 12.

Force of its Weight 161 English, or 15 12 Feet of Paris, in the space of one Second. Which Velocity of Descent, deduced from the pendulary Experiments, agrees notably with the faid Author's Experiments about falling Bodies; and therefore is to be accounted for certainly true.

Coroll. (8.) Therefore the perpendicular Line of a falling Body being given for any fpace of time, the Line of Defcent, whether Perpendicular or Oblique, is given for any other Space of Time; as being always in the duplicate Proportion of the Time. Thus in a direct Fall, as ten Seconds Square = 100 is to the Square of one Second = 1; fo is 1610 English Feet defcrib'd in 10", to 16L1 Feet defcrib'd in one Second. And in oblique Defcents

scents it is not much otherwise. For the Lines of Descent in an oblique Plane are by the same Reafon one to another as the Squares of the Times; all the Difference is, that the Force of Gravity which is the Caufe of the Descent, is to be diminish'd in this Case in the Proportion of the Perpendicular Line to the oblique. (See Fig. 1. Plate 4. Coroll. 1. Pr. 6.) For fince the oblique heavy Body. as we shew'd before, descends in the same time through the Line E A, that the Perpendicular doth through the Line EG; it is manifest, that the moving Force is every where in the fame Proportion. Therefore if we put the Cafe, that the heavy Body descends along a Plain so very oblique, that EG is only a third Part of EF; what we have to do is only to diminish the Force of the Gravity in the fame Proportion, fo that the Body be suppos'd to descend along a Line of 51 27 Feet only in the space of one Second, and the Calculation will be the fame as in the direct Fall.

November 27, 1704.

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#### LECT. XI.

VIII. L L Projectiles, not perpendicular to the Horizon, defcribe . Parabola's, fo far as they are not hindred by the reliftance of the Air.

(Fig. 2. Plate 4.) Let any Body be fuppofed at T, and let it in any given time tend by the Force of the Horizontal Projection according to the Tangent T E, from T to a, fo far as it is not kinder'd by fome other Force: Then let the Force

Force of Gravity furpervene, which acts in the Line TK perpendicular to the Horizon, or any of the Lines parallel to it, (a l, b m, c n, d o, e p) for by reason of the great diffance of the Center of the Earth, those Lines are to be accounted for Parallel. Since therefore, the Force of the Projection begets an uniform equable Motion, according to the Direction Te or Fl, Gm, Hn, Lo, K p, which are parallel to the fame TE, and the Velocity of this Motion, ac-

cording to the primary Direction, fuffers nothing from the Force of Gravi-

ty that supervenes; the Body will in the end of the first time be found somewhere in the Line (a 1); of the second Time somewhere in the Line b, m, of the third fomewhere in c, n, of the fourth fomewhere in do, as being Lines parallel and equi-distant. Then let us consider the Force of Gravity as supervening, from which alone in the mean while, that the Body would by the proje-Aile Force alone describe the Line Ta, it is carried downwards according to any little Line TF or al, fo that if there were no other moving Force but this of Gravity, the Body would in the End of the first time come to the Line F1. Since therefore, the Velocity of this Force fuffers nothing from the combination of the projectile Force therewith, any more than the Velocity of this latter fuffers from it, the Body, notwithftanding the projectile Force will be found still in the end of the first time in the Line Fr. From whence it appears, that in the end of every time, it must by the Conjunction of these Forces be found in the Interfections of those Lines al and. FI, bm and Gm, Oc. to wit, in the end of the first time in the Point L, of the second in the Point M, of the third in n, of the fourth in o, and fo on. Since therefore, if al be of one Part, b m

b m is 4, c n is 9, d o is 16, e p is 25, and fo on, they being amongst themselves by Prop. 4. as the Squares of the Times or Distances Ta, TB. T c, T d, T e; and from the Nature of the com-pound Motion, TF is to TG, as F1 squared is to Gm squared, and so in the reft. And since, according to the primary Property of the Parabo-la, the Absciffes of any Diameter whatever TP and TG, are also betwixt themselves as the Squares of the Semi-ordinates Fl and Gm, it is manifest, that the Points I, m, n, o, p, are in the Curve of a Parabola, the principal Diameter whereof is TK; and TF, TG, TH, TI, TK, are the Absciffes, and FI, Gm, Hn, Io, Kp, are the Semi-ordinates. And feeing all the Things here demonstrated do alike belong to any Diameter whatever, how oblique foever the Tangent of the fame may be to it, as well as to the Axis; it is manifest, that the Trajectories of all Projectiles are univerfally truly parabolic; that is, fo far as they are not hindered by the Reliftance of the Medium. O. E. D.

Coroll. (1.) Hence we may learn the Foundations of the Art of Gunnery. For fince all Projectiles carried to any Inclination whatever, do defcribe Parabola's greater or leffer, or at leaft a greater or leffer Part of the fame Parabola, fo far as they are not hinder'd by the refiftance of the Air, and feeing the Air is of fmall or no moment for the retarding the Motion of thefe Projectiles, by reafon of their Solidity, and the Velocity of the Motion; it is plain that the principles of that Art are to be taken from the Nature of the Parabola. The Ufe of this Corollary extends it felf a great way, and will be illuftrated in the Sequel by many Examples taken out of the faid Art.

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Coroll.

Co.oll. (2.) 'The Velocity of a Projectile being given, whatfoever the Angle of Inclination may be, there will also be given the distance of the Focus of the Parabola, from the Point where the Projection begun. Let s (See Fig. 4. Place 4.) be the Point of the Projection, where the Proje-Etile being thrown along the Tangent sv, begins to move in a parabolic Curve, and let sv be the Line to be describ'd by the Body in any given. time by the projectile Force alone; let alfo vc, or s r be a little Line to be defcrib'd by the force of Gravity alone in the fame given time. In the end therefore of that time, the Projectile will be found in the Point of the Parabola; and by reafon of the given Force of Gravity as well as of Projection, there will be given allo, whatloever may be the Inclination of the Tangent to the Horizon, the Lines vc or sr and sv or cv, that is, the Abscils of the Diameter so, and the femi-ordinate of the fame ; the third Proportional of which two is the Latus rectum belonging to the Vertex s ; which therefore will necessarily be given when the former things are given. From whence also, the fourth Part of that Latus rectum, which is the diffance of the Vertex s from the Focus of the Parabola, will also be given. Although, therefore, from the fame Velocity of Projection, divers Parabola's will be defcrib'd in divers Elevations, yet the Foci of them all will be equidiftant from the Vertex or Point where the Motion began, and confequently will be placed in a Circumference, whole Center is in that Point, Q. E. D.

Coroll. (3.) The Horizontal Range is then the longest, when it is directed according to a Line which is in the middle betwixt the Horizontal and perpendicular Lines, or in an Angle of

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45 Degrees above the Horizon. For, feeing the principal Vertex of any Parabola whatever, defcrib'd by Projectiles, is in the greatest Height of the Projectile, under which in the Axis it felf the Focus DF is placed ; feeing the Diftance of the fame Focus from the Vertex is given : Seeing laftly, the Horizontal Range will be the longeft where sF, the diffance of the Vertex s, from the Focus, is measured by e g an Ordinate to the Axis paffing through the Vertex s; fince these things are fo, the Horizontal Range will certainly be the longest, where sF the distance of the Vertex s from the Focus, coincides with sg the Ordinate to the Axis; for otherwife, by reason of the given diftance of the Focus, 2 Fsg will be greater than s F doubled : But where s F coincides with sg, sg will be double to sF, and confequently, the Horizontal Range will be the longeft, where sF coincides with sg; that is, where the Angle vsh is half-right. For the Angle vsf comprehended by the Tangent vs and sF, the diftance of the Vertex s from the Focus is always equal to the Angle bso, comprehended by the fame Tangent, and so the Diameter of the Parabola. Τf therefore, the Angle b so be an half right one, v s F will also be an half-right one, and confequently the Angle os F will be a right one, and the Line sF will become sh, and will coincide with the Ordinage sg, and the Ordinate sg will be the longest Horizontal Range of all.

Coroll. (4.) Since therefore, the Tangent of the Parabola doth only in that Cafe comprehend with the Diameter an half right Angle, where it toucheth the fame at the Extremity of the principal Latus rectum, which passet through the Focus; it is manifest, that every longest Horizontal Cast will be comprehended betwixt that

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Part of the parabolic Curve that is placed above the latus retum, the Focus it felf being in the Horizontal Line; and that the higheft Altitude from the Horizon in this Cafe is T F, a Quarter of the principal Latus retum.

Coroll. (5.) If the Angle of Elevation differs equally from an half-right one, whether it be greater or lefs, the greateft Horizontal Range will egually be diminified. For because of the right Angle hso, and the Angles vs F, os b, which are equal one to the other, their Excels or Defect with reference to a right Angle, will be equal to the Angle Fsh, whether the Focus be above the Horizontal Line, as in the greater Elevation, or beneath it, as in the lefs Elevation. But the acute Angle Fsh, and the right one Fhs, and the Side Fs being given, there is given withal the Side sh the femi-ordinate of the Axis, and sg the ordinate determining the Horizontal Range. Thus in two Projections of equal Velocity, where the Angles of Elevation are one of 40, the other of 50 Degrees, the Horizontal Range will be equal on both Sides, and in like manner in the Degrees 20 and 60,20 and 70, or, as is well known to them that practile this Art. Coroll. (6.) The Horizontal Diffances produced in a given Velocity, in divers Angles of Elevation, are as the right Sines of the double Angles of Elevation. To wit, as gs is every where, fo is h s the half thereof; but in the right Angled, Triangle F h s, because of the given Radius F s, and the Angle h F's, double to b's o, the Angle of the Tangent and perpendicular s h will be every where the right Sine of that Angle ; and confequently, the Horizontal Diffances will always be betwixt themselves as those Sines. يجددونه رلها

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Coroll.

Coroll. (7.) The Times of every Horizontal Range from a given Velocity in divers Degrees of Elevation, are betwixt themfelves as the right Sines of the Angles of Elevation. Let one Body be thrown (see Fig. 4. Plate 4.) according to the Angle of Elevation 1 c d, and another according to the Angle L A D. I fay, that the time in which the first Body reacheth through the Parabolic Arch c T I to the Point I, fituate in the fame Horizontal Plane with the Point c, will be to the time in which the latter Body reacheth along the Arch At L (Fig. 5. Plate 4.) to the Point L, fituate in the fame Horizontal Plane with the Point A. as the Sine of the Angle d c1 is to the Sine of the Angle DAL. For let there be in these Figures, as taken together,  $\triangle A$  equal to dc: Ae also (because of the Equality of the Time, in which the Bodies together would defcribe the equal Lines d c and  $\triangle$  a by the proje-Aile Force alone) will be equal to d1. But by the Nature of a Parabola before declared, D L is to  $\Delta e$  as DA squared is to  $\Delta A$  squared; or as DL squared is to DI squared. Therefore DL,  $\Delta$ i,  $\Delta$ e, are three Lines continually proportional. And fince (Prop. 4.) the Lines DL,  $\Delta e$ are in the duplicate Proportion of the Times,  $A_{i}$ , and  $\triangle e$  will themselves be in the proportion of the Times ; therefore, the former Time will be to the latter, as  $\triangle$  e, or d 1 is to  $\triangle$  i; that is, as the Sines of the Angle of Elevation d c 1 and D  $\triangle$  L. Q. E. D.

Coroll. (8.) The greatest Heights of Projectiles in a given Velocity in divers Angles of Elevation, are one to another as the Squares of the right Sines of the Angles of Elevation. To wit, as d 1 or  $\triangle e$  fquared is to  $\triangle 1$  fquared, to are the greatest Altitudes d l or  $\triangle$  c to D L. Q. E. D. Coroll.

Coroll. (9.) The greatest Altitude of Projectiles in a given Velocity, is where the Projection is perpendicular to the Horizon, and is the 4th Part of the Laws rectum, which in a given Velocity is always given ; as in this Cafe the Parabola ending in a right Line, the Vertex T (Fig. 4. Plate 4.) coincides with the Focus F, and the highest Altitude becomes equal to F s a quarter of the given Laws rectum; and confequently, which is to be noted by the way, half of the longest Horizontal Range, as we shall prefently demonstrate.

Carell. (10.) The Angle of Projection being given, but the Velocity being chang'd, both the highest Altitude, (Fig. 4. Plate 4.) that is, the principal Vertex of the Parabola, and the longest Horizontal Range, or the Ordinate sg will be changed in the duplicate Proportion of the Velosity. The former Part is manifest from what has . been before demonstrated, feeing the Altitudes of the Lines, or their Afcents, or Defcents, are always in the duplicate Proportion of the Velocities by Prop. 4. foregoing. And from this Part of the Proposition the other alfo follows; for, because of the likenels of all Parabola's, and of the like-nels of the Parts of like Figures on both Sides, if the Altitude Th be changed in the duplicate Proportion of the Velocity, the reft of the Lines alfo, as s h and sg will be changed in the fame Proportion. But the latter part of the Corollary may sho be cally deduc'd otherwife, and from of the Nature of the Parabola it felf; for let us put the Velocity to be twofold greater than it was before, in this Cafe, in what time the Projectile would describe before the Line s v, it will now describe the double of that Line ; but because of the Uniformity of the Force of Gravity, the Line

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Line v c or sr will not be changed. Therefore, as vc or sr given is Double to the Line's vy fo is that double Line to another Line, to wit, the Latus Rectum of the Vertex's, which is Fourfold of the Latus Rectum that belonged 3to that Vertex before. From whence the Fourth Part of this Latus Rectum, or sF, will be Fourfold alfo of SF the fourth Part of the former Latur Rettum and because of the Likeness of the Triangles sFh, sFh in both Cafes, the Lines sh and sg will be Fourfold also of themselves, sh and sg; and fo in the reft. The longeft Horizontal Range therefore in a Velocity Twofold greater, is Fourfold greater, in a treble Velocity Ninefold, and to in infinitum. Nay, indeed it is generally to be afa firmed, that all the Lines in a Parabola which are like, and in the like manner placed, are always increas'd and diminish'd, in the duplicate Proportion of the Increase and Diminution of the We- . locity; as may be gathered from what has been: 14 CODE walteb een ja already faid."

Coroll. (1,1.) The longeft Horizontal Range of every Parabola, is equal to half the Latus Refum belonging to the Vertex that terminates the principal Latus Refum. For in that Cafe, Fs is equal to sh; but Fs is the fourth Part of the Latus Refum belonging to the faid Vertex; and sg is double to sh; from whence sg is half the longeft Horizontal Range of that Latus Refum.

Coroll. (12.) Hence we may determine the longeft Horizontal Range belonging to every Degree of Velocity. For let it be made thus, as the Line sr, defcrib'd in one Second of Time by the Force of Gravity = 16|r English Feet, is to the Velocity of the Projectile sv, or rc, to be computed in the fame Time; fo is that Velocity to a fourth Number, which will give the Latus Retum

**Refum** of the Vertex s in the fame Feet: But the half of this Number will give the longeft Horizontal Range, as is abundantly manifeft from what hath been faid. Thus, if the Projectile Velocity be fo great, that it would carry the Ball One Thoufand English Feet in one Second, fay 16[1:1000:1000:62112 the Latus Refum of the Vertex s in English Feet. The longeft Horizontal Range therefore is 31056 Feet; beyond which Diftance, as nothing can be reach'd, fo within the fame it may reach any affign'd Place whatever, as will be shewn in the next.

Coroll. (12.) Probl. (1.) To reach by a Proiectile, in a Given Velocity, any Place whatever affign'd in an Horizontal Plane, not distant above half of the Latus Rectum of the Vertex s. Let the Place be at the Diftance of 20,000 English Feet, and the Velocity be the fame which was fuppos'd in the foregoing Corollary. Because therefore the Velocity is given, there is given the Latus Rectum of the Vertex where the Projectile will begin its Motion in a Curve, and confequently the 4th Part thereof, or the Line Fs, to wit, of 15528 Feet. But according to what was faid before. sh is 10000 Feet. From these things therefore, let there be found the Angle hsF by this Analogy, as sh is to sF, or as 10000 is to 15528; fo will the Radius be to the Secant of the Angle Fsh, to be found by the Table of Secants; to wit, 49°. 47'. Which Angle being taken out of a right one, or superadded thereto, will give the Sum of the equal Angles vsF and osb; the half of which vsF, or osb, will determine the Angle which the Tangent v b ought to comprehend with the Perpendicular so ; to wit, of  $90^{\circ}$ .  $-49^{\circ}$ . 47'.  $=40^{\circ}$ . 12'. (or  $90^{\circ}$ . × 49°. 47'. = 139°. 47'.) the half of which is 20°. κ 6...

6'. 20". or 69°. 52'. 20". to wit, according as we would have the Elevation greater or leffer than the Mean. If therefore a Leaden Ball be thrown with the given Velocity in the faid Angles, it will defcribe the Parabola required, and confequently reach the Place affign'd, without any other Declination from the Mark, than what the Refiftance of the Air makes, which indeed is fo fmall, that it ought not to be regarded. Thus we have folv'd the Problem, and taught how, with a given Vecity, to hir any Mark in any Horizontal Plane; which is not too far diftant.

Coroll. (14). Problem (2.) To reach by a Projectile Motion, in a given Elevation, any Place affign'd in an Horizontal Plane ; that is, from the Diftance of the Place given sg, and the Angle hsv given to determine the Velocity s v. Here the Quadruple of s F will give the Latus Rectum belonging to the Vertex st. That sv therefore may be found, vc or sr is to be drawn into the Quadruple of sf; and from thence will arife a Rectangle equal to the Square of vs or cr: the Ouadratick Root therefore being extracted but of that Rectangle, there will be found v s or c r. that Semi-ordinate which the Projectile ought to describe in one Second of Time. As for Example : Let sg, the Distance of the Object, be as before, to wit, 20000 English Feer, and the given Angle h s v 69°. 52'. 30". The Angle Fsv, or osb, will be of 20°. 6'. 20". and the Angle Fsh of 49°. 47'. Whence from the Tables of Sines, the Proportion of the Line sh to Fs will be that of 10,000 to 17,928 From whence Fs will be given, and the Latus Return of the Vertex s of 62112 Feet; which Number being drawn into v c or sr = 16| r. Feet, there will arife the Rectangular Number 1000,000, the lauare

Iquare Root whereof is 1000, which affords the Number of Feet in the Line sv. If therefore in the given Angle, the primary Velocity of the Projectile be fuch, as to carry it One Thousand Feet in one Second of Time, it will reach the Mark g placed in the Parabolick Curve s T g, if it is not a little retarded by a very small Resistance of the Air. And the Computation is altogether the same, if the Angle F s v, or os b, were suppos'd to be 69°. 53'. 30". as will easily appear from what was faid in the last Corollary foregoing.

Coroll. (15.) Hence alfo, from the Elevation given, or from the Velocity given, any Place whatever, as 1 may be reached that is placed out of the Horizontal Plane; that is, if in the fame Parabola produc'd, if there be Occasion, we note fome other Point, as g, placed in the Horizontal Plane. For the fame Cast which reacheth to the Place g, will alfo reach unto 1, or any other Place fituate in the fame Parabola.

Coroll. (16.) The Velocity of a Body which describes a Parabola, is every where as a right Line drawn from T, the Vertex of the Parabola unto the Middle of the Semi-ordinate; or as part of the Tangent drawn betwixt the Point of Contact m, and the Axis; that is, as the Secant of the Angle of Elevation above the Horizon. (See Fig. 2. Plate 4.) For the Line to be describ'd in the given Time, is as the Diagonal of the Parallelogram m Q R P; the Side whereof m Q is always given, and in P is equal to bm doubled. or to SG. Therefore the Velocity in the Point m, is to the original Velocity in the Point T, as mR is to PR, or as Sm is to GM; and thus it is every where. Therefore the Velocity in m, any Point of the Parabola, is to the Velocity on. which is any other Point of the Parabola, as the K 2 Pare

Part of the Tangent Sm is to Part of the Tangent X4, both being taken betwixt equi-diftant Diameters b m and TG; that is, as the Secants of the Angles of Elevation. Q. E. D.

Coroll. (17.) The leaft Velocity of all therefore is in T, the Vertex of the Parabola; and the Velocity is always by fo much the greater, by how much the greater the Diftance is from that Vertex.

Coroll. (18.) If therefore the Velocities of Bodies, thrown in divers Angles, be in the Proportion of the Secants of the Angles of the Elevation above the Horizon, they will all of them deforibe the fame or an equal Parabola, that is, Parts of the fame or an equal Parabola; Parts greater, where the Angle of Elevation is greater, and leffer where the faid Angle is lefs. But if the Velocities be in fome other Proportion, they must needs defcribe divers Parabola's, or Parts of divers.

Decemb. 4. 1714.

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#### LECT. XII.

#### A Lemma to the 9th and following Pro. politions.

H E Quantity of the Centripetal Force of Bodies mov'd round in Circles, is to be effimated from two things conjunctly, to wit, the Curvature of the Arches defcrib'd at the fame time, and the Velocity of the Motions in that Curvature. For

For fince all Motion is in it felf Rectilinear, and Bodies can be mov'd in Curves only by a centripetal Force impress'd from without; it is reafonable that the Velocity being given, we should determine the Curvature, which is generated from an extrinsic centripetal Force only, proportional a the fame Force. And on the other Side, fince to greater centripetal Force is requir'd to the making the fame Curvature, where the Velocity of the Projection, or of the equable Original Motion is the greater, and a lefs centripetal Force is requir'd where the faid Velocity is the lefs: It is reasonable that the Curvature being given, we fhould determine the centripetal Force, proportional to the fame Velocity. As therefore in comparing Rectangles, we determine their true Proportions by those of their Longitudes and Latitudes; to it is likewife to be done in the comparing of centripetal Force; by defining, to wit, their true Proportions in every given Time, from the Proportions of their Curvatures and Velocities conjunctly. Let it therefore be laid down for a certain Truth, That the Proportions of centripetal Forces are every where to be estimated from the Proportions of their Curvatures and Velocities conjunctly.

Scholium, For the Understanding the true Proportions of Curvature and Velocity, we are to observe, that the Curvature is every where equal in the leaft equal Angles, if the Subtenfes of the Angles of Contact be one to another, as the Radii or Diftances from the Center; as the Proportion of like Figures doth altogether require : And if the Curvature doth recede from that Proportion of the Diftances, the Proportions. of the Excels or Defect are to be taken for the true Proportions of exceeding or deficient Curvature,

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ture, which are afterwards to be computed. But as for the Velocity, it is every where to be comfider'd according to that Degree in which it ferves to promote the Angular Motion, and confequently every where in a Line Perpendicular to the Radius, or, which comes to the fame, it is to be effimated in the leaft Circular Arch. For wherefoever the Direction of the Motion is either upwards or downwards, by how much the Velocity is increas'd, by fo much is the Curvature always diminifh'd; and fo on the contrary : The Quantity which arifeth from the conjunct Forces of the fame, being yet in no wife chang'd; which is to be obferv'd every where.

PX. If two Bodies do in equal (Fig. 5. Plate 3.) Times run over Two whole unequal Circomferences b d g e, and BDGE, with an equable Motion, the centripetal Force in the grezter Circumference will be to that which is in the lefs, as the Circumferences are one to another directly; or, which is the fame, as their Diameters or Radii.

For becaufe of the Curvature given on both Sides, to wit, that of a whole Circle; the centriperal Force in the greater Circumference will be to that which is in the Lefs, as the Velocities of the Bodies; that is, as the Circumferences of the Cffcles, directly. Q. E. D.

Corbilary : If the periodic Times of Bodies revolving in Circles be Equal, then both the Velocities; and the centriperal Forces proportional to the fame, will be one to another, as the Circumferences or Diameters of the Circles directly; and on the contrary, if the centriperal Forces of Bodifferences or Radii of the Circles directly;

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rectly; then their Velocicies will also be in the fame Proportion, and the periodic Times will be equal.

Cordi (21) If the Force of fome central attracrive Body be directly as the Diffances from the fame Center; the periodic Times of all Bodies revolving in Circles about the central Body will be equal. And the fame Thing is to be afferted of Ellipfes, fince their entire Curvatures be equal to the entire Curvature of rany Circle whatever; and their Circumference an intermediate one; as it were betwist the Circumferences of Circles; taken on this Side and on that. From whence from the Equality of periodical Times in Circles; whether greater or leffer than Ellipfes, it will be obvious enough to affribe the fame Equality of periodic Times, to the intermediate Ellipfes alfo-about their Centers.

X.IIIf: two Bodies revolve in the fame, or equal Circles' with unequal Celerities, but both with an equable Motion, the centripetal Force of the Swifter will be to that of the Slower, in the Proportion nof the Gelerities duplicated; or as the Sausses of the Arches described together. For becaule of the given Gurvarure of equal Circles in ehund Arches; increating together with the Velotier, the Curvarand allo will increase in the fame Proportion ; therefore the centriperal Force, which is to be effinanced from the Garvature . and the Velocity conjunct, will be in a given Time in Proportions of the Arch, no the Arch defcribed at the fame Time, upon Account of the Velocity; and in Proportion of the fame Arch, deserveur at the fame Time, consider'd in refeeof-the Cutvature ; from whence by the Coninterent of both Propertions, the Icentripetal Force will be ( the Rehangle being reduo'd to g **G**vun Square,) K 4

Square,) in the duplicate Proportion of the Arches describ'd at the same Time; or as the Squares of those Arches. Q. E. D.

Coroll. (1.) Since the Periodic Times in equal Circles are reciprocally Proportional to the Velocities, the centripetal Forces will be reciprocally in the duplicate Proportion of the periodic Times, or as the Squares of the periodic Times reciprocally; fo that by how much the Square of the periodic Time is the greater, the centripetal Force is fo much the lefs; and by how much that Square is the lefs, fo much the greater is the centripetal Force, and in the fame Proportion.

Coroll. (2.) If many Bodies be mov'd in Circles about many Centers, each about their own, and this at the fame Diftance every one from the Centers; the attractive Force of the central Bodies will eafily appear, fince they are amongst themselves as the Squares of the Times reciprocally; and the Velocities also will eafily appear, fince they be in the reciprocal Proportion of the periodical Times.

XI. If two Bodies be carried in unequal Circles with equal Velocity, their centripetal Force will be in the reciprocal Proportion of their Circumference or Diameters; fo that in the leffer Circumference there will be the greater centripetal Force, and in the greater the lefs.

For becaufe of the given Vélocity, the centripetal Force in the given Time, will be as the Curvature of equal Arches, that is as the Circumferences, the Diameters or Radii directly; Q. E. D.

Coroll. (1.) Since the periodic Times in Bodies equally fwift, are betwixt themfelves in the fame Proportion, as the Circumferences to be defcribed; if the periodic Times of Bodies mov'd

mov'd in divers Circles, be directly as the Circumferences of the Circles, the centripetal Forces will be as those Circumferences or Radii reciprocally; and on the other Hand, if the centripetal Forces be as the Radii or Diffances reciprocally, the periodic Times will be as the Radii directly.

Coroll. (2.) If the Force of any attractive central Body be reciprocally as the Diffances of Bodies from the Center; fo that by how much the nearer Bodies approach thereto, fo much the greater the central Force is; and by how much they are more remov'd, fo much the lefs is that Force: The peroidic Times of Bodies, plac'd at divers Diffances, will be as those Diffances directly, and the Velocities of those Bodies will be equal.

XII. If two Bodies be mov'd in unequal Circles, with an unequal Velocity, in the fub-duplicate Proportion of the Circumferences, Diameters, or Radii, the centripetal Forces will be equal every where, and neither increas'd in the Accefs or Recefs.

For becaufe of the greater Velocity in the greater Circle, and this in the fub-duplicate Proportion of the Circumferences, the centripetal Force is to be increas'd in the greater Circle in the fame Proportion. And becaufe of the greater Curvature in the leffer Circle, and this in the fub-duplicate Proportion of the Circumferences, the centriperal Force is to be increas'd reciprocally in the leffer Circle. It is therefore manifelt, That the centripetal Forces are to be increas'd on both Sides in an equal Proportion, and confequently that they are full equal on both Sides. Q. E. D.

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As for Example, Let the Radius of the greater Circle be Fourfold of the Radius of the leffer Circle, or as 4 to 1; and let the Velocity in the greater be to the Velocity in the leffer, in a fubduplicate Proportion of the Radii, or at 2 to 1. Seeing now the Curvature of the greater is equal to the Curvature of the leffer, in like Arches, and in Arches equal, is to the fame reciprocally as the Radii ; it is neceffary that in a double Arch; which the double Velocity in the greater, will describe in the given Time, the Curvature should be half of the other. The Velocity therefore of the former Body is to that of the latter, as 2 to I : and the Curvature of the latter to the Curval ture of the former, as 2 to 1." From whence the Quantity of the centriperal Porce in the former, will be to the Quantity of the fame in the latter, as a Rectangle made of the Velocity of the former, drawn into the Cutvatute thereof, or as 2 x 1, is to a Rectangle of the Velocity of the latter. and the Curvature of the fame conjunctly, or as 1 x 2 ; that is, in the Proportion of Equalit ty: And thus every where. disc it to doord en Coroll. (iz) Since the periodic Times be in this

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Coroll. (22) Since the periodic Times be in this Cale one to another, in the fub-duplicate Proport tion of the Circumferences, Diameters, 601 Rat dil, the Squares of the periodic Times will be betwise themfolves as the Circumferences, 1641. If therefore the Squares off the periodic Times be one to another us the Circumferences, 1642, end centripath Force will be equal in all Diffances; and the Melocines in the fub-duplicate Proportion of the fatter Circumferences, 1640. Anit on the pence Hand, if the centriperal Poirces be equal in II Diffances, the Segmares of the period dic Times, will be as the Diffances or Radii ; and the

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the Velocities still in the sub-duplicate Proportion of the same.

Coroll. (2.) If the centripetal Force of any attractive central Body be altogether the fame in all Diftances, the Velocities will be in the fubidu plicate Proportion of the Diftances; and the Squares of the periodic Times will be to one alnother, as those Diftances or Diameters, of Circumferences.

XIII. If two Bodies be movid in unequal Circles, with an unequal Velocity, in the fub-duplicate Proportion of the Circumferences, *crc.* reciprocally; fo that in the greater Circle the Velocity be the leffer, and in the leffer Circle the greater, and this in the faid fub-duplicate reciprocal Proportion, the centripetal Force will be reciprocally as the Squares of the Radii or Diftances.

For becaule of the leffer Curvature in the greater Circle, and this in the felquialteral reciprocal Proportion of the Radii; and becaule of the leffer Celerity also in the greater Circle, and this in the fub-duplicate Proportion of the Radii, the centripetal Forces to be derived from these conjunct Proportions will be in the duplicate reciprocal Proportion of the Radii, or reciprocally as the Squares of the Radii, Q. E. D.

For Example : Let the Radius of the greater Circle be Ninefold of the Radius of the leffer, or as 9 to 1; and let the Velocity in the greater be to the Velocity in the leffer, in the Subduplicate Proportion of the Radii reciprocally, or as 1 to 21 Seeing the Curvature of the greater is to that of the leffer as before, that is, in like Arches equal, and in equal Arches reciprocally as the Radii . It must needs be, that in an Arch equal to the Third Part only of the other, which the Third Part of the

the other's Velocity will describe in the Given Time. should be in the Greater only one 27th Part of the other, or as 1 to 27. Therefore the Velocity in the greater Circle is to that in the leffer as 1 to 2, and the Curvature in the greater to that in the lefs, as I to 27. From whence the Quantity of the Centripetal Force in the greater. will be to the Quantity of the fame in the leffer. as a Rectangle of the Velocity and Curvature in the greater conjunctly, or as  $I \times I = I$  is to the Rectangle of the Velocity and Curvature in the less conjunctly, or as  $2 \times 27 = 81$ ; that is, as the Square of the Radius of the lefs 10 = 1 is to the Square of the greater = 81. And fo every where, For the Periodic Times will be to one another. as 27 is to 1; that is, in the Selquialteral Proportion; for 27 is a Geometrical Mean Proportional betwixt 9 and 81; and confequently the Proportion of 27 to 1, contains the Proportion of 9 to 1; and the halved Proportion of 81 to 9, or the Subduplicate of 81 to 27, [1:2:9:27:81: ...] and thus every-where.

Coroll. (1.) Seeing the Periodical Times in this Cafe are one to another in the Sefqui-plicate Proportion of the Radii, the Squares of the Periodical Times will be betwixt themfelves as the Cubes of the Radii. If therefore the Squares of the Periodic Times be betwixt themfelves as the Cubes of the Radii, the Centri-petal Forces will be betwixt themfelves as the Squares of the Radii reciprocally; and the Velocities ftill in the Subduplicate Proportion of the Radii reciprocally. And on the other hand, if the Centri-petal Forces be inverfly as the Squares of the Diffances or Radii, the Squares of the Periodical Times will be betwixt themfelves as are the Cubes of the Radii,

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and the Velocities still in the Subduplicate Pro-

Coroll. (2.) If the Centri-petal Forces of any Central attractive Body whatever be in divers Diftances from their Center, as the Squares of those Diftances reciprocally, the Velocities of Bodies revolved in the divers Diftances will be in the Subduplicate Proportion of those Diftances reciprocally; and the Proportion of the Periodical Times duplicated, will be equal to the Proportion of the Diftances triplicated, or the Squares of the Periodical Times will be betwixt themselves as the Cubes of the Radii.

Coroll. (3.) If the Motion be in an Ellipsi, then let the Middle Distance betwixt the greatest and the least be taken; and then also the Squares of the Periodical Times will be as the Cubes of the Radii as well as in Circles.

XIV. If two Bodies be carried in unequal Circles with an unequal Celerity; fo that by how much greater the Radius, Diameter or Circumference is, fo much the lefs the Velocity is; and by how much the lefs the Radius is, fo much the greater is the Velocity, and this in the Reciprocal Proportion of the Radii, the Centri-petal Forces will be as the Cubes of the Radii reciprocally.

For because of the leffer Celerity in the greater Circle, and this in the Reciprocal Proportion of the Radii; and by reason also of the leffer Curvature in that Circle, and this in the Duplicate Reciprocal Proportion of the Radii, the Centripetal Forces to be deriv'd from those Conjunct Proportions will be in the Reciprocal Triplicated Proportion of the Radii, or as the Cubes of the fame reciprocally.

As for Example: Let the Radius of the greater Circle be Twofold of the Radius of the lefs, or as 2 to 1. And let the Velocity in the greater be to the Velocity in the lefs Reciprocally as the Radii. or as 1 to 2 : In this Cafe the Curvature of the greater will be to the Curvature of the lefs in the Given Time as 1 to 4: Therefore the Velocity in the leffer, is to the Velocity in the greater. asia to 1; and the Curvature in the lefs, to the Curvature in the greater, as 4 to 1. From whence the Quantity of the Centri-petal Force in the lefs, will be to the Quantity of the fame in the greater : as the Rectangle  $2 \times 4 = 8$  is to the Rectangle 1 x 1 = 1, or as the Cubes of the Radii reinprocally; and fo every-where. 

Coroll. (1.) Since the Periodic Times be in this Cafe in the Duplicate Proportion of the Radii) if the Squares of the Periodic Times be berwixt themfelves, as the Biguadrate of the Radifferent which is the fame, if the Periodic Times themfelves be one to another as the Squares of the Radii ; the Centri-petal Forces will be betwixt themfelves as the Cubes of the Diffances or Radii inverfely, and the Velocities inverfely as the Radii. And on the other hand, if the Centri-petal Forces be inverfely as the Cubes of the Diffances, the Periodic Times will be betwixt themfelves as the Squares of the Radii ; and the Velocities fill as the Radii themfelves inverfely.

Coroll. (2.) If the Centri-petal Forces of any Central attractive Body whatever be in divers Diftances from their Center, as the Cubes of those Distances reciprocally, the Velocities of Bodies revolved in divers Distances will be in the Reciprocal Proportion of the Distances; and the Periodic dic Times in the Duplicate Proportion of those Diffances.

Coroll. (2.) All the fame Things concerning Times, Velocities, and Centri-petal Forces, whereby Bodies defcribe like Parts of any Curves whatever which are like, and have their Centers in the like manner posited, do follow from the Demonstrations of the foregoing Matters which were applied to Circles in particular, as applied to the other Cases.

Scholium. Since the Cafe of Proposition 12th hath place in the Celeftial Bodies; to wit, that the Squares of the Periodic Times are every-whete one to another as the Cubes of the Diftances ; and that confequently the Centri-petal Forces are as the Squares of the Diftances reciprocally, and the Velocities in the Sub-duplicate Proportion of those Distances reciprocally: Since, I fay, this Cafe hath place in the System of the World, and this alone, as our Countrymen, Sir Christopher Wren, Dr. Hook, and Dr. Halley have feverally Collected ; and that the fame is now generally received amongst Astronomers; This most noble Cafe requires to be more largely and diligently confi-der'd in what follows; while the Confequences of the reft are but more lightly and curforily touch'd upon. 11 51

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XV.

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#### LECT. XIII.

HE Area's, which revolving Bodies do defcribe by Radii, drawn unto an unmovable Center of Force, do both lie in immovable Planes, and are proportional to

the Times, and fo in any given Time everywhere equal; the Velocity of Motion in the leffer Diftance, and the Slownels thereof in the greater fo tempering the Defcription of the Area's, that from those various Diftances no difference of the Spaces run over in the given Time doth ever arise.

For let the Time be divided into equal Parts, and in the first small Time let the Body by its innate Force, or by a projectile Motion defcribe in the first Part of Time, any right Line, as A B. (See Fig. 6. Plate 4.) The fame Body in the fecond equal Part of Time, if nothing hinder'd, and no other Force was impress'd on it, would go ftraight forward to B c, describing the Line B c, equal to A B, fo that the Area's A S B, B S c, made by the Radii drawn to the Center S, would be equal. But when the Body comes at the Point B, let a centripetal Force, whether it be an Attraction, or fome Preffure tending to the Center S, at upon the fame Body by one fingle Impulfe, which is to the projectile Motion as any Right Line, as Bg to Bc; this new Impulse will make that the Body should decline from the Right Line Bc, and go forwards in another Right Line, to

to wit, BC the Diagonal of the Parellogram Bg Cc; fo that the 2d equal Part of Time being compleated, the Body will be to be

Law of Mot. 22. foregoing.

found at the Point C, in the fame Plane with the Triangle ASB. Join SC, and the Area defcrib'd by the Radius drawn from the Center to the Body, that is, the Triangle SBC will be equal to the Area of the former; that is, to the Triangle SBc, (by I. 27. of the Elements,) and confequently to the first Triangle SAB, to which, by what was faid before, SBc is equal. By the like Argument in the 2d equal Part of Time, the Body would reach by the Projectile Force (which being once impress'd doth still endure) from C to d ; fo that the Line Cd, which is to be described, would be equal to CB that was last describ'd. But if any Centripetal Force whatfoever, whether greater or lefs than the former, should again act upon the Body in the Point C, the Body in the End of the 2d Time will be found fomewhere in the Line Dd parallel to SC, and would be carried along CD the Diagonal of a certain Parellogram h D d C to a certain Point D; fo that the Triangle SDC would be equal to the Triangle SdC, and confequently to the Triangles SCB, SBA, which are equal one to the other; and by the fame Reafon, if the Centripetal Force acts fucceffively in D, É, F, making that the Body should in each equal little Portion of Time describe a several Diagonal, all these right Lines will lie in the fame Plane, and the Triangles S E D, S F E will be defcribed equal to the former. Now let the Number of the Triangles be increas'd, and their Latitude decreas'd infinitely, their ultimate Perimeter ADF will become a Curve Line, the Sides of a Polygon ending in a Curve. Τ.

Curve, and by reason of the continued and never-ceasing Action of the Centripetal Force, the Bodywill perpetually be drawn back from the Tangents of the Curve, and the Areas likewise by the fame Reason as before, will still be described in an unmoveable Plane, and be proportional to the Times. Q. E. D.

Coroll. (r.) Therefore the Velocity of a Body revolved about a Center, which is effimated according to a Line perpendicular to the Radius, will be in the reciprocal Proportion of the Diftances; for otherwife the Equality of the Areas could in no wife be kept.

Coroll. (2.) The Angular Velocity of a Body about the Center of Force, will likewife be in the reciprocal Duplicate Proportion of the Diffances. For fince the true Velocity is in the fimple reciprocal Proportion of the Diffances, as we have feen already, and the Diffance of the Center is the greater by how much the Motion is the flower, and in the fame Proportion; it is manifeff, that that angular Velocity with respect to the Center of Force, is in the duplicate Proportion of the Diffance reciprocally.

Coroll. (3:) Where the Position of the Tangent is perpendicular to the Radius, or Distance from the Center, and the Velocity of the projectile Motion makes the Centrifugal Force exactly proportional or correspondent to the Centripetal Force of the Central Body; the Body will neither approach to the Center, nor recede from it, but will be carried perperually with a circular Motion about that Center.

Coroll. (4.) But where the Position of the Tangent is oblique to the Radius, although the Velocity of the Projectile Motion be proportionate and correspondent to the Centripetal Force, that Centripe-

tripetal Force will fomewhat increase even the least descending Motion by conspiring together with it, and something diminish even the least afcending Motion by oppofing it; until at length the increas'd Motion exceeds the Centripetal Force, and the Body which before descended comes now to afcend ; and the diminish'd Motion at length yields to the Centripetal Force, and the Body which alcended before doth again defcend.

Coroll. (5.) From fuch like Circumstances ought to arife the Motions of Bodies revolved in Ellipses about any Center whatever. For although the Body be suppos'd to be now situated in the Course of its Revolution, at the leffer Axis of the Ellipsi, the Central Body possessing the Focus, or at the mean Diffance, and the Velocity of the Projectile Motion be supposed also in that Place to correspond exactly to the Centripetal Force; yet notwithstanding, because of the oblique Polition of the Tangent in the fame Place, the Motion will become not Circular but Elliptic; whilft the Body, as it is in defcending, doth by little and little acquire a new Force, by which it may afterwards alcend; and as it is in alcending, doth by degrees lofe fome Force by which it ascended before; until at length the Centripetal Force overcoming it, be compell'd to defcend ! And thus perpetually. From whence it is manifeft, by what means an Elliptic Motion may arife from a Motion impress'd according to an Oblique Line; in the mean while that the very fame Motion impress'd, according to a perpendicular Line; would have produc'd a Motion altogether Circular.

Corol. (6.) If in Vacuo the Area's defcrib'd be not proport

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portional to the Times of the Description, the Forces do not tend unto a Concourse of the Rays. For if they tended thither, the Areas would neceffarily be proportional to the Times. Which is contrary to the Hypothesis.

Coroll. (7.) In all Mediums, if the Description of the Areas be accelerated, the Forces tend not unto the meeting together of the Rays, but confpire rather with the Projectile Motion; if the Description of the Areas be retarded, that is, more than the Resistance of the Medium requires, the Forces tend not unto the meeting-together of the Rays, but are rather opposite to the Projectile Motion.

XVI. Every Body-which is mov'd in a Curve Line, and doth by a Radius drawn to fome Point, either immoveable, or going forwards uniformly with a Rectilineal Motion, defcribes Areas about that Point proportional to the Times, is urged or impres'd by a Centripetal Force tending to the fame Point.

Cafe (1.) For becaufe of the Equality of the Triangles Sc B, SC B (fee Fig. 6. Plate 4.) defcrib'd upon the fame Bafe SB, the Points C and and c will be (by I. 29. of the *Elements*) in the Line C c parallel to the Bafe; and confequently the Figure B c C g will be a Parallelogram, in which the Sides B c and B g reprefent the Forces, and B L is the Diagonal: And therefore the Body placed at B, is incited by the Force B g tending to S the Center of the Forces; and fo likewife in all the Points, C, D, E, F. 2. E. D.

Cafe(2.) And it is the fame thing, whether the Plane in which the Body defcribes the Curvilinear Figure doth reft, or whether it be mov'd together with the Body, the Figure defcribed, and its Central Point S uniformly ftraight forward. Where-

Wherefore the Demonstration of the former Cafe will hold in this alfo.

Scholium. A Body may be urged by a Centripetal Force, which is compounded of divers Forces, (as for Example, the Force of heavy Things towards the Center of the Earth, is compounded of Forces tending to all the Parts of the Earth, as will appear afterwards;) And in this Cafe the Senfe of the Proposition is, that that Force which is compounded of all, when it is reduc'd to one, tends to the Center of that Force.

Coroll. (1.) Seeing therefore in the System of the Primary Planets the Areas described by Rays thawn to the Center of the Sun, are always proportional to the Times; as is well known to Astronomers, the Planets are perpetually urged by a Force tending to the Center of the Sun; and in the same manner must we reason concerning the Secondary Planets, as revolved about their Primary ones, Saturn, Jupiter, and the Earth.

Coroll. (2.) As the Velocity of divers Bodies about a Center of Force ; where that Force is ast the Squares of the Diffances inverfely, is in divers Circles in the Subduplicate Proportion of the Diftances inverfely, as we demonstrated before : to from this and the foregoing Proposition it folhows, ithat the Velocity of the fame Body defcribing any Eccentric Orbit, taken as placed in its divers Distances from the Center, let the Condision of the Centripetal Force be what it will, is aso is the Diftance inversity; i. a if the Velocitype effimated by a Circular Arch, or in a Line perpendicular to the Radius as before : The Caule of swhich divers Velocity is this, that in divers "Circles the Areas in that Cafe are not equal on both Sides, but according to the Greatness of the Diftance greater, and in the fame Proportion of the L 2 . . . .

the Magnitude allo greater; when notwithstanding, in the Revolution of the fame Body, the Equality of the Areas doth altogether require a. Velocity reciprocally proportional to the Diffance. Thus, if two Planets are revolved about the Sun in divers Circles, the Radii of which do exceed one the other in the Quadruple Proportion, the remoter Planet would be carried with a Velocity which is only double to that of the other: But if the same Planet, performing its Circuits in a very Eccentric Ellipsis, be placed sometimes at a greater, somerimes at a lesser Distance ; and the fame, as before, exceeding and falling thort by turns in the Quadruple Proportion, the Velocity will be in the reciprocal Proportion of the Diffances, and in the leffer Diftance exactly Quadruple of the other; and fo in any Distances whatever Which thing ought to be kept in mind, in the Contemplation of the whole Planetary System. XVII. Every Body, which by a Ray drawn to the Center of any other Body howfoever mov'd, describes Areas about that Body proportional to the Times; is urged by a Force compounded of a Centripetal Force tending to the other Body, and of all the accelerative Force wherewith the other Body is urged. For, if first of all the Plane, and the Center of Forces in the Plane ido rist, the Areas will be proportional toother Finies; and if both Bodies be accelerated with the fame Celerity according to parallel Lines, the Areas will ftill remain proportional to the Times sFrom whences fince by the Hypothefis the Areas remain proport tional to the Times, both the Centripetal Force, the Gaule of them, will remain, and the acceled rative Force will remain every where the fame" Statistics of the man set of the Caractery 1. Jach ... Coroll. .....

Coroll,

Coroll. (1.) If any Body whatever doth with a Ray drawn to the Center describe Areas proporcional to the Times, and there be subducted from the whole Force wherewith the former Body is . nrged, whether Simple or Compound, the whole accelerative Force wherewith the latter Body is urged, all the remaining Force wherewith the former Body is urged, will tend unto the other Body, as to the Center.

Coroll. (2.) And if those Areas be nearly proportional to the Times, the remaining Force will tend to the other Body very near.

Coroll. (2.) And on the other hand, if the remaining Force doth come very near to the tending to the other Body, those Areas will very nearly be proportional to the Times.

Corall, (4,) If a Body doth with a Ray drawn to another Body defcribe Areas, which, when compared with the Times, are very unequal thereto; and that other Body doth either reft, or is moved uniformly straight forwards, the Action of a Centripetal Force tending to the other Body, is either none at all, or is mingled and diffurbed by other Forces. And the whole Compound Force, if it be fuch a Force, will be directed to fome other Center, whether immoveable or moveable, the Areas described about which will be proportional to the Times. The fame thing holds where another Body is mov'd with any Motion whatever, if so be the Centripetal Force be taken to be that which remains after the Subduction of the whole Force, which acts upon that other Body.

Scholium (1.) Because an equable Description of. Areas is a Token of a Center, which that Force. wherewith the Body is affected doth respect, and L4

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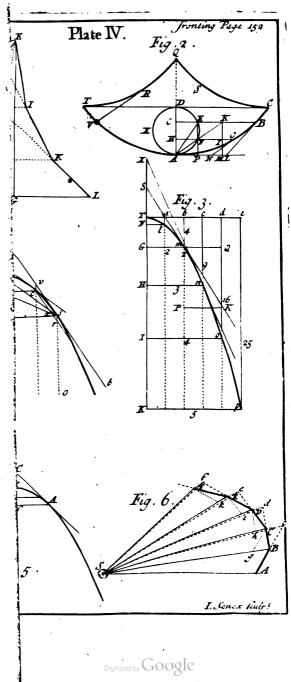
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the Body by this Centripetal Force is retain'd in a Curvilinear Orbit; and all Curvilineal Motion is rightly faid to be made about that Center, by the Force of which the Body is drawn back from the Rectilinear Motion, and perpetually rotain'd in its Orbit: In what follows, we fhall make use of that equable Description of Areas, as the Index of a Center, about which the Motion which is in a Curve is perform'd in free Spaces.

Scholium (2.) This 17th Proposition, with its Corollaries, appertains to the understanding the true Syftem of the World. For although all Planetary Motions are to be derived from a Projectile Motion once imprefied according to Tangents, and a Centripetal Force continually urging; yet those Centers unto which the Centripetal Forces tend, are also mov'd themselves, together with the Bodies that are revolved about them. Thus the Circulations of the Circumfaturnian, and Circumjovial Planets, and of the Moon, do proceed from a Projectile Motion once impress'd upon each, and from a Centripetal Force tending to the Centers of Saturn, Jupiter, and the Earth respectively; albeit, in the mean while those Central Bodies, together with their whole Satellitium, be mov'd about the Sun, the common Center of all the Primary Planets.

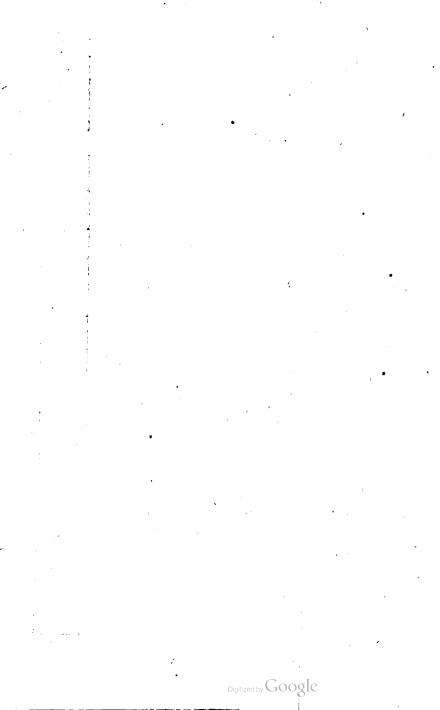
XVIII. A Problem. There being given in any Three Places whatfoever, the Velocity wherewith the Body defcribes a given Figure, by a Force tending unto fome common Genter or Point, to find that Center.

Let Three Right Lines, PT, TQV, VR, (fee Fig. 1. Plate 5.) touch a defcribed Figure in fo many Points, P, Q, R; whilft they meet together in T and V. Let there be erected PA, QB, RC, perpendicular to the Tangents in the Points



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Points of Contact, and let them be reciprocally proportional to the Velocities of the Body in those Points; that is, let PA be to QB as the Velocity in Q is to the Velocity in P; and Q B to R C as the Velocity in R is to the Velocity in Q. At the Extremities of the Perpendiculars A, B, C, let AD, DBE, EC be drawn at right Angles to the Perpendiculars, or parallel to the Tangents, and meeting together in D and E. Let TD and VE be drawn interfecting each other in S; and from the Point E, let E r and Ev be parallel to the Perpendiculars CR and BQ refpeaively. And in like manner from the Point S, lerDp and Dx be parallel to the Perpendiculars A P & BQ respectively. Then laftly, from the Point S let Ss, St, Sq be parallel to the fame Perpendiculars respectively, or perpendicular to the Tangents: I fay, that the Point 5 is the Center which is fought. For fince the Body revol-Prop. 1 5. fore ving, and placed fucceffively in the going Schol. Points P and Q, doth by Rays drawn to the Center of Force in ments. equal Time describe equal Areas, or equal Triangles; fince also those Triangles toge-

equal Friangles; fince allo thole Irlangles together defcribed, are as the Velocities, or as the Lines together defcribed in P and Q drawn into the refpective Perpendiculars, let fall from the Center to the Tangents PT, QT: Thole Perpendiculars will be reciprocally as the Velocities, and confequently as the Perpendiculars D p and D x directly. But becaufe of the Likenels of the Triangles TDx, TSt, and TDp, TSq; as is Dp to Dx, fo is the Perpendicular Sq to the Perpendicular St. And by the like Reafon, as is EV to Er, fo is the Perpendicular St to the Perpendicular Ss. And feeing this can be true only in this Point S the Concourfe of the Lines TD&VE,

it is manifest, that S is the Center of the Centriperal Force Q. E. D.

Jan. 29. 1705.

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# LECT. XIV.

F a Body be movid in an Ellipfis about the Center of the fame, the Centripetal Force will be directly as the Diftance of the Body from the fame Center.

For the Curvature every where in like Arches is in the quadruplicate Proportion of the Diffance; but the Velocity is in the fimple Proportion of the fame Diffance inverfly. From whence the Curvature, defcrib'd in a given Time, will be in the duplicate Proportion of the Diffance, and the Velocity in the fimple Proportion of the Diftance inverfly, and the Centripetal Force, which is to be effimated in this Cafe by the Excefs of the Proportion of the Curvature above that of the Velocity, will be directly as the Diffance. Q.E.D.

Corollary. If an Ellipfis, the Center thereof paffing away infinitely, be turn'd into a Parabola, the Body will be mov'd in this Parabola, and the Force now tending to a Center infinitely diffant will become equable. This is the Theorem of Galileo demonstrated by us above in Prop. 8, another Method. And if a Parabolic

Prop. 8, another Method, And if a Parabolic Section of a Cone, the Inclination of the Plane to the Cone that is cut thereby being chang'd,

chang'd, be turned into an Hyperbola; the Body will be moved about its Center in the Perimeter of the Hyperbola, the Centripetal Force being turn'd into a Centrifugal one, and that Force being greater in a leffer Diftance, and leffer in a greater Diftance; as the Nature of fuch Force doth altogether require.

Coroll. (2.) If the Centripetal Force of any attractive Body whatever be directly as the Diftance, to that in a greater Diftance the Attraction be in the fame Proportion also greater, and in a leffer Diftance lefs; the Body will be mov'd.

in the Ellipsi about a Central Body Corol.3,4,5, placed in the very Center of the Ellip- of Prop. 15. fis, or perchance in a Circle which the

Ellipsi may pass into; for, according to the Situation of the Tangent, of which before, the Body will be mov'd either in a Circle, or in an Ellipsi.

Coroll. (2.) And the Periodic Times of Revodutions made in all Figures about the same Center will be equal, as we also shew'd before. Coral.2, P.9. XX. If a Body be mov'd in a Spiral Lines which cits all the Radii in the fame Angle, the Centripetal Force will be reciprocally as the Cube of the Diftance from the Center of the Spiral For in the divers Parts of the Spirals, the Curvature of like Arches is equal; and that of equal Arches is reciprocally as the Diftance. But whilf Bodies revolve in Spirak, the Celerity will be every where reciprocally as the Diftance; and from thence also the Curvature will, in the given Time, be reciprocally in the duplicate Proportion of the Diffance. Therefore the Centriperal Force; which proceeds from the Proportions of the Cuntanire and Celerity conjunctly, will be in the triplicate Proportion of the Diftance recipro-1:00 S cally,

cally, or reciprocally as the Cube of the Diftance.

Corollary. If the Force of any attractive Body be in the triplicate Proportion of the Diftances from the Center reciprocally, all Bodies, the Directions of the Projectile Motions whereof are not perpendicular to the Radii, with what Velocity foever they go forth, will be mov'd in a Spiral which cuts all the Rays in a given Angle; and if the firft Body afcends, it will afcend infinitely; if it defcends, it will defcend to the Center in a Space of Time eafily to be found from the Quantity of the Spiral Area.

Scholium. If there were any regular Curve, the Curvature whereof from any Central Point whatever were in a duplicate Proportion of the Diftance directly, any Body whatever would revolve in it, if fo be the Centripetal Forces were amongst themselves in the reciprocal Proportion of the Diftances. For if the Curvature, in equal Angles, were according to the Hypothefis in the duplicate Proportion of the Diftance directly, the Curvature would in a given Time be always equal to it felf in all Diftances; and fince the Velocity is always reciprocally as the Diftance, the Centripetal Forces, to be estimated from the Curvature and Velocity conjunctly, would be as the Diffance reciprocally, and the Body would be movid in that Curve. D. E. D. Stan and 

So likewife if there were any regular Curve, the Curvature whereof from any Central Point whatever, were in the triplicate Broportion of the Diftance directly, any Body whatever would revolve in it, if fo be the Centripetal Force were in all Diftances equal. For if the Curvature in equal Angles were, by the Hypothefis, in the triplicate Proportion of the Diftance directly, the Cur-

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Coroll

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Curvature in the given Time would always be directly as the Diftance; and fince the Velocity is always as the Diftance reciprocally, the Centripetal Forces, by reafon of the Equality of the direct and reciprocal Proportions, would always be equal, and a Body would be mov'd in that Curve.

XXI. If a Body be mov'd in an Ellipsi about its Focus, the Centripetal Force will be every where in the duplicate Proportion of the Diftance from the same Focus reciprocally. For, as we nored above, the Curvature with refrect to the Focus in divers Parts of Ellipses, Parabola, and Hyperbolz, is every where in like Arches directly as the Diffance, and in equal Parts always equal: Now the Velocity every where is in the reciprocal Proportion of the Diftance. Therefore in Arches described in the same time, the Curvature is reciprocally as the Diftance from the Focus, and the Celerity is likewife in the fame reciprocal Proportion: From whence the Centripetal Force, to be estimated from the conjunct Proportions of the Curvature and Celerity, will be in the duplicate Proportion of the Diftance from the Focus reciprocally. Q. E. D.

Coroll. (1.) If the Force of any attractive Body whatever be in the duplicate Proportion of the Diftances from the Center reciprocally; all Bodies, at leaft where the Directions of the Projectile Motions are not perpendicular to the Radii, whatfoever Velocity of Motion alfo they may have, will be mov'd in Ellipfes, one of whose Foci will be possed by the Central Body: unless the Velocity of the Projectile Motions be fo great, as to be able to turn the Ellipfis into Parabola's, or even Hyperbola's,

Coroll. (2.) If a Body, according to the Law of the Centripetal Force here affign'd, be mov'd in an Ellipsi about one of the Foci, the Periodic Time of the Body, moving in the Ellipsi, will be to the Periodic Time of a Body mov'd in a Circle. the Radius whereof is in the Middle betwixt the greatest Distance and the least, or equal to the greater Semi-axis, in the Proportion of Equality. For fince the whole absolute Curvature of the Ellipsis is equal to the Curvature of the Circle. and the Sum of the absolute Velocities in equal Arches above and below the mean Diftance, becaufe of the equal Change of the Motion in an equal Arch, is always equal to the Velocity in a mean Circle; it is manifest, that the Centripetal Force is equal; and confequently that the Periodic Times are also equal one to another. Or we will rather demonstrate it thus: Let the fame absolute Velocity be supposed in the mean Diftance, which is in a Circle defcrib'd with the fame Semi-diameter; the Angle then, according to the Conic Properties, or Area described in the Circle, will be to the Angle or Area defcrib'd at the fame time in the Ellipsi, as the greater Semi-axis is to the leffer : and in the fame Proportion alfo, according to the Conic Properties, is the entire Area of the Circle to the entire Area of the Ellipfis. From whence, becaufe of an equable Defcription of Areas on both Sides, the Periodic Times also will be on both Sides equal.

Coroll. (3.) Therefore the Periodic Times in Ellipses are between themselves in the Sesquialteral Proportion of the greater Axes, as well as in Circles.

Coroll. (4.) Confequently the greater Axis being given, there is given withal the Periodic Time.

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Coroll.

Coroll. (5.) Seeing the Proportion of the Curvature and Celerity in a Parabola and Hyperbola is the fame, with respect to the Focus; by the fame Reason as before, a Body will be mov'd in a Parabola and Hyperbola about the Focus.

Scholium. Having now dispatch'd in a more easy Method, the Demonstrations of the fundamental Propositions of Sir Isaac Newton; I will take Liberty for a Conclusion, to adjoin another Demonstration of the last Proposition, which is the most Noble of all, and most of all accommodated to the Mundane System; which Demonfitration comes more up to Geometrical Rigor, and is that which I once transcrib'd from a Manufcript of Sir Isaac Newton's himself.

# A Proposition.

If any Body whatever be attracted towards the Focus of an Ellipfis, and if the Quantity and Proportion of the Attraction be fuch, that they make the Body to revolve in the Elliptic Perimeter; the Attraction in the leaft Diftance will be to the Attraction in the greatest, both Diftances being taken at the greater Axis, as the Squares of the Diftances of the Body in those Points from the Focus of the Ellipfis reciprocally.

Let (in Fig. 2. Plate 5.) A E C D be an Ellipfis; A and C the Extremities of the greater Axis: F the Focus whither the centripetalForce tends; and A F E, C F D, those Areas which a Body doth by Radii drawn to the Fo-Prop. 15. cus, describe in an equal Space of Time. Now those Areas are equal one to another, as being proportional to equal Times; that is, the

Scholium to I. 41. Elements. VI. 14. Elements, the Rectangles  $\frac{1}{2}$  AF  $\times$  AE, and  $\frac{1}{2}$  FC  $\times$  DC, are equal to each other; that is on the Hypothesis that the Arches AE and CD are taken so shall, that they may fafely enough be reckon'd for Right Lines.

Therefore A E is to C D, as F C is to F A. Let us now suppose the Right Lines A M and C N to touch the Ellipsi in the Points A and C, and the little Lines E M and DN [ to be supplied in the Figures ] to be from the Points E and D perpendicular to those Tangents. Now because the Curvature of Ellipses (that is, if we confider it in general, and in equal Arches with refpect to the Center of the fame ) is equal at both Extremities, these Perpendiculars E M and DN will be betwixt themfelves, (Coroll. 4. Prop. 2.) as the Squares of the Arches A E and C D. EM therefore is to DN, as FC square is to FA square. But in the fame time, in which the Body will from the Force of Attraction describe the Elliptick Arches A E and CD from A to E, and from C to D: the fame without that Attraction would have defcribed the Tangents AM and CN equal to those Arches. The Forces of the Attractions therefore which draw the Body back from the Tangents to the Curve, to wit, from M to E, and from N to D, are also betwixt themselves as those little Lines ME and ND subtending the Angles of Contact, which are describ'd at the fame time. The Attraction therefore at the Point A is to the Attraction in the Point C, as the little Line ME to the little Line ND; that is, by the Things already demonstrated, as FC square is to FA fquare; or as the Squares of the Diftances reciprocally. Q. E. D.

This

This Demonstration respects only the Extremities of Ellips; those which follow will apply the same Proposition to any Parts of Ellipses whatsoever.

Lemma. If a right Line touch an Ellipsi in any Point whatsoever, and if a Line be drawn through the Centre of the Ellipsi parallel to the Tangent, which may intersect a 3d Line drawn through the Point of Concact and either of the Foci; that Part of the 3d Line which is posited betwixt the Contact, and the Intersection, will be equal to half the greater Axis. Let (Fig. 2. Plate 5.) APCQ be an El-

Let (Fig. 2. Plate 5.) APCQ be an Ellipfis: AC the greater Axis: O, the Center: Ff the Foci; P the Point of Contact: OG the Line parallel to the Tangent; and PG that Part of the Line FP, which lies betwixt the Contact and the parallel to the Tangent. I fay that PG is equal to CO, or to half the greater Axis.

For join the Points PF; and draw the Line f H parallel to OG. Because the Lines Ff and FH are bisected in the Points O and G, AC will be equal to the Sum of the Lines PF and Pf, that is, to the Sum of the two Lines PF and PH, (by the Conics) or to the double of the Line PG. And therefore the half of AC, that is CO, is equal to PG. Q. E. D.

Another Lemma. Any right Line whatever drawn through either of the Foci of the Ellipsis to the Periphery, is to the Diameter of the Ellipsis which is parallel to the fame, as the fame Diameter is to the greater Axis of the Ellipsis.

Let APCQ (Fig. 2. Plate 5.) be an Ellipfis: AC the greater Axis : Ff the Foci : O the Center : PQ any Line drawn through the Focus F: M VOS

VOS the Diameter of the Ellipfis parallel to PQ. Here PQ. VS. AC, will be  $\ddagger$ . For let f p be drawn parallel to QFP, cutting the Perimeter of the Ellipfis in the Point p. And join the Points P p by the Line P p cutting VS in the Point T. Then draw the Line PR which may touch the Ellipfis in the Point P, and cut the Diameter VS produced in the Point R. There will now be by the Conics OT: OS: OR  $\ddagger$ . But OT is half the Sum of FP and fp, or of FP and FQ : and confequently OT doubled is equal to PQ. OS alfo doubled is equal to VS, and by the Lemma before demonftrated, OR, or PG doubled, is equal to AC. Wherefore PQ: VS: AC are  $\ddagger$ . Q. E. D. Coroll. AC x PQ=VSq=4OSq.

Lemma. (3.) If the right Line FP be drawn from either of the Foci of the Ellipfis to any Point in the Perimeter thereof : And to the Point P the Tangent of the Ellipfis Px; and if the little Line xy (See Fig. 2. Plate  $\varsigma$ .) fubtending the Angle of Contact, parallel to the Line PQ; the Rectangle of the fubtending little Line, and of the fame Line produc'd to the remoter Part of the Perimeter, is to the Rectangle of the greater Axis of the Ellipfis, and of the firft Line which was produc'd alfo to the Perimeter of the Ellipfis, as the Square of the Perpendicular Diftance betwixt the fubtending little Line, and the firft Line is to the Square of the leffer Axis.

For let AKBL be an Ellips: AB the greater Axis: KL the leffer : G the center : Ff the Foci : P any Point defign'd in the Perimeter : F P the first Line drawn through the Focus F to P : PQ the fame Line produc'd unto the Elips: P x the Tangent : x y the little Line fubtending the Angle of Contact : xI the fame fubtence pro-

produc'd to the Remoter part of the Perimeter : vz the Perpendicular Diftance of the Subtenfe and the first Line. These things supposed, I say that the Rectangle y x I is to the Rectangle **ABxPQ**, as is y z Square to KL Square. For let V S a Diameter of the Ellipsis be parallel to the first Line, and GH another Diameter parallel to the Tangent S x, or the conjugate Diameter to the former Diameter. The Rectangle yxI will then.by the Conics, be to P.x Square, or the Square of the Tangent, as the Rectangle SCV is to the Rectangle GCH; that is, as SV Square is to G H Square : Now, by the Conics, all Parallelograms describ'd about the conjugate Diameters of every Ellipsis are also equal betwixt themselves. From whence it follows, that the Rectangle of the double of P E drawn into G H will be equal to the Rectangle of the Axes ABxKL: And confequently ( by VI. 14. of the Elements ) GH is to KL as AB, that is, by the first Lemma, as the double of PD is to the double of PE; or, because of the likeness of the Triangles yzP and **PED** (when the Point v coincides with the Point P) as Px is to yz. Therefore Px is to GH as yz to KL : And confequently Px Square is to G H Square, as y z Square is to K L Square, (VI. 22. of the Elements). But by what hath been already affum'd, P x Square is to GH Square, as the Rectangle yxI to SV Square: And SV Square (by the Corollary of the 2d Lemma) is equal to the Rectangle of  $AC \times PQ$ . Therefore the Rectangle y x I is to the Rectangle AC×PQ, as yz Square is KLq. Q. E. D. Coroll. (1.) If yz be given, and confequently

yz Square, yx Square will also be given, and confequently yx it felf: That is, if the least Perpendicular Distance of a Point taken in the M 2 Elliptic

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Elliptic Perimeter from the Line drawn through the Focus, be given in divers Diftances from that Focus whatfoever; there will also be given the vanishing little Line, fubtending in the fame Place the Angle of Contact. For by what hath been just now demonstrated, since y z is given by the Hypothesis, and KL is also given; and since, as the Rectangle yx x XI is to the Rectangle AC'x PQ, fo is yz fquare to KL fquare : It will alfo be, the Line x I ending at last in the Line PQ, as  $yx \times PQ$  is to  $AC \times PQ$ , fo is yz fquare to KLq: But as y x × P Q is to A C x P Q, (VI. 1. Elem.) fo is y x to A C. Therefore, as y x is to A C. fo is y z fquare to K L fquare; and by inverting, as KL square is to yz square, so is AC to vx: fince therefore the reft of the Things are given, the Subtenfe y x will also be given. Q.E.D.

Coroll. (2.) I may also in this Place infer, that the Curvature of an Ellipsi with respect to the Focus is every where in the Proportion of the Diftance from the Focus directly. For fince  $\mathbf{y}$  z the vanishing Subtense of the Angle of Contact in a given Perpendicular Diftance, in all Diftances from the Focus is the fame; y x in Diftances pro-portional to the Radius F P at equal Angles, will be ( by Coroll. 4. Prop. 2. ) in the duplicate Proportion of the Radii. From which duplicated Proportion, there being taken away, as it ought to be, the Proportion of the Radius, the Proportion of the Curvature in divers Diftances will be left; the fame, to wit, with the direct Proportion of the Radii. Although therefore the Curvature of divers Circles in the fame Angles, is with refpect to the Center every where equal; yet in Ellipses, on the contrary, it is continually changed in divers Distances from the Focus, and in a greater Diftance becomes greater, in a leffer Diftance lefs; and

and this in the Proportion of the Increase or Dimunition of the Distance; as we noted before.

Coroll. (2.) To conclude, we may transfer both. the foregoing Corollaries to a Parabola and Hvperbola: For what Things have been once demonstrated of an Ellipsi, are to be understood to agree to a Parabola alfo; because of the Coincidence of Ellipses infinitely long with Parabola's : And then the Affections of Ellipse and 'Parabola's, becaufe of the mutual Agreement of all Conic Sections, are to be applied to Hyperbola's, changing those Things which the nature of the Line requires to be changed. Wherefore I may now affert, that the vanishing Subtense of the Angle of Contact at all equal Perpendicular Diftances from the Radius, is in every Conic Section always equal to it felf; and that the Curvature confequently in equal Angles is in the direct Proportion of the Diflances.

Feb. 5. 1703.

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#### LECT. XV.

*Cholium.* By almoft the fame Reafoning whichSir*IJ.Newton* made use of for finding out the Proportions of the vanishing Subtens with respect to the Focus of the Ellips; I may undertake to determine the Proportions of the fame Subtenits in Ellips with respect to the Center, M 3 by

by what he hath demonstrated. Princ. Mathe. viz.that y z fquare, drawn into SC Book 1. Pr. 10. fquare, (Fig. 4. Plate 5.) and •then applied to the Line y x; is equal to the double of KC fquare drawn into CB fquare, and then applied to the Line SC; or yzq x SC cube  $=_2 K C q \times C B q \times y x$ . If therefore z y be given, and confequently zy square, because of the Solid 2 K C q × C B q which is also given; y x will be every where as SC cube, or in the triplicate reportion of the Diftance directly. If therefore zy be the n as it ought to be; because the Subtense of the Angle of Contact, is in the duplicate Proportion of the Arch, the Subtenfe y x will be in the quintuplicate Proportion of the Diftance; or the Proportion of the Diftance being Subducted, the Curvature it felf will ftill be in a quadruplicate Proportion of the Diftance directly; or as the biquadrate of the Diftance directly.

Another Proposition. If a Body be drawn to one of the Foci of an Ellips, and from that attraction revolve in the Elliptic Perimeter, the Forces of the attraction will be every where, as the Squares of the Diftances from the fame Focus directly.

For (in Fig. 4. Plate  $\varsigma$ .) Let P be the Place of a Body revolving in an Ellipfis at any Moment of Time, and PX the Tangent of the Ellipfis in that Point; along which Tangent the Body would go forward with an uniform Motion, if it were affected with no attraction: Let the Point X be the Place whither the Body would reach in a given very fmall Space of Time; and let Y be the Place in the Perimeter of the Ellipfis, whither it doth from both Forces actually reach in the given Time. Let the Time be divided into very fmall qual Parts,

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that

that they may be accounted Physical Moments or Points : Let the attraction now act not continuedly, but by Intervals, but those very little ones; once to wit, in the beginning of every Physical Moment; fo that the first Force of the Attraction may Act at the Point P; the 2d at Y; and fo at equal Intervals perpetually : So that the Body may be mov'd along the Chord of the Arch PY, and then along the Chord of the following Arch, and fo on. Now becaufe the attraction in the Point P is directed towards the Point F, and draws down the Body from the Tangent PX unto the Chord PY; the little Line X Y produc'd by that Force of the attraction in P will be proportional to that Force, and parallel to its Direction, that is to the Line PF. Produce the Lines XY and PF unto the Elliptic Perimeter in I and Q; join the Points F, Y: and let YZ be let down Perpendicular to FP. (Fig. 4. Plate 5.) Let A B be the greater Axis of the Ellipsi, and KL the lesser. And by Lemma the 2d, the Rectangle Y X I will be to the Rectangle AB×PQ, as is YZ fquare to KL fquare. And confequently the Line  $\mathbf{Y}\mathbf{X}$  will be equal to the Solid of  $AB \times PQ \times YZq$ , applied to the Solid XI × KLq. In the fame manner, if py be the Chord of another Elliptic Arch, which the Body defcribes in the given Physical Moment of Time equal to the former; and px the Tangent of the Ellipsis in the Point P; and xy the Subtenfe of the vanishing Angle of Contact parallel to the Line pF; and if xy and pF cut the Perimeter of the Ellipsi in q and i; let y z from the Point y be let down perpendicular. to pF; the Subtenfe yx will by the former Reason be equal to the Solid made of AB into pq x yz fquared, applied to the Solid made of хi M A

x i x K L fquared; that is, becaufe A B and K L are given and ftanding Quantities, as  $\frac{PQ}{XI}$  Y Z fquar'd to  $\frac{pq}{xi}$  yz fquared. But becaufe the Lines P Y, py, are definited by a revolving Body in equal Times, the Areas definited, or Triangles PYF, pyF are equal: And the Rectangles P E xYZ, and pFxyz are double those Triangles they are equal: And YZ is to y z, as pF to PF; and confequently  $\frac{PQ}{XI}$  Y Z fquared, is to  $\frac{pq}{xi}$  yz fquar'd, as  $\frac{PQ}{XI}$  pF fquared is to  $\frac{pq}{xi}$  pF fquared.

Therefore, as YX is to yx, fo is  $\frac{PQ}{XI} pF$ fquared to  $\frac{Pq}{xi} PF$  fquared; that is, the Attraction in P is to the Attraction in p, as  $\frac{PQ}{XI} pF$  fquare is to  $\frac{Pq}{xi} PF$  fquared. Now, fuppofe the Times taken infinitely fmall to be equal, in which a Body defcribes the Subtenfes PY and py; fo that the Attraction may be continued, and the Body revolv'd in the Perimeter of the Ellipfis:

In this Cafe the Lines PQ and XI coincide; but p q and x i have been already fuppofed equal, therefore the Quantities  $\frac{PQ}{XI}$  pF fquared, and  $\frac{pq}{xi}$ PF fquared, become pF fquared and PF fquared. Therefore the Attraction in P, or the Line XY, will be to the Attraction in p, or the Line xy, as pF fquared is to PF fquared; or reciprocally, as the Squares of the Diffances from the Focus. Q. E. D. And

And the fame Proposition may be applied to the Parabola, as the Extreme of Ellipses, and alfo to the Hyperbola: But fince there are no Cœleftial Bodies that we know of, that are carried about in Hyperbola's; I shall not fearch out for a particular Demonstration of them.

XXII. The Velocity of a Body moving in a Parabola about a Body placed in the Focus, the Force whereof is in the reciprocal duplicate Proportion of the Diftances, is every where to the Velocity of a Body revolving in a Circle in the fame time, in the fubduplicate Proportion of the Number, Two to Unity; or as the Diagonal of a Square to its Side; that is, as 10 to 7 nearly.

For, fince the Diftance of a Body from the Central Body was fuppoled every where the fame, the Force of Attraction, or Line fubtending any Angle of Contact, will be always equal in any given Space of Time: And the Velocity in **2** Parabola will be to the Velocity in the Circle, as the Tangent of the Parabola to the Tangent of the Circle; to wit, where the Subtenfe is every where equal.' But the leaft Tangent in the Parabola, by the Conics, is equal to the fquare Root of the Rectangle of the Subtenfe drawn into the Latus Rectum of that Vertex. And the leaft Tangent in the Circle is equal to the fquare Root of the Rectangle of the Subtenfe drawn into the Diameter

of the Circle. But becaufe both III. 36. Elem. Subtenfes are given, and the Latus

Rectum of the Vertex of the Parabola is, by the Conics, double to the Diameter of the Circle; or as two to one: The first Rectangle will be double to the last, or as 2 to 1; from whence the Tangents or square Roots will be among themselves, as the square Root of the Number two to one, or as the Diagonal

Diagonal of a fquare to its Side ; that is as ten to feven nearly.  $\mathcal{Q}$ . E. D.

Coroll. (1.) Since therefore the Velocity in a Parabola is to the Velocity in a Circle, at the fame Diffance from the Focus, in a given Proportion; to wit, as  $\Upsilon$ . 2 to 1. And fince the Velocity in divers Circles is in the fubduplicate reciprocal Proportion of their Radii, the Velocity of a Body defcribing a Parabola at divers Diffances from the Focus will also be in the fubduplicate reciprocal Proportion of the Diffances.

Coroll. (2.) The Velocity of a Body revolving in an Ellipfis, is lefs than in a Parabola; and in an Hyperbola, greater at the fame Diffance from the Focus: From whence the Velocity in an Ellipfis, will be to the Velocity-in a Circle in a lefs Proportion than  $\Upsilon$ . 2 to I; and in an Hyperbola; in a greater at the fame Diffance from the Focus.

Coroll. (2.) Therefore the Velocity of a Body, at any Diftance from the Focus, being known, the Figure of its Trajectory may be allo known; to wit, whether it be a Circle, Ellipfis, Parabola, or Hyperbola: And from a more accurate Calculus, if it be an Ellipfis, or Hyperbola, what Species of those Figures it is that a revolving Body ought to defcribe.

*Coroll.* (4.) It follows from what was juft now demonstrated, that if any Body be moved according to any right Line whatever (unless it tends directly to the Focus) with any Velocity, and be acted upon at the fame time by a Centripetal Force reciprocally proportional to the fquare of the Diftances from the Center; the Body will be moved in one of the Conic Sections, having the Focus in the Center of Forces: To wit, if the Line, according to which the Projectile Motion of the Body tends, be perpendicular to the Radius, and the Velo-

Velocity equipolient to the Attraction : that is, if the Velocity in any given very fmall Space of Time, be equal to the fquare Root made of the Subtense of the Angle of Contact of that Circle drawn into its Diameter, that Body will be moved in a Circle. But if the Velocity be equipollent to the Attraction, and the Line of Direction oblique to the Radius, the Body will be moved. in an Ellipsi, the periodic Time whereof wou'd be equal to the periodic Time of the Circle in which it will move. But if the Velocity be greater or lefs than the Velocity before affigned, to neverthelels where it is greater, that it be not increased above the Proportion of the square Root of the Number Two to Unity, that Body will be moved in an Ellipsi, in the first Cafe greater, and in the laft leffer than the Circle. But if the Velocity be to the Velocity in the Circle, as the square Root of 2 to 1, the Body will be mov'd in a Parabola. If, laftly, the Velocity be greater, the Body will be mov'd in an Hyperbola.

XXIII. Problem. The Centripetal Force being reciprocally proportional to the fquare of the Diftance from the Center, to define the Times in which Bodies in falling ftrait down wou'd reach the Center. (See Fig.  $\varsigma$ . Plate  $\varsigma$ .)

Upon the fame principal Axis or transverse Diameter, A B, let there be described the Extreme Ellips; to wit, the Circle A D B, and the Right Line A B; from the Equality of these Transverse Diameters, the Periodic Times will be equal on both Sides; and consequently the Times of the half Revolutions will be equal to one another. (Coroll. 4. Prop. 21.) That is, the Time of Descent by the Diameter, is equal to the Time of Revolution along the Semi-circumference. Since therefore by what hath been demon-

monftrated before, it is easy to determine that Time of the half Revolution, it will also be easy to define the Time of the direct Descent. As for Example : The Time of the Lunar Half-period contains 1967115 Minutes; where, to wit, the Diameter of the Orbit is double to the mean Diftance from the Center of the Earth. And this Time is to the Time of the Half-period for half the Diftance, which is the thing we now enquire for, in the fequi-alteral Proportion of the Diftances; that is, almost as 2828 is to 1000, or as 19671| 5 is to 6955] 5. From whence the Time of the Half-period, in half the Diftance (or in cafe the Diameter of the Orbit was but half of what it is.) that is, the Time of the direct Fall to the Center of the Earth of a Body placed at that Diftance from the Earth in which the Moon is really placed, will contain 695515 Minutes; or 4 Days, 19 Hours, 55', 30". In this Space of Time would the Moon, if the Motion thereof was ftopp'd, and the Earth remain'd unmoveable. fall to the Center of the Earth. And by the fame Reasoning, the Time of the Fall of any Planet may eafily be determined, as is actually done in the following Scholium.

Scholium. Since therefore the Time of the Half-period of every Planet diminish'd in the Proportion of 1000 to 2828, is the Time of the direct Fall to the Center, the following Table, which is built upon that Foundation, will shew the Times of the Fall of all the Planets to their respective Centers.

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Mer-

Day. Hours. Mercury ١٢ 12 Venus 17 39 would fall to The Earth 64 14 the Sun in the Mars II Space of 12 Fupiter Saturn 1900

Of the Planets about Jupiter.

The inmost _ would fai	1 to ("	00:7
The inmoft The Second Would fail The Third Space of The Fourth	the)	00:15
The Third (Space of		1:6
The Fourth ) Space of	L	2:23

Of the Planets about Saturn.

The Inmost)	C	0:8
The Second (		0:I2
The Second The Third The Fourth Saturn in		0:19
The Fourth Saturn III	)	2:20
The Fifth		14 <b>:</b> I

The Moon, as above, would fall to the Earth in the Space of 4 Days, 20 Hours.

Feb. 19. 1703

#### LECT. XVI.

XXIV. A A Centripetal Force is reciprocally proportional to the fquare of the Diftance, to define the SpaceswhichBodiesfalling right

downwards would describe in any given Time.

If the Body doth not fall perpendicularly, it will defcribe fome Conic Section, the nether Focus whereof (because of the Descent of a proiectile Motion which is here fuppos'd ) will agree with the Center of Force, as is manifest from what goes before, Prop. 21. Let that Conic Seetion (fee Fig. 5. Plate 5.) be the Ellipsis ARBP : where the Velocity of the Projection is to the Velocity wherewith the Body would revolve in a Circle at the fame Diffance, in a lefs Proportion than is the square Root of the Number Two to Unity, (Coroll. 2. Prop. 22.) Let S be the nether Focus of the Ellipsic, and upon the greater Axis of the Ellipsis AB, let there be defcrib'd a Semi-circle A D B. And the right Line DPC being suppos'd to pass through the falling Body perpendicular to the Axis, and the Lines DS and PS being drawn to the Center, the Area ASD will be proportional to the Area ASP, and confequently to the Time. For (by VI. 1. Elements,) as CD is to CP, fo is the Area of the Triangle SCD to the Area of the Triangle SCP. And according to the Conics, as the fame CD is to the fame C P, fo is the Circular Area C A D to the Ellip-

Elliptic Area CAP; and confequently ASD the Sum of the former Areas, will be to ASP the Sum of the latter, as CD to CP, (V. 12. of the Elements;) or as the greater Axis of the Ellipsis to the leffer Axis of the fame; and confequently in the given Proportion, or proportional to the Time. Now A B the greater Axis of the Ellipr fis, or Diameter of the Circle remaining, let the Latitude or leffer Axis of the Ellipfis be perpetually (diminish'd; here, by the Force of what hath been already demonstrated, A S D will still remain proportional to the Time. Yea, let the Latitude be diminish'd infinitely, fo that the El-liptic Orb A P B may at length coincide with the Axis A B: and the Focus S with B the Term of the Axis: the Body will defcend in the right Line AC: and the Area ABD will become in this Cafe also proportional to the Time. From whence, if a right Line perpendicular to the Axis as CD, be fuppos'd to be mov'd downwards always parallel to it felf, fo that the Area ABD fhould be every where proportional to the Time, the Point C will determine the Place, unto which the Body would reach that falls downwards to the Center in the same given Time.

As for Example. Let A B the mean Diftance of the Moon from the Center of the Earth, be as before about 1257 696 000 feet; it is required that we should determine the Place of the Moon falling straight down, in the first Day of the Fall. It is known from what hath been already demonfirated, (Corol. 7. Prop. 2.) that if the Motion of the Moon should cease, it would fall about 16[I English Feet in the Space of one Minute. From whence (by Coroll. of Prop. 5. Select Prop. out of Archimedes,) the Circular Area A B D will be of about 39 483 812 704 000 square Feet, [as being equal

equal to the Rectangle of cd drawn into the half of A B. From whence, feeing there are 1440 Minutes in a whole Day, the Circular Area ABD belonging to the whole Day will be of about 128856690292760000 square Feet; but the given Time is one whole Day, or 1440'. If therefore we can define the Point D, fo that the Area ABD should be of 128856690292760 000 fquare Feet, the Sine of the Arch A D, that is, CD will determine the Line AC which is describ'd in that Time, as being the versed Sine of the fame Arch. But that Area is equal to the Rectangles  $\frac{1}{2}CD \times OB$ , and  $\frac{1}{2}AD \times OB$ , or to the Rectangle  $\frac{1}{2}CD \times \frac{1}{2}AD \times OB$ . If therefore the given Area be divided by the Semi-diameter OB, the Quotient will give the half of CD and OB. From the Table of Sines therefore that Arch is to be fought, the half of which being fuperadded to the half of its own Sine, will yield that Quotient. But that Quotient is by Calculation of 204909 120 Feet; or by reducing it to a Circle, whole Radius is of 10000000, will contain 3258484 of those Parts. And if in the Tables of Sines, we look for the Sine of 19 Degrees and almost 51 Minutes, the Sine of one Minute multiplied by 11302909 × 1130, will give 2287170 Parts, as belonging to the Arch AD, which is of 18 Degrees and 50', the Sine of which Arch is of 2228165 Parts; fo that the Sum of both will be of 6515225'; the half whereof is 3257667, which agrees with the first Number 3298484 exactly enough. The Line CD therefore is the Sine of 18° 50', and the Line describ'd in that Space of Time is the verfed Sine thereof, which is 525382 Parts long; that is, by reducing it to the Semi-diameter of the Moon's Orbit 33667390 Feet long, that is, 6276

6,376 Miles and 2,110 Feet. And in the fame manner it will be defin'd, in what time the Moon would defcend to the very Center of the Earth. But because we have deduced that before by another and more easy way of Computation, we shall not profecute it any further here.

Coroll. If the Figure R P B be not an Ellipfis, but an Hyperbola or Parabola, the thing will be dispatch'd in the same manner by a Rectangular Hyperbola, or any Parabola; but by reason of the Difficulty of the Practice, and that it is not necessary, we shall pass it by here.

Coroll. (2,) The Times wherein any Bodies would fall to the Center from divers Diftances, are betwixt themselves in the Sesqui-alteral Proportion of those Diftances directly. For the Line Ac, that in a given Time is produc'd at divers Diffances, is reciprocally in the duplicate Proportion of these Distances. From whence cd the least Sine, will be in the subfesquiplicare Proportion of the Diftance reciprocally; and the Area icd x'A B describ'd at the same time, in the fub-duplicate Proportion of the Diftance directly. From whence, fince the entire femi-circular Area ADB is in the duplicate Proportion of the Diftance directly, the Time proportional to the fame will be in the fefquiplicate Proportion to the Diffance directly. Q. E. D.

As for Example. Let another A B be double to this A B; then the vanishing Subtense of the Angle of Contact, or the little Line A c, will be only a 4th Part of the other A c: And the Sine c d will be sub-sefuplicate of c d, or as the Side of the Square to the Diameter; that is, 7 to 10 almost: the Area also  $\frac{1}{2}$  cd x A B will be to  $\frac{1}{2}$  cd x A B well-nigh, as  $2 \times 7 = 14$  is to x 10 = 10; From whence, the Area described in the N greater

greater Diftance will be to the Area in the.lefs. but which is described in the same time, nearly as 14 to 10, or as the Diameter in a Square is to the Side. But the entire Area to be defcribed by the greater Line AD in the Descent, is to the Area to be defcribed by the leffer Line BD in the Descent, as 4 to 1, or 40 to 10. Therefore the Time of the Descent in the greater Distance, will be to that in the leffer, in that Proportion in which the Ratio of 40 to 10 exceeds that of 14 to 10. But the Proportion of that Excels is the fame as that of 40 to 14, or as the Diameter of a Square is to the Quadruple of the Side. From whence the Lines are betwixt themfelves, as the Diameter of a square is to the Quadruple of the - Side; that is, in the fefqui-alteral Proportion of the Diffances directly. 2. E. D.

Coroll. (3.) If therefore we fuppofe any one of the primary Planets, as also of the Secondaries of Saturn and Jupiter, to fall to that Center of its Orbit, and have the Times of that Descent already defin'd and computed; it will be easy from the known Distances of the rest, to define the Times also of their Descent; which thing we have perform'd before upon another Ground, and therefore shall not repeat it again.

therefore fhall not repeat it again. Coroll. (4.) Since therefore the Velocity in an Ellipfis in a mean Diffance from the Focus; that is, the Velocity of a Body falling to O, the Center of an Ellipfis when it ends in a right Line, is equal to the equable Velocity of a Body revolved in a Circle, the Radius whereof is BO; it is manifeft, that the Velocity of a Body falling in O, the very middle of the Space, is equal to the Velocity of a Body revolved in a Circle at the fame Diftance. From whence it also follows, that the Velocity of a Body falling at a remoter Diffance,

is

is less than the circular Velocity, and at a nearer Diftance greater.

XXV. A Problem. If the Centripetal Force be proportional to the Altitude, or the Diffance of Places from the Center directly, to define the Times in which Bodies falling down will defcribe any given Spaces.

If the Body doth not fall perpendicularly, it will describe fome Conic Section, the Center whereof will agree with the Center of Force, as appears from what hath been already faid, Prop. 19. Let (Fig. 6. Plate  $\varsigma$ .) the Conic Section be the Ellipsis A R P B: Let O be the Center thereof; and upon AB the greater Axis of the Ellipfis, let there de described the Semi-circle ABND. and let the right Line DPC pass through the falling Body perpendicular to the Axis. Which done, and the Line DO and PO being drawn to the Center, the Area AOD will, by the Conics, be proportional to the Area AOP, and confequently to the Time. For, as before (by VI. I Elements) as CD is to CP, fo is the Area of the Triangle OCD to the Area of the Triangle O C P. And alfo, by the Conics, as the fame C Dis to the fame C P, fo is the circular Area C A D to the Elliptic Area CAP; and confequently, the Sum of the former Areas AOD will be (V. 12 Elements) to AOP the Sum of the latter Areas. as CD is to CP; or, by the Conics, as the greater Axis of the Ellipsis is to the lesser Axis of the fame; and confequently in a Proportion given, proportional to the Time. Now A B the greatest Axis of the Ellipsi or Diameter of the Circle remaining, let the Breadth of the Ellipsi, or its leffer Axis, be continually diminish'd; and by the Force of what hath been already demonstrated, the Area A O D will remain proportional to the N 2

the Time. And let that Breadth be diminish'd infinitely; fo that the Elliptic Orb A R B P may now fall in with the Axis, the Body will defcend in the right Line A C, and the Area A O D will in this Cafe also be proportional to the Time. From whence, if a right Line perpendicular to the Axis, as C D, be suppos'd to be mov'd always downwards parallel to it felf, fo that A O D may be every where proportional to the Time, the Point C will determine the Place unto which the Body in falling down will reach in the given Time.

Coroll. (1.) Becaule of the Equality of the circular Area that is every where to be described in equal Time about the Center of the Circle; the Motion of the Point D will always be equable, and will describe equal Arches in a given Time.

Coroll. (2.) The Times therefore of Bodies falling and defcribing what Spaces foever, as AC, are betwixt themfelves as the Arches themfelves AD; and the Spaces defcribed AC, are as the verfed Sines of those Arches.

Coroll. (3.) But the Velocities produced in any Places whatever, as in C, are as the right Sines of the Arches A D. For let the Line c d be drawn parallel to C D, at a Diffance infinitely fmall, and let d D the Tangent of the Circle be drawn. Whilft therefore the Point D defcribes the Tangent d D, the falling Body defcribes the little Line c C equal to d e; and becaufe of the given Velocity of the Point D, d D will alfo be given in length for the given Time. In the Triangle therefore d e D, d D the Radius of the Circle will be given, and de the right Sime of the Angle d D e. And becaufe of the Likenefs of the Triangles d e D, C O D, the Radius in that

that Place will be OD, and the Line CD the right Sine of the Angle AOD; Therefore the Velocity in all Points whatever, as C, is as the right Sine of the Arch AD. Q. E. D.

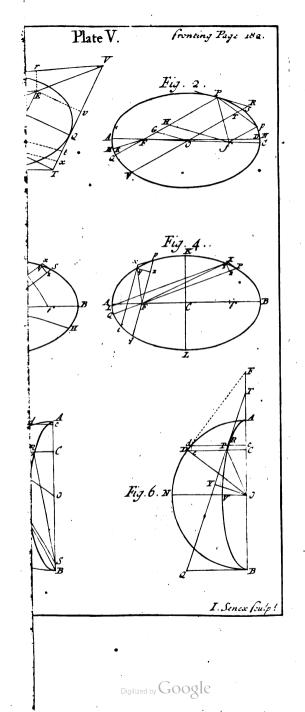
Coroll. (4.) The Times in which Bodies fall from any Places whatever to the Center, are always equal. For fince, by the Hypothefis, the accelerating Force, and confequently the Velocity arifing, is as the Line to be defcrib'd; it is manifeft, that the Times of Defcent are every where equal. 2. E. D.

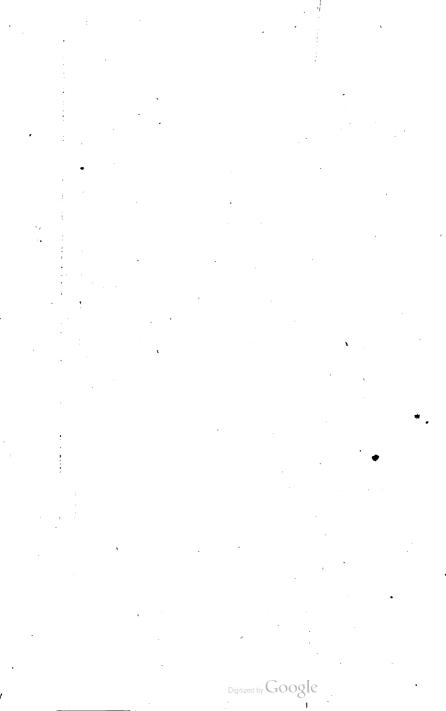
Coroll. (5.) Since, by what hath been demonftrated before, (Corol. 3. Prop. 19.) the periodic Times of all Bodies revolving about the Center of Ellipfes are equal, the Quarters alfo of the periodic Times A BPV will be equal. And fince this is true in all Ellipfes whatever, it will be true alfo in the Extremes of Ellipfes on this fide, and on that; to wit, in the right Line A O, and the Quadrantal Arch A N; that is, the Times wherein one Body in falling comes from any Place whatever, as A unto O, and another in revolving defcribes a Quadrantal Arch, will be equal every where. Q. E. D.

Scholium (r.) Since therefore the periodic Time of the Moon about the Earth, is (per Schol. Prop. 14.) to the periodic Time of any other Body revolving about the Center of the Earth at the Diftance of a Semi-diameter of the Earth, in the fefqui-alteral Proportion of the Diftances; and fince within the Surface of the Earth, the Centripetal Force is every where in the direct Proportion of the Diftance, as will hereafter be demonftrated : It will not be unpleafant to produce an Example of the foregoing Reafoning : and to fhew by Calculation, in what Space of Tine heavy Bodies would defcend to the Center, down N 2

fome empty Hole or Pit which reaches thither. For the finding therefore, according to what hath been already demonstrated, the Quarter of the periodic Time at the Surface of the Earth, as being the Time of the Descent of Bodies from the Surface to the Center of the Earth: Let it be made thus; as is the Cube of the Moon's Diffance.  $60 \times 60 \times 60 = 216000$ , to the Cube of the Semi-diameter of the Earth  $1 \times 1 \times 1 = 1$ : fo is the square of the Moon's Period 29242'x 29 242' = 1 547 871 649 to the square of the Period in the Surface of the Earth  $= 7 \, 166_{1} \, 07$ , the fquare Root whereof 8416 will yield the Minutes in Time, in which a Body or Planet at the Diflance of the Semi-diameter of the Earth from the Center would perform its whole Period about The Quarter of which Number 21115, will it. fhew the Space of Time in Minutes, in which all heavy Bodies whatever would fall thro' the Semi-diameter of the Earth to the Center of the And fince in all Diftances whatever, the fame. Time of falling is the fame, as hath been already fhew'd, (Coroll. 4. foregoing); it is manifest, that all Bodies would defcend, and fall from any Place to the Center in 21 Minutes, and 15 Seconds.

Schol. (2.) But if the Time of the Fall through any given Space whatever be requir'd to be found without the Use of Algebra; as for Example, thro' a 4th Part of the Semi-diameter; feek in the Tables of Sines, what Angle that is, the versed Sine whereof is a 4th Part of the whole right Sine; to wit, the Arch A D which is of  $41^{\circ} 25'$ . From whence the Time of the Fall (see Fig. 6. Plate 5.) along A C, the 4th Part of the Semi-diameter, will be to the Time of the whole Fall to the Center, as the Arch A D is to the Quadrantal Arch A N, (Corol. 2 foregoing) or as  $41^{\circ} 25'$  is to  $90^{\circ}$ .



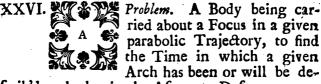


And fince  $90:41^{\circ}25':::21' 15$  horary Minutes: 9' 197 or 9' 58": It is manifeft, that any Body whatever would fall down a 4th Part of the Semidiameter of the Earth in 9 Minutes, and 58 Seconds. And that the Velocity in the Point C is to the greateft Velocity, or that which would be at the Center it felf, in the Proportion of the Right Sine CD to the whole Sine ON, (Corol. 2. foregoing) or as 66152 is to 100000, as is moft manifeft from what hath been just now demonftrated.

April 7. 1705.

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#### LECT. XVII.



fcrib'd, whether in the Ascent or Descent.

Let F ( in Fig. 1. Plate 6.) be the Focus of the Figure, T the principal Vertex, Tl or Ts the given Arch defcrib'd, or to be defcrib'd: Tt or Tq the Abfcifs of the fame Arch, which is alfo given, the Arch being given: tl or qs the Semiordinate, which alfo is given, the Arch being given. The Time wherein the Arch Tl or Ts is defcribed, is requir'd: The Parabola being given, the Latus Rectum of the fame, and confequently TF the Fourth Part thereof is given. From the Centripetal Force of the Central Body, there N 4.

is also given the Velocity of the Body at the principal Vertex; or that which (by Prop. 22. foregoing) is to the Velocity of a Body defcribing a Circle, the Radius whereof is TF, as the fquare Root of the Number Two is to Unity. From whence also there will be given the least Area to be defcrib'd by the Radius TF in any the leaft Time which is given. But the Area FTI or FTs. is equal to two thirds of the Rectangle Tt x tl, or  $Tq \times qs$ . To which, if there be added the Triangle F t l in the former Cafe, and in the latter the Triangle Fqs be taken away from the fame, there will also be given the Area Ftl or Fts; which being divided by the least Area described at the Vertex T, in any very small Time given, will give the Time fought.

As for Example. Let the Parabola given be that which the Comet that was feen see Newton, in Europe in the Year 1680, describ'd p. 494,& 498. in the End of that Year, and the Beginning of the next. Let Fq be equal to the Semi-diameter of the great Orb, to wit, of 10000 equal Parts, fuch as the Latus

Rectum contains 22618; and confequently FT of sol 2 Parts, and the whole Absciss T q of Parts 10,0592. Let us also suppose the Comet to have been in the Vertex of the Parabola, or in its Perihelium T, December 8. Four Minutes after Noon. For the finding the Velocity of the Comet in the Vertex of the Parabola, let there be found in the first place the Velocity of a Planet revolving in a Circle at that Diffance, by this Analogy: as is the square Root of the Distance FT, of  $59|_2$  Parts =  $7|_7$ , to the square Root of the Diftance F q, of 10,000 Parts = 100; Prop. 13. fo is the annual Velocity of the

Earth, to the Velocity of a Planet descri-

describing a Circle, the Radius whereof is FT. Then, as the square Root of the Number two =11414 is to Unity; fo will the Velocity of a Comet in the Vertex of its Parabola be, to the Velocity of a Planet in a Circle at the fame Diftance. But the Earth, by the mean Velocity of the fame. describes | 1195 Parts in the Space of one Minute. and 717:100::1195:1|552. From whence the Velocity of the Comet in its Perihelium, will be that which in the Space of one first Minute describes  $\frac{1|414}{1}$  1|552 = 2|19 Parts, such as the Semi-diameter of the great Orb contains 10000. and fuch that 59/2 of them are contain'd in the least Distance of the Comet. The Area therefore describid in that given Time by the Comet, with a Radius drawn to the Center of the Sun, is equal to the Rectangle of  $\frac{1}{2}$  59[2 x 2[99 =64 824 square Parts. That therefore we may at length find the Space of Time wherein the Parabolic Arch Ts, where Fq is equal to the Semidiameter of the Orbis magnus, is describ'd, we will compute the Area TsF, and compare it with the former Area which was describ'd in one Minute.

Therefore, as TF of  $\varsigma_9|_2$  Parts is to Tq of  $100\varsigma_9|_2$  Parts; fo let the fquare of FH of 118|4 Parts = 14018|\varsigma\_6, be to 2382018|61 fquare Parts; the fquare Root whereof = 1543|3 will, by the Conics, be equal to the Semi-ordinate qs: Which being drawn into half the Diftance Fq 1543[3 ×  $\frac{1}{2}$  10000 = 77165co, the Area of the addititious Triangle Fqs will come forth. But the whole Parabolic Area Tqs is equal to two 3ds of the Rectangle under Tq of 10059|2 Parts, and Sq of 1543[3 Parts, or to the fquare Parts of  $\frac{1}{2}$  15524363[36 = 10349575[57. From which Number, let there be deducted the Triangle Fs q

of 7716500 fquare Parts, the remaining defcribed Area will be of 2633075 57 fuch Parts; which being divided by the Parts of the Area belonging to one first Minute, there comes forth  $\frac{26330755}{641824}$ 

the fpace of Time fought; or that in which the Comet would defcribe the Arch Ts = 4061[9] $= 28^d$ . 4<sup>h</sup>. 59'. Wherefore the Arch Ts will be defcrib'd in 28 Days, and almost 5 Hours. And the Comet posses of the Point s on the 5th of January, about 4 hours Afternoon. Which also exactly agrees with Sir Isac Newton's Scheme deduc'd from Observations.

If therefore we, from fuch Calculations, fhall once have the Times rightly defin'd, wherein any Comet defcribes any Arches whatever, as T s of a Parabola, or rather an Ellipfis, fo eccentric that it may fafely be reckon'd for a Parabola, by the inverfe Method we may be able alfo to define exactly enough, the Arches belonging to any given Times; I mean the fame way of working by which, in *Kepler*'s Hypothefis and Tables, we are wont to find the Ceoquate Anomaly of the Planets from their mean Anomaly in a given Ellipfis.

Coroll. (1.) Seeing therefore the ablatitions Triangle Fs q vanisheth away in the Point h, the Area to be computed at that Time will be equal to two 3ds of the Rectangle of TF drawn into Fh; or  $\frac{2}{3}$  5912 × 11814 = 467618, and confequently the Time belonging to this Area will be equal to  $\frac{467618}{64[824]}$  = 1 h. 12'. 9". So that T h, the Arch betwixt the principal Vertex and the Ordinate to the Axis through the Focus, was defcribed in 1 h. 12'. 9". And the Comet poffes'd the

the Point s December the 8th, 17 Minutes after One in the Afternoon.

Coroll. (2.) Hence also the Space of Time wherein any given Arch is defcribed, doth eafily become known; viz. by computing the Time from the Perihelium to both Places, and taking away the fhorter Time from the longer: For by that means the Interval of Time belonging to the given Arch will become known. Thus the Time agreeing to the Arch Th = 1 h. 12'. 9". being deduc'd out of the Time agreeing to the Arch Ts =  $28^d$ .  $4^n$ . 59', the Remainder is the Interval of Time belonging to the Arch h s =  $28^d$ .  $2^h$ . 46'. 51''. And fo every where.

Coroll (2.) Hence also is deriv'd the Method of finding the defcribed Arch from the given Time. For feeing that at the Point h the ablatitious Triangle Fqs, and the addititious one Ftl, doth always vanish away; and confequently the Area in that Place may be eafily computed, as being in our Example of 4667 4516 square Parts: Since also in that Place TF is half FH: and fince, laftly, the Abscifs TF doth always increase in the fame Proportion, in which the square of its Ordinate FH increaseth : any Time whatever, or the Area proportional thereto, being given, the Arch belonging to the fame will alfo be given; if that Quantity of proportional Increments or Decrements be taken, that  $\frac{1}{2}$  qs x Fq being taken out of  $\frac{2}{3}$  qs × Tq the Remainder be the Quantity of the given Area. Thus, that I may find the Arch T of  $28^{d}$ .  $14^{h}$ . 59'' = 40619', that is, that which belongeth to the Area of 2633075 57 square Parts; I seek in the Tables of fquare Numbers, if I would work without Algebra, where fuch a Number is to be found, (the Line TF being taken for Unity, and the Area FT

FT h for the fquare of Unity; or for 1 TF × Fh = a c62d Part of the whole Area; and Fh being taken for the Number two : ) That the Numbers to be added to Unity being proportional, the fquare of the Numbers to be added to two  $\frac{1}{2}$  g s x q F being taken out of  $\frac{1}{3}$  qs x T q, the Remainder may be the given Area = 562: Which Number will occur no where elfe, but where F a is to FT as 10000 to 59/2, or as 167 to 1 nearly. From whence it is manifest, that the fought Arch is no other but that of which Tq of 10059 2 Parts is the Abscis. But fince this Method confifts in making Essays, and is indirect, it is not fo artful. However, what hath been here delivered contains enough in it, to fhew in some measure the Origin and Method of compiling Tables.

Scholium. Note, That Sir Isaac Newton's Geometrical Method doth fhew directly, from the given Time, the describ'd Arch; that is, if F T be made to ty, as the Time belonging to the Area Th F is to the given Time, the Point t posses the Middle of the Line T F, and ty being drawn perpendicular to T F; the Distance from the Focus F will be equal to ys: From whence the

See Newton, Book I. Prop. 30.

Circle describ'd from that Radius will determine the Point. But fince this Method is not fo fit for Calculation, we pass it by in this Place.

Scholium. Hitherto we have chiefly expounded the Motions of Bodies attracted unto an unmoveable Center, fuch as fcarce is in the whole Compass of Nature. For Attractions are wont to be unto Bodies, and the Actions of the Bodies attracting and attracted, are always mutual and equal, as we shew'd before, (Law of Motion  $\varsigma$ .); fo that neither can the Attrahent rest, nor the Attracted,

tracted, if they be two Bodies; but must both, as it were, by a mutual Attraction, where the projectile Motion of both is duly impress'd upon them, be revolv'd about the common Center of Gravity. And if there be more Bodies, ( which are either attracted by a fingle one, or attract each other mutually) these ought to be so mov'd among themselves, that the common Center of Gravity fhould either reft, or be mov'd uniformly in a ftraight Line, as we shew'd before, ( Law of Motion 25.) For which Reason, we proceed to fet forth the Motion of Bodies, as mutually drawing each other; confidering the Centripetal Forces as Attractions, altho' indeed speaking physically, they may perhaps be more rightly called Impulses. For we look not here fo much at the physical Caules, as at the Effect it felf, confidering and measuring things in a Mathematical Way, and using an easy and familiar Term, though in the ftrict Notion of it perhaps it may not agree.

XXVII. Two Bodies attracting one another, defcribe like Figures both about the Common Center of Gravity, and about one another; that is, whilft they really defcribe like Figures about the Common Center of Gravity, the Eye being placed in either of the two, and not perceiving, its own Motion, or that of the Center of Gravity, a Figure like to the fame will feem to it to be defcrib'd.

For the Diffances from the Common Center of Gravity are reciprocally Proportional to the Bodies, and confequently in a given Proportion to each other; and by Composition in a given Proportion unto the whole Diffance betwixt the Bodies. But these Diffances are carried about their Terms, with a common Angular Motion, because

because lying always in a ftreight Line, they do not change their Inclination to one another. But Right Lines which are in a given Proportion to one another, and are carried about one another with an equal Angular Motion, describe altogether like Figures about the same Points (in Plains which either rest together with these ( Points, or are mov'd with any Motion which is not Angular.) And therefore the Figures which are describ'd by these revolving Distances are equal. Q. E. D.

Scholium. Thus the Earth and the Moon are carried by a Monthly Motion about the Common Center of Gravity of both. But to us placed on the Earth, to whom neither the Motion of our own Seat, nor of the Center of Gravity, as being an Invisible Point, is perceptible, the Moon alone seems to be carried about; and so it must needs happen in all the rest of the Systems of the Planets.

XXVIII. If two Bodies attract one another with any Force whatever, and be in the mean while revolv'd about a common Center of Gravity; a Figure like and equal may be defcrib'd by the fame Force about either Body unmov'd, to the Figures which the Bodies fo mov'd defcribe "about one another.

In Fig.2. Plate 6.Let S and P be revolv'd about C, the common Center of Gravity; S from S to T, and P from P to Q; from a given Point s let s p and s q be drawn, equal and parallel to S P, T Q. Here the Curve pqv, which the Point p defcribes about the unmoved Point s, will be like and equal to the Curves, which the Bodies S and P defcribe about each other; and confequently, by our laft Proposition, like to the Curves S T and P Q V, which the fame Bodies defcribe about C, the

the common Center of Gravity; and this fo becaufe the Proportions of the Lines S C. C P, and S P, or sp to one another are every where given.

Cale (1.) That common Center of Gravity C. either refts, or is uniformly mov'd ftraight forward, by the 25th Law of Motion. Let us first fuppofe it to reft; and in s and p let two Bodies be placed, the unmov'd one in s, the mov'd one in p; and let them be respectively like and equal to S and P. Then let the right Lines PR and pr touch the Curves PQ and pq in P and p; and let CQ and sq be drawn forth unto R and r. Here becaufe of the likeness of the Figures CPRQ and sprq, RQ will be to rq, as CP is to sp; and confequently in a given Proportion : Therefore if the Force wherewith the Body P is attracted to the Body S, and confequently towards the intermediate Center C, fould be to that wherewith the Body p is attracted towards the Center s in that fame given Proportion; these Forces would in equal Times always draw Bodies from the Tangents PR, and pr to the Arches PQ, pq by Intervals Proportional to those Forces R Q. r q; and confequently the latter Force would make that the Body p fhould be turn'd round in the Curve p q v, which would be like to the Curve PQV, in which the former Force makes the Body P to be turn'd about ; and the Revolutions would be compleated in the Tame Times. But because these Forces are not to one another in the Proportion of CP to sp; bat (by reason of the Similitude and Equality of the Bodies 6 and s, P and p, and the Equality of the Diftances P, s p) equal to one another, the Bodies will be attracted equally from the Tangents; and therefore that the latter Body p fhould

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fhould be drawn thro' the greater Interval r g, greater Time is requir'd, and this in the fubduplicate Proportion of the Intervals; becaufe the Spaces described are in the duplicate Proportion of the Times by Proposition 4. Therefore let the Velocity of the Body p be fuppos'd to be to the Velocity of the Body P in the fub-duplicate Proportion of the Distance s p to the Distance CP: to that in Times which are in the fame fub-duplicate Proportion, the Arches PQ, pq may be described, which are in the entire Proportion like to one another : In this Cafe the Bodies P, p which are always attracted by equal Forces, will describe about the quiescent Centers C and s. like Figures PQV, pqv. the latter whereof pq v is like and equal to the Figure which the Body P defcribes about the moved Body S. Q. É. D.

Cale (2.) Let us now suppose that the common Center of Gravity, together with the Relative Space in which the Bodies are moved amongst themselves, goes forwards uniformly in a Right Line; by the 26th Law of Motion, all the Motions will be perform'd in the Space as before; and confequently the Bodies will describe about one another the same Figures as before; which therefore will be like and equal to the Figure p q v. Q. E. D.

Coroll. (2.) The periodic Time about the unmov'd Body s, will be greater than the periodic Time about the moved Body S, or rather that which is about C the Center of Gravity: and that in the reciprocal Proportion of the Angles defcribed at the fame time; that is, in the fubduplicate Proportion of the Radii gp and  $CP_3$ that is, in the fubduplicate Proportion of the Bodies S + P to the Body S. Thus, if the Moon p fhould

fhould be moved about s, which is the Earth unmov'd, at the fame Diftance that it is : And fince the Quantity of the Matter in the Moon is about one 40th Part of the Quantity of Matter in the Earth; the periodic Time of the Moon would be greater than that periodic Time of the fame, which is at prefent, nearly in proportion of the Number 40 to the Number 29 498. For it is 40: 29 498: 29 ..... From whence, fince the periodic Time of that Planet is now 27d. 7h. 42', or 29242'; in the other Cafe it would be 29841, or 27d. 16h. 1'. · . . • 7 <u>۱</u>

Coroll. (2.) Hence two Bodies drawing one another by Forces directly proportional to their Diftances, defcribe (fee Prop. 19.) both about the common Center of Gravity, and about each other Ellipfes Concentrical, and which have their Centers in the Centers of the Forces. And on the contrary, if fuch Figures be defcribed about the Centers of Ellipfes, the Centripetal Forces are directly proportional to the Diftances from those Centers.

Coroll. (2.), Two Bodies drawing one another by Forces reciprocally proportional to the Square of their Diftances, (fee Prap. 21:) do both about the Common Center of Gravity, and abour each, other, describe Conical Sections, which have their Foci in the Center, about which the Figures are described. And on the contrary, if such Figures be described about the Focus of Conic Section, the Centripetal Forces are reciprocally. proportional to the Squares of the Diftances. Coroll. (4.) Any Two Bodies revolving about a common Center of Gravity, (see Prop. 15.) do by their Rays drawn to that Center, and to each. other, describe Areas proportional to the Times,  $\mathbf{O}_{\mathbf{A}} = \mathbf{O}_{\mathbf{A}} + \mathbf{O}_{\mathbf{A}} +$ by.

by reason of the perpetual Direction of the Rays, or Centripetal Forces to those Centers.

May 14. 1705.

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#### LECT. XVIII.

XXIX. F Two Bodies S and P, (fee Fig. 2. Plate 6.) which are revolv'd about a common Center of Gravity, do attract each other with a Force reciprocal-

ly proportional to their Diftance from the Center : the Transverse Axis of the Ellipsis, which either of them, as P describes about the other S, will be to the Transverse Axis of the Ellipsi, which the fame Body P might describe in the same periodic Time about the other at Reft, as the Sum of the two Bodies S + P is to the first of two Proportionals betwixt this Sum, and that other Body S. For if the described Ellipses were equal one to the other, the periodic Times would be, by the last Proposition, in the subduplicate Proportion of the Body S, to the Sum of the Bodies S + P. Now, if the periodic Time be diminish'd in this Proportion in the latter Ellipsi, the periodic Times will become equal: And the Transverse Axis of that Ellipsi will (by Prop. 12.) be diminish'd inthe Proportion, of which this Subduplicate is the Selquiplicate; that is, in the Proportion of which the entire Proportion of S to S + P is triplicate; and confequently will come to be to the tranfverse Axis of the other Ellipsi; as the first of the two

two mean Proportionals betwixt S + P, and S is to S + P. And inverfly, the transverse Axis of the Ellipsis, described about the moved Body, will be to the transverse, Axis of that described about the Body unmov'd, as S + P is to the first of the two mean Proportionals betwixt S + P and S. Q.E.D.

Thus, if the Moon's mean Diftance from the Earth : that is, half the transverse Axis of the Ellipfis described in the Supposition of the Earth's being unmoved, be of 60 Semi-diameters of the Earth, in Proportion to the given periodic Time; that Diftance will be greater than 60 Semi-diameters of the Earth, on the Supposition of the Circumrotation both of the Earth and the Moon about a common Center of Gravity; and that in the Proportion of the Sum of the Earth and the Moon to the first of the two mean Proportionals betwixt the Sum of the Earth and of the Moon, to the Earth; or in-the Supposition of the Moon's being a 40th Part of the Earth, as 40 is to 29 66. For 29: 29 22: 29 66: 40. From whence, fince the Diffance of the Moon in the Hypothesis of the Earth's being unmoved. is put to be of 60 Semi-diameters of the Earth; it will be, in the other Hypothesis, of 601 Semidiameters.

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Coroll. From what was just demonstrated, it follows, that if two Bodies drawing each other by any Force whatever, and which are not moved from any thing elfe, nor impeded, be moved in any fort whatever; their Motions will be the fame in effect, as if they did not attract each other, but were both attracted with the fame Force by fome 3d Body placed in the common Center of Gravity: And the Condition of the attracting Force will be the fame, in respect of the Distance of the Bodies from that common Center, and in O 2 respect

respect of their whole Distance betwixt themfelves. For that Force wherewith the Bodies draw each other, because it tends to the Bodies, tends to the intermediate common Center of Gravvity; and the Distances from the Center of Gravity, are every where proportional to the Distances of the Bodies; and consequently the Forces are the soft and do in the fame Proportion increase or decrease, as if they proceeded from the intermediate Body in the Center of Gravity.

XXX. Many Bodies, whole Forces are proportional to the Quantity of Matter, and in the direct Proportion of the Diffances, may be mov'd in divers Ellipses about their Centers, in such fort that their Motions may continue perpetually without any Perturbation, and that the common Center of Gravity of them all may in the mean while reft:

In the first place, let the two Bodies T and L (fee Fig. 2. Plate 6.) be supposed to have D for the common Center of Gravity. If a projectile Morion be once impressed in due Proportion according to parallel Lines situate in the same Plane, but according to Directions contrary to both, these Bodies will describe like Ellipse, having their Centers in D the common Center of Gravity, as we shewed above, Prop. 19.

Now let S, a 2d Body, draw the two former T and L; with the accelerating Forces S T and SL, and be reciprocally drawn by them. The Force ST may, by the 22d Law of Motion, be refolv'd in the Forces S D, D T; and the Force S L into the Forces S D, D L. But the Forces D T, D L, which are as T L the Sum of them: [For fince the Proportions of the Parts T D and D L do always remain the fame, the Proportion of the whole also T L will remain the fame in all the Diffances

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of the Bodies T and L. ] And the accelerating Forces of the Bodies themfelves T and L, are as the Diftances TL; and the additional Force arifing from the Body S, and tending according to the Line TL, is, as we have already feen, as the fame Diffances TL. Therefore the Sum of the Forces TD and LD, respecting the Center of Gravity, are as the Diftances D T and T L. Bur these are greater than the former Forces, and consequently will make that those Bodies should defcribe Ellipses, either like to the former, with a fwifter Motion, if the projectile Force be accelerated in Proportion to the additional Centripetal Force; or of another Species, if that projectile Motion remain given. The remaining accelerating Forces SD and SD, whilft they draw those Bcdies equally, and according to the Lines TI, LK, parallel to DS, do nothing at all change their Situations each to other, but caufe that they should equally approach to the Line IK perpendicular to SD. But that Accels to the Line IK will be hindred, by caufing that the Syftem of the Bodies T and L : that is, that D the Center of Gravity of the Two on one Part, and the Body S on the other, should be revolv'd with due Velocities in the given Plane about C the common Center of the Three, according to parallel Lines. The Body S by fuch a Motion (because that the Sums of the Motions, being on both Sides directly proportional to the Diffance SD, and confequently to those CD and CS, do draw the Bodies towards the Center C:) The Body S, I fay, for this Reafon will (by the faid Motion) defcribe an Ellipsis about the fame Point C: and the Point D will describe an Ellipsi on the contrary Part; in the mean while that the Bodies T and L go on to defcribe their Ellipses, as before, about the moveable Center D. New, Q 3.

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Now, let a fourth Body, as V, be added, and it will be concluded by the like Argument, that this and the Point C may deforibe Ellipfes about B the common Center of Gravity of them all; the Motions of the former Bodies TL and S about the Centers D and C ftill remaining, but fomething accelerated: And the thing will be the fame in cafe of more Bodies.

Coroll. (1.) The Case of a System of Bodies revolving about other Bodies, where the Centripetal Forces are directly as the Distances, affords exact Ellipses, such as are in no wife disordered by the Addition of more Bodies. But by how much the more the Quantities of the Centripetal Forces depart from this Proportion, the Bodies must neceffarily, ceteris paribus, the more disorder and difturb one another's Motions.

Coroll. (2.) But if the Centripetal Forces be reciprocally as the Squares of the Diftances, and a System of two or more leffer Bodies revolving about a common Center of Gravity placed in the Focus of the Ellipsi, be pressed on one Side by a Body far greater than any of them, but sufficiently remote; and be preffed in fuch fort, that the common Center of Gravity of them all becomes not far distant from the Center of the greater Body, the common Center of Gravity of the Syftem of the leffer Bodies will defcribe an Ellipfis about the greatest Body, or rather about the common Center of Gravity of them all. But divers Inequalities will arife in the Motions of the leffer Bodies, which we shall explain in what follows. Such indeed as Aftronomers have noted in our Moon from most certain Observations.

Coroll. (3.) But the greatest Disorder of all will arise in the leffer System, if the greatest Body

dy should attract divers Parts of that System unequally at equal Diffances; that is, if the divers Kinds of the various Bodies should gravitate unequally, or in divers Degrees towards the greateft Body; especially if the Inequality of this Proportion should be greater than the Inequality of the Proportion of the Diftances from the greatest Body. For if the accelerating Force, whilft it acts equally, and according to parallel Lines. does nothing at all diffurb the Motions of Bodies amongst themselves, a Disturbance must neceffarily arife from the Inequality of the Action; and must be greater or leffer, according to the greater or leffer Inequality. The Exceffes of the greater Impulse, whilst they act upon fome Bodies, and not upon others; or act lefs upon fome than others, will necessarily change their Situation amongft themselves. And this Difturbance being added to the Difturbance which arifeth necessarily from the Inclination and Inequality of the Lines, will make the whole Difturbance the greater. Coroll. (4.) From whence, if the Parts of the leffer System should be moved in Ellipses about the Focus, or in Circles about the Center, without any other Difturbance of their Motions, than what proceeds from the Inclination and Inequality of Lines drawn from the greatest Body; it is manifest, that the accelerating Forces of all the Parts of the System towards the greatest Body, are in equal Distances equal; and that all the Bodies comprehended in the leffer Syftem, do equally gravitate towards the greatest Body at equal Distances.

Coroll. (5.) Hence it is also manifest, that the Parts of that lesser System are either urged by no accelerating Force, but what tends to the greatest Body, except it be very lightly and infensibly; or O A at

at leaft are urged very nearly with equal Impreffions, and according to parallel Lines. All which things it were easy to apply to the Systems of the Earth and Moon, of *Jupiter* and his Satellites, of Saturn and his Satellites revolving about the Sun.

XXXI. If a primary Planet revolving about the Sun carry a Moon along with it, this will be formov'd about the Primary, that it will perpetually be accelerated from the Quadrature with the Sun, unto the Conjunction or Opposition next following; but from the Conjunction to the Quadrature, it will be retarded; and confequently will be carried more fivility about the Conjunction and Opposition, but more flowly about the Quadratures.

In Fig. 4. Plate 6: let Q beithe Sun, S the primary Planet revolving in ESE its annual Orb. Por p.a. Moon describing its.own mentitual Peri--od A D B C about the Primary ; in which Orbit, let the Points A and B defign the Syrigies with the Sun: that is, the Comunction and Oppolition: C and D the Quadrantes, that is, the Points diftant on this Side and on that a Quarter of a Circle from the Conjunction & Opposition. Further, lerQS, or QK; or Qk, the mean Diftance of the Moon or Satellit from the Sun, represent the Quantity of the accelerating Arraction ; that, to wie, whereby the Secondary Plance tends to the Suni; where it is placed at the fame Diftance from the Sun, as the Primary; and Prop provident to be the Place of the Satellit ur its own Orbie EAnd let QL or 101 betaken in the hipe PQ pmp Q, produced if need be, which QL or Ql let be to QK or :Ok in the duplicate Proportion of QK or Qk to OP or Qp; that is, fo that it may be thus, PQ: CK:OR: QL #, or Qp:Qk:Qr. QI#: Thefe Things being thus, this Line QL or QL which

which was last found, will express the accelerative Attraction of the Moon placed at L or 1 towards the Sun lin Q. Then let S, P or S, p be joined, and L M or L m drawn parallel to it, and meeting QS in M or m. Here the accelerative Attraction in QL or Ql, by the 22d Law of Motion, will be refolved into the Attractions LM. and LF, or MQ; or into 1 m, and 1 f. or m Q: and this with the Directions of those Lines. Of which Attractions, that which is represented by MQ or mQ is reduced to the Attraction mS: by taking away the Attraction as QS. which is common to the Primary and its Satelles; and which confequently brings in no Anomalies. By which means the Attraction of the Sarelles, tending according to the Direction SQ, which ought to be reckoned in this Place, is reduced to the Attraction M S in the Place P, fo much as the Satelles is more attracted to the Sun than the Printary. From whence MS in the former Cafe, and mS in the latter, will defign the Difference of the Attractions tending along S.Q. And confequently the Satelles, by this means, is urged with a threefold Attraction, or rather with fuch an Attraction as may very well be refolv'd into three. The first and chief being that wherewith the Primary S draws this Secondary P or p; the fecond that which is proportion nal to LM or 1m, with the Direction of the Line L M or l m; that is, with the Direction of PS or p.S. parallel to L.M. or Im: From whence the whole Force, compounded of thefe two Attra-Ations, when it respects the Center of the Primary S, will make that the Body P or p, if it were impress'd with these alone, would even yet defcribe Areas about the fame Center Suproportional to the Times, by Prop. 15. But the Satelles is alfo

alfo acted on with a third Force, with one that is as MS or mS, and with the Direction from M or m towards S: that is, from L or I towards F or f. This, in the Polition P, tends more to the Sun than its Primary; and that according to the Direction parallel to QS by the Excels MS. And in the Polition p, it tends lefs to the Sun than its Primary; and this according to the fame Direction parallel to QS by the Defect mS. Which will come altogether to the fame, as if we fhould reckon the Excels MS from L towards F: and the Defect m S from f towards 1: or the Excels from M towards S, and the Defect from m towards S; or as if the Satelles were diffurbed on this Side and on that, by a double Sun placed opnosite each to other. For when the Primary is drawn back from its Secondary towards the Sun by a true Excels of Attraction, there will be altogether, as to the Primary, all the fame fenfible Effects, (and those alone are what we are now fearching after ) as would be if the Primary being unmov'd, the Secondary were drawn away by the same Difference of Attraction unto the Part opposite to the Sun. But now, fince this third Force which arifeth from the Difference of " Attractions parallel to S Q, dorh not tend to the Center S, neither doth the total Porce compountded of these three Attractions, that, to wit, wherewith the Satelles is mov'd, tend unto the faid Center. Wherefore, by Prop. 17. and 18. the Satelles will not describe equable Areas about the Center of the Brimary, or fuch as are proportional to the Times. But the Force represented by MS or mS, will difturb the equable Defcription of Areas. In the Quadrant CA, of the Semi-circle CAD, fuppoling the monthly Motion to be perform'd from West to Eaft

East through A, D, B, C, the faid Force accelerates the Motion of the Satelles about S, from C to A, by confpiring together with it; but after the Conjunction in A in the Quadrant AD, it retards the Motion by being opposite to it. But the Satelles being come unto the Quadrature about D. the 2d Force MS or ms vanisheth away; (because OK or Qk: QP or Qp; and confequently QL and Q1 alfo are then equal.) And therefore the Force expressed every where by the faid MS, can have no Effect in this Place. Therefore the Satelles, which about the Quadratures is urged by the reft of the Forces, and those only tending unto the Center of the Primary, will describe equable Areas by Rays drawn to the Center, or proportional to the Times. But whilst the Satelles goes over the Quadrant DB, Qm falls fhort of QS; and if we refer the diffurbing Force to the Satelles alone, it will tend from m to S, and will again accelerate the Motion thereof by confpiring together with it: But after the Opposition in B, the Force will still tend from m towards S; but will now retard the Motion of the Satelles by being contrary thereto; untill again about the Quadrature C, m S vanisheth away, and confequently its Effects cease. Again, seeing the Force MS or mS, which diffurbs the Area in the Passage of the Satelles from C to A, and from D to B, is continually increas'd, and in A and B becomes the greatest; and from these Points again is continually diminish'd, whilst the Planet is carried from A to D, and from B to C, until it at length vanisheth away in the Points D and C; it is manifeft, that the Motion of the Satelles, as beheld from its Primary, is the swiftest, cateris paribus, in the Conjunction and Opposition A and B, and floweft in the Quadratures C and D. Q. E. D. Coroll

Coroll. (1.) From hence we may falve that Inequality in the Motion of the Moon, named by Aftronomy, The Variation; wherein the Moon is carried more fwiftly in the Conjunction and Oppofition, than in the Quadratures; and this fo, that in moving from a Sprigy to an Octant, it gains about 26 Minutes above the middle Motion; and lofeth again the fame Quantity in its Motion from an Octant to a Quadrature; and fo perpetually. And the like Anomaly is to be expected in the little Moons of *Jupiter* and Saturn; although by reason of their great Diffance from the Sun and from us, and of their fhort mensfrual Periods, it is not fensible to us.

Corell, (2.) Hence allo it follows, that the Orbit of any Moon, exteris paribus, will be more curve in the Quadratures than in the Conjunction and Opposition. And confequently, if it be in it felf Ciecular, it will become fomething Elliptical, in such fort that the leffer Axis will be always placed in the Conjunction and Oppolition, and the groater in the Quadratures. But if the Orbit be of it felf Elliptic about its Primary pla--ced in one of the Foci, it will partake more of that: Figure; than if it were not affected with this Anomaly, Carres was the first that I know of, that affigned this oblong Figure to the Orb of the Moon, which he did only by way of Hypathefis and Conjecture: But in the mean while he fell into a great Error, when he determin'd what the Moon comes nearer to the Earth in all Conjunction and Opposition, and departs further off in the Quadratures ; when on the contrary, by the proper Eccentricity of that Orb, the Line of the Apfides being put, the Conjunction and Opposition, the Moon is more remote from the Earth in the highest Apsis, than in the Quadratures ;

tures; that Inequality which we have been fpeaking of notwithstanding. But the great Dr. Halley was the first who, from Observations, attributed this oblong Figure to the faid Orbit; or at leaft the first that communicated the same to the Publick ; and from thence fhew'd, that the Lunar Theory was to be corrected : But as to the Demonstration of this Corollary, it is easily deduced out of the Proposition. For Bodies which are swifter, do decline less from the right Path than flower ones: And besides, the disturbing Force MS or mS in the Conjunction and Opposition, is not only the greatest in it felf, but is alfo directly contrary to that Force wherewith the Central Body S draws the Body P or p; and confequently diminisheth that Force by being contrary thereto. But the Body p or P will decline less from the right Path, when it is less urged towards the Central Body S; and confequently will be more carried in an Oblong Elliptic Path about its Primary.

May 21. 1705.

#### LECT. XIX.

XXXII. F by reafon of the Diffance betwixt the Sun, and a primary Planet increas'd and diminifh'd by turns, the Action of the Sun be alternately

increas'd and diminish'd; the Radius of the Orbit of the Satelles will withal be increas'd and diminish'd, and the periodic Time of the Satelles about its Primary will be chang'd alternately in that

that is, will be increas'd when the Radius is increas'd; and on the contrary, diminish'd when it is fo.

The Force wherewith the Primary draws its Moon, is increased when the Moon is in the Quadratures C and D, by the Addition of the Force SP or Sp; the Force SM or Sm vanishing away ; and is diminish'd when the Satelles is in the Conjunction and Oppolition, by the taking away of the Force SM or SM. And because the Force S m or S m in the Conjunction and Oppofition, is twofold of SP or Sp in the Quadratures, where the Point R or r falls in almost with the Point B or A; the attractive Force of the Primary will be more increas'd than diminish'd inevery Synodical Month and confequently is to be reckon'd for abfolutely increas'd. Therefore the Force of the Sun being increas'd about the Perihelion of the Syftem, the attractive Force of the Primary will be more languid, and the Orbit will be enlarg'd; but the Force of the Sun being diminish'd about the Aphelion of the System, the attractive Force of the Primary will be more ftrong, and the Orbit will be contracted. But the periodic Time of the Satelles will be increas'd with the enlarging of the Orbit; and on the contrary, diminish'd with it : and thus every Year the middle Motion of the Satelles will be greater and leffer by turns ; and is to be accounted truly mean, only in a mean Diftance from the Sun.

Coroll. (1.) Hence we may folve that annual Inequality in the Moon, which refpects the middle Motion thereof; that, namely, in which the middle Motion of this Planet doth alternately exceed and fall fhort of the true middle Motion, by an Excels or Defect of 12' almost; exceeding it in the Passage of the Earth from the remoter Apsis to the mean Distance; and falling fhort thereof from the mean Distance to the near-

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er Apfis; and again, falling fhort from the nearer Apfis to the mean Diftance, and exceeding it from the mean Diftance to the remoter Apfis: and fo perpetually. And the fame thing is to be judged of the Moons about Saturn and Jupiter, in their Proportion. Albeit this Inequality in these, and the reft likewise, is fo very small, that it may very well be neglected in most Cafes.

Coroll. (2.) The truly original and primitive periodic Time of every Moon, that is, that Time in which it would revolve about its Primary, if it were without the reach of the Sun's Action, is a little florter than the middle periodic Time; and the orginal Diftance from its Primary, a little lefs than the prefent. Namely, because if the Force of the Sun, which debilitates the Force of its Primary, were taken away, it would approach nearer to its Primary; and thus the periodic Time would be the florter.

Coroll. (3.) Hence also we may infer, with the Famous Dr. Gregory, that if any Primary Planet should, through the Accession of new Matter, become greater than it was, and from thence its Attraction become proportionably greater; its Moon would revolve about it at a lefs Distance, and in a shorter time. As on the contrary, by the Diminution of the Matter of the Primary, the Orbit and periodic Time of its Moon would be enlarged. And the same thing would happen in any Primary, in case the Sun was increas'd or diminish'd.

Coroll. (4.) Since therefore it is manifest from the most ancient Astronomical Observations, as compared with the latter, that the periodic Times of the Primary Planets about the Sun, and of the Moon about the Earth, are the same in this Age, as they were 2000 Years ago; it is certain, that the Quan-

Quantity of Matter both in the Sun and in the Earth, is the fame that it was then, and hath had no fenfible Addition or Diminution.

Coroll. (5.) But if the Quantity of Matter in the Earth be fuppos'd to have been increas'd by Noab's Deluge, or by any other means, the Quantity of the periodic Month of the Moon must neceffarily have been diminish'd thereby.

XXXIII. If a Secondary Planet defcribes an Elliptic Orbit about its Primary, which is placed in the Focus of the Ellipfis; the greater Axis of this Ellipfis, or the Line of Afpes, will, by an angular Motion, go forward and backward by turns; but it will go forward more than it goes back; and in each Revolution of the Secondary, by the Excefs of the Progreffion, it will be carried towards the confequent Signs: That is, In the Conjunction & Oppolition with the Sun, it will go forwards; & in the Quadratures it will go backwards.

For the Force wherewith the Secondary Planet P or p is urged towards its Primary about the Quadratures, where the other Force M'S or mS is vanished away, is compounded of the Force L M or I m, and the Centripetal Force of the Central Body S. The former Force, if the Diftance be increas'd or diminish'd, is increas'd or diminish'd almost in the same Proportion directly; fo that in the greater Diffance from the Primary, the Attraction towards the Center becomes. greater, and in a leffer Diftance lefs. But the latter Force arifing immediately from the Primary, in a greater Diftance becomes lefs, and in a leffer Distance greater; and is always in the duplicate Proportion of the Diffance reciprocally. And, confequently the entire Force, or the Sum of the Forces towards the Center of the Primary, doth, upon the Increase of the Distances, increase in a leffer

leffer Proportion than the duplicate Proportion of the Diftance is; that is, it is not fo much diminish'd in a greater Distance, nor is it fo much in creas'd in a leffer Diftance, as the Motion about the Focus of the unmov'd Ellipsis doth require. But in the Conjunction and Opposition, the Force where with the Secondary is urged towards its Primary, is the Difference betwixt the Force wherewith the Secondary is drawn by the Primary, and the Force K L or k1; or in this Cafe S M or S m. And that Difference, because the Force SM or Sm is increas'd nearly in the very Proportion of the Diftance directly, decreaseth in less than a duplicate Proportion of the Diftance; and confequently is greater in a teffer Diftance, and lefs in a greater, than fufficeth for the defcribing an unmoved Ellipsis. But if the Centripetal Force decreafeth in more than a duplicate Proportion of the Diffance, as it comes to pass about the Conjun-Ation and Opposition; this is a little like the Cafe of the Decrease of the Centripetal Force in the triplicate Proportion of the Diftance; from whence a Motion in a fpiral Line, without any Change of the Tangent to the Radius, would follow. The Satelles therefore will revolve in some moveable Ellipsi; or a greater angular Motion will be requird, that the Tangents oblique to the Radius fhould become perpendicular to the fame, that is, that the Satelles should come to its Apsides, than would be required if the Forces were in the duplicate Proportion of the Diftances reciprocally; that is, the Line of the Apfides will go forward. And on the contrary, if the Centripetal Force decreaseth in a lefs than duplicate Proportion of the Distance, as it happens about the Quadratures, the contrary Cafe follows; and the Motion of the Secondary will arife from a Motion different from P that

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thatin a Spiral, which keeps the Angle of the Radius and Tangent : So that that Angle should be fooner chang'd, and fooner come unto a right Line, than it would come if the Force were in the very duplicate Proportion of the Diftance reciprocally; that is, the Line of Apfides will go back. But in the intermediate Places, betwixt the Conjunction or Oppolition, and the Quadratures, the Motion of the Aplis depends upon both Caufes conjunctly; fo as to make that it fould go forwards or backwards, according to the Excels of this or that. From whence, fince the Borce K L. or k1 in the Conjunction and Opposition, as we, lately noted, is twice as great as the Force L M or 1m in the Quadratures; the Excels in every: whole Revolution will be on the Side of the greater Force K I or k1; and will transfer the Apfiseach Revolution towards the confequent Signs.

Coroll. (1.) Hence we may folve that Inequality, or progreflive and regreflive Motion of the Lunar Aplis, in which the Apogeum is fo mov'd, that in its Conjunction and Opposition it goes forwards more swiftly, and in its Quadratures. goes back more flowly; and by the Excels of the progreffive Motion above the regreffive every Month, it is carried towards the confequent Signs about three Degrees; and thus goes over a whole Circle in the Space of Ten Years, or a little fooner. In the Moons of Jupiter, which are mov'd almost in Circles, the Apsides are none at all, or infensible at most, and confequently this Demonstration appertains not to them. In these of Saturn it will have place, if at any time some Eccentricity shall be discovered in their Paths; but by reason of the Shormels of their periodic Times, and their vaft Diffance from the Sun, and confequently the fmall Force of the fame, . . . the

the Change of the Apogeum will be fo very fmall that it cannot fall under our Observation. much less be brought under Computation.

Corolli (1) Since therefore the Progress or Regrefs of the Aplides depends upon the Decrease of the Centripetal Force, which is made in a Propopulon greater or lefs than the duplicate Proportion of the Diffance SP or Sp in the Transit of the Body from the nearer Apfis to the remoter ; as likewife upon the like Increase in the Return to . the nearer Apfe; and confequently is greateft where the Proportion of the Force in the higher Applis to the Force in the nearer, is most remore from the duplicate Proportion of the Diftances inverted ; it is manifelt, that the Aplides in their Conjunction and Opposition, by the ablatitious Force KL of SM-L M, or Sm-Im, will go forward more fwiftly; SP or Sp being at that time the least of all; and SM or Sm the greatest of all; and SP or Sp, or rather the Sum of them on both Sides being the least of all in the Quadratures From whence, in each Revolution of the Satelles, whilft the Apfides are about the Conjunction and Opposition, they will go forwards most swiftly in the Conjunction and Opposition of the Satelles, and go back very flowly in the Quadratures thereof; and confequently the Excels , of the progressive Motion above the Regressive will be the greatest of all, and the Apris will be moved very fwiftly towards the confequent Signs.

Coroll. (2.) But if the Apfides be about the Quadratures, then contrary Caufes will produce contrary Effects; and the Apfides will go forwards more flowly than before, when the Satelles . is in the Conjunction and Oppolition, and go back more swiftly in the Quadratures of it; yea, it may come to pals in the laid. Polition of the Ap.

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Apfides in fome particular Revolution of the Satelles, that the Regress of them in the Quadratures of the Satelles may furpais the Progress of the fame, which is when the Satellis is in the Conjun-Ation and Opposition. But because the ablatirious Force SM or Sm, that eaufeth the Progress of the Aplides in the Conjunction and Oppolicion, is, cateris paribus, about twice as much as the adjectitious Force which brings in the Regress of the Apfides in the Quadratures of the Satelles; and because the Apsides do also tarry longer in the Conjunction and Opposition than in the Quadratures; fince they move, in the former Place, towards the confequent Signs with the Sun, they, go forward, and confequently do accompany him longer ; but in the latter Place, moving to the antecedent Signs, they fooner pals the Square of the Sun, which moves in the mean while towards. the confequent Signs: From these Reasons it appears, that the Apfides go forward more fwiftly; and longer in their Conjunction and Opposition. and go back more flowly, but not fo long in their Quadratures; and that they by the Excels of the Progress above the Regress in one entire Revolution of them to the Sun, i. e. in the Space of about Thirteen Months, are still carried towards the confequent Signs. Thus, in the Orbit of the Moon, the Apogeum thereof is moved fo unequally, that it is to be brought under Rule by an Equation amounting to 12 whole Degrees and a Ouarter, as is to be feen in the Lunar Tables.

XXXIV. If a Satellite be mov'd in an Eccentric Orb about its Primary, the Eccentricity will be changed twice in every Revolution, and will be the greateft, when the Secondary is in the Conjunction and Opposition with the Sun; and the leaft, when it is in the Quadratures; and confequently

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quently will be increased continually in the Paffage from the Quadratures to the Conjunction and Oppolition; and in the contrary Paffage continually diminith'd

For fince it appears by what hath already been demonstrated, that the Centripetal Force towards the Primary removid at a great Diffance, doth fometimes decrease in a greater than the duplicate Proportion of the Diffance, fometimes in a lefs: and finde the Motion of the Satellite in an immoveable Orbit, and with one certain Eccentricity, depends upon the Decrease of the faid Force in the dublicate Proportion of the Diftance it felf; from the Change of this Proportion the Species of the Orbit must necessarily be changed. Thus. if the Centripetal Forces increase or decrease in more whan a reciprocal duplicate Proportion of the Diftance; it is manifest, that the Satellite in its Descent from the highest Apis to the lowest, being perpetually impell'd towards the Center by the Acceffion of that new Force, will incline more to that Center, than it would have done if the Increase of the Centripetal Force had been only in the duplicate Proportion of the Diftance diminished; and confequently will describe an Elliptic Orb inferiour to the former, and at the lowest Apfis approach nearer to the Center than it did in the higheft; and thus the Orb, by the Oc-. cafion of this new Force, is made more Eccentric. And now, if in the Return of the Satellite from the lowest Apsis to the highest, the faid Force should decrease by the fame Degrees by which it did before increase, the Satellite would return to the former Distance, keeping the Eccentricity lately obtain'd; whereas, if the faid Force doth decrease in a greater Proportion than that in which it increas'd before; the Moon being in this Cafe lefs attracted, will ascend unro an high-P

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er Diftance, and fo the Eccentricity will be fill more increased.

In like manner, if the Satellise in its Descent from the highest Apsis, be urged with a Force which is increas'd by Jefs than the studicate Proportion of the Diffance diminishid ; is is manifest, that it will defcribe an Elliptic Orb exterior to the former ( than that I mean, where the Centripetal Force was reciprocally as the Square of the Diffance;) and confequendy an Orb lefs excentrical ; and that this Eccentricity is fill more diminish'd, if in the Ascent the Contributal Force decreases tels or more flowly thanithad in+ creafed before. If therefore the Propertion of the Increase and Decrease of the Septripetal Force be increased in each Revolutions, the Eci centricity likewife will be increased ;sand on the contrary it will be diminish'd; where the Tame Proportion decrealeth. Seaing therefore in every Revolution, that Force decreafech in the Conjuni aion and Opposition of the Satelles dif a greater Proportion than thas which, is duplicate of the Diftance increasidy, and in the Quadratures of the fame in a lefs, as is manifest from what hach been already faid, it appears, phat about the Conjunation and Oppolition of the Satelles the Eccentricity, of the Orb describ'd is perpetually increas'd, and diminish'd about the Quadratures. And fince in many Revolutions compar'd amongst themfelves, there is the greatest Proportion of Decrease in the Conjunction and Opposition of the Apfides, and the leaft in the Quadratures of the fame, it is also manifest, that the greatest Eccentricity of the Orbit is when the Apfides are in: the Conjunction and Opposition, and the least, when they are in the Quadratures;; and confequently that the Eccentricity is diminish'd perpetually in the passing of the Apfides from the Conjunction

junction and Opposition to the Quadratures of the Sun; and are perpetually increas'd in the patting of the fame from the Quadratures to the Conjunction and Opposition.

Corollary. Hence we may folve that Eccentricity of the Lunar Orbit which is divers, and daily changing, as being greater in the Moon's Conjunction and Oppolition, lefs in the Quadratures; and likewife continually increasing in the passing of the Apogeum from the Conjunction and Oppolicion to the Quadratures, and in the contrary Cale continually decreasing. For in Astronomical Tables we find fo great a Diversity alligned to this Eccenttricity, that the Distance betwixt the Focus and the Center of the Ellipsis describd by this Planet, which we call the Eccentricity of the Orbit, is fome-66782 Parts, fometimes of 43319 times of 1000000 2000000 only; Juch Parts we mean that 1000000 are contain'd in the mean Distance of the Moon. So. that the Difference of Eccentricities is found to arife unto above half of the whole least Eccentricky.

June 4. 1705.

XXXI



#### LEC.T. XX.

F the Satellite be revolv'd about the Primary in an Orb, the Plane whereof is inclin'd to the Plane of the Primary, the Line of the Nodeswill be moved

with an angular Motion towards the antecedent Signs, but with an unequal Velocity: Moft fwiftly indeed when the Nodes are in the Quadratures, afterwards by Degrees more flowly, until  $P_A$  that

that be placed in the Conjunction and Opposition, and they wholly reft; and thus being always either Retrograde or Stationary in each Revolution of the Satellite, they move back. As likewise in the fame Revolution they go back more swiftly, cateris paribus, when the Satellite is in the Conjunction and Opposition, than when it is in the Quadratures.

For amongst the disturbing Forces, of which we have spoken so dr, the Force LM or 1m which is parallel to SP or Sp, that is always fituate in the Plane of the Orbit of the Satelles. and can induce no Change of the Plane of the •Orbit. The other Force also MS or mS, fituate in the Plane of the Elliptic, when the Nodes are in the Conjunction and Opposition, will be alfo placed in the Plane of the Orbit, as being posited at that time in the common Intersection of both Planes. But when the Nodes are not in the Conjunction and Opposition, this latter and greater Force, which is always in the Plane of the Elliptic, will not be in the Plane of the Orbit; and confequently will affect the Motion of the Satellite, as to Latitude, and make the Line of Nodes to go back towards the antecedent Signs. For let the Nodes be supposed to be placed in the Quadratures, this latter Force, which always acts parallel to the Elliptic, will perpetually draw back the Satelles whilft it is paffing the Nodes on either Side, and about to go forwards in its own Orbit, from the fame Plane; fo that the Place of the Interfection which is to be next, will be at fome Diftance from the former Interfection, and towards the antecedent Signs. But when the Nodes are betwixt the Conjunction and Opposition and the Quadratures, this latter Force will fometimes move them towards the confequent Signs, sometimes towards the antecedent; but will

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will always in an entire Revolution of the Satelles, by the Excess of the same Force towards the antecedent Signs, carry them back towards that Part : From whence, in the Conjunction and Opposition of the Nodes, they will remain immoveable; in their Quadratures, they will go back most fwistly; and partaking in the intermediate Places of both Conditions, they will go back more flowly; and confequently will always, in a compleat Revolution, be carried back towards the antecedent Signs, notwithstanding their being Retrograde and Stationary in particular Places of the Period. But it is to be noted, that when the Orbit is placed without the Conjunction and Oppolition, and Quadratures, whilft the Satellite goes forward from the alcending Node to the descending, and, vice versa, the Nodes go back more flowly, fo long as the Force MS or mS respects that Side of the Plane on which the Satellite is placed ; and go forward fo long as that Force respects the oppolite Side. Thus the Line of Nodes being placed in an Octant of the Sun, after its having been placed in the Quadratures, or about R and r, the Satellite having pass'd the Plane of the Ecliptic about R, is then towards the Sun: but the difturbing Force from R to the Quadrature C, tends to the contrary Part by an Odant of a Circle; which Force vanishing away in the Quadrature, the Force tending to the Sun takes the Place of it, and continues throughout the three reft of the Octants: So that the Line of Nodes of the moreable Orbit doth first go forward a little, then goes back a little more ; and fo likewife in the other Semi-circle ; until the fame Line coming to the Conjunction and Opposition, the Progress and Regress are in a manner equal; but both of them very fmall, and of very short Continuance, by reason of the near Coincidence of the Situation of the Plane with the Direction of the diffurbing Force. But that the Nodes

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Nodes in the fame Revolution of the Satellite, go back more fwiftly, cateris peribus, when the Sacellite is in the Conjunction and Opposition, than elfewhere, is manifeft, by reason the diffurbing Force is greater in that Place; and confequently will obtain a greater Effect.

XXXVI. The fame things being fuppofed, the Inclination or acute Angle of the Plane of the Orb of the Satellite to the Plane of the Ecliptic, is perpetually changed; and is then greateft; when the Nodes are in the Comunction and Oppolition with the Sun; and the leaft, ceteris pari-Jus, when they are in the Quadratures : And is diminish'd continually in the Passage of the Satellite from the Quadratures to the Conjunction and Oppolition, and increas'd continually from the Conunction and Opposition to the Quadratures. From whence it comes, that the Satellite being in the Conjunction and Opposition, the Inclination of the Planes becomes the leaft; and returns to the former Magnitude nearly, when the Moon comes to the next Node. And this Inclination of the Planes is diminish'd, whilst the Nodes are carried from the Conjunction and Oppolition to the Quadratures, and becomes the leaft of all, cueteris paribus, when the Nodes are in the Quadratures; then it increafeth by the fame Degrees whereby it had decreas'd before; and the Nodes being again return'd to the Conjunction and Oppolition; it returns to the former Magnitude. If the former Proposition be rightly understood, this will not for much require a particular Explication. For like as, whilft the Body goes forward by the former Motion from L to F, if an attracting Force, parallel to the Line L M, do supervene, which attracts towards M, and is represented by the Line L M, the Body will go forwards in the Diagonal LQ, and the Angle of Inclination MLQ will be lefs than MLF the former-

-former Angle of Inclination : Or thus, like as whill the Body goes forward from L to F by its proper Motion, it the like attracting Force parallel to the fame L'M' fupervenes, which attracts the contrary way, but nevertheless is repreferted by an equal Line, the Body will go forward in ambeher Diagonal, and the Angle will be greater than the former Angle: So it mult happen in like manner in our present Case, i. e. that a divers Inclination of the Plane will follow upon the Motion of the Nodes. For when the Nodes are in sthe Quadratures, that Motion of them which perpetually draws back the Satellite from the Plane of its Orb, diminisherh the Inclination of the Plane. in the mean while that the Satellite paffeth from the Quadratures to the Conjunction and Opposition; and increaseth the fame in the convery Transit; from whence it comes, that the Satellite being placed in the Conjunction and Opposition, the Inclination becomes the least of all; and returns to its former Quantity nearly in the Accels of the Woon to the next Node. But if the Nodes be found in the Octents next after their having been in the Quadratures, that is, sbour P and p; in this Cafe, according to what hath been faid already, the Inclination of the Plane) is perpenally diminish'd from either of the Nodes unto the goth Degree from thence ; then it is increased for the Space of 47 Degrees, or in the Transit whro the next Quadratures; and afterwards again is diminish'd for the other 45° or unto the next Node. So the Inclination is dimipift'd more than it is increas'd; and to is always lefs in the fublequent Node than in the foregoing. And by the like Reasoning, the Inclination is increased more than it is diminish'd, when the Nodes are in the other Octants, or about R and r. Therefore the Inclination is the greatest of all, when the Nodes are in

in the Conjunction and Opposition. In their Tranfit from the Conjunction and Opposition to the Quadratures, it is diminish'd in the Access of the Satellite unto them; and becomes the least of all when the Nodes are in the Quadratures, and the Satellite in the Conjunction and Opposition; then it increases by the fame Degrees by which it had decreas'd before; and the Nodes coming to the next Conjunction and Opposition, it returns to its former Magnitude. Q. E. D.

Corollary. From this and the former Proposition, we may folve the most known Phanomena of the Moon; I mean the annual Regress of the Nodes confisting of about 19 i Degrees, & that Mutability of the Inclination of the Orbit of this Planet, in which when the Nodes are in the Quadratures, the Angle of Inclination contains only 4°. 59'. 25": But when they are in the Syrigies, the fame Angle is found to arife to about 5°. 17'. 20".

XXXVIII. All the Inequalities which are in the Motions of the Secondary Planets revolving about their Primaries, are fomething greater in the Conjunction of the Satellire with the Sun; than they are in the Opposition.

For fince QS bears a greater Proportion to QA than QB bears to QS, by reason that SA, SB, cateris panibus, are equal, and that QS is greater than QA; the duplicate Proportion of QM to QS; will be greater ftill than the duplicate Proportion of QS to Qm. And confequently the Difference MS will be greater than the Difference mS; and LM greater than the million From whence the Effects derived from this Force, will be greater than those which are derived from the other.  $\mathcal{Q}_{3}E. D.$ 

But it is to be noted, that the Diffance of the Earth from the Sun is to vafily great, that the Difference of the Forces about the Conjunction of

of the Moon with the Sun, and about the Oppofition of the fame, is very fmall, and hath fcarce been diffinguished yet by any Observations. From whence it is not to be wonder'd, that Aftronomers have taken no notice of this Diffunction.

XXXVIII. The absolute Force of the Sun in the diffurbing the Secondary Planets, and the Effects thereof in divers. Diffances from the Sun; is in the triplicate Proportion of those Diffances inversly.

. For let the Diftance of the Satellite from the Sun be altered ; let the Radius of the Orbit of the Satellite be in the fame Proportion to the other Radius. In this cafe, the Diftance of the Secondary from its Primary, will be in a given Proportion to its Diffance from the San : From whence, according to this Hypothesis, the abfo-Inte diffurbing Force will be as the absolute Force of the Sun, or in that duplicate Proportion. Thus the thing would be, if the Radius of the Secondary System had increas'd or decreas'd in the fame Proportion, as the Diftance of the Sun increas'd or decreas'd; fo that they should still keep the same Proportion to one another, as before. But fince the Radius doth in no wife decrease by the Access of the Sun, or increase by the Receis thereof, that duplicate Proportion will be to be increas'd again by the other Proportion of the Distance of the Secondary from its Primary. From whence the entire compound Pro-

portion will be triplicate of the former. Q. E. D. As for Example : Let the Sun be fuppos'd as near again to the Earth, as it was before; or as 50 to 100. And let A B the Diameter be equal to two Parts, the Quantity of the abfolute Force of the Sun at S in the leffer Diffance, will be Fourfold of the Quantity of the famo Force in the greater Diffance; But the Force SM in the leffer Diffance,

Diftance, will be about Eightfold of the fame Force in the greater Diffance. For 19 × 49 = 2401 ; and 50 x 50 = 2500. From whence 2500 - 2401 = 99. And 991 × 99 = 9801 ; and 100 × 100 - 10000. From whence 10000 -9801 = 199. Therefore the Difference of the able have Force is almost in the double Proportion. or as 199 is to 99. And the mean abfolute Forces themfelves are in the Quadruple Proportion or as 4 to 1. Therefore the entire absolute Forcer compounded of them, is  $4 \times 2 = 8$  to  $1 \times 1 = 1$ . or in the reciprocal triplicate. Proportion of the Diffance nearly ... And fince the apparent Diameter of the Sun is almost in the triplicate Propertion of the Diftance, and the Force of the Central Body is also nearly the fame; the Sun's Force whereby he disturbs the Sacellite, and the Effect of it. | will be in the direct triplicate Proportion of the Sun's apparent Diameten very nearly

Scholium (1.) In the fame manner wherein the Sun placed without the Orbit of the Secondary Planet difturbs the Motion thereof; the fuperior Planets will diffurb the Motion of the lower. and Comers will diffurb the Motion of all the Planets. And the Actions of Planets and Comets. upon other Blanets, will produce the like Effects; though far lefs indeed, by reafon of the Smalness of their Bodies, if compared with the Sun, and the waft Diftances. But fome Effects there will be [ yea, of the Actions also of the inferior Blanets upon the fuperior] which if they cominue, and be for the most part directed the fame way, will at length become fensible. As for Example : The Apfis of the Osbit of the Earth will, after many Years, be moved towards the confequent Signs; although this Motion must necessarily be very fmall, if compar'd with the Motion of the Aphtes 1.5 • ,•

Apfides of the Moon the fame way. Thus, indeed, the Eccentricity of the Orbit of the Earth muft be fubject to fome Mutation; which, neverthelefs, is fo finall, that it can fcarce be collected from any Phanomenon.

Scholium (2.) And thus the fuperior Planets will fenfibly disturb the Motions of one another. if they be great ones, and tarry long about their mutual Heliocentrick Conjunction, they being then placed at the least Diftance from one another. Thus the Action of Jupiter up-on the Secondarias of Saturn, and of Saturn upon those of Impues, the mutual Gravitation of all the Planets one to another, which we have already proved, being supposed, is in no wife to be. flighted, at what time they are feen from the Sun. in Conjunction. For they are great Bodies, and far exceeding our Earth in Magnitude, and are. nean enough at that time, to make the Effects of their disturbing Forces become sensible. And that they are indeed fensible to us, will be. thewn hereafter from Aftronomical Observations. Scholings (2.) It is easy to estimate the divers. Quantities of the Sun's diffurbing Force in the, System of Impiter and that of Saturn, from the, known Quantity of the fame Force in the Ano-

malies of our Moon. For from the known Proportions of the Diffances of the Earth, and  $\mathcal{J}_{H\rightarrow}$ piter and Saturn from the Sun; and the known Effects of the faid Force in the Moon, by a certain Proportion of like Effects on both Sides, obferved by Sir Ifanc Newton; the Effects of that Force, in the Systems of Jupiter and Saturn, may be determin'd without much Difficulty.

XXXIX. A Broklew. To find the Proportion betwixt the Force whereby the Motion of a San tellite is diffushed by the Surf cande the Force where

whereby a Satellite is retained in its own Orb about its Primary, which is called its Gravitation towards its Primary.

For the whole diffurbing Force is compounded of the diffurbing Forces LM or 1m, and SM or Sm: And also by reason of the vast Diftance of the Sun, the Line LQ or 10. is almost parallel to the Line MQ; and confequently the Force L M or 1m is very near equal to its mean Quantity, or to the Radius of the Satelles S P or S p: And likewife by reason of the Sun's vaft Diftance S M or Sm, or L P and I p. are equal to treble the Line K P or k p. From whence, fince in the Triangle SKP or Skp. which is Rectangular at K or k, the Angle KSP or kSp is the Diffance of the Satellite from the Ouadrature ; and the Side K P or kp is the right Sine to the Radius SP; the differbing Force SM or Sm will be to the diffurbing Force LM or 1m. as is the Radius to the treble of the Right Sine of • The Distance of the Satellite from the next Quadrature. From whence, if the Proportion of the diffurbing Force SP or Sp to the Centripetal Force of the Primary, or to the Force of Gravity, were once known, the diffurbing Force S M or S m would eafily become known. Which therefore we find out by this Method. The diflurbing Force SP. or Sp is to the Centripetal Force of the Primary towards the Sun, as the Line Sp or Sp is to the Line SQ; or as the Diftance of the Satelles from its Primary, is to the Diftance of the Sun from the fame Primary. But the Centripetal Force of the Primary towards the Sun, is to the Centripetal Force of the Secondary towards its Primary, as the Squares of the periodic Times drawn into the Radii of the Circles; or as SQ is to SP or Sp; and as the Squares of the 2

the periodic Times together. From whence, by Équality of Proportion, the Quantity of the diflurbing Force will be to the Force of Gravity (the former Proportion of SP or Sp to SQ deflroying the other reciprocal Proportion of SQ to SP or Sp,) as the Squares of the periodic Times  $\mathcal{Q}$ , E. D.

Coroll. Since therefore the periodic Time of the Moon is 39343'; and the periodic Time of the Earth about the Sun is 525969'; the difturbing Force SP will be to the Force of Gravity towards the Earth, which is at the Moon, as 29243 x 29343' is to 525969' x 525969'; that is. 25 1547871649 is to 276642288961; or as I is to  $178^{11}_{12}$ . And fince the Force SM or Sm in its greatest Quantity, or in the Conjunction and Opposition, is to the former Force as 2 to 1; the Force SM or Sm in the Conjunction and Oppofition, will be to the Force of Gravity as 2 is to 178<sup>11</sup>/<sub>14</sub>, or as 1 is to 5918. Therefore that diffurbing Force of the Sun SM or Sm, is in the Conjunction and Opposition about a 60th Part of the whole Force of Gravity in the Moon towards the Earth. Or rather the Force SP or Sp being taken away in this Cafe from the Force SM or Sm, as may very well be done, the whole diffurbing Force in the Conjunction and Opposition will be to the Force of Gravity as 1 is to 891, or a 90th Part of the fame nearly. And in other Places, the Force SM or Sm will be to the Force of Gravity, (the whole Sine being put to be equal to Unity) as treble the right Sine of the Diftance from the next Quadrature is to 1781.

XL. If many fluid Bodies, either diftinct, or gathered together into one Fluid, be moved about a primary Planet; each Part of the Fluid in inperforming its Motion about the Primary after the manner

manner of a Satellite, will come nearer to the Primary, cæteris paribus, and be moved more swiftly in the Conjunction and Opposition of the fame, and of the Primary, than in the Quadratures. And the Nodes of this Ring, or its Interfections with the Plane of the Ecliptic, will reftin the Conjunction and Oppolition. But out of Conjunction and Opposition, they will be carried towards the antecedent Signs; and this most swiftly in the Quadratures, more slowly in other Places. The Inclination of the Ring alfo will be varied; and the Axis thereof will be moved to and fro in each monthly Revolution; and the Revolution being compleated, it will return to that Polition which it had before; fo far as it is not carried about by the Preceffion of the Nodes. All these things do follow of their own accord, from what hath been already demonstrated; and fo do not require a peculiar Demonstration.

Corollary, From hence fome of the Phanomena of the Ring of Saturn, if fo be it be a Fluid, may eafily be understood. Yea, indeed, if it be folid, the Nodes of the fame, its Interfections I mean with the Ecliptic, will reft in their Conjunction and Opposition, when the Sun is found in the Plane of the Ring, as well as in that of the Ecliptic. But out of the Conjunction and Opposition they will go back, and this most fwiftly in the Quadratures, and more flowly in other Places. The Inclination of the Ring will also be varied, and the Axis thereof in each Revolution about the Sun will nod, and twice vary its Inclination towards the Ecliptic, and twice return to its former Polition, only it will be carried about by the Precession of the Nodes, as is manifest from what has been already faid.

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XII. If a Fluid be contain'd in a Channel form'd in the Surface of any Planet, Primary or Secondary, and be uniformly revolv'd together with the Planet with a diurnal periodic Motion; each Part of this Fluid will be accelerated and retarded by turns, in its Conjunction and Oppofition; or at Noon day and Midnight, will be fwifter; in the Quadratures, or at the 6th Hour Evening and Morning, it will be flower than the contiguous Surface of the Globe ; and thus it will flow in the Channel, and return back by turns perbetually. For the Fluid will be diffurbed by the unequal Attraction of the Sun, because the Attraction of the nearer Parts will be greater, and that of Parts more remote lefs; while the Force LM or 1 m will draw the Fluid down in the Quadratures, or at the 6th Hour in the Evening and Morning; and make that the Parts of it, which are placed there, should de-Icend unto the Conjunction and Oppolition, or unto the Noon and Midnight; and the Force SM and Sm will draw the fame upwards in the Conjunction and Opposition, or ftop the Descent of it, and cause it to ascend unto the Quadratures ; and thus perpetually.

Coroll. Hence we learn the Caufe of the Flux and Reflux of the Sea. If we allow the diffurbing Force of the Moon, as well as of the Sun, and duly apply what hath been alreadly demonftrated to the prefent Cafe. But this fo well known and flupendious Phænomenon of Nature, will come to be treated of afterwards more largely and diffinctly; to which Place therefore we refer our Reader.

Octob. + 22, 1709.

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#### LECT. XXI.

XLII. F a Solid Ring be put about a Globe perfectly fpherical at the Equator of the fame, and flick to it; there will indeed be no Motion of Flux and Reflux,

but the vibrating Motion of Inclination, and the Precession of the Nodes, will remain. Let the Globe have the fame Axis with the Ring, and compleat its Revolution in the fame time; and with its Surface touch the Ring inwardly, and cleave to it; by its participating of the Motion thereof, the whole Frame will vibrate to and fro, and the Nodes will go back. For the Globe, as above shew'n, is indifferent to receive all Impres-The greatest Angle of Inclination of the fions. Ring without the Globe, would be where the Nodes are in the Conjunction and Opposition. In their Progrefs from thence to the Quadratures, the Ring endeavours to diminish its Inclination, and by that Endeavour impresseth its Motion upon the whole Globe. The Globe retains the Motion imprefs'd, until that the Ring by a contrary Endeavour takes away this Motion, and impresses a new Motion upon the contrary Part. And thus the greatest Motion of the decreasing Inclination is in the Quadratures of the Nodes, and the leaft Angle of Inclination is in the Octants after the Quadratures. Then the greatest Motion of Inclination is in the Conjunction and Opposition, and the greatest Angle in the next Octants. And the

the Cafe is the fame with a Globe without a Ring, which either is fomething higher in the Parts about the Equator than about the Poles, or confifts of a more denfe Matter. For that Excefs of Matter in the Parts about the Equator fupplies the Place of the Ring.

Coroll. (1.) For the fame Reason that the redundant Matter of the Globe causes the Nodes to go back, and consequently by the Increase thereof causeth the Regress to increase, and by the Diminution thereof that the same Regress should be diminished, and by its being taken away that the Regress should cease; it will come to pass, that if more than the redundant Matter be taken away, or, which comes to the same, if the Globe be more depress'd, or of a rarer Substance towards the Equator than towards the Poles, the Motion of the Nodes will be forward, or towards the consequent Signs.

Coroll. (2.) Hence allo, from the Motion of Nodes, the Conflitution of a Globe may be gathered: To wit, if the Globe conftantly keep the fame Poles, and the Motion be towards the antecedent Signs, the Matter about the Equator is redundant, but if towards consequent ones, deficient. Let us suppose a Globe uniform, and perfectly spherical, first to rest in a free Space, and then by fome Force, whatever it be, impress'd on the Surface, to be driven forwards, and from thence to acquire a Motion partly circular, partly ftreight forward. Becaufe the Globe is indifferent to all Axes paffing through its Center, and is no more determin'd to one Axis, or one Situation of the Axis than to another; it is manifeft, that it will never change its Axis, or the Inclination of the fame, by any Force of its own. Now, let the Globe be impell'd obliquely in that fame Q 3 Part

Part of the Surface, as before by some new Impulfe; fince the Impulse, whether it be fooner or later, makes no Alteration in the Effect; it is manifest, that these two Motions impress'd fucceffively, will produce the fame Motion, as if they had been impress'd at the same time: that is. the fame as if the Globe had been impell'd at first with a fimple Force compounded of both Impulfes ; and confequently a fimple Motion about an Axis of a given Inclination. And the fame is the Reason of the Second Impulse made in any other Place of the Equator of the first Motion, as of the first Impulse made in any Place whatever in the Equator of that Motion, which the fecond Impulse without the first would produce; and confequently of Impulles made upon any Places whatever. These will generate the fame circular Motion, as if they had been impress'd at one and the same time upon the Place of the Intersection of the Equators ' of those Motions, which they had severally generated, if they had been impress'd afunder. The Homogeneous and perfect Globe therefore doth not retain more distinct Motions ; but compounds all the impress d ones, and reduceth them to one; and is in it felf perpetually revolv'd, by a fimple and uniform Motion, about a fingle Axis of a given Inclination, as being always invariable, Nor can a Centripetal Force, tending rowards any extriniick Body whatever, change the Inclination of the Axis, or Velocity of the Rotation. If a Globe be understood to be divided into two Hemispheres by any Plane whatever passing through its Center, and through the Center unto which the Force is directed, that Force will always urge both Hemispheres equally, and so the Globe, as to the Morion of Circumrotation, will incline to neither Part. But let new Matter be when the state of the ٤ added 

added fomewhere betwixt the Pole and the Equator, heaped up in the Form of a Mountain: This will both difturb the Motion of the Globe by the perpetual Endeavour of departing from the Center of its Motion, and will make the Poles to wander over its Surface, and to describe Circles about it self, and its opposite Point. Nor will that Irregularity be corrected, but either by placing the faid Mountain in one of the Poles, in which Cafe, as was faid before, the Nodes will go forwards; or in the Equator, and then the Nodes will go back; or by adding fome new Matter on the other Part of the Axis to counterpoife the Mountain in its Motion. And thus the Nodes will go forwards or backwards, as the Mountain : and the new Matter added on the opposite Part. are nearer to the Pole, or to the Equator.

Coroll. (3.) Since therefore it is manifest from Aftronomical Observations, that the Nodes of the Equator of the Earth do perpetually go back about 50" in every Year; which Regress is called the Precession of the Equinox; it follows, that the Equatoreal Parts of the Earth are higher than the Polar. And, vice versa, fince from the Diurnal Motion the Figure of the Earth is, as will be shew'd asterwards, that of an oblate Spheroid, (the Polar Parts being more depress'd than those about the Equator;) it is manifest from thence, that the Nodes of the Equator must go back yearly.

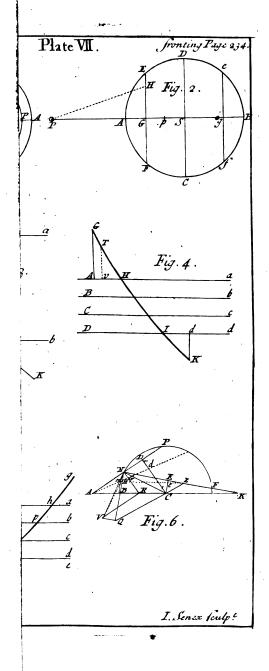
Coroll. (4.) From what hath been faid, it is also manifest, that the Axis of the Earth will vibrate to and fro yearly; and in every annual Revolution be inclin'd twice towards the Equator, and twice return to the former Position. It is mamifest also, that the greatest Motion of the decreasing Inclination of the Plane of the Equator, Q 4 and

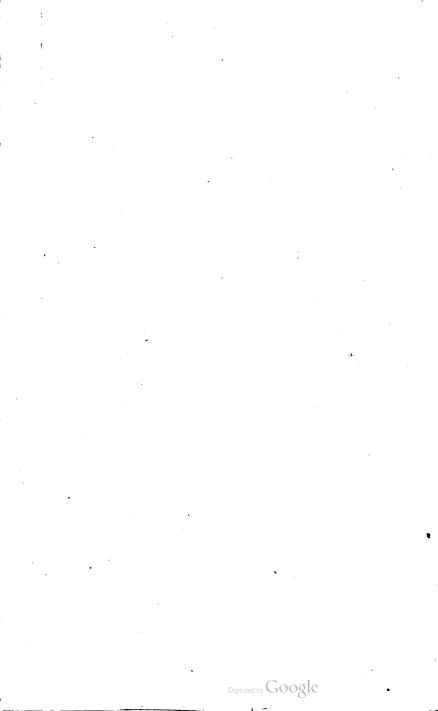
and of the Ecliptic, doth happen in the Quadratures of the Nodes; and that the leaft Angle of Inclination falls in the Octants after the Quadratures, or about the Middle of Leo or Aquarius. Laftly, that the greateft Motion of Inclination falls in the Conjunction and Opposition of the Nodes, or in the Equinoxes; and the greatest Angle of Inclination in the next Octants, or about the Middle of Taurus or Scorpio. But by reafon of the Smalness of these Motions, these Effects will be altogether infenfible, and to be discover'd by no Observations of Astronomers. But it is to be noted, that contrary Effects were to be attributed to our Earth, if so be the Parts about the Equator were more depress'd than the Polar.

Coroll. (5.) And from hence the Evalion devifed by the Famous Dr. Gregory, to thew that the annual Parallax of the Fixed Stars is built upon a weak Foundation, and that neither the Diftance of the Fixed Stars, which are observed, nor the annual Motion of the Earth, can be certainly concluded from thence; this Evasion, I fay, falls to the Ground. Let us produce in this place the Words of Dr. Gregory, and spare so much time as to debate this Matter with him particularly. Mr. Flamsteed's Method of observing the Parallax of the Fixed Stars, hath been explain'd by us in our Aftronomical Lectures, to which I refer you. Now, from this Method rightly underftood, it is manifest, that the Polar Star, for Instance, is more distant from the Pole about the Summer, than about the Winter Solftice; and this by a very fenfible Difference, as being about 40" or 45". From whence Mr. Flamfteed concludes, that the Earth must certainly be mov'd about the Sun, and that the Fixed Stars are fubject

ject to a Parallax sensible enough, and that their Diffances confequently may be gathered from thence. Now, what doth Dr. Gregory fay to this? Doth he deny the Observation it self? No, in no wife. Doth he affign for the Caufe of the faid Difference of Distance, that very small Nutation of the Axis of the Earth, by which he supposeth. with Mr. Flamsteed, that the Inclination of the Ecliptic to the Equator is leffen'd about the Solflices, and increas'd about the Equinoxes? No, not this neither. For Mr. Flamsteed had shew'd, that that very fmall Nutation doth rather confirm than weaken his Opinion. What therefore he attributes the faid Phanomenon to, as its Caufe, let us hear his own Words, Page 275. " This " Method, faith he, supposeth the Axis of the " Earth to be always most exactly parallel to it " felf, when it is in the opposite Points of its " own Orbit, where the Observations are made. And why should it not suppose this, or that it is parallel to it felf exactly enough for the prefent Purpose ? But he goes on : ] " Although that "" fmall Nutation of the Axis, of which we fpoke " juft now, doth in no wife hinder Mr. Flamfteed's " Observation; yet there is another Nutation of " it, which may produce the Diversity of the " Distance of the Polar Star from the Pole; that " is, if the Southern Hemisphere of the Earth " be of a more dense Frame than the Northern " ( whether it be from hence that that hath lefs " Summer than this, and therefore more cold; " or from the Inequality of the Continents about " the Poles, or from fome other Caufe unknown " to us.) fince at the Winter Solftice, the Sou-" thern Pole inclines to the Sun, and is withal " 'nearer to it than the Northern, and in the Sum-" mer Solftice this latter inclines to the Sun; the ' Axis

Axis of the Earth will be more inclin'd to the Plane of the Ecliptic in Winter Time, than in 55 " the Summer; and the Angle whereby the Polar " Star is diftant from the Pole, would be less at 66 the Winter than the Summer Solftice, altho' " the Pole Star were placed at an infinite Di-55 ftance, and the Lines drawn from thence to the Magnus Orbis were to be reckoned for pa-~~ " rallel. Since therefore the whole that can be 46 made of Mr. Flamsteed's Observation is this. that the apparent angular Diftance of the Po-66 " lar Star from the Pole, is lefs in the Summer " Solftice than in the Winter; and this may 66 arife from two Causes, either from the Con-" course of right Lines drawn from the Earth to " the Polar Star, in divers Situations of the " Earth to that Star, if the Earth's Axis in one " of the Observations be parallel to it self in 66 another, which Mr. Flamfteed supposeth; or " from the Concourse of right Lines coinciding " with the Axis of the Earth in its divers Situa-" tions, the Polar Star being suppos'd to be infi-·**6** ( nitely diftant; the Parallax of the Fixed Stars " cannot be certainly concluded from that Obfer-"vation. Becaufe the whole Observation may " confift, and the right Lines drawn from divers " Places of the Earth in its Orbit to the Pole Star " infinitely diftant may remain parallel; tho' the " Parallax of the great Orb, with respect to that " Star, be fuppos'd to be none at all. Yea, this "Obfervation (faith he) doth not fo much as έc prove immediately the annual Motion of the ¢¢ Earth. For although the Earth remains in the middle (making by its Rotation about its Axis, 66 çc as in the Semi-Tychonic System, the apparent " Diurnal Motion of the Stars, )the Sun when pla-" ced in the Southern Signs may fo attract the Southern Hemif-





" Hemisphere of the Earth, which is then nearer, ¢¢. and is perhaps more dense, that the Distance " of the Polar Star from the Pole at the Time of " the Winter Solftice, fhould be lefs than the " same Distance is when the Sun is placed in the *Northern* Signs, where it is more remote from the Hemisphere that is then turned to it; from " which and the lefs Denfity of the fame Hemifphere, which is perhaps conjoin'd therewith, . . . it is no wonder that this Hemilphere should be d. lefs attracted. Thus Dr. Gregory, who deviseth the like Evalions for the reft of the Observations of Mr. Flamfteed and Dr. Hook on this Subject. But I answer,

(1.) That, as to the affigned Caules of the Nutation of the Earth, the lefs Summer, to wit, of the Southern Hemisphere, and greater Cold, or the Inequality of the Continents about the Poles; if this Learned Man would derive that Denfity of the Southern Hemisphere above the Northern from these Causes, which may suffice to the moving of the Earth fo many Seconds from its former Pofition; he might as well go about to move Mount Cauca (us from its Place with a Leaver. I do admire at his Ungeometricalness in this Business. that he would not first estimate in some fort the Force and Quantity of these Caufes, before he attributed fo huge Effects to them. But his Prudence is to be commended, that he added, or from fome other Caule unknown to us: For he knew very well, that an unknown Caufe cannot be computed. But in the mean while I will fpeak freely and openly, that there can be no Caufe affign'd of this divers Denfity of the Hemispheres of the Earth, which he supposes, but what is contrary to the Mechanical Formation of the Planets, and the modern Phznomena of Nature. For,

(2.) If one Hemisphere of the Earth was a little higher or denser than the other, that Nutation of the Earth which he hath devised, would in no wise follow from thence. For in this Case, the Axis of the Globe would nodd indeed, but so that the Angle of the Inclination would twice in a Year return to its greatest, and twice to its least Quantity; and this so that that Angle would be of the same Quantity in both the Solffices, which plainly undermines the Foundations of his Hypothesis.

(3.) From this unequal Altitude or Denfity of the Hemilpheres of the Earth, if fo be it exceeds the Altitude or Denfity of the Equator. the Progress of the Equinoxes would follow: Whereas it is a thing certain, and acknowledg'd by Mr. Gregory himfelf, and every one, that they continually go backwards, and not forwards. But if he affign the Inequality to be fuch only, as not to infringe the greater Altitude or Denfity of the Equator; fo that fo much as the Parts about one Pole do exceed the Equatoreal Parts in Altitude. fo much the Parts about the other Pole fall fhore thereof: Neither will this be any Help to his Cause. For because of the Defect of Force in one Hemilphere, which compensates the Excess in the other; the Forces on both Sides will be in a poize, and there will be no entire Force which fhould move the Axis, and cause any Nutation. So that neither from that unequal Altitude or Denfity supposed, will his supposed Nutation of the Axis in any wife follow.

(4.) If we should, for Disputation's Sake, suppose that Nutation of the Axis, neither yet would this Learned Man attain his Aim. For he suppose that Nutation, as would reduce the Axis in one of the Solftices unto the least Angle of

of Inclination, and unto the greatest in the other. Now from the Principles of Sir Ilaac. Newton before laid down, which are Dr. Gregory's Principles likewife, it would follow, that the greateft Angle of Inclination of the Axis will be in the Octants after the Conjunction and Opposition of the Nodes, and the least in the Octants after the Quadratures of the fame; fo that as we faid before, in both the Solftices themfelves, which are in the Middle betwixt the greatest and leaft Angle, no Diversity at all of the Angle of Inclination is to be expected. From whence alfo. which is to be noted by the way, both Mr. Flamsteed himself, and Dr. Gregory who follows him. are altogether mistaken, when they suppose that Nutation of the Axis, to which the Preceffion of the Equinoxes is owing, can have any place here.

(5.) If, laftly, we fhould be minded to fuppofe the Nutation of the Axis, to be in the Time, and to the Parts affign'd by Dr. Gregory; the Quantity of Inclination would be far less than to produce Mr. Flamsteed's Parallax. Let us grant to him, that the Axis of the Earth doth vibrate to and fro every Year; let us grant alfo, that in one of the Equinoxes this Nutation is to the one Part, and in the other Equinox to the contrary; fo that the greateft Difference poffible should arife from thence. Yet, how very small will this Difference be: To wit, according to the Calculation we made formerly, ( fee Left. Aftron. Page .) it is manifest, that this huge Nutation which arifeth from the fensible Altitude of about 17 Miles, whereby the Semi-diameter of the Equator exceeds half the Axis, did only amount to a Part of one Second. What therefore is this Minute Difference to the Parallax, which arifeth

to

to three whole Quarters of one Minute? This Caufe therefore is in no wife fufficient to that Effect. To conclude, it is most certain that this Evasion of Dr. Gregory's, whereby he would shew that the annual Motion of the Earth doth not follow upon Mr. Flamssteed's Observations, is no small Error of his, and leaves a blemish upon a Work otherwise valuable for Demonstrations strictly geometrical, a Beauty not to be met with often elsewhere in Physical Tracts.

Scholium. But it is to be noted, that the Famous Mr. Flamstead hath not ordered his Reasonings altogether rightly in this Place, which the French have lately noted; and hath fometimes deduced the Parallax of the Fixed Stars from Phanomena in no wife proving it. But yet when I looked more narrowly into this Matter, Eleven of Fifteen remarkable Observations, which the French allow to be true, and agreeing with their own, do even yet thew the Parallax of the Fixed Stars; and of those Four that seem to disagree with it, there is only one of that Quantity as to give us any Trouble in this Busines; which therefore it is reasonable to think to be owing to some Miftake, whether in the observing or in the writing. Efpecially fince the like Parallax feems manifeftly to appear from the accurate Observati-ons of Dr. Hook. But these Things we leave to the further Diligence and Scrutiny of Aftronomers.

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Oftob. 29. 1705.

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#### Lест. XXII.

XLIII. F each particular Body of any Syftem, as A and B feverally confidered, draws all the reft of the Bodies with accelerative Forces, which are as the

Squares of the Diffances from the attracting Body reciprocally, the absolute Forces of all those Bodies will be one to another as are the Bodies themselves.

Let the Body A, by its accelerating Force reprefented by a, draw the Body B; and becaufe of the Diftance which is on both Sides the fame, let B reciprocally draw A by the Force reprefent-ed by b. The Quantity of Motion is on both Sides equal, because of the Reaction that is on both Sides equal to the Action : And that Quantity of Motion doth altogether arife from the Velocity drawn into the Quantity of the Matter, Therefore the Rectangle A x b is equal to the Rectangle B × a. And confequently the accelerating Force of the Body B will be to that of the Body A, at equal Distances, as the Body B is to A. And confequently the abfolute Forces of the Bodies will be one to another, as the Bodies themselves: To wit, the Sum of equal Forces tending every where unto equal Parts, at equal Distances. Q. E. D.

Scholium. By fuch like Propositions, we are led unto the Analogy betwixt Centripetal Forces and Central Bodies, to which those Forces are directed. For it is reasonable to think, that the Forces which

are directed towards Bodies, should depend upon the Nature and Quantity of those Bodies, as it comes to pals in Magnetics. And as often as these Cases happen, the Attractions of Bodies are to be effimated by affigning to each Part of them its proper Force, and fo gathering the Forces into one total Sum. But as for the Word Attraction. we use it here generally for any Endeavour whatfoever of coming unto another, which is found in Bodies, whether that Endeavour be from the Action of Bodies, either of themfelves tending to one another, or by mutual Emiffion of Spirits acting one upon the other; or whether it arife from the Action of the Ether, or Air, or any Medium whatever, corporeal or incorporeal. which forces the Bodies floating in it towards one another. In the fame general Senfe, we use the Word Impulse; not confidering in this place the physical Species and Qualities of the Forces, but their Mathematical Quantities and Proportions; as we propos'd above in the Definitions. In which Confideration of them, the Quantities of the Forces are to be fearched out and defin'd, and those Proportions which follow upon any Conditions whatever that are suppos'd But when we defcend unto Physics, these Proportions are to be compared with the Phænomena, that it may be known what Kind of Force it is which agrees to each Kind of attractive Bodies. And then we may at length, and not till then, fafely difpute concerning the Species, Caufes, and phyfical Reafons of Forces. Let us fee therefore by what Forces Spherical Bodies, fuch as are commonly the greater Bodies of the World, the Sun, Fixed Stars, Planets, and Comets, confifting of attra-Aive Particles in the manner just now defign'd, ought

Mathematical Philosophy. ought to act one upon another; and what Sort of Motions will follow from thence.

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XLIV. If towards each equal Points of a Spherical Physical Surface of equal Thickness every where, but which Thickness is fo fmall that it is not to be regarded, there be a Tendency of equal Centripetal Forces decreasing in the duplicate Proportion of the Diftances from the fame Points; any Corpuscle placed any where within this Surface, will not be attracted unto any Part by the faid Force ; but will either reft, or continue that Motion which is begun without any Disturbance, and in the fame manner as if it were acted upon with no Force at all from that Surface.

In Fig. c. Plate 6. let HIKL be that fpherical Surface, and P a Corpuscle placed within it. Through P let there be drawn to the Surface any two right Lines, intercepting the very fmall Arches HK, IL. And here because the Triangles HPI, LPK are similar [for the Arches HI and K Liare for fmall, that they are to be taken for right Lines; and the Angles vertically opposite at P are equal; and the Sides containing their equalAngles are (by III. 25. withVI. 14. and VI. 6. of the Elem.) on both Sides proportional] therefore those Arches will be proportional to the Diftances HP and LP; that is, PH will be to PL, or PI to PK, as IH is to KL. And any little Portions of the fpherical Surface at HI and KL, bounded on every Side by innumerable right Lines paffing through the Point P, whether they be Polygons or Circular, will be fimilar Figures, and confequently in the duplicate Proportion of those Arches or Diftances from the Corpuscle. And the whole attracting Forces towards the contrary Parts will, by reafon of the nearer Situation of the leffer Surface, and the remoter Situ-R

Situation of the greater, counterpoife and deftroy each other. And by the fame Argument, all Attractions throughout the whole Surface will be deftroy'd by the contrary Attractions. And confequently the Body P will be impell'd to no Part by these Attractions. Q. E. D.

Coroll. (1.) Since therefore any Sphere, which hath a Concave, Concentric, Spheric Space within, may rightly be diffinguifhed into innumerable fuch like fpherical Surfaces of an inconfiderable Craffitude; and fince from the Force of this Demonstration, no one of these Surfaces can attract a Body placed within it unto any Part: It is manifest, that the whole Sphere can impress no Force upon the Corpuscle within it. But that this Corpuscle, if it was in Rest before, will still rest; or if it was in Motion before, of what fort foever it were, it will still continue that Motion; any Attraction which may be in the exterior Sphere notwithstanding.

Coroll. (2.) And fince this thing may with Parity of Reafon be demonstrated concerning any Corpufcles whatever, compounding what Body or Mass of Matter foever; it appears, that all Bodies whatever, placed within such a Concave Sphere, are uncapable of receiving any Impression from any attractive Force of that Sphere.

Coroll. (2.) If therefore our Earth, as made of fuch Spherical Surfaces compos'd of attractive Particles, hath a Spherical Central Cavity, Animals placed there are affected with no Force of Gravity from those Surfaces, and perform their Motions with the fame Liberty, as they would do if there was no fuch thing as Gravity in Nature. And the fame is to be faid of the Planets and Comets, and of the Sun, and the Fixed Stars.

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XLV. The

XLV. The fame Things fuppos'd as before, a Corpufcle placed without the fpherical Surface will be attracted to the Center of the Sphere by a Force reciprocally proportional to the Diftance from the fame Center.

In the double Fig. 6. Plate 6. let there be two equal Surfaces, (or rather the fame Surface put :twice) one Mark'd with great Letters, the other with finall, AHKB, ahkb defctib'd from the Centers S, s with equal Diameters AB, ab; and let P p be two Corpuscles, (or rather one and the fame Corpufcle placed at divers Diftances from the fpherical Surface; ) placed without in the Continuation of those Diameters. Let the right Lines PHK, PII: phk, pii be drawn from the Corpuscles, cutting off from the greatest Circles A T B, a tb equal Arches, HK, hk: and ITI, iti differing, the latter from the former, as little as may be. And let the Perpendiculars SD,sd be let fall to PK,pk; and SE; se to PI, pi; and IR, ir to PK, pk. Of which, let SD, sd cut PI, pi in the Points F and F. Let there be let fall also to the Diameters the perpendicular Lines IQ, iq; and because of the Equality of the Lines DS, and ds; ES, and es; and of the most small vanishing Angles DPE, dpe; the Lines PE, PF, and pe, pf (the Difference FE, fe, and the little Lines DF, df vanishing) may be accounted for equal; as having their last Proportion, those Angles DPE. dpe. and DSE, dse vanishing away, the Proportion of Equality. These Things being thus, in the like Triangles PRI, PDF, and pri, pdf, PI will be to PF, as RI is to DF; and pf will be to pi, as DF or df is to ri: And both the equal Proportions being compounded into one, the Rectangle PI x pf will be to the R 2 Rea-

Rectangle PF into pi, as the Rectangle R Ix df is to the Rectangle DF x ri; that is, as RI is to ri: that is, in the last smilar Triangles IRH. irh (because of the right Angle at R and r; and the Angle R H I agreeing to the Angle r hi, if the equal Circles were applied to each other ) as the vanishing Arch I H is to the vanishing Arch Again, in the like Triangles PIQ, PSF: ih. piq, psf, Pil is to PS as IQ is to SE, and ps is to pi as SE or se is to iq. And both the equal Proportions being compounded, the Rectangle PI x ps will be to the Rectangle PS x pi. as the Rectangle IQ × s e is to the Rectangle SExiq; that is, as IQ is to iq. And both the principal Proportions being compounded, the Quantity PI × PI × pf x ps will be to the Quantity p i x p i X PF, XPS; that is, PIq x pf x ps will be to pi 9 × PF×PS., as the Rectangle IH×IQ is to the Rectangle i h x i q; that is, as the Circular Surface or Ring which the smallest Arch IH will defcribe in the Circumvolution of the Semicircle AHTB about the Diameter A B. is to the Circular Surface or Ring which the fmalleft Arch i h will describe in the Circumvolution. of the Semicircle a h t b about the Diameter a b. And the Forces wherewith these Surfaces do attract the Corpufcles P and p, are, by the Hypothefis, as the Surfaces themfelves, fo far as the Squares of the Diftances do not increase or diminish the the fame Forces; and confequently those Forces are as the Surfaces themfelves applied to the Squares of their Diffances from the Bodies; that  $\frac{PI^{q} \times pf \times ps}{PI^{q}}$  is to  $\frac{pi^{q} \times PF \times PS}{pi^{q}}$ is, as or as pfxps is to PF x PS. These entire Forces likewife are to their oblique Parts, which by 613 6 a Re-1.1.1

a Revolution made of the Forces tend to the Centers according to the Lines PS and ps, as PI is to PQ; and as pi to pq; that is, (because of the like Triangles PIQ PSF; and pig. psf:) as PS is to PF, and as ps to pf. From whence, by Equality of Proportion, the Attra-Etion of this Corpufele P towards the Center S. will be to the Attraction of the Corpufcle p towards the Center s, as  $\frac{PF}{PS}$  pf × ps is to P s PFxPS; or as PFxpfxpsxps is to  $pf \times PF \times PS \times PS$ , or allo as  $ps \times ps$  or psqis to  $PS \times PS$  or PSq; that is, as the Squares of the Diftances from their Centers reciprocally.

And by the like Argument, the Forces wherewith the remoter Surfaces defcribed by the Circumvolution of the remoter Arches HL and hl draw the Corpufcles, are as the Squares of the Diftances from their Centers reciprocally. And the Forces of all the like circular or annular Surfaces into which both the fpherical Surfaces may be diffinguish'd, by taking always equal Arches, as HK, hk and ITI, it i; or, which is the fame, by taking the perpendicular S D equal to s d, and SE equal to se; the Forces of all these annular Surfaces, I fay, are in the faid Proportion. And from thence, the Sum of the Forces, or the Force of the whole fpherical Surfaces, will be exerted upon the Corpufcles in the fame Proportion. 2. E. D.

Coroll. (1.) Since therefore every entire Sphere may be rightly diftinguish'd into innumerable such like concentrical fpherical Surfaces; and fince from the Force of this Demonstration any one of the Surfaces may fo attract that Corpufele, that the Force of Attraction towards the Center is in the duplicate Proportion of the Diftance reciprocally ; R 2

cally; it is manifest, that the whole Sphere alfo doth fo attract that Corpufcle, that the Centripetal Force is in the duplicate Proportion of the Distance from the Center reciprocally.

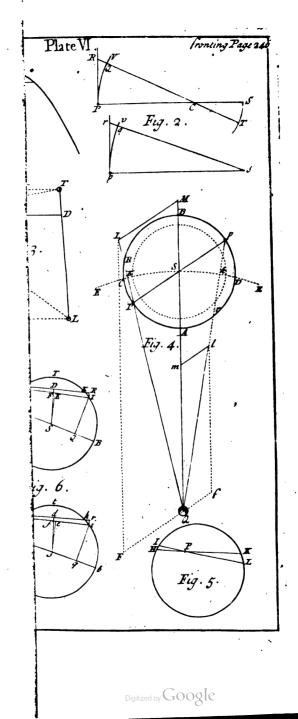
Corol. (2.) And fince the reft of the oblique Forces IQ, iq, eftimated from the oppofite Hemifpheres, are oppofite to, and deftroy each other; the entire Force exercis'd upon the Corpuscle, will be altogether equal to that Force tending towards the Center.

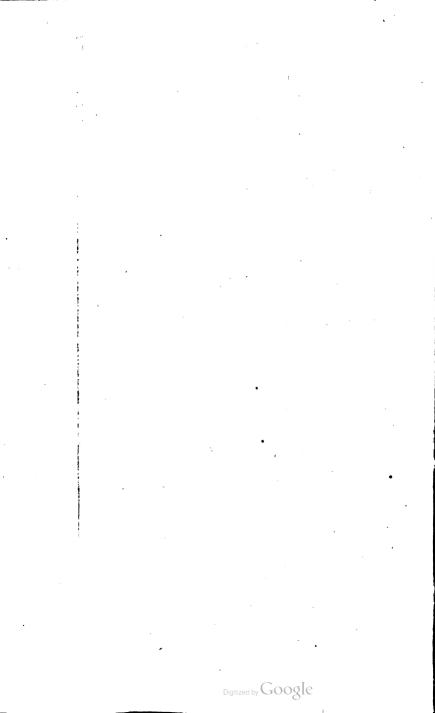
Corol. (2.) And feeing the Demonstration would proceed in the like manner, if instead of one Corpuscle, any Body compounded of those Corpuscles were supposed (for what agrees to one Particle must, by the same Reason, agree to every one, and consequently to the Sum of them;) it appears, that every Sphere consisting of Particles equally attractive, doth so attract every Body, that the Quantity of the Attraction is in the duplicate Proportion of the Distance from the Center of the Sphere reciprocally.

Corol. (4.) Therefore the Attraction of the Sphere is in the fame manner, as if the whole of the Forces tending towards the Center was gathered together in the Center it felf, and united and propagated it felf on every Side round about from that one Point.

XLVI. If unto each Point of any Spheres which are Homogeneous, or of the fame Denfity, equal Centripetal Forces do tend, decreafing in the duplicate Proportion of the Diftances from the Points; and the Proportion of the Diameters of the Spheres to the Diftance of the Bodies from the Centers of the fame Spheres be given : the Forces wherewith the Bodies are attracted being compar'd amongft themfelves, will be

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be found proportional to the Semidiameters of the attracting Spheres.

That is, The Forces of the Spheres are as the attracting Particles themfelves, or, as the Spheres themfelves; that is, in the triplicate Proportion of the Semi-diameters, to wit, at equal Diftances. But when the Diftances are fuppos'd unequal, and unequal in the very Proportion of the Semi-diameters, the Forces will be diminish'd in Proportion to the Diffances; that is, by the Hypothesis in the duplicate Proportion of the Semi-diameters of the Spheres. The remaining Forces therefore, which are to be estimated from the Excess of the triplicate Proportion above the duplicate, will be in the fimple Proportion of the Semi-diameters directly. Q. E. D.

Corol. (1.) Hence, if any Bodies be revolv'd in Circles about Spheres confifting of Matter equally attractive; and the Diftances from the Centers of the Spheres be proportional to the Diameters or Semi-diameters of the fame; the periodic Times will be equal. For the Equality of the periodic Times follows from the Forces in the direct Proportion of the Diffances : as we have fhew'd before.

Corol. (2.) And the Inverse of it is also true : if the periodic Times be equal, the Diftances of the revolving Bodies from the Spheres, if fo be the fame be Homogeneous, or of the fame Denfity. will be proportional to the Semi-diameters of the Spheres.

Corol. (2.) And from the periodic Times given, together with the Diftances of the Bodies from the Spheres, the Densities of the Spheres will alfo be given : To wit, by computing what periodic Times would follow from thence at Diftances proportional to the Semi-diameters of the Spheres, and

and by determining from the Excels or Defect of the periodic Times, the Defect or Excels of Denfities reciprocally proportional to the fame. Examples of which, in the Sun, Jupiter, Saturn, and the Earth, will be produc'd hereafter.

XLVII. If unto each Point of fome given Sphere, which is Homogeneous, or of equal Denfity every where, there be a Tendency of equal Centripetal Forces decreasing in the duplicate Proportion of the Distances from the Points; a Corpuscle placed within the Sphere, is attracted with a Force proportional to its Distance from the Center thereof.

In the Sphere ABCD (of Fig. 1. Plate 7.) described from the Center S, let the Corpuscle P be placed; and from the Center S with the Interval SP, conceive an inner Sphere to be defcribed, to wit, PEQF. It is manifest, by Prop. 44. That the Concentrick Spherical Surfaces, of which the Difference of the Spheres is composed, the Attractions in one part being every where deftroy'd by the contrary Attractions, do not act at all upon the Corpufcle P: There remains only the Attraction of the inner Sphere PEQF. Therefore the Centripetal Force decreafeth, by reason of the leffer Sphere which attracts in the triplicate Proportion of the diminish'd Distance from the Center; but increaseth in the invers duplicate Proportion of the Diftance, because of the Accels to the Center. Therefore the remaining Force, to be estimated from the Excels of the triplicate Proportion above the duplicate, will be in the direct Proportion of the Diftance from the Center. 2. E D.

Corol. (1.) If fuch a fort of Sphere be bored through the Center, all Bodies let fall from all Distances, whether little or great, will defcend unto

unto the Center in an equal Space of Time; in the Space, to wit, of 21'. 9". in our Earth, as we observed before.

Corol. (2.) And if there be no Medium, which refifts the Motion of the defcending or afcending Bodies, every Body let fall will, when it hath paffed the Center, afcend as far beyond the Center as it before defcended to it; and fo will, by a perpetual Afcent and Defcent, imitate the Motions of pendulous Bodies vibrating in a Cycloid. And thefe Vibrations, if we may fo call them, will be perform'd in equal Times.

Corol. (3.) But if, as many very fmall Intervals as you will, Concentrical to fuch a Kind of Sphere, be fuppoled to be interpos'd betwixt any fpherical Surfaces whatever, and any Bodies whatever be underftood to be revolv'd in thefe Intervals about the Center, like fo many little Planets; the periodic Times of all thefe Planets will be equal every where. That is, every Period will be perform'd in the fame Space of Time, in which any Body whatever being let down would perform the whole Vibration compounded of Going and Returning: Thus, in our Earth, the faid circular Periods would be performed in 1 h. 24'. 36''. As may eafily appear from what hath been demonstrated before.

Scholium. It is to be noted, that those Surfaces, of which we suppose folid Bodies to be compos'd, are not purely Mathematical, or void of all Thickness; but such thin Orbs, that their Craffitude is as nothing. In like manner by Points, of which we say Lines are compos'd, and from thence Surfaces and Solids, Particles of equal Magnitude, but which is so fmall that it is not to be regarded, are to be understood.

Nov, 19. 1705.

#### LECT.

#### LECT. XXIII.

XLVIII T T Sphere is attracted with a Force reciprocally proportional to the Square of its

Diftance from the Sphere. For let the Sphere be diftinguish'd into innumerable Concentrical Spherical Surfaces; the Attractions of the Corpuscle arising from each of the Surfaces, will be reciprocally proportional to the Square of the Distance of the Corpuscle from the Center, by *Prop.* 45. And likewise by compounding, the Sum of Attractions, or the Attraction of the whole Sphere, will be in the same Proportion. Q. E. D.

Corol. (1.) Hence, in equal Diffances from the Centers of Homogeneous Spheres, the Attrations are as the Spheres themfelves; or as the Cubes of the Diameters are one to another. For, by *Prop.* 46. if the Diffances be proportional to the Diameters, the Forces of the Spheres will be as the Diameters : Let the greater Diffance therefore be diminifuid in that Proportion; and thus the Diffance being now made equal, the Attration will be increasid in that duplicate Proportion, and confequently will be to the other Attraction in that triplicate Proportion of the Diameters, that is, in the Proportion of the Spheres themfelves.

Corol. (2.) In any Diffances whatever, the Attractions will be as the Spheres applied to the Squares of the Diffances.

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Corol.

Corol. (3.) If a Corpuscle placed without an Homogeneous Sphere, be drawn with a Force reciprocally proportional to the Square of its Diftance from the Center, and the Sphere in the mean time confists of attractive Particles; the Force of every Particle will decrease in the duplicate Proportion of the Distance from that Particle.

Corol. (4.) Since therefore all the Planets, both Primary and Secondary, are attracted to the Sun; all the Secondaries about *Jupiter* are attracted to the Center of *Jupiter*; all the Satellites of Saturn to the Center of Saturn; and the Moon to the Center of the Earth: Every one to its own Center in divers Diftances, with a Force reciprocally proportional to the Squares of the Diftances re-Ipectively; the Force of every Particle compofing the Body of the Sun, Jupiter, Saturn, and the Earth, decreafeth in a duplicate Proportion of the Diftance from the fame Particle.

XLIX. If unto each Point of a given Homogeneous Sphere, there be a Tendency of equal Centripetal Forces decreasing in the duplicate Proportion of the Distances from the Points; every other similar Sphere will be attracted with a Force reciprocally proportional to the Square of the Distance of the Centers.

For the Attraction of every Particle is reciprocally, as the Square of the Diftance thereof from the Center of the attracting Sphere, by Prop. 45. and therefore it is the fame, as if the whole attracting Force lay in one fingle Particle fituate in the Center of this Sphere. But this Attraction is as great as the Attraction of the fame Corpufcle wou'd be, if fo be it were attracted by each Particle of the attracted Sphere with the fame Force wherewith it attracts them. But this Attraction of

of the Corpufcle would be by the laft Prop. reciprocally proportional to the Square of the Diftance thereof from the Center of the Sphere; and confequently the Attraction of the Sphere, which is equal to the fame, is in the fame Proportion.  $\mathcal{Q}$ , E. D.

Corol. (1.) Attractions of Homogeneous Spheres towards other Homogeneous Spheres, are, as it is in those of Points, or the most minute Corpuscles, as the attracting Spheres applied to the Squares of the Distances of their Centers from the Centers of those which attract.

Corol. (2.) The fame thing holds, where the attracting Sphere doth alfo attract it felf. For each Point of this will draw each Point of the other with the fame Force, whereby it is interchangeably drawn by them. And confequently fince in all Attractions, both the attrahent and the attracted Body are urged or acted upon; the Force of the mutual Attraction will be doubled, keeping the Proportions.

Corol. (3.) All those things which have been demonstrated above, concerning the Motion of Bodies, about the Focus of Conic Sections, do hold where the attracting Sphere is placed in the Focus, and the Bodies are mov'd without that Sphere.

Corol. (4.) But those things which were demonstrated concerning the Motion of Bodies about the Center of Conic Sections, do hold where the Motions are performed within the Sphere; to wit, where a Sphere not perfectly Concave, but full of Concave Parts, is supposed; as we observed before.

L. Prop. If Spheres which are diffimilar in the Process from the Center to the Circumference, (as to Density of Matter, and the attractive Force,)

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Force,) are nevertheless altogether fimilar in their Progress in a round in every given Diffance from the Center; and the attractive Force of every Point decreaseth in the duplicate Proportion of the Diffance of the attracted Body: the whole Force wherewith one such Sphere draws the other, is reciprocally proportional to the Square of the Diffance of the Centers.

For fuch a fort of Sphere may always be divided into fimilar Concentrick Spherical Surfaces. And fince it hath lately been demonstrated, that every Surface feparately confidered, doth fo draw all other Surfaces feparately confidered, that the whole Force wherewith fuch a fpherick Surface draws any other, is reciprocally proportional to the Square of the Distance from its Center; the Proposition will appear manifest of entire Spheres compounded of fuch Surfaces. Q. E. D. Corol. (1.) Hence, if many fuch like Spheres

Corol. (1.) Hence, if many fuch like Spheres being like to one another in all things, do attract each other; the accelerating Force of each upon each will be at equal Diftances of the Centers, as the attracting Spheres themfelves; or as the Quantities of Matter contain'd in the fame.

Corol. (2.) And in all unequal Diftances whatever, as the attracting Spheres applied to the Squares of the Diftances betwixt the Centers of the Spheres.

Corol. (3.) But Moving Attractions, or the Weights of Spheres in Action upon or towards Spheres at equal Diffances of Centers, are as the attracting and attracted Spheres conjunctly; that is, as the Contents of the Spheres produc'd by Multiplication. For fince the attracting Body, becaufe of Reaction, which every where is equal to Action, the tending to the contrary Part, is mov'd towards the Body attracted with the like Quan-



tity of Motion, that is, with a Celerity recipro-cal to the Bodies; and this would be if there were properly no attractive Force of the Body which is attracted : And fince to those who inhabit any Sphere, the whole Velocity of Spheres approaching to one another is neceffarily referr'd to the other Sphere, and not to the Sphere they upon; becaufe they cannot difcern dwell their own Motion; hence it comes to pass, that all the Centripetal Force of the other Sphere. that, to wit, wherewith it approacheth to their own, or that rather wherewith they are both of them carried with the fame Tendency towards the mutual Concourse, and which is called the Weight of the other; is not only proportional to the attracting Sphere, but to both Spheres taken together. This is that which goes by the Name of the Weight of any Body towards the Earth, which makes that that Body and the Earth are carried towards each other with a relative Velocity of Approximation. Thus we fhew'd before (Prop. 22.) that the Gravity of the Moon towards the Earth, is of that Quantity that it would fall to the Center of the Earth in the Space of 4 h. 20'. almost. Not that all that Velocity is to be referr'd to the Moon it felf; but that if all the respective Velocity of approaching, arising from the Motion of both Bodies, were to be referrich to the Moon alone, as those that inhabit the Earth are wont to refer it ; this would make that the Moon would fall to the Center of the Earth in that Space of Time.

Corol. (4.) In unequal Diffances Moving Attractions, or the Weights of Spheres towards Spheres, will be as those Contents applied to the Squares of the Diffances betwixt the Centers.

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Corol.

Corol. (5.) The fame Things hold alfo, a fortiori, where the whole Attraction ariseth from the attractive Virtue of both Spheres mutually exercifed upon the other Sphere. For the Attraction will be doubled, the Proportion being kept.

Corol. (6.) If fome fuch Spheres be revolv'd about others quiefcent each about each; and the Diftances betwixt the Centers of the revolving ones, and of the quiefcent, be proportional to the Diameters of the quiefcent; the periodic Times will be equal.

Corol. (7.) And, vice versa, if the periodia Times be equal, the Distances will be proportional to the Diameters.

Corol. (8.) All the fame Things which we demonstrated above, concerning the Motion of Bodies about the Foci of Conic Sections, do hold where the attracting Sphere of what Form and Condition foever, which hath been already defcrib'd, is placed in the Focus.

Corol. (9.) As also where the attracting Spheres revolve of what Condition foever, that have been already describ'd, *i. e.* either wholly Homogeneous, or in the same Diffances from the Centers so.

LI If to each equal Point of Homogeneous Spheres, there be a Tendency of Centripetal Forces equal at equal Diftances, but at divers Diftances directly proportional to the Diftances of the Points from the Bodies attracted; the Force compounded of the Forces of all the Parts, wherewith the two Spheres do mutually draw each other, will be as the Diftance betwixt the Genters of the Spheres.

Cafe I. In Fig. 2. Plate 7. Let A C B D be a. Sphere compounded of these attractive Forces: S the Center of the same: P a Corpuscie attracted:

ed: PASB the Axis of the Sphere paffing through the Center of the Corpuscle: EF and ef two physical Planes of a Thickness not to be regarded. wherewith the Sphere is cut, which are also perpendicular to the Axis, and on this fide and on that equally diftant from the Center of the Sphere. The Points G and g, the Interfections of the Planes and the Axis: and H any physical Point in the Plane E F. The Centripetal Force of the Point H upon the Corpufcle P, exercifed along the Line PH, is, by the Hypothesis, as the Diftance it felf PH; which by Refolution of Forces is to be divided into GH, GP. From whence the Force along the Line PS, that is towards the Center S, is as the Length it felf PG: IHG, or one Part of the Forces being deftroy'd by the Force of the Point which is equally diftant from the Axis, on the other Side of the Ax is directly opposite in the fame Plane. ] Therefore the Force of all the Points in the Plane EF: that is, the Force of the whole Plane, whereby the Corpufele P is drawn towards the Center S. will in like manner be as the Number or Sum of the Points drawn into the Diftance PG: that is. as the Content under the Plane EF, and the Difance PG. And in like fort the Force of the Plane e f, whereby the Corpuscle P is drawn towards the Center S, is as the equal Plane drawn into the Diftance PG. And the Sum of the Forces of both Planes is as the Plane EF, drawn into the Sum of the Diffances  $PG \times Pg$ ; that is, as that Plane drawn into PS the double of the Diftance of the Center and the Corpufcle : [Becaufe of the Lines PG, PS, Pg, which are Arithmetically proportional; and from thence the Sum of the Extremes equal to the Double of the Mean. ] That is, as the Double of the Plane EF; or the Sum

Sum of the equal Planes drawn into the Diffance PS. And by the like Argument, the Forces of all the Planes in the whole Sphere, equidiffant on this fide and on that from the Center of the Sphere, are as the Sum of the Planes drawn into the Diffance PS: that is, as the whole Sphere drawn into the Diffance of the Center thereof from the Corpufcle. And becaufe the Sphere is given in every Diffance, the entire attracting Force will be as PS, the Diffance of the attracted Corpufcle from the Center of the Sphere. Q. E. D.

Cafe (2.) Now let the Corpuscle P draw the Sphere, to wit, all Points of it, with a Force directly proportional to the Diffance of the Points from the Corpuscle: And by the same Argument it will be prov'd, that the Force wherewith that Sphere is drawn will be as the Diffance P S.  $\mathcal{Q}$ . E D.

Cafe (2.) Then let another Homogeneous Sphere be compounded of innumerable Particles, as P, attractive in like manner; that is, in the direct Proportion of the Diftance. And because the Force wherewith every Particle is drawn, is as the Diftance of the Corpuscle from the Center of the first Sphere drawn into the fame Sphere; and confequently is the fame, as if the whole proceeded from one fingle Point in the Center of the Sphere: The whole Force wherewith all the Corpufcles will be drawn in the 2d Sphere, that is, that wherewith that whole Sphere is drawn, will be the fame as if that Sphere were drawn by a Force proceeding from one fingle Corpufcle placed in the Center of the first Sphere. And therefore it will be proportional to the Diftances betwixt the Centers of the Spheres. Q. E. D.

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Cafe (4.) Now let the Spheres draw one another; and the doubled Force will still preferve the former Proportion. Q. E. D.

 $Ca[e(\varsigma)]$  Next, let the Corpufcle p be placed within the Sphere ACBD: and here, becaufe the Force of the Plane ef upon the Corpufcle will be as the Content under that Plane, and the Difance pg, or as ef x pg; and the contrary Force of the Plane EF, as the Content under that Plane and the Diftance pG, or as EF x pG: or as e f x p G: Therefore the attracting Force will be as the Difference of the Contents, that is,  $ef \times pg - pG$ ; or as the double of ef drawn into half the Difference pg-pG, =2 ef x  $\frac{1}{2}pg$  $-\frac{1}{2}$ pG: That is, because SG, Sg are equal, as the Sum of the equal Planes drawn into half the Difference of the Diftances, or into pS the Diftance of the Corpuscle from the Center of the Sphere. And by the fame Argument, the Attraction of all the Planes in the whole Sphere, as E F, e f, equidistant on this side and on that from the Center; that is, the Attraction of the whole Sphere will be as the Sum of all the Planes, or the whole Sphere drawn into p S, the Diftance of the Corpufcle from the Center of the Sphere. Q. E. D.

Cale (6.) And if a new Homogeneous Sphere placed within the former, be compounded of innumerable Particles as p; it will be prov'd, as before, that the Attraction, whether it be fimply of one Sphere into another, or mutual and on both fides, will be as the Diftance of the Centers p S. Q. E. D.

LII. If Spheres in the Progress from the Center to the Circumference (as to Density of Matter, and attractive Force) howfoever diffimilar, are nevertheless in the Progress round about at every

every given Diftance from the Center every where fimilar; and the attractive Force of every Point is directly as the Diftance of the attracted Body: The whole Force wherewith two fuch Spheres will mutually draw each other, will be proportional to the Diftance betwixt the Centers of the Spheres.

For fuch a Sphere may always be divided into equal Circles EF, ef, and in the fame Diffances from the Centers G, g, into Homogeneous ones; and fince from the Force of what hath been already demonstrated, every circular Perimeter, of which every whole Circle is compounded, doth afford a Force proportional to the Diffance from the Center of the Sphere; the whole Force will also be in the direct Proportion of the Diffance from the Center.

Corol. What has been above demonstrated in the Corollaries to the 50th Prop. concerning the Attractions of Spheres where the Law of Attraction was in the duplicate Ratio of the Distance inversity, may be every where applied to this Place, mutatis rite mutandis. But especially those which we have formerly demonstrated, concerning Bodies moving about the Center of Conic Sections, obtain where all the Attractions are made by the Force of spheric Bodies of the same fort that was just now described; and the attracted Bodies are Spheres of the same Kind.

Scholium. We have now explained the twonotable Cafes of Attractions; to wit, where the Centripetal Forces decreale in the duplicate, or increase in the fimple Proportion of the Diffances : Causing the Bodies in both Cafes to be revolved in Conic Sections, to wit, by the first Law about the Focus, by the fecond about the Center; (and the first Cafe agreeing to Bodies placed without S i the

the Spheres, but the latter agreeing to the Bodies placed within them :) And compounding the Centripetal Forces of fpheric Bodies decreasing by the fame Law in their Recess from the Center, or increasing with the Bodies. Which is well worthy to be taken notice of, and is very ufeful to folve the Phænomena of the Solar Syftem. It wou'd be tedious, and of little Use, strictly to examine the other Cafes in this Place, which would exhibit Conclusions lefs elegant, and more remote from the Constitution of the World. Therefore, fince we have already explained the Attractions of fpheric Bodies, we may go on to the Laws of Attraction of other Bodies confifting of like attractive Particles: But we don't design to handle them in particular. Therefore we shall only fubjoin fome more general Propositions of the Forces of Bodies of the like fort, and of the Motions arifing from thence, which will be of Use in Physicks.

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#### LECT. XXIV.

III F two fimilar Mediums be feparated from each other by a Space bounded on both fides with parallel Planes; and a Body in its Paffage thro this Space be attract-

ed or impelled perpendicularly to either Medium, & be neither accelerated nor retarded in its way by any other Force; and if moreover the Attraction be every where the fame, at equal Diffances, from both Planes taken along the Line of its Morion; the

the Sine of the Angle of Incidence upon the one Plane, will be to the Sine of the Angle of Emergence out of the other Plane, in a given Proportion ; i. e. what foever the Angle of Inclination shall be, the Proportion of those Sines will always be found to be the fame.

Cafe (1.) Let (in Fig. 2. Plate 7.) Aa Bb be two parallel Planes. Let a Body fall upon the former along GH, and be in its whole Paffage attracted or impell'd towards the Medium of Incidence; and by this Action defcribe a Curve HI. and then emerge by the Line IK. Let the perpendicular IM be erected upon the Plane of Emergence Bb, meeting both the Line of Incidence GH produc'd in M, and the Plane of Incidence Aa in R. And let the Line of Emergence KI produc'd meet HM in L. Thus GM will be the Tangent of this Curve in the Point H; and the Line LK the Tangent of the fame in the Point I. Then, from the Center L with the Interval LI. let a Circle be describ'd, which may cut as well H M in P and Q, as M I produc'd in N. And now in the first place, if the Attraction or Impulse be suppos'd to be uniform, that Curve H I will, according to what hath bern demonstrated before, be a Portion of a Parabola; (fee Prop. 8.) and the Line LV, perpendicular to both Planes, will be one of the Diameters thereof; and the right Line H I, bifected by the fame L'V in the Point C, will be an Ordinate of that Diameter. Now it is a Property of this Figure, that the Re-Atangle contain'd under the Latus rectum, belonging to the Vertex H, (which in this Cale is always given (by Corol. 2. Prop. 8.) by reason of the Velocity of the Bodies here Juppos'd to be given ; )' and the Absciss HD or I M, which is equal to it; is equal to the Square of the Semiordinate S 2

ordinate DI or HM, which is equal thereto. The Tangent alfo of this Parabola HM will be bisected in the Point L: (For in the like Triangles HCL, HIM, as HC is to HI, so will HL be to HM. But HC is half of HI; therefore HL is also half of HM.) From whence, if you let down I O perpendicular to MI, MO and OR will always be equal; and the Equals IO, ON being added, the Wholes M N, I R will likewife be equal. Since therefore IR, the Diftance of the Planes, is given in all Inclinations whatever: MN allo, which is equal to the fame, will be given in all Inclinations. And confequently the Rectangle NM x MI will be to the Rectangle under the Latus rectum belonging to the Vertex H and MI, as the given NM is to the given Latus rectum, or in a given Proportion. Now the Re-Etangle under HD or MI, and the Latus rectum, is equal to DIq or HMq. And therefore the Re- $\alpha$  angle NM + MI is to HMq in a given Proportion. But NM + MI is equal Corol. 1. p. 30. to PM + MQ; that is, to the Book 3. of the Difference of the Squares of ML Elem.

and P L, or of M L and L1: And HMq hath a given Proportion to L Mq, a fourth Part of it felf. Therefore the Proportion of M L q-LIq to L Mq is given; and by dividing the Proportion of LIq to L Mq, and the fubduplicate Proportion of the fame Line L I to L M. But in every Triangle, the Sines of the Angles are proportional to the oppofite Sides, (Corol. 1. Prop. 20. Book 2. of the Elem.) Therefore the Proportion of the Sine of LMR, or of AHG, the Angle of Incidence, to the Sine of the Angle of Emergence MIK or LIR; or of the Angle LIM the Complement of the fame unto two right Angles is ftill given. [For the Sine of the Angle LIR, and

and of LIM, the Complement of the fame unto two right Angles, is the fame.] Q. E. D.

Cale (2.) Then let a Body pais successively thro' divers Spaces bounded with parallel Plains, as (in Fig. 4. Plate 7. ) A a b B, B b c C, C c d D, &c. and be mov'd with a Force uniform in each Sphere taken a part, and divers in the divers Spaces: Here, by what hath been demonstrated already, the Sine of Incidence on the Plane Aa. will be to the Sine of Emergence from B b in a given Proportion; and this Sine, which is the Sine of Incidence on the Plane Bb, will be to that of Emergence from the 2d Plane Cc in a given Proportion; and the fame is to be faid of this last Sine to the Sine of Emergence from Dd; and so infinitely. And by Equality of Proportion, the Sine of Incidence on the first Plane will be to the Sine of Emergence from the last, in a given Proportion. Then let the Intervals of the Planes be diminish'd, and the Number increas'd infinitely; to the end that the Action of Attraction or Impulse may be continual, according to any Condition which may be affign'd: And the Proportion of the Sine of Incidence on the first Plane to the Sine of Emergence from the laft. will still be given. Q. E. D.

LIV. The fame Things being fuppos'd, the Velocity of a Body before the Incidence will be to the Velocity of the fame after the Emergence, as the Sine of Emergence to the Sine of Incidence.

In the former Figure, let AH, Id be equal, and let the Perpendiculars AG, dK be erected, meeting the Lines of Incidence and Emergence GH, IK in G and K. In GH let TH be taken equal to IK, and Tv be let down perpendicular to the Plane Aa. And let the Motion of the S 4 Body

Body be diffinguish'd into two, one perpendicular to the Planes Aa, Bb, Cc, Dd, the other parallel to the fame. The Force of Attraction or Impulse, whilst it acts along the perpendicular Lines, will not at all change the Motion along the Parallels; and therefore the Body by this Motion will, in equal Times, go over those Intervals of the Parallels, which are betwixt the Line A G and the Point H, and betwixt the Point I and the Line DK: that is, will in equal Times describe the Lines GH and IK. And therefore the Velocity foregoing the Incidence will be to that which follows the Emergence, as GH is to IK or TH; as AH or ID is to vH; that is, (in respect of the Radius TH or IK) as the Sine of Emergence to that of Incidence. Q.E.D.

LV. The fame Things fuppos'd; and that the Motion before the Incidence is fwifter than that afterwards; a Body by the inclining of the Line of Incidence will at length be reflected, and the Angle of Reflection will be equal to that of Incidence.

For (fee Fig. c. Plate 7.) imagine that a Body doth, betwixt the parallel Planes A a, B b, C c, Dd, &c. describe parabolick Arches, as above; and let HP, PQ, QR, &c. be those Arches. And let the Obliquity of the incident Line GH to the *i*ft Plane A a be fuch, that the Sine of Incidence is to the Sine of the right Angle; that is, to the Radius of the Circle, in that Proportion which the fame Sine of the first Incidence hath to the Sine of the Emergence from the laft Plane Dd, into the Space express'd by Dd Ee. And , here, because of the Sine of Emergence now made equal to the Radius, or Sine of the right Angle, that Angle of Emergence will be a right one; and confequently the Line of Emergence will

will now coincide with the Plane D d. Let the Body come to this Plane in the Point R : Here, because the Line of Emergence coincides with the fame Plane, it is manifest, that the Body cannot move farther, or towards the remoter Plane express'd by E.e. Nor indeed can that go forwards in the Line of Emergence R d, because it is perpetually attracted or impell'd towards the Medium of Incidence : Therefore it will return betwixt C c and D d, defcribing Q R q the Arch of a Parabola, the principal Vertex whereof will be the Point R : and it will cut the Plane Cc in the fame Angle q, as it did before in Q. Then, in going forwards in the parabolic Arches op, ph. dr. like and equal to the former QP. PH, it will cut the reft of the Planes in the fame Angles in p and h, as before in P and H; and will at length emerge in the fame Obliquity in h, with which it fell upon H. Conceive now the Intervals of the Planes A a, B b, C c, D d, &c. to be diminish'd infinitely, and their Number increas'd; fo that the Action of Attraction or Impulse may be continual according to any Condition whatfoever; and the Angle of Emergence, which was before always equal to the Angle of Incidence. will remain still equal to the fame. Q. E. D.

Scholium. Not unlike to thefe Attractions, as it feems, are the Refractions and Reflections of Light made according to a given Proportion of Secants, as Snellius difcover'd; and by confequence according to a given Proportion of Sines, as Des Cartes has expounded it: [For fince every Sine is to the Radius, as the Radius is to the Secant of the Complement; and the Angle of Incidence betwixt the Radius and the Plane, called by Snellius the refracting Angle, is the Complement of the Angle called by Cartes, that of Incidence

dence betwixt the Radius and the Perpendicular; the Proportion of Secants with Snellius, will wholly agree and fall in with the Proportion of Sines us'd by Cartes ] For that Light is propagated and returneth from the Sun to the Earth in about 7 or 8 Minutes Space, is what is now manifest from the Phænomena of the Satellites of 74siter. confirm'd by the Observations of divers Astronomers; and the Rays of Light which are in the Air, (as Grimaldus fome while ago difcover'd, by letting in Light through a Hole into a dark Chamber; and Sir Isaac Newton hath more fully experimented) in their Passage near the Angles of opake or transparent Bodies, are bowed about the Bodies, being, as it were, drawn towards them: And of these Rays, those which in that Passage come nearer to the Bodies, are the more bowed ; being, as it were, the more attracted, as Sir Isaac himfelf diligently observed, and hath lately fet forth at large in the 2d Book of his Opticks. Now, fince there is fuch an Incurvation of the Rays in the Air without the Knife, the Rays alfo which fall upon the Knife must have been bended in the Air before they touched the Knife : And there is the like Reason for those which fall upon Glass. The Refraction therefore of the Rays of Light is not made in the Point of Incidence, but by little and little in the continual Incurvation of the Rays, which happens partly in the Air, and before they touch the Glass, and partly as it should feem in the Glass it felf after they have enter'd into it. Nor is the thing otherwise in Reflexions, as Sir Ilaac Newton hath fhew'd most accurately in the Book just before cited; whither we refer our Reader, who is fludious of these things. Now, because of the Analogy which is betwixt the Propagation of the Rays of Light and the Progress of

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of Bodies, it seem'd proper to premise the three foregoing Propolitions, as preparatory to true Optics. But we must note by the way with Sir Ilaac Newton, that fpheric Figures are more accommodate to Optic Ufes than any of the Conic Figures are. And according to his Opi-nion, if the Objective Glaffes of Perspectives were made of two Glaffes fashion'd spherically, and fitted to contain Water betwixt them; it might come to pais, that the Errors of Refractions, which happen in the extreme Surfaces of the Glaffes, would be corrected exactly enough by the Refractions of the Water. And he judgeth. that fuch Objective Glaffes are to be preferr'd before Elliptic and Hyperbolic ones; not only becaule they may be fashion'd more easily and exactly, but also because they refract the Pencils of the Rays, fituate without the Axis of the Glass. more accurately. However, the divers Refrangibility of divers Rays will ever hinder Optics from being brought to Perfection by Figures of Glaffes, either spherical or any other whatever. And unless the Errors arising from the foresaid Head can be corrected, all our Labour will be employ'd to finall purpole in correcting the reft. But as concerning all these things, the Famous Author himfelf is to be read in that noble Optic Work mention'd before.

Schol. (2.) But fince it hath feem'd good to that great Man, to propole certain Propolitions in that Book without their Demonstrations; it will be worth our while to bring in in this place the Demonstrations of them, which have been either lately found out, or elsewhere delivered by the fame Author; that fo there may be nothing in that Famous Treatile, which Beginners may fumble

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them.

LECT. XXV.

Prop. (1.)

ETACB be a reflecting ipherical Surface, the Center whereof is E. Let the Radius EC be bilected in the Point T. And if the

Points Q and q be marked in the Line E C, on the fame fide of the Point T: So that T Q, T E and Tq, be Lines continually proportional; and the Point O be the Focus of the incident Rays, the Point q will be the Focus of the Reflex Rays. For by the Hypothesis QT: TC:: TC:: Tq. and by compounding QT+TC = QC:QT: CT + Tq = Cq:CT = ET; that is, QC:OT:: Cq; ET. And by alternating QC: Cq::QT:ET. But by V. 19. of the Ele-ments, QT:ET::QE:Eq. Therefore by Equality QC: Cq:: QE: Eq. From whence, in the Triangle, the Bale whereof is Q q, and the Vertex in the fpheric Surface A C B, fo near to the Point C, that the greater of its Sides should be nearly equal to QC, and the leffer to qC; the Base Qq will be so divided by a Line drawn from the Point of the Sphere to the Center E, that the Parts Q E and E g fhould be one to another in the Proportion of the Sides QC and qC. And

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And confequently, by VI. 2. of the Elements, the Line drawn from the Vertex of the Triangle through the Center E, will biled the vertical Angle of the Triangle, and make equal Angles on both fides. From whence the Radij paffing through Q, because that the Angles of Incidence and Reflection are equal, will be reflected to the Point q, and conversity. Q. E. D. Prop. (2.) Let A C B be the refracting Sur-

Prop. (2.) Let A Q B be the refracting Surface of a Sphere, the Center whereof is E. In B the Radius E C produc'd on both

fides, let the Points T and t be Lat. page 8. mark'd; fo that as well E T as C t Caje 3. (which are equal one to the other)

may be to the Radius EC, as the leffer of the Sines of Incidence and Refraction is to the Difference of those Sines. Then let the Points Q and g be marked in the fame Line; fo that TO may be to E T or C t, as E t is to t q. But let the Places of the Points be such, that the Line t q may be on that fide of the Point which is contrary to that fide which the Line TQ is on, with respect to the Point T. Now, if the Focus of the incident Rays be in the Point Q, the Focus of the Refracted ones will be in q. For by the Hypothesis, as TQ is to TC, so is ET to t q. And by compounding, TQ is to TQ+TC=QC, as is ET = Ct to Ct + tq = Cq; and by alternating, TQ:Ct::QC:Cq. And by compounding and inverting TQ + Ct = QE: TQ::QC+Cq=Qq:QC. Or Qq: QC:: QE:: QT. From whence, together with what Hugens bath demonstrated in his Dioptrics, page 26. Szc. the Proposition is manifest. Prop. (2.) Let ACBD ( Page 8. Cafe 4.) be a refracting spherical Glass on both fides Convex or Concave, or at least Planc-convex or Planoconcave,



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concave, the Axis whereof ( or Line which cuts both Surfaces perpendicularly, and paffeth thro' the Center of the Sphere) is CD. In the Axis, let the Points F and f be the Foci of the refracted Rays, found as above : Those, to wit, which would agree to the Radii, parallel on both fides to the Axis, if there were only one refracting Surface. Let the Line F f be bilected in the Point E : and from the Center E, let a Circle be defcrib'd, with the Radius E F or E f. Now, let any Point Q O be the Focus of the incident Rays. Let QE be drawn interfecting the former Circle in the Points T and t, and let the Point q be marked in the fame Line, fo that the Line to may be to t E, as the fame t E or its Equal T E is to TQ. Then let the Line t q lie on that fide of the Point t, which is contrary to that which TQ lies on, with respect to the Point T. Thus the Point a will be the Focus of the refracted Rays; of those, to wit, which are near to the Axis, of which alone any Regard is to be had in these Cafes. For by the Hypothesis, TQ:TE::tE ! t q. Therefore by compounding TQ:TE + TQ, =QE::tE:tE+tq = Eq. From whence by V. 12. of the Elements, TQ :: QE :: TQ + t E = QE : QE + Eq = Qq. From whence, together with what Hugens hath demonstrated in Page 67. Oc. of his Dioptrics, the Proposition is manifest.

Prop. (4.) The Mixture of the Rays of the Sun in pt, a refracting Glafs (Book 1. Page 46.) is to the Mixture of the Rays of the Sun patting through an empty Hole, as the Breadth of the fame Glafs is to the Difference of the Breadth and Length of it, or as a g is to gm. For let a h be to a m, as a g is to A G. In this Cafe the Space a h will be equal to all the Areas of the lef-

fer Circles, as being, in the duplicate Proportion of the Rays on both fides: And the Mixture of the Rays would be equal, if fo be all the leffer Circles did meet together in that Space. But fince they are difpers'd through the Space pt, the Mixture will be as g h to g m. From whence, fince the Mixture of the Rays in the Glafs P T, is to the Mixture of them as paffing through the empty Hole, as A G is to G M, or as a g to g h; and the Mixture in the Glafs pt, is to the Mixture in the Glafs PT, as g h is to g m; by Perturbation of Equality, the Mixture in the Glafs pt will be to the Mixture which agrees to the Rays paffing without Refraction, as a g is to g m.  $\mathcal{Q}$ . E. D.

Prop. (5.) If a Body in any given Velocity whatever, fall upon a Space (see Book 1. Page 57.) of an inconfiderable Breadth, and terminated on both fides with parallel Planes, and in its Paffage towards the remoter Plane, be attracted or impell'd perpendicularly; in fuch fort, that the attracting or impelling Force is either every where the fame, or at leaft the fame at equal Diftances from that Plane; the perpendicular Velocity of the Body, which hath now passed that Space, will be equal to the Sum of the Squares of the former Velocity, and the square Root of the Velocity acquir'd in passing through. But if the Body be retarded in its passing through, the Difference of the same Squares is to be taken instead of the Sum of them; and thus the Proposition will hold good. It follows from Newt. Mathem. Principl. Prop. 29. Probl. 27. Corol. 2.

**Prop.** (6.) If any Bodies or Rays of Light paffing through a Space, fuch as before, and bounded with parallel Planes, be acted upon in their paffing with a like Force, but which is formetimes

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times greater, fometimes lefs; the Book 2. Page Motion last acquired will be in the fubduplicate Proportion of the Force, which begets the fame; in fuch fort.

that the Squares of the Motions will always determine the true Proportions of the fame. Let AB be a refracting Surface, or let it reprefent a Space of a Thickness not to be regarded, bounded with parallel Planes, which is of a refractive Force. And let I C be a Ray of Light falling very obliquely upon the refractive Plane at the Point C, fo that ACI the Complement of the Angle of Incidence may be indefinitely fmall. And let CR be the refracted Ray. From any Point B, let the perpendicular BR be erected, cutting the refracted Ray in R. Where if CR represent the Motion of the refracted Ray, which is refolv'd into two Motions CB and BR; that Part of the Motion CB will be parallel to the refracting Plane, and BR will be perpendicular to the fame; and fince the Motion along the Plane A B, is in no wife chang'd by the Force perpendicular to the fame, C B will be given, by reafon. of the given Velocity of the Rays, which here fuppos'd. And the Line BR will be a Motion produc'd by the Refraction in a given Time; and it will be in the fubduplicate Proportion of the Force which produceth it. For, because of the given Latitude of the refractive Space, the Times of the Transit in which the refracting Force would act, will be reciprocally as the Velocifies produced, or as the producing Forces reciprocally; and becaufe of the Velocity, the Force being given in the Proportion of the Times, the Line BR would be as the producing Forces reciprocally; and the Time being given, as the fame Forces directly. Therefore neither being given. enty :

given; the Line B R will be in the fubduplicate Proportion of the Forces; for then the Times and Forces being reduc'd unto an Equilibrium, neither will the Force preponderate the Time, nor the Time the Force; which no otherwife could answer each to other. Thus, indeed, if the Forces be put to be in a quadruple Proportion; a double Velocity will be produc'd in half the Time; or the Line B R will be double of the fame Line, and thus every where. Q. E. D.

**Prop.** (7.) In the Solution of the Rainbow, the Arch QF and the Angle AXR will be the greateft, where ND is to CN, as  $\sqrt[7]{II-RR}$  is to  $\sqrt[7]{2}$  RR. In which Cafe alfo, 2 R : I :: NE, ND; and the Angle AYS, which is made by the Rays AN and HS; will be the leaft, where ND is to CN; as  $\sqrt[7]{II-RR}$  is to  $\sqrt[7]{8}$  RR. In which Cafe alfo, 3 R : I :: NE : ND. Which twofold Proposition we will demonstrate with the Famous Sir I are Newton, in his Manuscript Optic Lectures.

**Problem.** If Rays, whether parallel or inclin'd towards foms common Point, do fall upon a Sphere to be refracted there; to defign the Concourle of the refracted Rays without the Axis, which are next to one another, and lye in the fame Plane with the incident Rays. Let AN (in Fig. 6. Plate 7.) be an incident Ray. Let AN (in Fig. 6. Plate 7.) be an incident Ray, NK the refracted Ray thereof; and NV in the Plane of the Triangle ANK, a right Line touching the Sphere in N. To AN draw NR perpendicular, and meeting the Axis AC in R: and RV parallel, and meeting the Tangent NV in V. Likewife, draw NQ perpendicular to NK, and VQ paralel to the fame, meeting it in Q; and draw QC meeting NK in Z, Z will be the Concourfe of the Rays neareft to AN. And let An be

another of the incident Rays, infinitely near to the former AN, and meeting NR in G. Draw nZ meeting NQ in H; and to AN and NK; let down from C the Conter of the Sphere the Perpendiculars C DI and C E. meeting An and nZ in d and e. Now, fince AN is fuppos'd to be infinitely near to An, the Arch N n, which is infinitely fmall, may be reckoned for a right Line coinciding with the Tangent NV; and the Triangles NGn, NRV, as alfo NHn, NQV may be accounted for like. Wherefore it is, DC:: Dd:: (NR:NG::NV: Nn::NQ:NH):: EC: Ee. And converfly, DC:(DC--Dd) dC::EC:(EC---Ee) eC; and alternately, DC:EC::dC: e C. But DC is to EC, as the Sine of Incidence is to the Sine of Refraction, because NK is the refracted Ray of AN; and confequently allo, dC is to eC as the Sine of Incidence to that of Refraction; and therefore, fince the Angles DAd and EZe be infinitely fmall, and confequently Cd is perpendicular to An, and C e to n Z, or at least equipollent to perpendicular, nZ will be the refracted Ray of An. 2 E. D.

Corol. (1.) ND: N E (or NP: NF):: NR: NQ. For NC being drawn, becaule of the Triangle NDC, like to the Triangle NRV; and the Triangle NEC like to the Triangle NQV: it is ND: NR (:: NC: NV) :: NE : NQ: and alternately, ND: NE: NR: NQ. Hence refults a more ready Solution of the Problem; to wit, Erect NR, NQ perpendicular to the Rays AN, NK; of which two Perpendiculars, let NR meet the Axis AC; and let the other NQ be to NR, as NF is to NP. Then draw QC, which will meet with NK in the fought Point Z. Corol.

Corol. (2.) It is also thus,  $AN \times DC \times NE$ :  $A D \times EC \times ND$ :: NZ: EZ. For AD: A N:: DC: NR. And from thence NR =  $A N \times DC$ . A D. Likewife, ND: NE:: NR: NQ.

And from thence,  $NQ = \frac{A N \times D C \times NE}{A D \times N D}$ 

And confequently,  $AN \times DC \times NE$ :  $AD \times ND \times EC$  (:: NQ: EC) :: NZ: EZ.

Corol. (2.) If A the Radiant Point be infinitely diffant, or fend forth parallel Rays; it being put thus, I: R : the Sine of Incidence: the Sine of Refraction; it will be I  $\times$  N F : R  $\times$  N P :: NZ : E Z. For in this Cafe, A N and A D, fince they are infinitely long, ought to be reckon'd for equal; and confequently by Corol. 2. of this, D C  $\times$  N E : E C  $\times$  N D :: NZ : E Z. But by the Hypothefis, D C : E C :: I : R; and confequently I  $\times$  N E : R  $\times$  N D (:: N Z : E Z) :: N P : N F. But it is to be noted, that the Refolution of the Problem is eafily accommodated to any Cafe whatever, mutatis mutandis; whether the incident Rays decline from fome one Point, or incline to the fame, or fall parallel.

Problem (2.) From parallel Rays refracted unto a Circle, to determine that Ray; Part whereof being included in the Circle, hath a given Proportion to that Part of the fame Ray refracted, which is included in the fame Circle. Let A N be the incident Ray: (fee Fig. 1. Plate 8.) N K the refracted: N P and N F, the Parts of them included in the Circle : C D and C E Perpendiculars let down to those Parts from the Center of the Circle; and B C a Semi-diameter drawn parablel to AN. And let it be C D: C E :: I : R and N P; N F :: p : q. These Things being put, T a

that the Point N may be found, which determines the Rays A N and NK ; creet at BC the Perpendicular B X, and let the Square thereof be to the Square of BC, as <u>qq pp</u> is to <u>II-RR</u>. Thus рр C X being drawn, will eut the Circle in the int N. For by the Hypothesis p:q NF::) ND:: NE. And I:R:: : C.E. Wherefore  $\frac{q}{p}$  N D = N E; and  $\frac{R}{I}$  $\rho = CE$ . Furthermore, fince ND q + CD a  $= N \cup q$  = N E q + C E q; take from hence NDq + CEq, and there will remain C D q + C E q = N E q - N D q; that is, by substituting the Values of C E and N E, just now found,  $CDq - \frac{RR}{II}CDq := \frac{qq}{pp}NDq - \frac{q}{P}$ - R R Η· N D q; and by making a Reduction -C D q, =  $\frac{q q - p p}{p p} N D q$ . Which being refolved into Proportionality, <u>qq-pp</u>: <u>II--RR</u>  $(::CDq:NDq)::BXq:BCq. \ Q.E.D.$ 

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April 8. 1706.

#### LECT. XXVI.

Problem (3.) HE Sun enlightning a transparent Sphere, to determine the greatest Inclianation to the Axis of the Rays emerging after one

Reflection. Let BNK (in Fig. 2. Plate 8.) be the proposed Sphere: BQ the Diameter, or the Axis parallel to the incident Rays: AN fome one of the incident Rays: NF the refracted Ray thereof: FG the reflected : and GR the again refracted; and thus the greatest Angle, which RG can make with the Axis BQ, is to be fought.

To which purpose it is to be observed, that in that Cafe alone, where RG is the most of all inclin'd to BQ, the Rays which are the nearest to A N, can emerge parallel to R G. For in other Cafes, of the emerging Rays nearest to it, some are continually more inclin'd to BQ, others lefs; and confequently are fomething inclin'd to one another.

It is also to be observ'd, that the Rays which meet at the Point of Reflection, will emerge parallel. For draw a n parallel to A N, and as near to it as may be; and let n f be the refracted Ray thereof, fg the reflected; and gr the fecond refracted one. And the Points F and fcoinciding. when the Angles NFn and GFg are equal; and the Refractions at N n and G g be like, the emerging Rays GR and gr will be parallel, as well as the incident NA and a n.

The Ray AN is therefore to be fought, the refracted Ray whereof concurs with the refracted  $\mathbf{T}$ one

one of the Ray an, which is nearest to it in the Point F. And indeed, by Corol. 2. Problem 1. (if C D and C E be let down from the Center of the Sphere perpendicular to the Rays, and it being put I:R::CD : CE.) If those Rays meet in any Point Z, it will be  $I \times NF : R \times NP$  $(:: N \mathbb{Z} : \mathbb{E} \mathbb{Z}) :: N \mathbb{F} : \mathbb{E} \mathbb{F}$ , (the Point Z, to wit, falling at F, according to the Hypothesis,) 1:2:1. Wherefore  $I \times NF = 2R \times NP$ ; and I:2 R:: NP: NF. The Proportion therefore of NP to NF is given ; and from thence, by the 2d Problem, the Point N will be given. To wit, Let the Tangent BX be drawn at the Top of the Circle, the Square whereof let it be to the Square of the Semi-diameter BC, as  $\angle R R - II$ is to II-RR, and let CX be drawn. For this will meet the Circle in N; and from N when found, the other things will eafily be determin'd.

Corol. (1.) Hence it comes to be thus, 2 R R: II--RR:: CNq:NDq. For fince it is, 4 R R --II:II --RR:: BXq:BCq, by compounding it will be 2 R R: II--RR (:: CXq:BCq): CNq:NDq.

Corol. (2.) It is alfo thus, I: 2 R:: ND: NE. For it was above I: 2 R:: NP: NF. And from these the Resolution of the Problem becomes more expeditious.

Scholium. With the greateft Inclination of the Radius RG, the greateft of the Arches FQ, bounded at the refracted ones NF, is allo given. For the Angle FCQ fubtended by FQ, is equal to the Angle which CF and AN comprehend, *i. e.* equal to half of the Angle comprehended by RG and AN, or BQ; and confequently that which is defin'd by the Ray AN falling upon the Point which is now found, is the greateft of the Arches Arches F Q, as well as of the Angles comprehended by R G and B Q.

**Problem** (4.) The Sun illustrating a pellucid Sphere, to determine the least Inclination to the Axis of the Rays emerging after two Reflections.

Let AN and a n be two incident Rays very near to one another; and let them, after two Reflections in Ff and Gg, emerge along HS and hs. Now it is manifest, that only in that Cafe where the acute Angle comprehended by BQ and SH is the leaft, those Rays HS and hs can be parallel, as was faid before of the Rays G R and gr: And where this happens, the Ray FG will be parallel to fg. From whence, double the Arch Ff(= to the Arch Ff + Gg = to the ArchFG-fg = to the Arch NF-n f) is = to the Arch n N --- F f; and confequently triple the Arch F f = is equal to the Arch N n. And fince N F is divided in Z, in the Proportion of those Arches, as is manifest; NZ will be = to  $_{3}ZF$ , or 2 E Z. Since therefore by Corol. 2. Problem 1.  $\mathbf{I} \times \mathbf{NF} : \mathbf{R} \times \mathbf{NP} : : \mathbf{NZ} : \mathbf{EZ} : : 3 : 1.$ therefore the Proportion of NP to NF is given; and from thence, by Prob. 2. the Point N will be given, by drawing B X, which may touch the Circle in the Vertex B; and the Square whereof is to the Square of BC, as 9 R R-II is to II-RR: and by drawing CX to meet the Circumference in N. N therefore being found, the other things are eafily determined.

Corol. (1.) Hence it is, 8 R R: II - R R:: C N q: N D q. for 9 R R - II: II - R R:: BX q: B C q. And by compounding 8 R R: II -- R R (:: C X q: B C q) :: C N q: N D q. Corol. (2.) It is alfo thus: I: 2 R:: N D: N E. Forafmuch as above it was, I: 3 R:: N P: N F.

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A Scholium. The greateft Inclination of the Ray K T to the Axis, when it emerges after three Reflexions, will be determin'd in the fame manner as the greateft of the Arches Q G; to wit, in that Cafe F G and f g will meet together in G, and the Arch F f (= to the Arch F g- f g= to the Arch NF - n f) will be equal to N n - F f: and from thence doubled, the Arch F f will be = to the Arch N n and N Z will be equal to 2 Z F, and confequently 4:1::N Z:E Z:: (by Corol. 2. Prob. 1.) I x N F: R x N P:: or I: 4 R::NP:NF. And confequently by Prob. the 2d, 16 R R- II: II-R R; B X q: BCq. From whence it follows, 15 R R: II-R R;: C N q: N D q; and I: 4 R: t N D: NE.

And fo if the leaft Inclination of a Ray energing after four Reflexions be enquir'd, it will be determin'd by making it thus; 25 R R - II: II-RR: : BXq: BCq. Or 24 RR: II-RR : : C Nq: N Dq. And I: 5 R:: ND: NE. And fo on in infinitum.

Scholium. From the Determination of the Bounds of the Rainbow by the Famous Newton, I will take occasion to folve a certain Phanomenon, or rather the Absence of a certain Phznomenon, which fometime hath feem'd to me very difficult, and almost infolvable. And it is this: Why we do not fee a Rainbow about the Sun, at the Diftance of about 26 Degrees; fince there the Rays come to our Eyes by a double Refraction without any Lofs in the Reflexion? For by Computation, there is in that Place a Conffipation of the Rays requilite and sufficient for affecting the Sight. And it increaseth the Doubt, that it feems probable at the first that this should be the principal Rainbow of all, and decked with the most lively Colours, as proceeding from a double

double Refraction, without any Lofs or Diminution of the reflected Rays. For, as the Primary Rainbow doth far exceed the Secondary, for that it proceeds from a double Refraction, and one fingle Reflexion, whereas in the Secondary there are two Reflexions; fo one would think, that the Rainbow which we spake of, should, in the Splendor of its Colours, as much exceed the Primary one of the two last mention'd, as that doth the Secondary; and fhould appear about the Sun like a most noble Crown, whenfoever the Air affords Spherical Drops in that Angle of 26 Degrees for making a Rainbow. And it is to be admir'd, that this Difficulty hath never been touch'd upon by those Philosophers that have treated of the Rainbow. But our Solution of it is this: We do therefore fee the Rays that are throng'd about the Limits F and G, rather than the other, because fo many of them, as AN an, which enter'd the Drop of Rain parallelwife, return back in the fame manner, as R G, rg: SH, sh, and fo enter the Eye together; whereas on the other hand, unless they did come forth also in that parallel manner, they would make fome Angle, and fo could not enter the Eye together, how thick and throng foever they might otherwife be about the Point F and G. From whence, feeing the Rays which come forth about the Point F, do not go forth parallel, but make a certain Angle; it is manifest, that they cannot enter the Eye together, and confequently cannot afford a Rainbow.

LVI. All the Parts of an Homogeneal Mathematical Fluid, [that is, a Body, the Parts whereof yield to any Force whatever impress'd, and in yielding are easily mov'd amongst themfelves, ] which Fluid is inclos'd in any unmov'd Yessel, and press'd on every Side, are (if you let, afide

afide the Confideration of Condensation, of Gravity, and of all Centripetal Force) equally prefs'd on all Sides; and remain every one of them in its Place, without any Motion arising from that Preffure.

Case (1.) Let a Fluid, contain'd in a spherical Vessel, be uniformly presid together on all Sides: No Part of the fame will be mov'd by that Preffure. For as if fome Part as D (fee Fig. 2. Plate 8.) were mov'd thereby, all the like Parts at the fame Distance from the Center must be mov'd in like manner : and this because the Preffure of them all is alike and equal, and we have excluded all Motion but what arifeth from the Preffure. But they cannot all come to the Center, but the Parts about the Center must be condens'd, which is contrary to the Hypothesis; nor can they recede farther from it, but there will be a Condenfation about the Circumference, which is likewife contrary to the Hypothesis : nor can they be mov'd in a Circle about the Center; for every Force which should determine the Motion of any one Part, or of them all laterally, and to either this Side or that, is here excluded; much lefs can the fame Part be mov'd contrary ways at the fame time. Therefore no Part of the Fluid will, in this Cafe, be mov'd out of its Place. Q. E. D.

Cafe (2.) Each of the fpherical Parts of this Fluid are equally prefs'd on every Side. For let E F be one of those Parts; and if it be not equally prefs'd on all Sides, let the leffer Preffure be ingreas'd, until the Preffure be every where uniform and equal; and then the Parts thereof, by the ift Cafe (which belongs as well to this little Sphere, as to one contain'd in a folid Veffel) will remain in their Places. But before that Increase, they would remain in their Places by the fame first Case;

Cafe; [for we treat here of fuch a Fluid, the Parts whereof, as we there demonstrated, do remain in their Places;) and by the Addition of a new Force, they will be mov'd out of their Places by the Definition of a Fluid : Which two Things are contradictory. Therefore it was falfly faid, that the Sphere E F was not equally prefs'd on every Side. Q. E. D.

Cafe (1, 2) Befides, the Preffure of the divers fpherical Parts will be equal. For the fpherical Parts prefs each other equally in the Point of Contact, becaufe that Reaction is always equal and contrary to Action. But alfo by the 2d Cafe, all the fpherical Parts whatever, are equally prefs'd on all Sides. Therefore any two fpherical Parts, not contiguous, are prefs'd with the fame Force, becaufe the intermediate fpherical Part toucheth both. Q. E. D.

Cefe. (4.) All the Parts of this Fluid are equally prefs'd on all Sides. For any two Parts may be touch'd by fpherical Parts in any Points whatfoever; and there they do equally prefs those fpherical Parts by the 3d Cafe; and because of Reaction, which is always equal to Action, they are reciprocally equally prefied by them. Q. E. D.

Cale  $(\varsigma.)$  Since therefore any Part whatever of this Fluid, as GHI, is inclosed in the reft of the Fluid, as in a certain Vessel, and is equally presed on every Side; and the Parts thereof do equally press one another, and are at reft among themselves; it is manifest, that all the Parts of any Fluid whatever, as GHI, do equally press one another, and are at reft amongs themselves.  $\mathfrak{D}$ . E. D.

Cafe (6.) Therefore, if that Fluid be inclos'd in a Vessel which is not rigid or unyielding, and be not equally press'd on every Side, the fame will

will give way to the ftronger Pressure by the Definition of Fluidity.

Cafe (7.) Therefore, in a rigid Veffel, a Fluid will not fuftain a ftronger Preffure on one Side than on another; but will give place to the fame: and that in a moment of Time, becaufe the ftiff Side of the Veffel does not purfue the yielding Liquor. But in yielding, it will prefs the opposite Side; and fo the Preffure will incline to an Equality on every Side. And becaufe the Fluid, as foon as it endeavours to depart from the Part which is more prefs'd, it is ftopp'd by the Refiftance of the Veffel on the opposite Part, the Preffure will on every Side be reduc'd to an Equality in a Moment of Time, without local Motion; and immediately the Parts of the Fluid will, by the 5th Cafe, prefs one another equally, and be at reft amongft themfelves. Q. E. D.

Corol. Hence the Motions of the Parts of fuch a Fluid amongft themfelves cannot be chang'd by a Preffure, which is upon the external Surface of the fame in any Place thereof, unlefs either the Figure of the Surface be fomewhere chang'd, or all the Parts of the Fluid, in preffing one another more vehemently, or more remifsly, flide more eafily or difficultly amongft themfelves.

Corol. (2.) But fince the Definition and Affections of this Mathematical Fluid do feem altogether to agree with the Nature and Phænomena of natural Fluids; it is reasonable, that the Demonstrations of these Cases be applied to our natural Fluids, to Water especially, and the like. From whence it will be manifest, that the rest of the interval Parts of a Fluid amongst themselves, doth in no wise contradict the Nature of Fluidity; and that all the Motion of the Parts of Fluids amongst themselves, is to be reckon'd as owing to Fermen-

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Fermentation, Heat, or other extrinick Caufes, rather than to the Nature it felf of Fluidity. For, if the Parts of a Body be either fpherical, or fpheroidal, and perfectly polifi'd, fo that they can never be join'd one to another, but rather do only touch another in phyfical Points; a Congeries of these Particles will compose a Body, such as is commonly call'd a Fluid, altho' those Particles be altogether at rest amongst themselves: A Fluid therefore seems to consist of Parts very moveable, but not necessarily mov'd.

June 2. 1706.

#### LECT. XXVII.

F all the Parts of a fpherical homogeneous Fluid, which are at equal Diffances from the Center, and lie upon a concentrical, fpherical Bottom, do e-

qually gravitate towards the Center of the whole; the Bottom suffains the Weight of a Cylinder, the Base whereof is equal to the Surface of the Bottom; and the Altitude is the same as that of the Fluid, which lies upon the Bottom.

Let (in Fig. 4. Plate 8.) d h m be the Surface of the Bottom, and a e i the upper Surface of the Fluid. Let the Fluid be diffinguished by innumerable fpherical Surfaces, into concentrical Orbs of equal Thickness; and imagine the Force of Gravity to act only upon the higher Surface of every Orb, and that the Actions upon equal ' Parts

Parts of all the Surfaces are equal. The uppermoft Surface therefore a e i, is prefs'd by the fimple Force of its own Weight, wherewith as well all the Parts of the supreme Orb, as the 2d Surface bfk (by Prop. 56.) are prefs'd. Belides, the 2d Surface bfk is prefs'd by the Force of its own Weight, which being added to the former Force makes a double Preffure. By this Preffure, and by the Force of its own Weight withal, that is, by a treble Preffure, the 2d Surface cgl is urg'd: And fo the 4th Surface is urg'd with a fourfold Preffure, the sth with a fivefold, and fo on. The Preffure therefore wherewith every Surface is urg'd, is not as the folid Quantity of the Fluid which lies upon it, but as the Number of Orbs unto the Top of the Fluid; and is equal to the Gravity of the lowest Orb, multiplied by the Number of Orbs ; that is, to the Gravity of the , Solid; the last Proportion whereof unto the before defin'd Cylinder ( if fo be the Number of Orbs be increas'd, and their Craffitude diminish'd infinitely; fo that the Action of Gravity be render'd continual from the loweft Surface to the uppermost) will become a Proportion of Equality. The lowest Surface therefore fustains the Weight of a Cylinder, the Bafe whereof is equal to the Surface of the Bottom: and the Altitude the fame as that of the Fluid lying upon it. Q.E.D.

And by the like reasoning the Proposition is manifest, where the Gravity decreaseth in any Proportion of the Distance from the Center, which may be assign'd; as also where the Fluid upwards is more rare, and more dense beneath. Q. E. D.

Corol. (1,) The Fluid therefore is not urg'd by the whole Weight of the incumbent Fluid; but fuftains only that Part of the Weight, which is defin'd defin'd in this Proposition ; the reft of the Weight being fustain'd by the arched Figure of the Fluid.

Corol. (2.) If an entire Sphere confifteth of fuch a Fluid to the very Center, the Center will fuftain no Weight; the whole Weight being upheld by the arched, or rather in this Cafe by the fpherical Figure of the Fluid.

Corol. (2.) But in equal Diftances from the Center, the Quantity of the Preffure is the fame, whether the Surface be prefs'd parallel to the Horizon, or perpendicular, or obliquely; and whether the Fluid, as continued upwards from the Surface prefs'd, arifeth perpendicularly according to a right Line, or creeps along obliquely by crooked Cavities and Channels, and those regular or never fo irregular, wide or narrow. That the Preffure is nothing chang'd by these Circumftances, is gathered by applying the Demonstration of this Proposition to every Cafe of Fluids.

Corol. (4.) By the fame Demonstration (and Prop. 56. foregoing) it is collected, that the Parts of an heavy Fluid acquire no Motion amongst themselves from the Pressure of a Body lying upon them; if so be the Motion which ariseth from Condensation be excluded.

Corol. (5.) And therefore, if another Body of the fame specifick Gravity, which cannot be condens'd, be immerg'd into this Fluid, it will acquire no Motion from the Weight of the Body lying upon it; it will not ascend nor descend; nor will it be compell'd to change its Figure: If it be spherical, it will remain spherical; if square it will remain so, the Pressure notwithstanding; and this whether it be hard or soft, or even the most Fluid; whether it float freely in the Fluid, or sink. For every internal Part of the Fluid hath the Nature of a Body immers'd; and the fame

fame is to be faid of it, as of all Bodies immerg'd, which are of the fame Magnitude, Figure, and fpecifick Gravity. If the Body immerg'd fhould become liquid, keeping its Weight ftill, and put on the Form of a Fluid, if it before alcended or defcended; or from the Preflure put on a new Figure, it would do the fame ftill; and that becaufe the 'Gravity thereof, and the other Caufes of Motion, do remain. But by the  $\mathfrak{sth} Ca/e$  of the former Proposition, it would now reft, and retain its Figure : Therefore it did fo before alfo.

Corol. (6.) Therefore a Body, which is fpecifically more heavy than the Fluid which is contiguous to it, will fink; and that which is fpecifically lightet will afcend, and will obtain fo much Motion and Change of Figure, as that Excefs or Defect of Gravity can caufe. For that Excefs and Defect hath the Nature of an Impulfe, wherewith the Body, otherwife conflituted in an Æquilibrium with the Parts of a Fluid, is urged; and may be compared with the Excefs or Defect of Weight in either Part of a Balance.

Corol. (7.) Therefore the Gravity of Bodies placed in Fluids is twofold, the one true and abfolute; the other apparent, vulgar, and comparative. The abfolute Gravity is that whole Force whereby a Body tends downwards, or would defcend in an empty Place. The relative and vulgar Gravity is the Excels of Gravity, whereby a Body doth tend more downwards than the Fluid that encompaffeth it. The Parts of Fluids, and of all Bodies, gravitate in their proper Places with the former Sort of Gravity; and therefore with their conjoin'd Weights, they compose the Weight of the whole. For every whole is heavy, as may be experienced in Vessels full of Liquors; and the Weight of the Whole is equal to the Weights

Weights of all the Parts, and therefore is compounded of the fame; for it can be deriv'd from no where elfe. With the other Sort of Gravity, which may be called apparent, vulgar, and comparative, Bodies do not gravitate in their own Places, or immers'd in their Fluids respectively, that is, compar'd amongst themselves, are not one heavier than another, but hindering the Endeavours of one another to defcend, they abide in their own Places, as if they were not heavy: Like as any heavy Bodies whatever, placed within a concave Sphere from the Equality of Gravitation every way, do appear in no wife to gravitate, as was observ'd above. Thus the things which are in the Air, and do not overweigh, or defcend at all in the Air, as Clouds and Vapours are judged by the Vulgar not to gravitate at all. What things overweigh the Air, and confequently descend in it, as not being suffain'd by the Weight of the Air, as Hail and Drops of Rain; these the Vulgar judge heavy. The vulgar Weights are nothing elfe, but the Excels of the true Weights above the Weight of the Air. From whence also those things are commonly efteem'd light, which are lefs heavy; and whilft they give way to the Weight of the Air, alcend upwards. And they may be faid to be comparatively light, but not absolutely, for that they would descend in a Vacuum. So likewise in Water, the Bodies which afcend or defcend by reafon of their leffer or greater Gravity, are comparatively and apparently light and heavy; and their comparative Levity or Gravity is the Defect or Excels wherewith their true Gravity either is exceeded by the Gravity of the Water, or doth exceed it. But what things do neither afcend nor descend, albeit they by their true Weights increase

creafe the Weight of the whole; yet comparatively, and in the Senfe of the Vulgar (and of Philosophers too of late) they do not gravitate in the Water: For the Demonstration of these Cases is the fame.

Corol. (8.) Those things which have been faid concerning Gravity, and the Force whereby Bodies descend to the Center of the Earth, in either an absolute, or reciprocal duplicate Proportion of the Distances, are to be understood to hold in all other Centripetal Forces both absolute, and increas'd or diminish'd, according to any Proportion whatsoever of the Diminution or Increase of the Distance.

Corol. (9.) And therefore, if the Medium, wherein any Body is mov'd, be urg'd either with its own Gravity, or any other Centripetal Force; fo that the Body, by means of the fame, is pufh'd on more forcibly than otherwife it would be; the Difference of the Forces is to be estimated from that moving Force which, in what goes before, we have confider'd as a Centripetal Force. But, if the Body be urged more lightly, the Difference of the Forces is to be reckon'd for a Force which tends from the Center.

Corol. (10.) Since Fluids, in preffing the included Bodies, do not change their external Figures; it is manifeft, by the Corollaries of the foregoing Proposition, that they will not change the Situation of the internal Parts amongst themfelves; and confequently will not hurt Animals immers'd in them, neither will they excite any Senfation in them, if Senfation depends upon the Motion of the Parts; only fo far forth as these Bodies may be condens'd by a Compression, which is on all Sides of them: And the fame is to be faid of any System of Bodies whatever, which

which is encompais'd with fome Fluid, and comprefs'd thereby. As to the Parts of the Syftem, and their Motions, it will be the fame as if it was in a Vacuum; and they will retain only their comparative Gravity, unless the Fluid do refift their Motions, or be necessary to the keeping them together by its Compression.

LVIII. Fluids, which are not in a Descent, do in every given Base press themselves, and other Bodies, as well those which are immers'd, . as the containing, in Proportion of the perpendicular Altitude, and not of the Quantity of the Matter; that is, the Preffure of a Cylinder of Water, of the Heighth of four Feet, where the Area of the Circle of the Cylindrical Column is only of one Inch square, is equal to the Pressure of a Cylinder of Water of four Feet high, where the Area of the Circle of the Cylindrical Column is of 200 or 1000 square Inches, and thus every where; to wir, if the Bale of the Water, communicating with the Water contained in the Tube, be in both Cafes equal. This is a most known Rule in Hydrostaticks, often found by Experiments, which I fhall endeavour to demonstrate as follows, it not having been demon-strated hitherto, as I suppose, physically or mathematically. It is well known, that the Quantity of any moving Force, or the Effect answering thereto, doth arife from the Quantity of the Matter multiplied into the Velocity; and confequently that whatfoever the Matter which is mov'd be, as to its Quantity, the Preffure will be the fame, if the Velocity be reciprocally proportional thereto. Thus the Forces of the Balance, Leaver, and other fuch mechanical Inftruments. is deriv'd from the Combination of these two things, the Matter and the Velocity; and you may

may move any Weight whatfoever, by any Force how fmall foever, if the Diftance of the Preffure. or fmaller Weight from the Hypomochlion, be reciprocally proportionated to that of Weights. Thus one Poundweight . at the Diftance of four Feet from the Hypomochlium, is as much as four Pounds at thr Diftance of one Foot ; for here the fingle Pound is mov'd with a Velocity, which is fourfold of that of the other ; and confequently is of equal Force in its Motion. Whilft it is moving. I fav, but not otherwife, as many feem to reckon. For if at any time the Machine refts, it is manifeft, that the Gravity, or Preffure, or Force, is now in Proportion to the Matter, and the four Poundweight is four Pounds, and the one Pound weight no more than one Pound. If at any time the Machine refts, I fay. For indeed, if we speak physically, or at least mathematically. no Body doth wholly reft; but every Body is then faid to reft, when the Quantity of the Motion is fo fmall, that it cannot be perceiv'd.] According therefore to what hath been faid, the Water contain'd, and the Veffel also which contains it, are always in some fort of Motion, and do never perfeetly reft; [ which indeed if they did, a Column of Water, as I suppose, that is, an hundred-fold greater, and confequently of an hundred fold greater absolute Gravity, would press the Veffel an hundred fold more : ] But the absolute Reft of the fame not being to be supposed, it is to be faid, that the Preffure of a Column of Water, the Area of the Bale whereof is of one Foot, and that of a Column whole Bale is 100 Feet fquare, is the fame; while the Excess, to wit, of the latter Column, in respect of an hundred-fold Magnitude and absolute Gravity, is compensated and counter-balanc'd by the hundred-fold greater Velo-



Velocity of Descent in the leffer Column: And that the Cafe is the fame here with that of an inverted Syphon of unequal Legs, where the Water keeps an Æquilibrium, by reason of the Velocities of Ascent or Descent, which are in both the Channels reciprocally proportional to the Ouantity of the Water.

Corol. (1.) Fluids therefore prefs, not in Proportion of the Quantity of the Matter prefling, but of the perpendicular Altitudes.

Corol. (2.) Therefore a Wooden Trencher. thrust down to the Bottom of a Bucket of Water, will rife to the Top, notwithftanding the Quantity of the Water which lies upon it is much greater than that which is under it; for by reafon of the Communication which is, by the fmall Interval that is betwixt the Edge of the Trencher and the Bucket, betwixt that little Cylinder of Water that is under the Trencher, and that Cylinder which is above it; the Weight of the incumbent Water will make that which is under, to lift up the Trencher with a Force equal to the faid incumbent Weight.

Corol. (2.) There is no Occasion therefore for the Hylarchic Principle of the Famous Dr. More, for the folving this Effect.

Corol. (4.) Thus it is with Fluids not actually descending. But if they, with the Vessel containing, do by the common Force of Gravity of them all actually descend; the Communication of the Preffure, as I suppose, paffeth away, and the Effect thereof ceaseth; but this fo, that even still, as in the former Cafe, the Pressure is according to the perpendicular Altitude, and is the fame where that Altitude is the fame, whatfoever the Column is otherwife, as to its Magnitude, whether little or great: So that at length we U 2 may

may make the Proposition more universal, and fay, without any Restriction, that Fluids do press according to their perpendicular Altitudes, and not according to the Quantity of the Matter.

November 11. 1796.

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# LECT. XXVIII.

LIX F the Denfity of a Fluid, compos'd of Particles which do flee from each other, be as the Compreffion; fo that if the preffing Force be two, or four, or eight-

fold, the Denfity thence arifing is fo likewife; the Centrifugal Force of the Particles is reciprocally proportional to the Diftances from the Center: And, vice versa, where the faid Force is reciprocally proportional to the Diftances from their Centers, the Particles which flee from each other compose an elastic Fluid, the Denfity whereof is proportional to the Compression.

Let a Fluid be underftood to be inclos'd in the cubic Space ACE, (fee Fig. c. Plate 8.) and then to be brought by Compression into a less cubic Space a c. Here the Distances of the Particles, by reason of their keeping the like Situation amongst themselves in both Spaces, according to the Nature of Fluidity, will be as the Sides of the Cubes A B, a b; and the Densities of the Fluid, reciprocally as the cubic Spaces ACE, a c e. Let the Square DP be taken in the Side ABCD of the greater Cube,

Cube, equal to d b the Square of the Side of the leffer Cube. And by the Hypothesis, the Preffure wherewith the Square DP urgeth the inclofed Fluid (or is urged thereby) will be to the Preffure wherewith the Square d b urges its inclofed Fluid, as the Dentities of the Medium are to each other; that is, a b cub. to A B cub. But the Preffure wherewith the Square B.D urgeth the inclosed Fluid, is to that wherewith the -Square DP urgeth its Fluid, as the Square DB is to the Square DP, or as ABq is to a bq. Therefore, by Equality, the Preffure wherewith the Square DB, urgeth its Fluid, is to that wherewith the Square d b urgeth its Fluid, as a b is to AB; or reciprocally, as the Diftance of the Particles. For the triplicate Proportion of the Sides a b and A.B. being substracted from the duplicate Proportion of the same; the simple Proportion of the Sides, or the Distance of the Particles, remains reciprocally proportional to the Preffure of the fame upon the Veffel containing. As for Example: Let the greater Cube be eight-fold of the leffer, or the Side of the greater double of the Side of the leffer. Then indeed, the Denfity of the Fluid, in the leffer Veffel, will be eight fold of the Density in the greater, by reason of the same Quantity of Matter contain'd in an eight-fold less Space. And when by the Hypothesis, the Compression into the given Space was made ex-actly proportional to the Density, the whole Compression, or the compressing Force, adequate to the fame in the leffer Cube, will be in the eight-fold Proportion of the Compression, or comprelling Force, in the greater : But the entire Surface, wherewith the Compression, or the Surface of every Square in the leffer Cube, is to the Surface of every Homologous Square in the U A greater, • · · · · · · ·

greater, in a fub-quadruple Proportion. The eight-fold Preflure therefore is to be compared with another Preflure of the fame Particle, difpers'd into a four-fold greater Space: Therefore, in a Space four-fold lefs, the fame Quantity of Matter, or the fame Particles of the Floid, fultain an eight-fold Preflure; wherefore every fingle Particle muft needs fultain a Preflure two-fold greater; or, that the Centrifugal Forces of the Particles floud be reciprocally proportionol to the Diffances of the fame. <sup>10</sup>Q. E. D.

Thus if, by the Planes FGH, fgh drawn through the midit of the Cubes, the Fluid be diftinguish'd into two Parts; hefe will mutually preis each other with the fame Force, wherewith they are preis'd by the Planes A C, ac, that is, in the Proportion of a b to A B; and confequently the Centrifugal Force, whereby these Pressures are fultain'd, are in the same Propertion. Becaule of the fame Number, and the fame Situa-tion of the Particles in both the Cubes, the conjunct Force with all the Particles exercife upon all, according to the Planes FGH, fgh, are as the Force which each exerciseth upon each. Therefore the Force which each exercifeth upon each, according to the Plane FGH in the greater Cube, is to the Force which each exercifeth upon each, according to the Plane f g h in the leffer Cube, as a b is to A B; that is, as we have already demonstrated, reciprocally as the Distances of the Particles from one another. Q. E. D. And, vice versa, if the Force of each Particle be reciptocally, as the Diffance of the Particles, i. e. reciprocally as the Sides of the Cubes A B, a b; the Sums of the Forces will be in the fame Proportion, and the Preffure of the Squares DB, db, as the Sums of the Forces; and the Preffure of the Square

Square DP, to the Preflure of the Square DB, abg; to ABq. And by Equality, the Preflure of the Square DP, to the Preflure of the Square a b, as a b cub to AB cub. for the Simple Proportion being compounded with the Duplicate, forms'a Triplicate one'; fo that the Force of the Comprefifon in the one, is to the Force of the Compreflion in the other, as the Denfity of the Fluid in the former, to the Denfity of the Fluid in the farter. Q. E. D.

Corol. (1.) Since therefore it is manifest by Experiments, that the Density of our Air, compress and rarified by Turns, is Proportional every where to the compressing Force, or the Compression it Telf; it seems that the Air con-sists of Particles which see from, or chase away one another in the inverse Proportion of the Distances. For altho' this Centrifugal Force may feem Diametrically opposite to the Universal centripetal Force or Gravity which we Ipake of, fo that it cannot could be the fame ; yet it may come to pals, that befides that general Law of Grayity which belongs to Matter as fuch, and without any Refeet had to the Figures, Forms, Circumitances or Motions of the fame; there may be other Laws, and natural Porces, whe-ther of attracting, or the contrary, belonging to the fpecial Figures, Forms, Circumstances or Motions of Particles of Matter, and in a peculiar Manner annexed to the fame, upon which many of the more difficult Phenomena of Nature may depend. Thus indeed it feems, that the Particles of Air, when they have acquir'd that peculiar Temperament, Figure or Form, from which they are fined to compole fuch an Elaffic Fluid as we call Air, are immediately fitbject to the new and special Law or Centrifugal Force belonging

belonging to fuch Particles, and fuch, only. For our most perspicacious Author doth justiy fuspect, that the greatest Part of the special Phenomena of Nature depend upon fuch a Force as hath been mentioned, and are owing to Caufes not yet known, whereby Particles of Matter are either driven upon one another, and fo cohere in regular Figures; or are driven away, and recede from one another ; which Force being unknown, it is no wonder that Philosophers hitherto have in vain attempted to explicate the Works of Nature; and which confequently being now by degrees difcover'd, or in the way to be fo, it is to be hoped that in Time, at least, we fhall come at length unto, the not the primary Caufes, yet the next to them, and fuch as will be as well accommodated to Geometrical Calculation, as Humane Ules, and only a concerning Scholium. But what hath been faid concerning

Sebolium. But what hath been laid concerning the Centrifugal Forces of the Air, and fuch like Fluids, is to be underflood of Juch Forces only, as are terminated in the next Particles, or diffus'd not much further: Examples of which we have in magnetic Bodies; the attractive Force of which is terminated almost in Bodies of their own Kind which are next to them. The Loadstone's Vertue is contracted by a Plate of Iron which is interpos'd, and almost terminated in the fame. For the more remote Bodies are not fo much drawn by the Magnet it, felf, as by the Plate. In like manner, when Particles chafe away other Particles of their Kind which are next to them, their Force in the mean while not reaching unto remoter Particles, of them fuch Fluids will be compos'd, as we have been treating of in this Proposition. It fills and the fame of the function primated as we have been treating of in this Proposition. It fills and the fame of the function of the properties of the form fuch Fluids will be composide the function of the function of the function of the properties of the form fuch Fluids will be composide the function of th

Corol. (2.) By the fame Reason it seems, that belides the general Force of Gravity, there are other attractive Forces peculiar to the Particles of fome Bodies, or to very small Distances, and other Circumstances of particular Bodies, from whence Phenomena, otherwife unaccountable, will naturally proceed. From fuch an Attraction as this, the Refraction or Inflection of the Rays of Light in Pellucid, or about the Angles of opake Bodies, which, to wit, do attract before the Contact, and the more forcibly at the leffer Diftance; as our Author hath observed in his excellent Optic Treatife. And from the like Cause, as he notes in his Latin Edition of the fame Work, the Spheric Figure of little Drops both of Quick-Silver, and the like Fluids feems to arife. For these Particles, as it seems when at a little Diftance from each other, attract ftrongly; and like as in the great Bodies the Planets, their Spherical Figure refults from an equal Gravitation of the Conftituant Parts one towards another; fo it is reafonable to derive the Spherical Figure of the little Drops which we were speaking of, from an equal Centripetal Force of the Particles that compose them, whilft they approach to one another; especially fince we fee that these Particles do fo quickly, and in a Moment, and so exactly cast themselves into the faid Figure; as is manifest in the known Phenomena of the Rainbow, which are wholly owing to an inftantaneous and most exact Conformation of the Particles into that Figure. 'And to the fame Caufe are fome other Phenomena of Fluids, which are otherwife most difficult to be folv'd, reasonably attributed, But this by the Way,

LX. The

LX. The Quantity of Matter in all Bodies, is exactly proportional to their Weight.

For when the Refiftance of the Air is taken away, as is done in Mr. Boyle's Vacuum; all Bodies, whether they feem most folid and heavy, or most rare and light, descend together with a common and given Velocity, as foon as they are let down from the fame Heighth. All Pendulous Bodies also whatever, where the Centers of Vibration are equidiftant from the Center of Sufpension, do even in the Air Descend and Afcend together, for a great Space of Time, if they begin at the fame Time to vibrate, in an equal Arch of the fame, or an equal Cycloid, or even an Arch unequal; and where an equal Arch is de-Icrib'd, they are mov'd altogether with the fame Celerity, whether they be hard or foft, folid or Jiguid; whether great or fmall, or of whatfoever Form and Figure. From whence it is manifest, that the moving Force is every-where in the fame Proportion with the Matter to be mov'd; or that the fame Force of Gravity doth equally affect all Bodies in the fame Diftance from the Center of the Earth. For that great Bodies do, cæteris paribus, descend something more swiftly, and keep their Motions fomething longer; this is from hence, namely, that the Surface of Bodies, according to which they are exposed to the Relistance of the Air, or any Medium whatever, is in Similar Bodies, in the duplicate Proportion only of the Diameters, or Homologous Sides; whereas the Solidity of the fame, according to which, both the Quantity of Matter, and the Force of Gravity is to be effimated, is in the tris plicate Proportion of the faid Sides : So that if the Diameter of a Ball of Stone be Three fold of that of another Ball of the fame Matter, the Surface

face of the fame, and confequently the Velocity being given, the Refiftance it meets with in the Air, will be only Nine-fold of that of the other : when, neverthelefs, its Solidity, and Qantity of Matter, and Gravity which is Proportional thereto, will be Twenty-feven-fold of the fame in the other. From whence it is no Wonder, that the Refiftance, which in Proportion of the Gravity, is fo much less in the greater Sphere, should affect and retard the fame Sphere in a leffer Proportion than it affects and retards the leffer. But that there is fo great a Difference of Velocity of Descent in the Air, betwixt Gold and Chaff, suppofe, this depends not only upon this Difference of Surfaces, but especially upon the Difference of Specific Gravity, whereby Gold doth far more exceed the Gravity of the Air, than Chaff doth : but the Excels of the Specific Gravity of a Body descending in the Air, above the Specific Gravity of the Air it felf; this alone is that Gravity, which forces a Body which is placed in the Air to defcend, as we shew'd above. And therefore it is not to be wonder'd, that Gold should fall in the Air much more fwiftly than Chaff, altho' in a Vacuum they are always observ'd to descend with equal Velocity.

Scholium. If the Velocity it felf of all Bodies in a Vacuum, upon the Surface of the Earth, be requir'd to be given in known Meafures; we are to know, that as well by the direct Observation of Bodies falling Perpendicularly, as by the Vibrations of Pendulums, and Computation made from thence by Hugens; it is with the Consent of Geometricians, determin'd to be of that Quanticy, that in one 2d of Time, Bodies descend 15 i Feet of Paris, or 16 I English Feet; or in the Space of one Hour, 208656000 English Feet, *i. e.* almost

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almost 40000 English Miles; as from the Calculation, according to which Bodies descend, in the Proportion of the Time duplicated, it will prefently appear.

LXI. In Pendulous Bodies, which are relifted only in the Proportion of the Velocity, when the Vibrations are in a Cycloid, whether the Arches defcrib'd be greater or leffer, are every where Isochronal.

The Truth of the Proposition, as spoken of in a Vacuum, where there is no Resistance of the Medium, hath been demonstrated above. And if the Resistance be as the Velocity, or as the Arch every where to be described; the rest of the Velocity shall likewise be in the same Proportion; and consequently the Time of Vibrating will be equally retarded on both Sides, and the Vibrations will still remain as before, *i. e.* of equal Time.  $\mathcal{Q}$ . E. D.

Corol., Therefore relifting Mediums, make the Time of each Vibration longer than it would be in a Vacuum. And Experience teffifies this of Pendulum Clocks; the Vibrations whereof, have been obferv'd to be perform'd fomething more quickly in a Vacuum, than in the Air. For the Refiftance takes off fomething from the Force of the moying Gravity, and confequently doth fomething refract or leffen its Effect, and the Velocity of the Motion.

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Nov. 25. 1706.

LECT.

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#### LECT. XXIX.

AXXII Odies mov'd with an unequal Velocity in a very Subtle Fluid, are refifted by the Fluid in the Duplicate Proportion of the Velocity.

For fince the Body which is mov'd the fwifter, doth both pass through a greater Part of the Medium, in Proportion of its Velocity, and meets every unequal Part of the Medium with greater Force, in Proportion of the same Velocity, the whole Resistance arising from both Causes conjoin'd, will neceffarily be in the Duplicate Proportion of the same Velocity. Which Proportion doth agree well with Experiments. Albeit the Defect, as to Slipperines of the Parts in the Air, which give Way, arising from Elasticity, and some Cohesion of the Parts of most Fluids, must needs something difturb that Proportion.

Corol. (2.) Since therefore, the Vibrations of Pendulums in a Cycloid, where the Refiftance is in the Simple Proportion of the Velocity, would be Ifochronal; but the Refiftance in the Air, and fuch like Mediums, is almost in the Duplicate Proportion of the Velocity; it is manifest that the Times of Vibrations in a Cycloid, and much more in a Circle, are, when the Vibration is in the Air, not altogether in divers Arches, but in the greater Arches, fomething greater, by Reafon of the too great Refiftance.

Conel.

Corol. (2.) Hencei it follows, that for the moft exactly obtaining the Equality of Times in Pendulum Clocks, it is requisite that the Pendulums should always describe the same Arches; otherwise. by Reason of the unequal Velocity, where the Arches described are greater, the Motion will be flower; where leffer, it will be fwister than it ought to be. From whence also the Cause may be shew'd, why greater Clocks, placed in a Ship, and toss'd up and down, do not so exactly shew the Hours, as those which are upon Land, and are at Reft. For by Reason of the Concussion, Arches are described fometimes greater, and sometimes leffer; and from thence fome Inequality must necessarily follow.

Corol. (2.) Shorter Vibrations, whether in a Cycloid, or in a Circle, are more Ifochronal than longer ; which is because of the Resistance of the diffurbing Medium; and the fhortest are perform'd in the fame Times nearly as in a Vacuum ; where also the Cycloid, and Circle, do just coincide, and the Vibrations in one, scarce differ from those in the other. From whence alfo, the Pendulum Clocks, which are govern'd by a long Pendulum, do fhew the Hours much more exactly, than those which have a shorter Pendulum; forasmuch as far lesser Arches are describ'd by those than these. But the Times of those Vibrations which are made in greater Arches, are something greater, because the Resistance whereby the Time is lengthened, is greater, in Proportion of the Length describ'd in the Descent, ( to wit, because of the greater Velocity ) than the Reliftance in the fublequent Afcent, whereby the Time is Contracted. But the Time of Vibrations, as well fhort as long, is alfo fomething lengthned by the Motion of the Medium.

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dium. For when Bodies become flower in their Motion, they are a little lefs refifted; and when they are accelerated, a little more, than when they are mov'd uniformly; becaufe whilft the Medium goes forward the fame way with the Bodies, by that Motion which it hath receiv'd from them, in the former Cafe it is the more agitated, in the latter lefs; and confequently doth more or lefs confpire with the Bodies moved in it. It refifts Pendulums therefore in their Defcent more, in their Afcent lefs; and from both Caufes the Time is lengthened.

LXIII. All Sounds, whether fmall or great, do go almost with one given Degree of Velocity; and this given Velocity is so fwist that it constantly goes about 1142 Feet in one Second of Time, which is Eight Miles in just 27", or about 68520 Feet; that is, near 12 Miles in an entire first Minute, or near 780 Miles in an Hour.

For all Sounds, whether great or fmall, muft go with the fame Velocity which a Stone, or other descending Body falling from half the Height of the Atmosphere, supposing it uniform, i. e. just fo high as by its Weight would reduce our Air on the Earth's Surface to its present Degree of Denfity; I fay all Sounds must go with the fame Velocity, which a Stone at last would acquire by falling from half that Height; or, which is the fame thing, that Sounds must go fo far as that entire Height comes to in the fame Time that the Stone would defcend from the one half of that Height; because the last Velocity acquir'd, if it had been uniformly continued all that while, would have gone twice as far as that Line which had been delcrib'd by an unequal Velocity, gradually increasing from Reft, as we have above from Galilæo demonstrated p.; and X there-

therefore in the fame time would have gone the whole Height of that uniform Atmosphere : Now this Stone falling from half that Height . descends to the Earth in about such a Space of Time as answers to the former Observations of the Velocity of Sounds. Now, that Sounds muft needs go with the fame Swiftness that a Stone would arrive at from half that Height beforementioned, is thus demonstrated. The descending Stone is urged downwards only by its own natural Gravity, or infinitely small Degree of Velocity uniformly impress'd upon it; and fo its Velocity in equal Time has an equal Increase, and becomes greater exactly in the Proportion of the Time of its Descent. The Atmosphere's Tenfion or Elafficity, which conveys the Sound with its own natural Degree of Velocity, or attempt to Motion, arifes also in this Cafe from the natural Gravity of each physical Part of the upper Surface uniformly augmented by the Addition of equal Parts of the inferior Surfaces quite down to the Bottom; fo that in both Cafes, the Velocity actually acquir'd in one is the fame with the Conatus ad motum, the Tenfion or beginning Velocity of the other : Just like two fmall or nascent Quantities originally equal, and which afterward are augmented in the fame Proportion, whofe last Quantities must therefore be equal alfo. This being fo, and the Tenfion of the Parts of the Air being almost the same, whether the Motion or Sound be great or fmall; 'tis plain, that tho' the Quantity of the Sound will be in Proportion to the Quickness of the Vibrations of the Sounding Body, and if that Velocity be increas'd by the Concurrence of the Wind, or diminish'd by its Opposition, the Sound will either be ftronger and reach farther, or be weaker and ftop fooner :

fooner; yet will the Velocity of the Sound itfel be always proportionable to the Tension of the Air which conveys it, and that Tenfion being nearly fix'd and certain', this Velocity of all Sounds must be nearly fix'd and certain. Now what Time is necessary for the Descent of a Stone from half the Altitude of fuch an uniform Atmosphere, as we have before suppos'd, will be thus computed. The specifick Weight of Water to that of Quickfilver, is known by many Trials to be as about 1 to 13  $\frac{3}{4}$ ; and when the Mercury is 20 Inches in Altitude, the specifick Gravity of Air to that of Water is about that of I to near 900, as has been found also by many Tryals. Nay indeed, confidering that most of the Elder and Foreign Experiments come nearer to that of I to a 1000; and that however the specifick Gravity of those Parts which are properly elastical Air, if they were freed from Vapors and other Bodies which are not elastical, and have nothing to do in the Conveyance of Sounds, would be then not the roooth Part; I shall chuse r to 1100 for the Proportion of Air properly speaking to Water: whence it will follow that true elaffical Air will be to Quickfilver as 1 to 15125, and 30 Inches of Quickfilver, which is a Balance for an equal Column of Air, will correspond to 29 times 15125 Inches of Air, or to 452750 Inches thereof; that is, to 27812 English Feet, or about 7 English Miles; which, if the Air were uniform in Density, would be its entire Altitude. But falling Bodies are known to defcend half that Altitude, or 18906 Feet in about 24 Seconds of Time. Whence Sounds ought to propagate themselves with such a Velocity, as will carry them 27812 Feet; the Altitude of the Air, if it were uniform in Density; in the Space of a little above X 2

above 24 Seconds, the Time of a Stone's Descent from half that Altitude; and by confequence will be propagated about 1142 Feet in one Second. about 68520 Feet, or near 12 Miles in one whole Minute, i. e. near 780 Miles in an Hour, agreeably to the best Observations. Sir Ilaac Newton, in his first Edition, calculated the Velocity of Sounds to be fomewhat lefs, by taking only I to 850 for the Proportion of Air to Water, which in this Cafe, as he now owns, ought rather to be taken as only 1 to about 1100: And it plainly appears, that the Observations of the Velocity of Sounds do generally make it greater than his first Numbers did allow. As to his Demonstration of the Conclusion we have here brought this Matter to, tho' it be extremely fubtil and ingenious. yet is it too long, too remote, and too intricate to be infifted on in this place; and therefore it was thought proper to make use of this more eafy and intelligible Method of Demonstration.

Coroll. (1.) If the Denfity of Air be increas'd or diminifh'd, the Sound it felf, or Violence of the Noife, will be increas'd or diminifh'd in the fame Proportion: Which thing doth well agree with the Experiments of Sounds made in rarified or condens'd Air.

Corol. (2.) If the Wind confpire with the Motion of the Air, the Sound or Noife will be increas'd and carried farther; as being now made up of the Sum of the Motions of the Sound it felf, and the Wind. If the Wind be contrary, the Sound will be diminifh'd, and fooner flopp'd, as now confifting of the Difference of the faid Motions only. Which neverthelefs is fo to be underftood, that the Veloci-

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ty of the Sound it felf, which was defign'd above, alter but very little For Sound depends not on the continual Motion of the Air, but of the Pulfations of the fame propagated after the manner of Waves by Vibrations, and a continual Viciffitude of Goings and Returnings, as will be fhewn afterwards. And of what Sort foever the Difference of the Noife is, which arifeth from the different State of the Sonorous Body, or of the Wind; yet the Denfity of the Air, and its Elafticity, do remain almost the fame; and fo the Effects of them, or the Velocity of the propagated Sounds, will remain likewife almost equal.

Corol. (3.) Sounds therefore, of what Kind foever, whether they be great or finall, are propagated through Air of a given Denfity and Elasticity almost with the fame Velocity; as the Experiments alfo, which have been made by Philosophers, do shew.

Corol. (4.) The Velocity of Sounds therefore in any Place whatever being given, or that whereby they go about 1142 English Feet in one Second; from the Interval of Time of Sounds given, there is given withal the Interval of Distance of the Sonorous Body. Thus, for Instance, if we number 10" of Time betwixt the Fire of a Cannon seen, and the Sound heard; it is manifest, that the Gun is 11420 Feet distant, or somewhat above two Miles. As likewise, if 5" pass betwixt our seeing the Flash of Lightning, and hearing the Thunder, we may reckon that the Thunder-Cloud is about 5710 Feet, or a little above one Mile distant from us.

Decemb. 2. 1706.

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#### LECT. XXX.

S the Refiftance of Fluids in divers Velocities is in the duplicate Proportion of the Velocity; fo in divers Denfities the Velocity being given, it is in

the direct Proportion of the Denfity it felf; but the Denfity and Velocity being given in the duplicate Proportion of the Diameters; and confequently the Refiftance in general is in a Proportion compounded of the duplicate Proportion of the Velocity, and the fame Proportion of Diameters, and the fimple Proportion of the Denfity of the Medium directly.

These Things are easy, and stand in no Need of Demonstration. For if two Spheres do exceed one the other, as to their Diameters, in the double Proportion, or be as 2 to I, and the greater be mov'd with a Velocity double to that of the other, and in a fluid Medium double to the other in Denfity; it is plain, that in any given Space of Time the whole Refiftance of the greater Sphere, or Motion loft, is to the whole Reliftance the lesser Sphere meets with, or its Motion loft, as 2 x 2 x 2 x 2 x 2 to 1 x 1 x 1 X 1 X 1, or as 22 to 1; and thus every where. But it is to be noted, that Refistance proceeds equally both from Fluids and Solids, cæteris paribus; unless fo far as in a very fluid Medium, when the Motion is fomewhat flow, the Medium it felf by a Circulation of Motion, and an Impetus thereby made on

on the hinder Part of the Body mov'd in it, doth fomething promote the Motion of the Body; which reciprocal Impetus of the Medium on the Body ought to be lefs in fwifter Motions of the Body, and in very fwift ones none at all; as our Famous Author found the thing to happen in very accurate Experiments, which he made about it.

Corol. (1.) The Mediums therefore in which Projectiles are carried the fartheft without any fensible Diminution of Motion, are not only very fluid, but much rarer than the Bodies moved in them; otherwise they would presently flop the Motion of the Projectiles, and bring it to reft.

Corol. (2.) From whence it follows, that our Air, or all the Matter contain'd in it, is very fmall, if it be compar'd with the Matter in Bodies, that are carried forward very far and fwiftly in it; and is fo far from the Cartefian Plenum, that it doth not indeed posses the 20000th Part of the containing Space.

Corol. (2.) And it follows alfo, that the Ether, or all the Matter contain'd in the Planetary Spaces, thro' which the Planes have revolv'd for fo many thousands of Years with fuch Velocity, and this without almost any Loss of Motion at all, is very small, if compared with the Matter contain'd in the Planets themselves; fo that, as will easily appear by Calculation, that Space ought rather to be counted a Vacuum than otherwife.

Corol. (4.) The whole Cartefian Philosophy therefore falls to the ground, which is entirely built upon a Plenum, and a Celestal Matter, X 4 which

which he calls His 1st and 2d Elements. Nor can that ingenious Fiction any longer fublift, when its Bafis is thus deftroy'd by our Author's Experiments, and what he hath demonstrated: Especially when He has not only taken away that Plenitude of Matter, but fhew'd alfo that there is nothing of the forefaid Matter in the Pores of Bodies. For by the Experiment of a very long Pendulum vibrating in the Air a long while, and by estimating the Loss of Motion, when compar'd with the Reliftance of the Air made upon its Surface, he found that either there was none at all, or a plainly infenfible Refiftance in the internal Parts. From whence it is rightly concluded, that there is either none at all, or a plainly infenfible Quantity of any fubtle Matter in the Pores of Bodies; whereas, from the Cartehan Plenitude, compared with the specifical Gravity of the Pendulum, it ought to be far greater than the grofs Substance it felf of the Pendulum.

LXV. No Rectilinear Preffure can be propagated through a Fluid, in right Lines only.

For fince the Parts of a Fluid are always in Motion every way, or are at leaft every way eafily moveable, and will upon any Occafion be actually mov'd; it cannot be, but that any Preffure whatever, which is first communicated in a right Line, must urge the contiguous oblique Parts more or lefs; and that these oblique Parts must urge others in like manner that are placed obliquely; and thus in infinitum. The Preffure therefore, as foon as it is propagated to Particles which do not lie in the right Line, will begin to divaricate, and be propagated obliquely for ever; and when fome Part of the Preffure is intercepted by fome Obstacle,

Obstacle, the remaining Part now, as well as before, will divaricate into all the Spaces beyond the Obstacle.

Corol. (1.) Hence the Reafon appears, why Sounds let into a Chamber, either by the interpos'd Walls, or through the Windows, fpread themfelves into all Parts of the Chamber, and are heard at all Angles, not only as reflected from the opposite Walls, but as propagated thro' the Air on every Side from the Window.

Corol. (2.) The Rays of Light which are propagated through the Ether, or Air, or any other Fluids whatever always in right Lines, are not Impulses or Modifications of that Fluid, as it is in Sounds, but real Corpuscles flowing from the Fountain of Light, and propagated by a true Motion through the Medium; as most of the other Phanomena of Light do also show.

LXVI. Every tremulous Body in an Elaftic Medium will propagate the Motion of Pulles on every Side forwards; but in a Medium not Elaftic, it will excite a circular Motion.

Cafe (1.) For the Parts of the tremulous Body, in their alternate going and returning, will in their going drive forwards, and confequently prefs and condenfe the Parts of the Medium next thereto; and in their returning will permit the faid Parts of the Medium to expand themfelves, and return to their former Situation. Which certain Parts of the Medium going and returning alternately, as doth the tremulous Body it felf, will act in the fame manner upon the Parts of the Medium next to them, as the tremulous Body did

did upon them, and will propagate the fame tremulous Motion to those further Parts of the Medium, and these last will propagate it to others more remote than themfelves; and thus in infinitum. And in every one of the defign'd Divisions of the Medium, the Parts will be alternately condens'd and relax'd ; in their Going condens'd, and in their Return relax'd, like as it is in the treinulous Body that began the Motion. Not that they all go and return at the fame time, but alternately; for the Expansion of the foregoing Division makes the Condensation of the 2d, and is at the fame time with it, as the Expansion of the ad is at the fame time that the Condenfation of the 2d is. But the Parts which go, and in going are condens'd, becaufe of their progressive Motion wherewith they strike Obstacles, are Pulses; and therefore successive Pulses will be propagated from every tremplous Body through an Elaftic Medium; and this at Diftances from each other nearly equal, becaufe of the equal Intervals of Time, wherein the Body doth by each Trémor • excite each Pulfe. D. E. D.

Corol. (1.) Altho' the Parts of a tremulous Body do go and return according to fome cerrain and determinate Direction, or Part; yet the Pulfes propagated from thence through the fluid Medium will, by the foregoing Proposition, fpread themfelves every way on the Sides; and will be propagated every way from the tremulous Body as the Center, according to Surfaces almost fpheric and concentric. Of which we have an Example in Waves; which if they be rais'd by a tremulous Finger, will not only go forward, according to the Direction of the Motion of the Finger, but will prefently be propagated on all Sides,

#### Mathematical Philosophy. 315 Sides, and encompass the Finger in the Form of concentric Circles; for the Gravity of the Water supplies in a fort the Place of Elasticity.

Corol. (2.) Hence we may collect, that the Number of the propagated Pulfes is the fame with the Number of the Vibrations of the tremulous Body, and is not multiplied in the Progrefs. For every phyfical little Line, as foon as by the Expansion it hath return'd to its firft Place, would reft there, were it not urged with a new Motion by the Force of the tremulous Body it felf, or the Pulfes propagated from it. And therefore it will actually reft, when Pulfes cease to be propagated from that Body.

Corol. (3.) Hence the Reafon appears why Sounds, when the Motion of the fonorous Body ceafeth, do prefently ceafe; and are hear'd at a great Diftance no longer than at a leffer: For the Caufe ceafing, the Effect muft needs ceafe alfo.

Corol. (4.) Hence we may understand the Caufe of the Increase of Sounds, in the Stenterophonick Tubes. For a reciprocal Motion is wont, in each Recourse, to be increas'd by the Cause that produces it: For the Motion in the Tube, which hinders the Dilatation of the Sound, is reverberated more ftrongly; and therefore is the more increas'd from the new Motion impress'd in each Reflexion. And fince all that Force of the fonorous Body, or Voice, which otherwife must in the fame time have been propagated through an entire Sphere, which hath the Length of the Tube for its Radius, is now that up within the Hollow of the Tube, and goes out of the Aperture with a great Strength; it is evident, that the tremu-

tremulous Motion of the Air, or the Violence of its Pulles, is greatly increas'd from thence. and confequently ought to reach unto a much greater Diftance; but this fo notwithstanding, that the Velocity of the Propagation doth every where remain still the fame and unvaried. The Sound therefore, as I suppose, is increas'd in these Tubes in the Proportion of the whole spheric Surface aforefaid, to that Part of it which is contain'd within the Aperture of the Tube. But it would be worth the while that Experiments fhould be made about this Matter, to determine whether the Increase of Sounds in these Tubes be in that Proportion which hath been defin'd; that we may hereafter pronounce with more Certainty, and may be able to accommodate these Tubes more to the Use of Mankind. 1.1

Cale (2.) But if the Medium be not Elastic. because the Parts thereof which are pressed by the vibrating Parts of the tremulous Body cannot be condens'd, the Motion will be propagated in an Inftant to Parts where the Medium doth more eafily give way; that is, to Parts which the Body would otherwife leave empty behind it. The Cafe is the fame here, as with Projectiles in general in any Medium whatever. The Medium. in giving way, doth not go back in infinitum, but by a Circulation comes at length to the Spaces which the Body leaves behind it. Thus it is that the Medium gives way to a tremulous Body alfo. by a circular Retroceffion; and as often as the Body returns to its former Place, the Medium is repell'd from thence, and returns to its former Place.

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Corol.

Corollary. The Carteflans therefore are miltaken, who suppose that the Agitation of the Parts of the Sun, or any Flame, fuffices to a Preffure, which is to be propagated through the Ambient Medium in right Lines, fo as to conftitute the Rays of Light. For fuch a Preffure ought to be, not from the Agitation only of the Parts of the Flame, but from the Dilatation of the whole.

Decem. 9. 1706.

### LECT. XXXI.

THE a folid Cylinder, infinitely long, be revolv'd in an uni-I form and infinite Fluid about its own Axis, the Polition whereof is given, and the Fluid be mov'd round by the Impulse of this Cy-linder only; and every Part of the Fluid perseveres uniformly in its Motion; the periodic

Times of the Fluid will be as their Diftances from the Axis of the Cylinder directly; and the Velocities will be every where equal.

For let the Fluid be diffinguish'd into innumerable Cylindrical Orbs concentric to the Cylinder, and of the fame Thickness every where. And because the Fluid is supposed to be homogeneous,

neous, and the Cylinder, by its circular Motion. endeavours to put all the contiguous Parts of the Fluid, and through them the further Parts in infinitum, into its own angular Motion, and confequently into a Velocity of Motion that is in direct Proportion of the Distance, fo that each of chem shouldbe turn'd about in the fame periodic Time with it felf; it is plain, that every Orb doth then ceafe from further Acceleration, and that the Parts of them perfevere uniformly in their Motions, where the Reliftance or Impression on the Concave Part, is equal to the Refiftance or Impreffion on the Convex Part : (For otherwife the fironger Force prevailing, the Motion will be changed on that Part. ) Therefore. where the respective Velocity, according to which Refiftance will arife in the given Surface, shall be in the reciprocal Proportion of the Surface, the Impressions on both Parts will be equal; that is, in the prefent Cafe, where the angular Velocity is in the reciprocal Proportion of the Diftance it felf, or where the abfolute Velocity is always equal, the periodic Times alfor will be in the direct Proportion of the Diffance. Ø. E. D.

Covol. (r.) If the Fluid be not infinite, but contain'd in a Cylindrical Veffel; the exterior Cylinder alfo will be turn'd round, and its Motion will be accelerated until the periodic Times of both Cylinders, and of the inclos'd Fluid, be equal one to another. But if the exterior Cylinder be violently detain'd, it will endeavour to retard the Motion of the Fluid; and unlefs the interior Cylinders preferve its Motion by fome Force continually imprefs'd, it will make the fame to ceafe by Degrees.

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Corol.

Corol. (2.) But fince the periodic Times of the Planets are not in the Proportion it felf of their Diffances from the Sun, but in a Proportion which is fefqui-alteral of the fame; and confequently their abfolute Velocities are not every where equal, but in a fubduplicate Proportion of the Diffances, as all Aftronomers acknowledge; it appears, that the Confitution of fuch an Ethereal Fluid doth in no wife agree to the Solar Syftem; nor doth the Supposition of it any ways help the Cartefian Vortices.

LXVIII. If a folid Sphere, in an uniform and infinite Fluid, be revolv'd uniformly about its own Axis, the Position whereof is given; and by the Impulse of this alone the Fluid be turned round, and every Part of the Fluid perfeveres uniformly in its Motion; the periodic Times of the Parts of the Fluid will be as the Squares of the Distances from the Center of the Sphere.

Let the Fluid be diftinguish'd into innumerable Concentric Spherical Orbs of the fame Thicknefs: And, as before, the Fluid will then only perfevere in its uniform Motion without any Acceleration or Retardation where the angular Motions of the Parts of the Fluid about the Axis of the Globe be reciprocally as the Spheric Concentric Surfaces themfelves, or as the Squares of the Diftances from the Center of the Globe reciprocally; or laftly, as the periodic Times of the Parts which are reciprocally proportional to the angular Velocities themfelves; where these be as the Squares of the Diftances from the Center of the Globe directly.

Corol. (1.) If the Fluid be not infinite, but contain'd in a fpheric Veffel; the fpheric Veffel also will be turned round, and its Motion will be accelerated until the periodic Times of the Sphere, and

and Veffel, and inclos'd Fluid, be equal to one another. But if the fpheric Veffel be violently detain'd, it will endeavour to retard the Motion of the Fluid; and unlefs the Sphere preferve its Motion by fome Force continually imprefs'd, will make that the fame, as, in the former Cafe, fhould by Degrees ceafe.

Corol. (2.) But fince the periodic Times of the Planets are not in the duplicate Proportion of their Diftances from the San, as we have feen already; it is manifeft, that the Conftitution of fuch an Ethereal Fluid doth in no wife agree to the Solar Syftem; nor are the Cartefian Vortices in any wife help'd from the Supposition of the fame.

Corol. (2.) Since the Bodies, which being carried in a Vortex, go perpetually the fame round without confiderable Accefs to the Center, or Recefs from it; (as it is in all Planets, both Primary and Secondary;) they muft needs be of the fame Denfity with the Vortex, and be carried along together with the contiguous Parts: And fince this Sort of Vortices muft be fo mov'd, that the periodic Times fhould be in the duplicate Proportion of the Diftances (contrary to what happens in all the Planets;) it is manifeft, that the Planets are not carried along in Corporeal Vortices. Which alfo will be made more manifeft from the following Proposition.

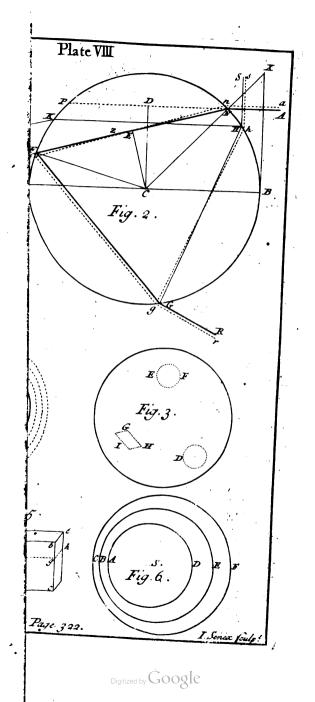
LXIX. The Velocities of all the Planets, whether Primary or Secondary, about their Central Bodies, by being in the reciprocal fubduplicate Proportion of the Diffances from their Centers, do wholly overthrow the Cartefian Hypothesis of Vortices.

For the Planets, as is now known every where, do revolve each of them about the Central Body in

in Ellipses; and this in fuch fort, that by Rays drawn to their Foci they describe Areas proportional to the Times; and that the Velocities should be in the reciprocal subduplicate Proportion of the Diftances. But the Parts of an Ethereal Vortex cannot be revolv'd with fuch a Motion. For (in Fig. 6. Plate 8.) let A D, B E, CF be three Primary Orbs describ'd about the Sun S. Of which let the outmost C F be Concentric to the Sun; and let the Aphelia of the two inner be A and B, and their Perihelia D and E. Therefore the Body which is revolv'd in the Orb CF will, by a Ray drawn to the Sun in describing Areas, which are proportional to the Times, be moved with an uniform Motion : But the Body which is revolv'd in BE will, according to the Laws of Aftronomy depending both upon Geometrical Demonstrations and Celestial Obfervations, be mov'd more flowly in the Aphelion B, and more fwiftly in the Perihelion C; when yet, according to Mechanic Laws, the Matter of the Vortex must to be mov'd more swiftly in the narrower Space which is betwixt A and C, than in the wider Space which is betwixt D and F; i. e. more fwiftly in the Aphelion than in the Perihelion. As for Example: In the Beginning of the Sign Virgo, where Mars's Aphelion now is, the Distance betwixt the Orbs of Mars and Venus is, to their Distance in the Beginning of Pisces, almost in the sesqui-alteral Proportion, or as 2 to 2: And therefore the Matter of the Vortex betwixt those Orbs in the Beginning of Pifces, ought to be carried more fwiftly than in the Beginning of Virgo, in the fame fefqui-alteral Proportion. For by how much the narrower or firaiter the Space is through which the fame Quantity of Matter paffeth in the Time of one Revolution, with fo much Y

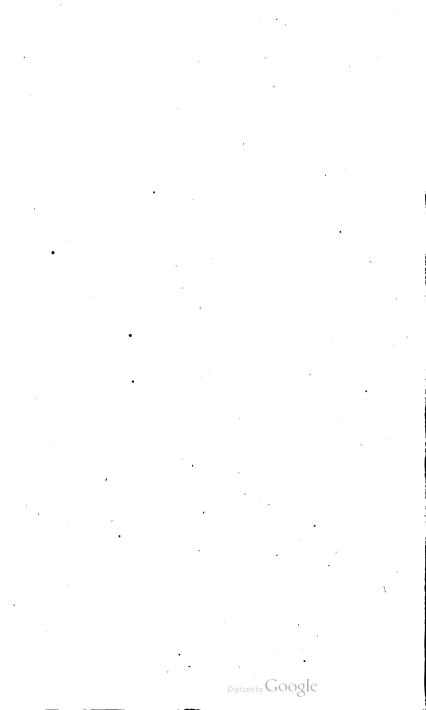
much the greater Velocity it must pass thro' it. Therefore if the Earth refting relatively in this heavenly Matter be carried along by the fame, and revolv'd about the Sun together with it; the Velolocity thereof, in the Beginning of Pisces, ought to be to the Velocity of the fame, in the Beginning of Virgo, in the fefqui-alteral Proportion. or as 3 to 2. From whence the apparent Motion of the Sun, in one Day's Time in the Beginning of Virgo, ought to be greater than 70', and in the Beginning of Pifces lefs than 48'; when yet (by the Testimony of Experience) that apparent Motion of the Sun is fwifter in the Beginning of Pifces, than in the Beginning of Virgo; and therefore the Vortex is mov'd more fwiftly in the Beginning of Virgo, than in the Beginning of Pisces. The Hypothesis therefore of Vortices doth wholly contradict the Aftronomical Phanomena; and Terves not fo much to explicate, as to difturb the Celeftial Motions.

Scholium. Hitherto we have delivered the Principles of Natural Philosophy out of our Famous Author; yet, speaking properly, we have delivered them not Philosophically or Physically, but rather Mathematically. Forafmuch as we have hitherto confidered the general Laws and Conditions of Motions and Forces, which chiefly belong to Aftronomy and Natural Philosophy, moftly in a Mathematical and Universal Method: Nevertheless, that our Work fhould not feem altogether dry and barren, we have every where illustrated it with Scholia, and Corollaries Aftronomical, Physical, Optical, and also Mechanical; and fo have prepared the Way to true Philosophy and Astronomy, that is, the Newtonian. It remains, that we come now at length to the Nature of Things, and to the Philo-



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Mathematical Philosophy. 323 Philosophical Causes of these Phanomena, both Aftronomical and Physical, and to the true System of the World; and that we fet before you the Frame and Constitution of the same System, fo far as it depends upon the Principles already laid down; omitting here, or only lightly touching upon those Things, which we had observ'd in the foregoing Scholia or Corollaries.

Jan. 29. 170%.

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#### LECT. XXXII.



LXX. HE Six Primary Planets, each with its own Satellites, if they have any, encompais the Sun with their Orbs, and revolve about it.

That Mercury and Venus revolve about the Sun. is manifeftly demonstrated from their Faces exactly imitating those of the Moon ; as is every where now known by Telescopic Observations. For fometimes they fhine with a full Face about the Conjunctions, but with the leaft apparent Diameters; they being then fituate beyond the Sun, and imitating a Full Moon: and then with an obscure Face about the other Conjunctions, but with the greatest apparent Diameters; they being then fituate on this Side the Sun, and imitating a New Moon. And they appear likewife of a gibbous or hollow Face about the Octants, and of an halved and dichotomous one about the Quadratures, like as Y 2 ic

it is is in the Moon : Sometimes' they pais thro' the Difcus of the Sun, and appear as Spots therein, inducing a partial Eclipfe; and fometimes they pass beyond the Body of the Sun, being in the mean while invisible to us. From whence it is certain, that these two Planets are revolv'd about the Sun, and not about the Earth. And altho' Mercury is fo rarely feen, as appearing to us only about its greateft Elongations, and when it passeth over the Sun, that all the faid Faces cannot be actually observ'd fo clearly in this Planet, as in Venus; yet notwithstanding, fince what Faces of Mercury can be feen, do exactly answer to this Position; and since those of Venus, a Planet of the fame Condition, lie open to our Observation, and do every where fully answer. the faid Polition ; there is no room to doubt concerning the reft as to Mercury. From the full Face of Mars alfo, near the Conjunction with the Sun, and the gibbous Face thereof in the Ouadratures, it is manifest, that it revolves about the The fame thing is also demonstrated con-Sun. cerning Jupiter and Saturn, from their Faces which are always full, as it ought to happen at fo great a Diftance. For albeit these Planets ought to have their Faces about the Quadratures fomething diminish'd; yet fince that Diminution of Light is fo very fmall that it can fcarce, or rather not at all be observ'd and seen by us, their full Face is to be reckoned to agree very well with the faid Polition. But that the Orbit of the Earth encompaffeth the Sun, is abundantly manifeft from the annual Parallax, which we have elfewhere explained.

Corollary. From hence with De Cartes, and the reft also of the Astronomers of the foregoing Age, we gather that the Ptolemaic System of the World, which

which alone was cultivated and celebrated for fo many Ages foregoing, comes to nothing. And we gather alfo, that the Tychonic Syftem, which was afterwards receiv'd and celebrated by fo many and great Aftronomers, doth wholly fall to the Ground; and doth not in any wife agree with the Phænomena, which have been observ'd of And laftly we gather, that the Copernilate. can Syftem, which hath for fo long a time been approv'd of and follow'd by most of the best Aftronomers, is the true System of the World, and is that alone which doth prefent to us that Order of all the Planets, which agrees to the Nature of Things, and to Aftronomical Obfervations. Therefore it may justly feem strange, that the Famous Dr. Gregory, that Excellent Interpreter of the Newtonian and Copernican Astronomy, a Man fo well skill'd in the true Mundane Syftem, should bestow fo much Time and Pains in delivering and fetting off those and other false and imaginary Hypotheses. When it is fo certain, that the Pythagorean or Copernican Order of the Planets is that alone which is True and Genuine; and that the reft of the Hypotheles are only fictitious, To what Purpole Mould we mix the Truth with mere Shadows, and difturb the Contemplation of the Nature of Things with manifest Falsities? Let therefore those, once indeed most Noble, most Famous Systems, be now banished for ever out of the Aftronomical World; and that only be admitted, cultivated, delivered and taught, which now, at length, we find to be the only one that corresponds to the true Order of Nature, and to real natural Caufes. But this by the way.

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#### LXXI. The

LXXI. The periodic Times of the fix Primary Planets, are in the fefqui-alteral Proportion of their mean Diffances from the Sun.

This Proportion, which was first found out by *Kepler*, the Parent of the *Newtonian* Philosophy, is now acknowledg'd by all. The Measure of the periodic Times is agreed upon amongst all Astronomers; but as for the Magnitudes of the Orbs, the fame *Kepler* and *Bullialdus* have exceeded all others in the Diligence they have us'd for determining the fame: And the mean Distances which answer to the periodic Times, do not fensibly differ from the Distances which they have found, and are for the most part in the Middle betwixt them; as may be seen in the following Table.

#### The mean Distances of the Planets from the Sun.

Saturn. Jupit. Mars. Earth. Venus. Mercu. 951000.519650.152350.100000.72398.38585. 954198.522520.152350.100000.72398.38585. 953806.520116.152399.100000.72333.38710.

And now we will give the true Periods, as alfo the Diftances which come neareft to the Truth, from Mr. Flamfteed's Parallax of the Sun, viz. of 10".

Mer-

		D.	H.	<i>'</i> .
Mercury	) و س )	87	23	16
Venus	-E. 0	224	16	49
Earth & Moon		365	6	9
Mars	Le s	<b>68</b> 6	23	27
Jupiter	Sun in Volume	4332	12	20
Saturn		10759	6	36
Mercury	] <sup>2</sup> ]	32,000	,000	) %
Venu's •	E	59,000	,000	Miles.
Eartb		81,000	,000	ςΣ
Mars	ៅ ដី ៅ	123,000	,000	2
Jupiter		4 <b>2</b> 4,000	,000	1 20
Saturn		777,000	,000	١ <u>ٿ</u>

Now, as to the Methods of finding these Diftances, they are thus determin'd.

Of the Diftances of Venus and Mercury, as compar'd with that of the Earth, there is no room to doubt; fince thefe are gathered by plain Trigonometry from their greateft Elongations, known by eafy Obfervation. As for the Superiors, all manner of Dispute concerning their Diftances from the Sun, which are deduc'd from the Arch of Retrogradation, is taken away by the Eclipses of the Satellites of Jupiter reduc'd to an accurate Calculation, according to the other Distances, and which agree with the Observation, for by those Eclipses, the Position of the Shadow which Jupiter casts, is determin'd; and by this means Jupiter's Heliocentric Longitude is had, whilft his Geocentrical is had immediate. ly by Observation. Therefore in the Plane Triangle connecting the Centers of the Sun, Jupiter, and the Earth, all the Angles are given, and con-Y 4 fequently

fequently the Proportion of the Sides is also given; or the Proportion of the Diftances of *Jupiter* and the *Earth* from the Sun.

Corollary. Therefore the Proportion of the Diftances from the Sun, is given exactly in all the Planets; fo that if the Diftance of any one of the Planets was given in fome known Measure, as in Miles or Semi diameters of the Earth, we should withal have the true or absolute Diftances of all: But this is what is yet wanting.

LXXII. The fix Primary Planets do always, by Rays drawn to the Sun, defcribe equal Areas in equal Times, and in general Areas proportional to the Times. This Equality of the Areas in equal Times, which is another Foundation of the Newtonian Philosophy, is owing likewise to the Observation of the fame Kepler. Whilst the five other Planets are, in respect of our Earth, sometimes Progreffive, fometimes Stationary, and then Retrograde; they do always go forward, in respect of the Sun, and that with an uniform Motion nearly, fuch that it is fomething fwifter in the Perihelia, and flower in the Aphelia, to preferve the forefaid Proportionality of Areas. This Proposition, which is well known to Aftronomers, is demonstrated as Jupiter in a peculiar manner; viz. by the Calculation of the Eclipfes of its Satellites, which is built upon this Hypothesis, and is exactly agreeable to the Observation. For by these Eclipses, as we have said already, Jupiter's Longitude and Distance from the Sun are exactly determin'd.

LXXIII. The Moon, by Rays drawn to the Center of the Earth, defcribes in equal Times Areas almost equal; and in general, Areas almost proportional to the Times.

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This appears from the apparent Motion of this Planet, as compar'd with its apparent Diameter, which is in the general nearly reciprocally proportional to the Diffance. I faid in the Propofition, *almost* proportional; becaufe the exact Proportionality is fomething diffurb'd by the Sun's Force, as we have explained that Matter elfewhere: But taking away that Diffurbance, the Proposition would be as exact and full, as in the Primary Planets; and that for the fame Reafon.

LXXIV. The Satellites of Jupiter do, by Rays drawn to the Center of Jupiter, describe Areas nearly proportional to the Times: And their periodic Times are in the sefui-alteral Proportion of their Distances from their Centers.

Both Parts of the Propolition are manifeft from Aftronomical Observations. For their Orbs do not differ sensibly from Circles Concentric to Jupiter, and their Motions in these Circles are found to be almost uniform; And as for the Proportion of the periodic Times here meant, it is what all Aftronomers agree in. Besides, Mr. Flamfreed, who hath stated all Things most accurately by the Micrometer, and the Eclipses of these Satellites, hath, both by Letters written to Sir Isaac Newton, and by his Numbers themselves communicated to him, signified that that seful alteral Proportion doth hold here as exactly as possible, fo far as he can discover by Observation. Which will be manifest from the following Tables.

The Periodic Times. H. D. 27<sup>1</sup>/<sub>2</sub> 13<sup>3</sup>/<sub>3</sub> 42<sup>5</sup>/<sub>5</sub> 18 1 3 12 234 3 7 16 16 32 4



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		I	2	3	4	iter.
According to	Caffini, Borelli, Townl.bymicr. Flamf.bymicr. Flam.byEcl.Satel. Period.Times.	5231	8 8[78 8[85 8[876 8[878	14 159	241902	Semi-diam. of Jupi

The Distances from the Center of Jupiter.

LXXV. The Satellites of Saturn do, by Rays drawn to the Center of Saturn, dcfcribe Areas proportional to the Times: And their periodic Times are in the fesqui-alteral Proportion of their Diftances from the Center of their Primary.

Both Parts likewife of this Proposition are prov'd from Astronomical Observations: For their Orbs scarce differ sensibly from Circles concentric to Saturn, and their Motions are sound to be almost uniform in these Circles. And as concerning the Proportion of the periodic Times, this will appear to every one that will take the Pains to compute it from the following Table, which we here present the Reader out of Mr. Hugens's Cosmotheoros, Page 101, 102.

The Period.	Times.	Distances from the Center of To, both by Observ. and Period.
	_	

	D.	H.	•.	".	
Ţ	I	21	19		
2	2	17	41		2 14 3.2
3	4	I 3	47		3 14 2 0
4	15	22	4 <b>!</b>		4 4 655
S	1 79	22	4		5 12 75
	Novem	<b>b. 17</b> , 1	797.	-	T

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### LECT. XXXIII.

LXXVI. T H E Force whereby the Primary Planets are perpetually drawn back from right Lines, and retain'd in their Orbs, does refpect the Sun; and is

as the Squares of the Diftances from the Center of the Sun reciprocally.

For on account of the forefaid Proportionality of Areas, this Force must tend to the Sun; and becaufe the periodic Times are in the fefqui-alteral Proportion of the Diftances, the Quantity of the Force must be every where in the reciprocal duplicate Proportion of the Diftances; as we demonstrated before : But this 2d Part is also demonstrated most fully from the Figure of the Orbs. For, if the Planets were mov'd about the Sun in spiral Lines, cutting the Rays in a given Angle, the Centripetal Force would be in the triplicate Proportion of the Diftances, or as the Cubes of the Diftances, reciprocally. But if they were mov'd about the Sun in Ellipses, which have the Center of the Sun in their Center, the faid Force would be in the direct Proportion it felf of the Diftances: But when the Ellipses, in which they are mov'd, have the Center of the Sun not in their Center, but in one of their Foci, as the Cafe really is, and all Aftronomers do acknowledge; then the faid Force must needs be in the duplicate Proportion of the Diftances recipro-Which cally.

This is also demonstrated by the Quiescence of the Aphelia. For where the faid duplicate Proportion doth hold exactly, there the Aphelia must rest; when the faid Proportion approacheth to the simple direct Proportion, then the Aphelia must go back; but when it inclines to the triplicate Proportion, they must go forwards.

LXXVII. The Force wherewith the Satellites of Jupiter and Saturn are perpetually drawn back from right Lines, and retain'd in their Orbs, respect the Centers of Jupiter and Saturn respectively; and are as the Squares of the Diftances from those Centers reciprocally.

For on account of the aforefaid Proportionality the Areas about the Centers of Jupiter and Saturn, the faid Force must tend to those Centers; and because of the sesqui-alteral Proportion which the periodic Times have to the Diftances, the Quantity of that Force must be every where in the reciprocal duplicate Proportion of the Diftances, by what was in the foregoing Propolition mention'd to have been demonstrated by us before. But we can fetch no Argument to prove this latter Part of our prefent Proposition from the Figure of the Orbs; for that those Orbs, of which we speak at present, are Circles, or Ellipse pot senfibly different therefrom : Nor confequently from the Quiescence of any Aphelia; for in Circles where there can be no Line of the Apfides, there are no Aphelia.

LXXIX. The Force wherewith the Moon is perpetually drawn back from a Rectilinear Motion, and retain'd in its Orb, refpects the Center of the Earth; and is as the Squares of the feveral Diftances from the fame Center reciprocally.

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For on account of the Equality of Areas about theCenter of the Earth in equal Times, excepting fo far as the fame is difturb'd by the Force of the Sun; the faid Force must tend unto the Earth : And because of the Elliptic Figure of the Orbit, which hath the Center of the Earth in one of the Foci, the Quantity of the Force must be every where in the reciprocal duplicate Proportion of the Diftances. For altho' the Figure of the Lunar Orbit be not exactly Elliptic, and confequently the Center of the Earth is not placed exactly in one of the Foci of the fame Orbit; yet notwithftanding, fince all this Variety doth arife from the difturbing Force of the Sun only, the Figure is to be underftood to be in it felf, or primarily an exact Ellipsi, and to have the Earth placed in one of its Foci; and confequently to have the Centripetal Force in the duplicate Proportion of the Diffances reciprocally : Yea, whilft the thing is as it is, the very flow Motion of the Moon's Apogeum fhews, that that Force is in the faid duplicate Proportion very nearly, if not exactly. For by our Author's Calculation, it appears from the flow Progress of the Apogeum, that the Centring Force of the Moon towards the Earth, comes above fixty times nearer to the duplicate than to the triplicate Proportion. Which small Difference arising, as was faid, from the Action of the Sun, is to be neglected. It remains therefore, that this 2d Part of our Proposition holds good, as it was propounded. Which will alfo be more fully manifeft, by comparing the Centripetal Force of the Moon with the Force of Gravity upon the Surface of the Earth: Which will be done in the next Proposition.

LXXIX. The Moon gravitates perpetually towards the Earth; and by the Force of Gravity

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is always drawn back from a Rectilinear Motion; and retain'd in its Orb.

For by Experiments of Pendulums, which have been made as exactly as could be, it appears that the Force of Gravity upon the Surface of the Earth, is of the fame Quantity with the Centripetal Force of the Moon; which hath been fhew'd to be in the duplicate Proportion of the Diftances reciprocally: And confequently from the faid Experiments, that Quantity of the Moon's Centripetal Force is more fully demonftrated; and at the fame time it is fhew'd, that that Centripetal Force of the Moon is no other than that Force which we call Gravity. For if any should fay that it is different from it, it must be acknowledg'd however that that Centripetal Force of the Moon, whatever it is, would be felt upon the Face of the Earth; which Force therefore, as join'd with the Force of Gravity, would make Bodies to fall to the Earth as fwift again as they do, and in the Space of one Second of Time to describe 22 2 English Feet instead of 16 1. [Unlefs any one fhould fay, that this Force of Feet 16[1, in one 2d of Time, is indeed a Compound Force, compounded of that Force wherewith the Moon tends to the Center of the Earth. and that Tendency thither which Bodies upon the Face of the Earth would have without it.

LXXX. The Secondaries of Jupiter and Saturn gravitate towards Jupiter and Saturn respectively, and the Planets which are mov'd about the Sun, immediately gravitate to the Sun; and by the Force of Gravity are drawn back from Rectilinear Motions, and retain'd in their Orbs.

For the Revolutions of all these Planets about their respective Centers, are Phænomena of the fame Kind with the Revolution of the Moon about

about the Earth; and therefore ought to depend upon Causes of the same Kind: Especially when it hath been demonstrated, that the Forces on which these Revolutions depend, respect the Centers of *Jupiter*, Saturn, and the Sun; and that in departing from *Jupiter*, Saturn, and the Sun, they decrease in the same Proportion, as the Force of Gravity decreaseth in the Recess from the Earth.

Corol. (1.) Therefore Gravitation is towards all the Planets. For it is certain, that Venus, Mercury, and the reft of the Planets, are Bodies of the fame Kind with Jupiter and Saturn: But we note alfo in this place, that by the 5th Law of Motion Gravitation is reciprocal; and that as the Secondaries of Jupiter and Saturn gravitate towards their Primaries respectively, fo their Primaries gravitate respectively towards them; and the Earth towards the Moon; and the Sun towards all the Planets, both Primary and Secondary.

*Corol.* (2.) The Gravity which respects every **Planet**, is reciprocally as the Square of the Diftance from the Center thereof.

LXXXI. All Bodies gravitate towards each of the Planets; and their Weights towards the fame Planets, at equal Diffances from the Center of the Planet, are proportional to the Quantity of Matter in each.

The Descent of all heavy Things towards the Earth, if you set alide that unequal Retardation which ariseth from the Resistance of the Air, is in equal Times, as hath been observed now for a long time, and we also noted before; whether the descending Bodies be great or small, soft or hard, or of whatsoever Texture of Parts. Which exactly agrees with the Experiments of Pendulums

lums defcending in Arches, whether Circular or Cycloidal. For all Bodies being let down at the fame Diftance of the Center of Oscillation from that of Sufpension, and in equal Arches, make their Ascent and Descent in equal Spaces of Time, and vibrate for a long while. Therefore, fince the Obliquity of the Curvilinear Motion is. in this Cafe, every where like and equal; the fame Bodies let down together in a Vacuum would, in equal Times, describe equal Spaces in a perpendicular Descent; and confequently are impell'd with a Weight every where exactly proportional to the Quantity of the Matter. For where a double or treble Quantity of Matter is urged with a Force double or treble, and no otherwife : the Velocity of the Motion will always be equal: that is, where any equal Particle of any Body whatever is urged with an equal Force of Gravity, the Sum of all, whether in a great Body or a fmall, will be urged with a proportional Force of Gravity; and all, neitheir accelerating nor hindring one another's Endeavours, will always defcend with equal Velocity, and will in the fame degree gravitate towards the Earth. That the Thing is thus in the Experiments of Pendulums, we shew'd before ; and our Author try'd the Matter particularly in Gold, Silver, Lead, Glass, Sand, common Salt, Wood, Water, and Wheat. He took two wooden Boxes round and equal, and fill'd one with Wood; and the fame Weight of Gold he hanged, as exactly as he could, in the Center of Oscillation of the other. The Boxes hanging by equal Cords, of Eleven Foot each, made Pendulums altogether equal, as to Weight, Figure, and the Refistance of the Air. And being placed just by one another, they were found to vibrate equally, and to go and come together

gether for a long while. And in Bodies of the fame Weight, the Difference of the Quantity of Matter, which would fcarce amount to the 1000th Part of the whole, might, by these Experiments, be manifeftly discovered. But now that the Nature of Gravity towards the reft of the Planets, and towards the Sun it felf, is the fame as that towards the Earth, there is no reason to doubt. Which is also manifest from the Spherical Figure of all, which can fcarce be deduced from any Thing elfe, than an Equilibrium of all the Parts, mutually gravitating towards each other. Furthermore, let Terrestrial Bodies be suppos'd to be lifted up unto the Orb of the Moon, and being together with the Moon, depriv'd of all Motion, to be let down to fall to the Earth. Bv what hath just been demonstrated it is certain, that in equal Times they would defcribe Spaces, equal to those which the Moon it felf would defcribe; and confequently, that they are to the Quantity of Matter in the Moon, as their Weights to its Weight. Besides, because the Satellites of Ju-piter and Saturn are revolv'd in Times, which are in the Sefquialteral Proportion to their Diflances from the Centers of Jupiter and Saturn respectively; their accelerating Gravities towards Jupiter and Saturn will be reciprocally, as the Squares of the Diftances from those Centers: and therefore in all equal Diftances from Jupiter and Saturn, their accelerating Gravities will be-come equal, and will equally affect all Bodies. And therefore in falling in equal Times, from equal Heights, they would defcribe equal Spaces, like as it comes to pass in heavy Bodies on this And by the fame Argument, the our Earth. Planets about the Sun let down at equal Distances from the Sun, would in their Descent Z towards

towards the Sun, in equal Times, describe equal Spaces. Moreover, that the Weights of Jupiter and Saturn, and their Satellites towards the Sun are Proportional to the Quantity of Matter, is manifest from the Motion of the Satellites. which is most Regular; and their Orbits, which are almost Concentrical with their Primaries. For if some of these were more Attracted to the Sun in the fame Quantity of Matter than others are. the Motion of the Satellites would be difturb'd by the Inequality of the Attraction; and fo far disturb'd that if at equal Distances from the Sun, the accelerating Gravity of one of Jupiter's Satellites towards the Sun, were greater or leffer than the accelerating Gravity of Jupiter it felf towards the Sun, though it were but by one 1000th Part of the whole Gravity ; then, according to our Author's Computation, the Diftance of the Center of the Orb of the Satelles from the Sun, would be greater or leffer than the Diftance of Fupiter from the Sun, by a 2000th Part of the whole Diftance; or in a Sub-duplicate Proportion of the Diftance; that is, by a 5th Part of the Distance of the outmost Satelles, from the Center of Jupiter; which Occentricity of the Orb would be very fenfible. But the Orbs of the Satellites of Jupiter are concentrick to Jupiter, and therefore the accelerating Gravities of Jupiter, and his Satellites towards the Sun, are equal to one another. And by the fame Argument, the Weights of Saturn, and its Satellites towards the Sun, at equal Diftances from the Sun, are as the Quanricies of Matter in them. And the Weights of the Moon and Earth towards the Sun, are likewife exactly Proportional to the Mals of Matter contain'd in them. And the Thing is the fame, as to the Weights of each Part of every Planet, towards

wards any other whatever ; whether they be Internal Parts, or External : For if fome Parts did gravitate more, others lefs, than according to the Quantity of the whole Matter, the whole Planet, or Satelles, would, according to the Kind of Parts with which it most abounded, gravitate more or lefs than according to the Quantity of the whole Matter; which is contrary to Experience.

Nov. 24. 1707.

### L E C T. XXXIV.



Oroll. (1.) Hence the Weights of Bodies do in no wife depend upon their Forms and Texture.

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For if they were varied with the Forms, they would be greater and lefs, according to the Variety of the Forms in equal Matter; which is altogether contrary to Experience.

Coroll. (2.) Therefore all Bodies which are about the Earth, whether Wood, or Metals, or Stones, or Water, or Air, or Vapours, gravieate towards the Earth, and according to the Proportion of the Matter, are equally heavy. If Bark, or Wooll, or Air, be of the Weight of one Pound in a Vacuum; and Gold, or Silver,

or Brass, be of the same Weight there, the Quantity of Matter will be equal in them all.

Coroll. (2.) Therefore the Weight of all Bqdies whatever in a Vacuum, is the most certain Test of the Quantity of the Matter. For in Bodies equal in Bulk, there is wont to be so great Difference as to the Density, that from the apparent Magnitude, the Quantity of the Matter can in no wife be determin'd. But fince the Quantity of the same is every where Proportional to the same Weight, it may be determin'd most certainly from the fame Weight.

Coroll. (4.) Therefore there must needs be a Vacuum. For if all Spaces were full, the Specifick Gravity of that Fluid, wherewith the Region of the Air, yea, and the Vacuum of Mr. Boyle would be filled, by reason of the Denfity of the Matter, which is the greatest that can be, and most perfect or absolute, or rather infinite, would not fall below, but exceed the Specifick Gravity of Quick-Silver, or Gold, or any other Body, which is counted the denseft and heavieft. And therefore Gold it felf could not defcend in the Air, which is contrary to Experience. To omit here those Arguments which are brought to prove that there could be no Motion in a Plenum, which indeed feem folid enough in themselves to determine us to the same Side of the Queftion.

Coroll. (5.) Since therefore the Quantity of the Matter is every where known from the Weight, as well as the Refiftance; and fince it appears from the Weight, that almost all Bodies upon the Face of the Earth contain more void Space than folid Matter in them; fince alfo, from the very little, and almost imperceptible Refiftance of Planets and Comets, it appears,

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that

that the Heavenly Spaces are void of all Matter ; yea, that the Planets and Comets themfelves, and alfo the Sun and fixed Stars, are, as it were, Points in Comparison of the void Space : It is plain, that Nature is so far from abhorring aVacuum, as some have imagined, the Cartesians especially, that it feems to contain little in it besides a Vacuum: So little can Human Wit perform, in tracing out the Works of God, where Mathematical Reasonings, and Experiments, are wanting.

The most fagacious Mind of Cartes himself, too much destitute of these Foundations, was never able to find out the true Physical Causes of Things, and those which would agree to the later Discoveries.

Corol. (6.) The Force of Gravity is of a different Kind from the Magnetick Power. For the Magnetick Attraction is not in Proportion to the Matter attracted ; lince some Bodies are more, others lefs, others not at all attracted. And the Magnetick Force is far greater, according to the Quantity of the Matter, than the Force of Gravity, fince a very small Loadstone may exceed the attracting Force of the whole Earth it felf. and lift up an Iron Key from it. Nay, the Magnetick Force may be increas'd or remitted in the fame Body; and in the Recess of the Magnet. it decreaseth in more than a duplicate Proportion of the Distance, which yet is the perpetual-Proportion of Gravity; because the Force is much stronger in the Contact of the Surfaces, than when the Bodies are in the least separated from one another.

LXXXII. The Force of Gravity hath Place in all Bodies, all those at least, which are in the System of the Sun, and is Proportional to the Quantity of Matter in each. Z 3 That

That all the Planers do gravitate towards each other ; and that the Gravitation towards every one feparately confider'd, is reciprocally as the Square of the Diftance of Places from the Center of the Planet, we have already prov'd. If there should any Doubt arife here, it must cortainly be about the Gravity of one primary Planet towards another; for as for the common Gravity of all towards their Central Bodies, the Thing, bv what hath been before demonstrated, is plainer, than to be in any wife denied. But as for the other, we have a plain Proof of that alfo. For when fome Years ago, Saturn tarried along while near its Conjunction with Fupiter; and confequently, by reason of the Magnitude and Nearnels of its Body, could not but have some sensible Effect, in disturbing the Satellites of Jupiter, if so be Jupiter, with its Satellites, did gravitate towards Saturn, according to the general Law of mutual Attraction, the Thing was found to be indeed thus: For Mr. Flamfeed himfelf, who at first denied any fuch Disturbance in the Motions of the Secondary Planets of Jupiter, the Thing being better confidered, and the Observations being more exactly compared with the Calculations, ingeniously confess'd, that that Universal Law of Gravity holds in this Cafe alfo; and that those Motions did indeed appear difturb'd by the Neighbourhood of Saturn, and accordingly differ'd from the former Calculations. It follows therefore, by Prop. 81. and the Corollaries thereof, that every Planet gravitates towards every Planet, and that this Gravitation is Proportional to the Matter contained in them. Moreover, fince all the Parts of every Planet, as of Mercury for Instance, do gravitate towards every other Planet, as Venus for Inftance; and the Gravity of every Particle

is to the Gravity of the whole, as the Matter of the Part to the Matter of the whole; and fince alfo all Re-action, by the Sixth Law of Motion, is equal to Action; Venus will reciprocally gravitate towards all the Parts of Mercury; and the Gravity of Venus towards every Part, will be as the Gravity of the fame towards the whole, as the Matter of the Part is to the Matter of the whole.

Carollary. Therefore the Gravity towards every whole Planet arifeth from, and is compounded of the Gravity towards each Part; like as it comes to pass in Magnetick and Electrick Attractions, where by how much the greater the Attrahent is, so much the greater, cæters paribus, is the Attraction: For all Attraction towards the whole, arifes from the Attractions towards each Part; nor can the Thing be conceiv'd otherwise. This will be more easily understood in Gravity, if we conceive many of the leffer Planets, which attract all Bodies feverally, to meet together, and to make one great Planet. For the Force of the whole must be compounded of the Forces of the compounding Parts, and be the adequate Refult of the fame.

But now, if any one fhould in the fame Place object; That all the Bodies with us, on the Face of the Earth, ought to Gravitate thus towards each other; whereas fuch a Sort of Gravitation is never perceiv'd: The Answer is ready, namely, That although the Bodies now spoken of, do indeed Gravitate towards each other, yet since the Gravitation of any particular Body towards another, is to the Gravitation of that Body towards the whole Earth, at the same Distance, as the other Body is to the whole Earth; it must needs Z 4

be far lefs than to fall under the Notice of Senfe. Corol. (2.) The Gravitation towards each equal Particle of a Body, is reciprocally as the Square of the Diftances from the Particles.

LXXXIII. If the Matter of Two Globes gravitating each towards the other, on every Side in Places equi-diftant from the Center, be homogeneous, the Weight of either Globe towards the other, will be reciprocally as the Square of the Diftance betwixt the Centers.

After that our Author had found that the Gravity towards the whole Planet doth arife from, and is compounded of the Gravities towards the Parts, and is towards each Part reciprocally proportional to the Squares of the Diftances from the Parts; he yet doubted, whether that duplicate reciprocal Proportion would hold exactly in the whole Force compounded of the many Parts, or only very nearly. For it might be that that Proportion, in greater Diftances, might hold wellenough; but near the Surface of the Planet, by Reason of the unequal Distances of the Particles, and their unlike Situations, it might notably err. But at length, by Prop. 44 and 45, and their Corollaries, he understood that the fame Proportion holds exactly in fuch fpherical Bodies, as are equally denie every where at the fame Diftance from the Centers.

LXXXIV. A Prob. To determine the Weights of Bodies towards the Planets or the Sun, at given Diftances from the Centers of them.

Cafe (1.) To determine the Weights of Bodies placed without the Surface of the Planets at equal Diftances. Now fince the Weights, at equal Diftances, are as the Quantities of Matter in the Planets towards

towards which the Gravitation is; and fince that Weight or Quantity of Matter is known only by the Quantity of the Attraction, as the Caufe by the Effect; and fince, laftly, that Quantity of Attraction is directly proportional to the Squares of the Velocities in these equal Circles, or reciprocally to the Squares of the periodic Times; the Proportions of the Weights will eafily be known from the Squares of the Velocities. From the periodic Times therefore of the Planets that have others revolving about them, which Times were declared before; the Proportion of the Weights towards the Sun, Jupiter, Saturn, and the Earth respectively, will be as follows.

	Sun		229	600
The Weight	Jupiter		-	208 72
towards the	Saturn			97[328
towards the	The East	rtb	r	I
The Weight towards the <sup>2</sup>	The Mo	07	Ĩ,	O

Now the fame Numbers which fhew the Proportion of the Weight, fhew likewife the Proportion of the Quantity of the Matter. But as for reducing the periodic Times agreeing to the real Distances, to periodic Times agreeing to any given Diftance, it is eafily done by this Analogy; As the Cube of the real Diffance is to the Cube of the Diftance given; fo is the Square of the real periodic Time, to the Square of the periodic Time fought. The square Root therefore of this Number will give the periodic Time which is fought: And by this means the Proportions of the Weights and Matter in the Sun, in Jupiter, in Saturn, and the Earth, are obtain'd. And altho' the Moon, which hath no Satelles about it, doth afford no fuch Argument as this of a Satellit's Weight towards

wards it, or the Quantity of its own Matter; yet notwithftanding, fince it prefents to us another Argument of the fame Thing, to wit, in the Flux and Reflux of the Sea; we thought it not improper to fet down in this place, and by way of Anticipation, that Gravitation towards this Planet, which will afterwards be prov'd from that Flux and Reflux.

Cafe (2.) To determine the Weights of Bodies at the Diftances of the Semi-diameters of the Planets, or upon their Surfaces. This is done by the fame Method as in the former Cafe, and by the like Analogy accommodated to these particular Diftances. In which Calculation, if we take Mr. Flamsfeed's Semi-diameters of the Planets for the true ones, they will ftand thus:

I ne Lattor	in Diameter	763460 1 67870 81155 4444 7935
The Earth Moon Venus Mercury		7935 2175 7906 4240

The Weight therefore of equal Bodies upon the Surfaces of those Stars, is as follows:

The Sun	24
· · · · · · · · · · · · · · · · · · ·	r
Saturn	I 99
e of The Moon	0[515
F CSaturn	IL7

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April 26. 1708.

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### LECT. XXXV.

Problem. To determine the Densities of the Planers; Since we have the Quantity of the Matter in five Planets determin'd

in the former Cafe of the laft Proposition; and in the latter Cafe, we have the Diameters of the Planets determin'd according to Mr. Flamsted; it will be no difficult thing, from the given Quantity of Matter contain'd in the given Spheres, to compute the Density of the same Matter; which is done to hand in the following Table.

and the second	( The Moon	700
anda i	The Earth	3[87
The Denfity of	The Sun	ILOO
	JJupiter	0[76
Section 2	Saturn	060

LXXXVI. Gravity in proceeding from the Surfaces of the Planets downwards, decreaseth in the simple Proportion of the Distances from the Centers very nearly.

For if the Matter of the Planet were every where the fame as to Denfity, this Proportion would hold exactly by Prop. 47. And where it obtains not exactly, the Difagreement is no other than fuch as the unequal Denfity ought to produce.

Corollary. Therefore the Gravity of Bodies on the Surfaces of the Planets, is the greatest of all, and

and on both Sides decreaseth; and is upwards in the reciprocal duplicate Proportion of the Diftance, and downwards in the simple Proportion direct.

LXXXVII. The Motion of Planets and Comets may be maintain'd for a very long Space of Time in the Heavens.

For fince the Refiftance of Mediums, which alone can ftop or retard these Motions once begun. is diminish'd in Proportion to the Weight or Denfity of the Matter; fo that Water, which is near 14 times lighter than Quickfilver, doth refift lefs in the fame Proportion; and Air, which is almost a thousand times lighter than Water.doth refift less in the same Proportion: If we look bevond our Atmosphere, which doth it felf also wax more rare by degrees, as it were infinitely, into the Heavens, where the Weight or Denfity. of the Medium is vaftly diminish'd, above what it is in any Part of our Atmosphere; the Resiftance will be fo very fmall, that for fome thoufands of Years it can scarce become any whit fenfible; accordingly it is evident that it hath been insensible, because the Celestial Motions have endured from the Infancy of Aftronomy unto this Day, without any notable Change or Lofs of Motion.

Corollary. But fince, in an infinite Duration, that very fmall Refiftance, if there be any, muft needs retard and ftop all those Motions; it is manifest upon this Hypothesis, that the present State of the Heavens neither was eternal à parte ante, nor shall be so à parte post. And this will hold good upon another Account also, especially if with Sir Isaac Newton we suppose the Force of Gravity to obtain not only in the Solar System, but also thro' the whole Universe. For if the Fixed Stars,

Stars, or Suns with their Planets and Comets, of whatfoever Number they are, fo that it be not infinite, be subject to the Force of Gravity; In an infinite Time it would have come to pals thoulands of Years ago; that they would have been gathered together into one Heap, and have been reduc'd to reft in the Center of the Universe. Which thing also would, fome time or other, come to pass in an infinite Time yet to come, without the Interpolition of the Divine Providence. As therefore the prefent State of Things had a Beginning, which is owing to the good Will, Wifdom, and Power of Almighty God; fo at length it may and will have an End upon the foregoing Hypothesis, that is, according to the Natural or Establish'd Order of Things; unless it should please Almighty God, by his extraordinary Interpolition, to prevent it: Without whole continual Interpolition, on which this wonderful Force of Gravity wholly depends, it cannot last the least Space of Time.

LXXXVIII. The common Center of Gravity of the Earth, Sun, and all the Planets, either refts, or is mov'd uniformly in a right Line. This is manifest from what hath been demonstrated before : But indeed it appears by no certain Token, whether it refts or is mov'd. This only is to be concluded, That if it be mov'd, and with it the whole Solar Syftem, the Motion muft needs be very flow [ unless it be mov'd uniformly and evenly with the Centers of other Syftems.] For the Fixed Stars, which encompais us on every Side, neither appear greater nor less to us at this Day, than they did to the Ancient Aftronomers 2000 Years ago. Which Phanomena feems to shew the rest, or at least the very slow Motion of the faid Center.

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Corol. (1.) Hence the common Center or Gravity of the Sun, and all the Planets, is to be reckon'd for the Center of the Solar System, or Planetary World. For fince the Sun, and all the Planets, gravitate towards one another, and therefore are in perpetual Agitation, more or lefs, according to the Force of their Gravity, as it hath been shew'd under the foregoing Laws of Motion; it is plain, that their moveable Centers ought not to be reckon'd for the quiescent Center of the World. If that Body indeed is to be placed in the Center, towards which all other Bodies do most gravitate, and which is next of all to the unmoveable Center, as it is reasonable that we fhould efteem it ; that Privilege certainly is to he allow'd to the Body of the Sun; which therefore, speaking physically, is defervedly accounted the Center of the Planetary World. But, if we would fpeak accurately and mathematically, fince the Sun it felf is mov'd, and no fenfible Body doth reft in the Center; the Center of Gravity of the whole System is to be chosen for the real Center of our World, which Center doth indeed most probably rest, and the Center of the Sun comes very near to it. Upon the whole therefore, Physically the Sun, but Mathematically the Center of Gravity is the Center of our World.

\* Corol. (2.) There is therefore no perfect Reft of a real Being in Nature. For supposing that the common Center of the System doth reft, that is the only thing (if we may so call it) which doth reft; all the Parts of the Systems being in perpetual Motion. I faid real Being; because this Center of Gravity is not a physical Body, or any thing real, or other than a Mathematical Point, *i.e.* a plain Nothing: from whence, in confequence of our present Argument, it is to be faid that nothing, real

real doth reft, or that there is not any real and abfolute Reft in the whole Solar System.

IXXXIX. The Body of the Sun doth never reft, but is in a perpetual Agitation: tho' it never departs far from the common Center of Gravity of all the Planets. For fince the Quantity of Matter in the Sun is to the Quantity of Matter in Fupiter, 2s 229600 is to 208 72, or as 1100 to I; and the Diftance of Jupiter from the Sun is to the Semi-diameter of the Sun, as 424,000,000 is to 281,720, or as 1100 to 1; that is, in the fame Proportion or thereabouts; the common Center of Gravity of the Sun and Jupiter, which is placed at a Diftance reciprocally proportional to those Bodies, will fall upon the Surface of the Sun almost. By the fame Argument, fince the Quantity of Matter in the Sun, is to the Quantity of Matter in Saturn, as 229,600 is to 97[328, or as 2360 is to 1; and the Diftance of Saturn from the Sun is to the Semi-diameter of the Sun, as 777,000,000 is to 381,730, or in fomething lefs Proportion; the common Center of Gravity of Saturn and the Sun will fall upon a Point fomething below the Surface of the Sun. From whence the common Center of Gravity of Jupiter and Saturn, as placed on one Part, and of the Sun as placed on the other, will in no wife be diftant by a whole Diameter of the Sun from its Center. And in pursuance of the fame Argu-mentation, if the Earth and all the inferior Planets are, in the Libration, underftood to be fet on the fame Side of the Sun: By reafon of the Nearnels and Smalnels of those Planets; the common Center of Gravity of all will scarce be distant from the Center of the Sun one entire Diameter thereof. But in other Cafes, which commonly happen, the Diftance of the Centers 1 . . 1 is

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is lefs; and where the Planets placed on this Side and on that do counterpoize one another, none at all. Therefore, altho' the Center of Gravity be indeed fuppos'd to reft; the Sun, by Reafon of the various Situation of the Planets, will be mov'd a little towards all Parts; but will never depart far from that common Center of Gravity.

All the Primary Planets are mov'd in XC. Ellipses, which have a common Focus in the Center of the Sun; and by Rays drawn to that Center, they defcribe Areas proportional to the Times. This is true also in the Secondaies, as revolving about the Centers of their Primaries. We deduc'd these things above from Astronomical Phænomena; but now the Principles of thefe Motions being known and eftablished, from these we gather these heavenly Motions à priori. For from the Direction of Gravity towards the Centers of the Sun and primary Planets, the forefaid Proportionality of the described Areas doth follow; and from the Law of Gravity towards those Centers, which is in the reciprocal duplicate Proportion of the Diftance, that Elliptic Figure of the Orbs about those Centers placed in the Foci is neceffarily deriv'd, as we have demonstrated above out of our Author. And these things would be exactly thus, if the Sun and the Primary Planets refted from acting mutually upon one another. For their Orbs would be in Geometrical Srictness Elliptical; and the described Areas would be exactly proportional to the Times. However, those mutual Actions of the Sun and Planets upon one another are fo very fmall, that they ought not to be regarded. And the Motion of the Planets about the Sun as moveable, or any other Planet as fuch, is lefs difturb'd than it would be if the fame were unmoveable, as we observ'd before : From

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From whence, speaking physically, the Propolition is still to be accounted true. The Action indeed of Fapitien upon Saturn, and its five Satellites; and of Saturn upon Saturn, and its four Satellites; is not alcogether to be neglected. Since these Planets are great ones, and placed at a very great Distance from the Sun. From whence, by their mutual Attractions about their Heliocentric Conjunctions; which, by reason of the Slownels of their Motions; endure for no instill Time; some Inequalities will arife on both Sides, is well in the Figures of their Orbits, as in their Motions; but yet fcarce to be so much diftinguish'd in the unequal Motions of the Primaries themselves, as in those of their Secondaries, of those about Fapitar especially;

s muscholium: According to our Author's Computaron, the diffurbing Force or Gravity of Saturn : towards! Flepiten lisiste bthe Gravity of Saturn - 51 towards the San, about the Conjunction of those Planers, as i is to 204, or thereabouts. And the Différence of the Gravities of the Sun towards Saturn, and of Jupiter towards Saturn, is to the Gravity of Jupiter towards the Sun, as I to 1923. To which Difference the greatest diffurbing Force of Sation towards Jupiter is propori Hondla From whence the Difturbance of the Orb ficing stuptor is far lefs than is that of Saturn: But the Diffurbances, which are in the reft of the Orbs, are fo very small, that they are not to be regarded. XCI. The Aphelia and Nodes of the Orbs doiteft. Because of the Force of Gravity in the duplicate Proportion of their Diftances reciprocally, the Aples and Aphelia ought to reft of themfelves; as was noted before. And becaule the fame Force doth always respect a Point almost unmoveable, the Planes of the Orbs ought Valido general vali val al gracenti alfo

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allo to reft'; and when the Planes reft; the Nodes or Interfections of the Orbs mult real too, f. But it is to be noted, that in Sucdeffion of Time fome Inedhatities will arife from the Actions of Planets and Comets upon one another; buorizat they will be fo very fmall, that by reafon themof they are not to be regarded utit is allo to be noted, that we do in this place suppose, with all; Afronomers, the RefPof the Center of Graviey of the whole Syftem ; althours was hinted about in are not yet able to demanstrate that Reft. 23 Thefe Things Supposed, we shall deduce the following Corollaries. cuoixolA leuponu orb ni b'flingniff

Corol. (1.) The Pined Stars, reft; becaule that they keep their given Poficion's towards the Anhelia and Nodes which reflues This will feen a new Way of reafoning and Aftronomy, infer the Reft of the Fixed Starsifrom the Reft of the Syftems of the Planets owneres, on the contrary, 'we have hitherto been wone to determine the Moriohs of the Planets from the supposed Reft of the Fixed Stats. And thus it must needs have been, follong as our Famous Author's true Caufes of the Celeffial Motions were unknown. 01 Corol. (z.) Since the Marallan of the Fixed

dir. Stars, even the Annul sin to very fmall that it fcarce falls under the Oblervation of the molt accurate Obfervers : The Force of the fer Stars, by Realon of their inmense Diftance, gan, produce no fenfible Effects in but Syftem.1.

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Corol! (2!) -From whence in follows other, 7diciary Affrology, as it is called, I which I depends not 11 only dipon the Policions and Influences of the Planets, but of the Fixed Stars allor wants all fure Poundation ; fince is supposed the Forces of those Bodies roube exceeding ogreation which the foregoing Corollary has rightly observ'd, 216

are indeed very fmall, or rather none at all. But "we may add this allo me the prefent Cafe, that "the influential Force of all the reft of the Planers, excepting the Sun and Moon, which Aftrologers "talk for much of, is either by reafon of the Im-"menfe Diffance, or the Smallnels of their Bodies, "to very little in our Atmosphere, and about the Earth, that It can feared be by any fure Token differed? If far is it from being able to produce those great and wonderful Effects which they "Inpole. Those who, like Idolaters, conceive the Stars to be Gods, or that Gods possible and animate them, have fomewhat wherewithal they "who flave quitted for gross an Error as that, it is a Wonder how they floud come to adhere thus obstinately still, to those Aftrological Trifles and Abfurdicies.

May, 17. 1708.

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XCII. THE Diurnal Motions of the Planets are uniform and equable; and the Librations of the Moon arifes from its Diurnal equable Motion, as compar'd with its Menftrual Inequable, and perform'd according to an Axis inclin'd to its Orbit. Thefe Things are noted ellewhere : and there-

fore we need not make many Words about them now.But because the Dayof the Moon revolving A 2 Uni-

Uniformly about its own Axis, is a Month (I mean here the Periodic Month;) The fame Face of this Planet will always nearly respect the Superior Focus of the Ellipfis, but not the Earth. which is placed in the Inferior Focus; because the Angular Motion also about that Focus is almost equal, but about the Earth plainly unequal. And therefore, according to the Situation of the Superior Focus, it will decline commonly on this Side, and on that from the Earth, and will fhew to us sometimes more Easterly, sometimes more Westerly Parts; which is the Libration of the Moon as to Longitude. But the Libration of the fame as to Latitude, wherein fometimes more Northerly, and at other times more Southerly Parts are prefented to us, must arife from the Inclination of the Moon's Axis to the Plane of the Orbit ; as is manifest to him that confiders it.

Corollary. We may note also in this place, as we have done elsewhere, how exactly these two Motions of the Moon, which in no wise depend one upon the other, to wit, the Diurnal and Menftrual, do agree together; so that the one hath not been found for above these 2000 Years, to overgo the other in the least. Which could not be without the Providence of God.

XCIIL The Axes of the Sun and Planets, which are moved with a Diurnal Motion, are lefs than those Diameters which are Perpendicular to those Axes. Or the Figure of the Sun and Planets, which are revolv'd each about its own Center, is that of an Oblate Spheriod; that is, that of a Solid produc'd by the Revolution of an Ellipsis about its Leffer Axis.

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The Planets, and all the Celeftial Bodies whatever, if all Circular Diurnal Motion were taken away, must needs, by reason of the equal Gravitation of the Parts on all Sides, put on a Spherical Figure. But on the fame Diurnal Motion it will come to pass, that the Parts necessarily receding from the Axis of Motion, and thereby detracting from the Gravity about the Equator, must endeavour to ascend, where the Motion is the swiftest. And therefore in that Place the Matter of the Planet. unless it be very Solid, will by its Afcent unto the Equator increase the Diameters of the fame ; but will diminish the Axis at the Poles, the Gravity of the Parts being nothing diminish'd there. Thus the Diameter of Fupiter (the Observations of Calimi and Mr. Flamfteed agreeing thereto) is found to be shorter about the Poles than from East to West. And by the fame Argument our Earth ought to have its Axis leffer than the Diameters of the Equator. For unless the Thing were fo, and that our Earth were fomething higher at the Equator than about the Poles, the Seas, by reason of the greater Gravity there, would fettle downwards about the Poles, and in Afcending about the Equator would overflow all. But by reafon of the greater Velocity of the Diurnal Motion, and the leffer Denfity ; Jupiter ought to have a much more fenfible Difference of its Diameters than any other of the Planets, or than the Sun it felf. From whence Aftronomical Observers have hitherto been able to discover this Difference in no other Planet but this. But that our Earth is of this Figure, appears not only from the Argument just now produc'd, but also from the most accurate Experiments which have been made by Pendulums. For by how much the nearer Pendulum-Clocks, of the fame Length of the Pendulum, are brought to A a . 2 the

the Equator, fo much the more Slow their Vibrations are observed to be; and by how much the nearer they come to the Poles, their Vibrations are found fo much the quicker: because the Center of the Earth, which in the former Case is more remote, and in the latter nearer, doth promote the Acceleration and Retardation of the pendulous Bodies respectively; as it must neceffarily come to pass according to the present Proposition.

Scholium. If you would know exactly the Proportion of the Axis of every Planet unto the Dia-' meters of the Equator, you must go through the manifold Intricacies of our Author's Calculation. But if you would have the Benefit of this Calculation without the Trouble of the fame. take it thus, By Calculation our Author found that the Centrifugal Force of the Parts of the Earth under the Equator, a iling from the Diurnal Motion, is to the Force of Gravity upon the Superficies of the Earth, as r is to 289. From whence if (in Fig. 1. Plate 9.) A P B Q reprefents the Figure of the Earth, produced from the Revolution of an Ellipsis about the Lesser Axis PQ; and ACQ, acq be a Canal full of Water, reaching from the Pole Qq' to the Center Cc, and from thence going forwards towards the Equator A a; the Weight of the Water in the Leg of the Canal AC c a, is to the Weight of the Water in the other Leg QC c q, as 289 is to 288 almost. Because the Centrifugal Force arising from the Circular Motion, will fustain and take away one Part from the 289 Parts; and the Weight 288 in the other Tube will fuffain the reft. of the Parts. For the Thing is not only true in the Surface of the Earth, but in all the Parts of both the Tubes, because the Centrifugal Force,

Force, and the Gravity of the inferior Parts, as taken every where at proportional Diftances from the Genter, are diminifad in the fame Proportion in the Progress to the Center. And then, by continuing the Calculation, the Gravity towards the Earth in the Place Q will be to the Gravity in the Place A, as for is to goo; and the Centrifugal Force in will make that the Excels of Altitude in the Leg A C c a, fhould be a  $\frac{1}{687} = \frac{1}{225}$ th Part of the Alritude in the other Leg QC cq; or in our Earth, that the Semi-diameter of the Earth at the Equator, should exceed the Semiaxis or Semi-diameter at the Poles by about 17 b Miles, These Things, I fay, will be thus, in Cale that the Earth confilts of an uniform Matter. For if the Matter at the Center be more dense, as certainly it ought to be, than at the Surface; the Excels of Altitude at the Equator must be fomething greater: because that if the redundant Matter at the Center, whereby the Density is made greater, be subducted and considered feparately; the Gravity towards the reft of the Earth uniformly dense, will be reciprocally as the Diftance of the Weight from the Center; but the Gravitation towards the fame redundant Matter, will be reciprocally as the Square of the Di-france from that Matter nearly. Therefore the Gravity under the Equator, which is towards that redundant Matter, will be lefs than, the Gravicy was towards the Place of that Matter by the? foregoing Calculation; and therefore the Earth there, by Realon of the Defect of Gravity, will alcend lomething higher than was defin'd above. But now the French have found by Experiments, that the Length of Pendulums performing their' Vibrations in one Second of Time towards the Equator, is less than that in which they perform. Anell oil to nerrold Ara A or H () I the i.

the fame towards the Poles in a greater Proportion than the foregoing Calculation requires. And therefore the Earth feems to be much higher at the Equator, than the foregoing Computation makes it to be, and indeed no lefs than 21 ½ Miles: and accordingly to be denfer at the Center than in Mines near the Surface, as Reafon altogether requires.

Corol. (r.) If the Excels of Gravity in the Parts about the Poles, above that which is in the Equatoreal Parts, were once more exactly defin'd by more accurate Experiments made to that Purpole, we fhould at length have an universal Meafure determin'd; that, to wit, which would exactly define the due Length of a Pendulum for Seconds, in the feveral Places which lie betwixt the Equator and the Poles. From whence, as well an Equation of Time, which is now indicated by equal Pendulums in divers Places, as the Proportion of the Semi-diameters of the Earth, and of the Denfity of the fame at the Center, fo that the fame be fuppos'd to increase uniformly, will become known.

Corol. (2.) Since the Proportion is the fame in a Canal full of Water, as in one fill'd with any other Fluid; and the fame alfo as in the Earth, which is fuppos'd to be fluid within; while in the mean time in a folid Earth the thing is otherwife; fince alfo it is known by Experiments and Obfervations, that the Earth is indeed higher at the Equator than at the Poles: from thence it is manifeft, that either the whole Earth was fluid, when its diurnal Motion first began; or at least that if contain d a great Fluid within, which, by yielding, might give place to the Elevation at the Equator; and Depression at the Poles.

Corol. (2.) If the diurnal Motion of the Earth fhould

fhould be gradually retarded, unless it contain'd within it fome great Fluid, which would give way to the Change of its Figure; the Seas would defcend towards the Poles, and overflow all there.

Corol, (4.) If the diurnal Motion of a Planet, of a greater or leffer Magnitude, but of a given Denfity, be accelerated or retarded in any Proportion whatever, the Centrifugal Force will be increas'd or diminish'd from thence in the duplicate Proportion thereof; because of the Increase or Diminution, both of the Curvature and the Velocity in the fama Proportion : and therefore the Difference of Semi-diameters will be increased or diminished in that fame duplicate Proportion. But if the Denfity be increas'd or diminish'd in any Proportion whatfoever, because the Gravity is increas'd or diminish'd in the fame Proportion, the Difference of Semi-diameters will be increas'd or diminish'd in that same Proportion also: That is, the Difference of Semi-diameters will be in a Proportion compounded of the duplicate Proportion of the periodic Times, and the fimple Proportion of the Denfity, both reciprocal. From whence, fince the Difference of Semi-diameters in the

Earth is  $\frac{3}{687}$  Parts of the whole Semi-diameter;

and the Square of the periodic Time in Jupiter, which periodic Time is  $9^{h}$ , 56', is to the Square of 24<sup>h</sup>, the periodic Time in the Earth, as 5 to 29; and the Denfity of Jupiter is to the Denfity of the Earth, as 1 is to 5: the Difference of the Semi-diameters in Jupiter will be to the Difference of the Semi-diameters in the Earth, as  $\frac{29}{5}$  in  $\frac{5}{1}$  in  $\frac{1}{229}$  is to 1, or as One is to Eight-Therefore the Semi-diameter of the Equator of

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of the Equator of Jupiter, is to the Semi-axis as 9 to 8. From whence, by the way, It is no wonder that to great a Difference flouid lie open to Aftronomical Obfervation. But it is to be remark'd, that these Things are thus, where the Denfity of the Planer is uniform. But if the Matter of Jupiter be denfer at the Center than at the Circumference, as it was before obfervid in 1 general; the Difference of Semi-diameters will be greater ftill, and more easy to be obfervid. Let the Aftronomical Obfervers therefore take notice how far this Corollary agrees with the Diameters of Jupiter, which are measur'd by the Micrometer.

XCIV. The Increase of Weight in going forwards from the Equator to the Poles, is very near as the Square of the right Sine of Extitude; or, which is the fame, as the versed Sines of Latitude themselves.

Becaule the Weights of the unequal Legs of the Canal of Water ACQ qca are equal, and do poize one another; and the Weights of the Parts, like or fimilar to the whole Legs, and which are alike fituated, are to one another as the Weights of the Wholes, and confequently are equal betwixt themselves; the Weights of the Parts, which are equal, and alike, fituated in the Legs will be reciprocally as the Legs: that is, reciprocally as the Diffances of the Bodies from the Center of the Earth and the thing is the fame, in all homogeneous and equal Bodies, whatfoever, which are alike fituated in the Legs of the Canal. Bodies placed in the uppermost Parts of the Canals, or in the Surface of the Earth, will have their Weights in that Proportion to one another reciprocally, as their Diftances from the Center are : And the fame is to be faid of Weights,

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in any other Regions whatever, through the , whole Surface of the Earth. And the Increase of Weight in the Earth, which is a Spheroidal Oblate Body, as the Famous Dr. Gregary hath demonstrated (Astron. Book III. Prop. 52.) is as the Square of the right Sine of the Latitude of the Place, or, which comes to the fame, as the versed. Sine of Latitude nearly.

Coroll. Since therefore Dr. Gregory hath demonstrated in the same Place, that the Longitudes of Pendulums vibrating in equal Time, are betwixt themselves as the Distances from the Center of the Earth reciprocally; the Difference of the Length of Pendulums will be as the Square of the right Sine of Latitude: And thus every where.

XCV. The unequal Motions of the Satellites of *Jupite*, and *Saturn* are plainly 'like and analogous to the unequal Motions of the Moon, and arife from like Caufes.

I mean the Motion of the Nodes in Antecedentia, and of the Aples fometimes in Antecedentia, but more flowly, and fometimes in Confequen-tia more fwiftly, and by the Excels of this latter -Motion their being mov'd in Confequentia upon. the Whole; the Motion of Variation, and the reft of the like Motions, must be the fame in these Secondary Planets as in the Moon, and therefore do not require to be diffinely handled. It is true, that by Reafon of the Smallness of thefe Inequalities, and Slownels of thefe Motions in the other Secondaries, their Motions appear very regular, when compared with the Motions of the Moon; which hath made fome of the later Aftronomers to deny all Motion to the; Nodes of those other Secondaries. Nevertheles, Mr, Flamfteed, in conferring his Observations with thole

thole of Mr. Cassini, hath found that the Nodes of thole about *Jupiter* do indeed go back though more flowly; and it is not to be doubted, but that Time will more certainly and exactly discover and determine the same, and all the other mention'd Inequalities in the Satellites, both of *Jupi*ter and Saturn.

#### May 21. 1708.

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### LECT. XXXVII.

XCVI. THE Flux and Reflux of the Sea arifes from the Gravitation of the Water towards the Sun and Moon, or the Attractions of those Luminaries.

That the Sea in the Space of every Day, as well Lunar as Solar, ought to fwell twice, and twice to fettle and fall-back, is manifest from what hath been demonstrated above. But that the greateft heighth of the Water doth not fall just at the Appulse of the Luminaries to the Meridian, but follows the fame by the Space of about three Hours, is what we shall now undertake to explicate. That the thing is indeed fo, appears from the Observations of the Tides, as well as in the Atlantic Ocean, and the whole Eastern Tract of the Ethiopic betwixt France and the Cape of Good Hope, as upon the Coaft of the Pacific along Chili and Peru. In all which Shores, the High-Water fatts about three Hours after the Time aforefaid : unlefs

unless it be where the Motion is not retarded by its being propagated through Shallows. Now the Reason is this: When the Luminary is in the Meridian, the attracting Force is then certainly at the greatest, but the Effect of that Force is not on perfeveres uniformly until a contrary Motion defitions, or at leaft retards it yet come to its Height. For all impress'd Moti-Sea; or Ocean rather, which for the Six Morning : ··· Hours, if we may to call them when we speak of the Moon, is continually increas'd; and by its confpiring with the diurnal Motion, accelerated ; ought, by reason of this its greater Celerity, to go forwards still farther, and to accumulate the Waters more and more, until the fame Force, by tending afterwards contrary to the diurnal Motion, doth by degrees retard the Course of that Motion which is going forwards; and by and by to make the fame Waters to proceed with lo flow a Motion, that a Reflux of the Ocean follows : Which Retardation of the Motion ought 13 to be most notable about the Octants, or the third Hour. Examples of fuch like greagest Effects, as following some pace of Time after their greatest Causes, we have yearly in the greatest Heat of the Summer, and Gold of the Winter; which falls not in the Solffices themfelves, but about the Octants, if I may fo speak, about a Month and an half after; and in every Summer-day, in the greatest Heat of the Day; which happens an Hour or two after Noon, rather than at the Noon it felf. So in the prefent Cafe, whilft after the greatest Force of all, and that raising of the Waters which is thereby, Forces next to the greateft, and not yet turned to the contrary Part, do fill operate; the Forces which are less than the greatest, being super-added to the Motions which Were

vowere ftirr'd up by the greateft, and go forward by their own proper Tendency, must needs ob-- rain a greater, Effect, than Forces still increasing, - fuper-added to leffer Motions, could have. Then 1 it is to be noted allo, that the attracting Force it - felf, which lifes the Water directly upwards, doth nofearce fenfibly fall thort, of its greateft Quantity for an Hour or two after the Appulle of the Luminary to the Meridian, altho' the Direction of the Attraction which accelerates or retards the Waters, be directly upwards in the Meridian it - felf, and from thence is chang'd. The Waters ... therefore will be most accumulated, where the Barts, which have just now pass'd the Meridian with the greatest Velocity, do fall upon ather Parts which had before been retarded at the Quadrature ; and to by confpiring with the Endeavour of the other, do make the greatest. Flood of diall: which happens about the third Hour, For - in this place we intend not fo much the vulgar Hours, as those which are reckon'd from the Ap-In pulle of the Sun and Moon to the Meridian of er the Place, as well below as above the Horizon; f) and by the Hours of a Lunar Day, we underfland 10 24 Parts of shat Time in which the Moon, by its risiapparent dininal. Motion, is revolved to the Mebridian of lany Place.

XCVIL - The Tides which depend on the Force of the Sun, and on the Force of the Moon in feverally, do not make a double Tide, but a fingle mome; which is to be estimated from the Conjunandien of their Forces.

- W For like as any Body whatever, which is imfor like as any Body whatever, which is imfor prefield by a double Force, cannot go forwards in Ill fit wo Lings: but from the Conjunction of the Forsubcess will proceed in the Diagonal of a Paralleloto gram in the fame manner as if it had been acted area in the fame manner as if it had been acted area in the fame manner as if it had been acted

upon by one lingle Force', according to the Di-rection, of the Diagonal ; fo thole two Motions, which the two Luminaries do excite telpectively, will not appear leverally, but will make one will not appear leverally, but will inake one mix'd Motion. In the Conjunction and Oppoli-tion of the Luminaries, their Effects will be con-join d: and the greateft Floods of all will be made, as arising from the Sum of the Forces at that Time. In the Oiladfatures of these Lami-naries, the Sun will fift up the Water, whilf the that Time. In the Oldal atures of thole Lami-naries, the Surl will fift up the Water, whilf the Moon depression of the Water, whilf the Moon lifts it up i and there the Flood will bawhe least of all, as being the Refut of the Difference of the Forces only. And becaule as it appears by Experience, the Force and Effect of the Moon in the present Cala is much greater than that of the Sun, the greates for the greates that that of the Sun, the greates for the greates that that of the Sun, the greates for the greates that that of the Sun, the greates for the greates that that of the Sun, the greates for the greates that that of the Sun, the greates for the greates that that of the Sun, the greates for the greates that the of the Sun, the greates for the greates that the syzygies and not greater for the greates the the syzygies and to had annes, the greatest flood of all; which by offer Lunar Hour, and by the Solar Force Mone in the third Solar Hour, will, by the Composition of the Forces, fail upon fome intermediate Time, which will be much heard to the third Lunar boon that that to the Water will also precede the force of the Greates at which Time the shift be difference water will also precede the force of all, a first greates at which Time the shift be difference greates the third Lunar, the is first of all, a first greates the third Lunar, the is first of all a first greates for the third Lunar, the is first of all a first greates the water will also precede the first of all a first greates the third Lunar, the is first of all a first greates the brow the greates the follow the third Phinar Hour by the like In-the Aponr And the greates Heighth of the Water will follow the third Phinar Hour by the like In-the greates to the Stages Heighth of the Water will follow the third Phinar Hour by the like In-the greates to the Stages and this also by the greates to the Stages and this also by the greates to the Stages and this also by the greates the third Phinar the Ordents of the stages of the stages and the after th

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Moon. Thus will the thing be in the Oceani or open Sea: For in the Mouths of Rivers, the greater Floods, cæteris paribus, require the longer Time, and to come unto their Heighth a little more flowly.

XCVIII. The Tide ought to be different, according to the different Diffances of the Luminaries from the Earth, both every Year and every Month; and this in the criplicate reciprocal Proportion of those Diftances, or in the triplicate direct Proportion of the apparent Diameters.

This we have demonstrated before : Nor is it to be wondred at, that thele Effects should be greater at less Diftances, and lesser at greator, Wherefore the Sun in Winter-Time, when it is about the Perigee, will make, the Tides after the Syzygies to be fomething greater, becaule of the greater Sum of the Forces; and those after the Quadratures to be something less, because of the Difference of the Forces, than they will be in Summer-Time ; cateris paribus. And the Moon. every Month, when it is about the Perigee, will make greater Tides than is Days before and after, when it is in the Apogee. From whence, if the Perigee Situation of the Moon happens, about the Copjunction, the Day Elood will be increas'd, and the Night-Flood diminifh'd ; but if that Situation happens, about the Oppelition, the Night-Flood will be increased, and the Day-Flood diminify'd. From whence allo it comes to pais; that two Tides the greatest of all do not follow one anothen two Syzygies together. For if the Moon bein one of the Syzygies about the Perigee, and raileth the greateft Tide at that Time, by the Conjunction of its Force with that of the Sun ; in the other of the Syzygies it must needs be about the Aposee, and have less Force. XCfX: The

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XCIX. The Tides likewife ought to be divers, according to the divers Declination of the Luminaries from the Equator.

For if the Luminary were placed in either of the Poles, it would draw the Water conftantly without Intention or Remission of the Action : and confequently would make no Reciprocation of the Water. Therefore the Luminaries, in departing from the Equator towards either, Pole, will by degrees lofe their Effects; and therefore will raife leffer Tides after the Solfficial Syzygies, than after the Equinoctial. But after the Solfficial Quadratures, the Tides will become greater than after the Equinoctial; because the Effect of the Moon, which is now placed above the Equator, doth most of all exceed the Effect of the Sun. Therefore the greatest Tides fall after the Equinoctial Syzygies, and the least after the same Quadratures of the Luminaries; and the greateft Flood about the Syzygies is always attended with the least about the Quadratures, as Experience testifies. But by the lesser Distance of the Sun from the Earth in Winter Time than in Summer, it comes to pais, that the greatest Tides and the least do oftner precede the Vernal Equinox, than follow it; and do ofmer follow the Autumnal Equinox, than precede it.

C. Some Phanomena of the Tides, and Effects of the Luminaries, are divers, according to the divers Latitude of Places in the Earth; and effectially as to the Night and Day-Floods, which follow one another immediately.

In Fig. 2. Plate 9. let A p E P defign the Earth, covered on every fide with deep Water. Let G be the Center thereof. p P the Poles. A E the Equator. F any Place without the Equator. Ff the Parallel of the Place. D d the correfpon-B b dent

dent Parallel on the other fide of the Equator. H that Place of the Earth which is directly under the Moon's Place, which was Three Hours before, or the middle Point of the elevated Water. h the Place opposite thereto, or the Point of the Water in the other Part of the Earth. where the Water is most elevated. Kk, Places distant 90 Degrees from thence. CH, Ch the greatest Altitudes of the Sea, measured from the Center of the Earth; and CK, Ck the leaft Elevations. And if from the Axes Hh, Kk an Ellipsi be defcrib'd; and then by the Revolution of this Ellipfis about the greater Axis Hh, there be describ'd a Spheroid HPKhpk; this will describe the Ficure of the Sea nearly: and CF, Cf; CD, Cd will be the Elevations of the Sea in the Places F f and Dd. Moreover, if in the forefaid Revolution of the Ellipsi, any Point whatever, as N describes the Circle NM, which cuts the Parallels Ff Dd in any Places, as R, T, and the Equator AE in S, CN will be the Heighth of the Sea in all the Places R, S, T fituate in this Circle.1 Hence, in the diurnal Revolution of any Place whatever, as F, the Flood will be the greatest there, three Hours after the Appulse of the Moon to the Meridian above the Horizon ; afterwards the Ebb will be the greatest in Q, three Hours after the fetting of the Moon; then the Flux will be the greatest in f, three Hours after the Appulse of the Moon to the Meridian below the Horizon; and laftly, the Ebb will be the greatest in Q, three Hours after the rising of the Moon ; and the latter Flood in f. will be lefs than the former Flood in F. For the whole Ocean is diffinguish'd into two Hemispherical Floods; one : in the Hemilphere KHkC, which looks to the North; the other in the opposite Hemisphere Khkc. · ~ ~ ; .

Khkc, which looks to the South; which therefore we may call the Northern and Southern Floods. These Floods, which are opposite each to other. come by turns to the Meridian of each Place. with the Interval of about 12 Lunar Hours betwixt. For fince the Northern Regions do more partake of the Northern, and the Southern Regions do more partake of the Southern Flood; from thence there proceed Tides alternately greater and lefs in each Place without the Equator. But the greater Tide, when the Moon declines to-wards the Vertex of the Place, will fall about three a-Clock after the Appulse of the Moon unto the Meridian above the Horizon; and the Flood, when the Moon changeth its Declination, and recedes from the Vertex, will be changed into a lefs: And the greatest Difference of Floods will, for this Reafon, fall upon the Times of the Solftices, especially if the Moon's ascending Node be in the Beginning of Aries; that fo the Moon, when it is nearest to the Vertex, and the remotest from it, may have the fame diurnal Revolution. And this is confirm'd from Experience; by which it is found, that the Morning Tides do in Winter-Time exceed the Evening; and in Summer the Evening exceed the Morning Tides : At Plimouth, for Instance, by the Heighth of one Foot, and at Briftol of fifteen Inches, as appears from the Observations of Mr. Coleprefs and Mr. Sturmy. But that these Differences do not seem fo great as might be expected in Places for remote from the Equator, may be owing to fome other Caufe. The Motions describ'd hitherto are something chang'd by that Force of the Reciprocation of the Waters, wherewith the Tide, even though the Actions of the Luminaries should cease, might endure for some Time. This Confervation of B b 2 the Y

the Motion once imprefs'd, doth diminish the Difference of the alternate Tides; and makes the Tides next after the Syzygies greater, and diminishes those next after the Quadratures. For from hence it comes to pass, that the alternate Tides at *Plimoutb* and *Briftol* do not differ much more than by the Heighth of 12 or 15 Inches; and that the greatest Tides of all in the fame Ports, are not those which are next after the Syzygies, but the third Tides after them; which agrees exactly with what was faid before. For all these Motions are retarded, in their passing thro' Shallows; fo that the greatest Tides of all in fome Streights, and the Mouths of fome Rivers, are the fourth or even the fifth after the Syzygies.

Nov. 8. 1708.

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#### L É C T. XXXVIII.

CI HE Phænomena of the Flux and Reflux of the Ocean in particular Places, as Streights, Ports, Mouths of Rivers, Imall Seas, and which communicate

little or not at all with the Ocean; in those also which are far distant from the Equator, do recede more than a little from the general Laws of the Tide before fet down, and are commonly altered by those particular Circumstances.

As for Example; it may come to pais, that the Tide may be propagated from the Ocean thro' divers

vers Streights, and quicker through fome than others ; in which Cafe, the fame Tide being divided into two or more which come fucceffively, may make new Motions of divers Kinds. It may come to pais alfo through the Length of the Way, or the various Winding of the fame, or by means of Obstacles which are in the Way, that the Tide may be diminish'd and almost stopp'd. (From whence it comes, that where there be a great Number of Islands, as the Moluccoes, the Philippines, in the Mexican-Bay, the Antillæ, there is almoft no Flux, or far less than in the wide and open Ocean.) Again, a Tide which is in a mean State in the Ocean may become very great in Rivers, because of the Narrowness of the Passages, and the Heighth of the Shores. In fmall Seas alfo, there is none, or a very fmall Tide. For fince the greatest Tide ought to happen in the deep Ocean only, which is open to the Eaft and West for the Space of 90 Degrees; by how much less the Sea is, fo much the less the Acceleration and Retardation of the Waters, that is, the Flux and Reflux, muft needs be; Nor can it be great, unless the Sea doth communicate freely with the Ocean. For if it communicateth not, or but little therewith, as it is in the Mediterranean, a lefs Tide for that Reafon is to be expected. In those Seas also which are remov'd far from the Equator, where the Tide must be propagated in a less degree, especially if they have but little Communication with the Ocean; and fo will be but fmall, as it comes to pass in the Baltic and the Northern Which happens also in the Euxine and Ca-Seas. fpian Seas, not only by reason of their something Northerly Situation, and their small Communication with the Ocean, if they have any at all, but by reason also of the Smalness of those Seas. In Bb 2 Seas

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Seas which lie open, and extend themfelves a great way from the East to the West, as in the Pacific Sea, and the Parts of the Atlantic and Etbiopic Sea without the Tropics, the Water is wont to be elevated unto the Heighth of 6, 9, 12, or 15 Feet. And in the Pacific Sea, which is deeper and wider, the Tides are faid to be greater than in the Atlantic and Etbiopic. In the Etbiopic Sea, the Elevation of the Water betwixt the Tropicks is lefs than that in the Temperate Zones, by reafon of the Narrownels of the Sea betwixt Africa. and the Southern Part of America. In the Middle of the Sea the Water cannot ascend, but it must defcend at the fame time to both Shores, the Eastern and Western; when yet in our narrow Seas it ought to descend by turns unto the Shores. For this Reason the Flux and Reflux must be very fmall in Islands, which are very remote from the Continent. In fome Ports, as hath been very lately observ'd, where the Water paffing thro' shallow Places is forc'd to flow in and out with a great Violence, to fill and empty by turns narrow Bays, the Tides are greater than usual; as at Plimouth and Chepstow-Bridge in England, the Hills of St. Michael and the City of Avranches in Normandy, Cambaia and Pegu in the East-Indies. In these Places the Sea coming and going back with a great Velocity, fometimes overflows the Shoars. and then leaves them dry for many Miles. Nor can the Force of Flowing in and Reflowing be ftopp'd, until the Water be elevated or depress'd 30, 40, 50, or sometimes 60 Feet. And the thing is the fame, in fome measure, in long, shallow, and narrow Streights, as the Magellanic and that wherewith England is encompass'd. But in Shores which have a fleep Descent towards a deep and open Sea, where the Water may be rais'd and fettle

fettle freely without any accessional Force of flowing in and returning; the Tide, if we would determine the general mean Quantity, is to be reckon'd to arife to the Heighth of about 12 Feet, i. e. if we measure from the Low to the High-Water Mark. But of all Sea-Tides, that is the most to be admir'd, which Dr. Halley speaks of from the Observations of Mariners, as being in a Port of the Kingdom of Tunquin at Batham, in the Northern Latitude of 20°. 50'. There the Water, on the Day following the Paffage of the Moon over the Equator, ftagnates; then the Moon declining to the North, it begins to ebb and flow, not twice as in other Ports, but once only in a Day; and the High-Water falls at the fetting of the Moon, and the Low-Water at the rifing of the fame; and this Tide is increas'd with the Declination of the Moon until the feventh or eighth Day; and for the other feven Days, it decreafes by the fame Degrees by which it increafed before : and the Moon changing its Declination, it ceafeth; and from thence is prefently changed into a Reflux. For then the Reflux falls at the fetting of the Moon, and the Flux at the rifing, until this Plane doth again change its Declination. There is a double Entrance into this Port, and the neighbouring Streights; the one from the Chinese Ocean, betwixt the Continent and the Leuconian Island ; the other from the Indian Sea, betwixt the Continentand the Isle of Borneo. It feems probable, that two almost equal Tides do come into this Port from the different Tides of this Ocean ; the former of which precedes the other by the Space of almost fix Hours, and falls ; Hours after the Appulse of the Moon to the Meridian of the Port. When the Moon in this its Appulse to the Meridian is in the Equator, there will come at each fix Hours End equal Fluxes, which falling upon Bb 4 mutual

mutual Refluxes will make the fame equal to the Fluxes; and fo will caufe that for that Day the Water will feem to be moved with no Tide at all. When the Moon declines from the Equator, the Tides in the Ocean will become by turns greater and leffer, as we shew'd in the last Proposition but one; and from thence two greater Fluxes. and two leffer ones, will be propagated into this Port by turns. But the two greater Fluxes, by joining their Waters, will make the higheft Flux in the middle Time betwixt both; the greater Afflux and the lefs will make that the Water should afcend unto a mean Heighth in the middle Time betwixt them ; and betwixt the two leffer Fluxes. the Water will ascend unto the least Altitude. Thus, in the Space of 24 Lunar Hours, the Water will come not twice, as it is in other Places. but once only unto its greateft Altitude, and once unto its leaft; and the greatest Altitude, when the Moon declines to the Pole which is above the Horizon of the Place, will fall fix Hours after the Appulse of the Moon to the Meridian of the Place; and the Moon changing its Declination, it will be chang'd into a Reflux. Therefore one Tide coming in the Space of 12 Hours from the Indian Ocean, and the other in the Space of 6 Hours from the Chinele Ocean through those Streights respectively, which were before mentioned; and fo falling one at the third, and the other at the ninth Lunar Hour, feem to make those anomalous Tides. But these and fuch like particular Phanomena are every where to be left to the Observations of the neighbouring Shores and Seas.

Scholium. If we would decline the Intricacy and Tediousness of our Author's Calculation, and define only to know the Quantities of the Forces, they they

they are thus: The Sum of the Sun's Forces, as well in depreffing the Waters in the Places which are 90 Degrees from it, as in elevating them in the Places which are under it, and those opposite thereto, if they be taken conjunctly; or the whole Force of the Sun to move the Sea is to the Force of Gravity with us, as 1 isto 12,868,200. But fince the Centrifugal Force of the Parts of . the Earth arifing from its diurnal Motion, which is to the Force of Gravity as 1 to 289, doth make that the Heighth of the Water under the Equator should exceed its Heighth under the Poles, by the Measure of \$1820 Feet of Paris: The Solar Force of which we now treat, fince it is to the Force of Gravity as 1 to 12,868,200, and confequently to that Centrifugal Force as 289 to 12,868,200, or as 1 to 44,527; it will make, that the Heighth of the Water in the Places under the Sun, and opposite therto, should exceed the Altitude of it in the Places which are 90 Degrees distant from the Sun, by the measure only of one Foot of Paris, and a little above 11 Inches; according to this Analogy 44,527: 1::85,820: 112. and 1 of an Inch. Now the Force of the Moon for the moving of the Sea, which is the principal Force, is to be deduced from the Proportion which it bears to that of the Sun, and to bediftinguish'd by the Effects or Sums of the Motions in the Syzygies, and the Differences in the Quadratures: By this Computation, the Force of the Moon is to the Force of the Sun, when Observations are compar'd together, as 4[48 to I nearly, or in a round Number almost five-fold. · Coroll. (1.) Since therefore, as we have feen before, the Sun's Force ought to elevate the Water unto the Heighth of almost two Feet, the Moon's Force, which is almost five times as great; ought to elevate the Water into the Heighth 1

Heighth of about 9 Feet; and the Lunar Force and the Sun's conjoin'd, as in the Syzygies, will elevate the fame unto near 11; and when the Sun's Force is fubftracted from the Moon's, as in the Quadratures, will elevate it about 7 Feet. Now this Force doth abundantly fuffice to caufe all the Motions of the Sea, and doth very well agree with the Quantity of the Motions defin'd above; and by answering fo well to the fame, doth plainly confirm the Truth of that Caufe of the Tides which we have affign'd.

Corol. (2.) Since the Force of the Moon to move the Sea is to the Force of Gravity, according to what hath been demonstrated before, only as 1 is to 2871400; it is manifest, that that Force is far less than to be perceived in any Experiments of Pendulums, or in any Static, or Hydrostatic Experiments whatsoever. This Force can have a fensible Effect in the Sea only.

Corol. (2.) Forasmuch as the Force of the Moon to move the Sea, is to the Sun's Force upon the fame as near  $\varsigma$  to I; and those Forces are as the Densities of the Bodies, or the Quantities of Matter contain'd in equal Space, and as the Cubes of the Diftances or Diameters conjunctly; for Bodies equally dense are as the Cubes of the true Diameters directly, in respect of the fame Diftance; and the moving Forces in this Cafe are alfo as the Cubes of the Diftances reciprocally, or as the Cubes of the apparent Diameters directly; and confequently it is the fame thing whether the Sun be nearer or more remote, greater or lefs, fo that the apparent Diameter be certain and determinate: For these Reasons the Density of the Moon will be to that of the Sun, at is its Effect, or as 4[48 to 1; and as the Cube of the apparent Diameter of the Moon is to the Cube of the Sun's appa-

apparent Diameter, *i.e.* as 4L48 to I; and as 720 to 672 conjunctly =  $4L48 \times 720$  to  $100 \times 672$ , or as 67 to 32 almost; but the Density of the Sun to the Density of the Earth, is as 100 to 396. Therefore the Density of the Moon to the Density of the Earth, will be as 21 to 17, nearly, or almost as 5 to 4. Therefore the Body of the Moon is considerably Denser; and if I may use fuch an Expression, more Terressical than the Earth it felf, as we observed before, by way of Anticipation.

Coroll. (4.) From whence, fince the true Diameter of the Moon is to that of the Earth, as  $\varsigma$ to 18, or as 1 to  $3 \lfloor 6\varsigma$ ; the Mals of the Moon will be to that of the Earth, as the Cubes of those Numbers, compounded with the Proportion of Density, or as  $1 \times \varsigma$  to  $49 \times 4$ ; that is, as 1 to 40, very near.

Coroll. (5.) The Accelerating Gravity, or the Weight of equal Bodies on the Surface of the Moon, will be as the Quantity of Matter in the Moon, to the Quantity of Matter in the Earth, with the Reciprocal Duplicate Proportion of the Diftance from the Centers compounded; that is, as  $1 \times 13$  is to  $40 \times 1$ , or near a third Part of the Accelerating Gravity on the Surface of the Earth, as we noted formerly by way of Anticipation.

CII. The Figure of the Body of the Moon (abstracting from the Elevation of the Equatoreal, and the Depression of the Polar Parts, depending upon the Diurnal Motion,) is something Oval, or that of an Oblong Spheroid; the greatest Axis whereof produced, passed thro the Center of the Earth; and exceeds the Lesser Axes Perpendicular to the same by the Excels of about 187 Feet. If then the Body of the Moon were Fluid like our Sea, the Force of the Earth to elevate

up that Fluid in the hither and opposite Parts, would be to the Force of the Moon upon our Sea. as the attracting Force of the Earth is to that of the Moon ; Or as the Quantity of Matter in the Earth is to that in the Moon, by reason of the equal Diftances; if the Leffer Diameter of the Moon did not change that Proportion. That whole Force therefore from the Composition of those Proportions, will be in a Proportion compounded of 40 to 1, and 1 to 3165; or as 40 × 1 is to  $1 \times 3 \lfloor 65$ ; that is, as 40 is to  $3 \lfloor 65$ . From From whence, fince by what was before demonftrated, our Sea is lifted up about 9 Feet by the Force of the Moon, the Lunar Fluid ought to be lifted up about 92 Feet by the Force of the Earth. And for this Caule the Moon is of a Spheroidal Figure ; the greater Axis whereof being produced, would pafs thro' the Center of the Earth, and exceed the Diameters or Perpendicular Axes by about 187 Feet.

Corollary. And from thence perchance it is, that the fame Face of the Moon is turn'd more directly to the Earth than otherwife it would be. For the Moon cannot Reft in another Situation, but by Librating to and fro, will always return to this Situation. Neverthelefs, the Librations, by reafon of the Smallnefs of the Force in fuch a fmall Excefs of the Greater Axis above the Leffer ones, will be exceeding Slow; fo that the Face which ought always to look to the Earth, may look to the Other Focus of the Lunar Orbit, by reafon of the Equability of the Angular Motion about it, as was explain'd before, and not prefently be drawn back from thence and turned to the Earth,

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CIII. Co-

CIII. Comets are higher than the Moon; and are moved in the Region of the Primary Planets.

CIV. Comets are mov'd in Conic Sections, having their Focus in the Center of the Sun; and by Rays drawn to the Sun, describe equal Area's in equal Times, and in general Area's proportional to the Times.

CV. The Bodies of Comets are Solid, Compact, Fixed, and Durable, like the Bodies of the Planets; and they are commonly encompass'd with huge Atmospheres; and do always acquire Tails from their Neighbourhood to the Sun; but these fometimes longer, and fometimes fhorter.

These Propositions contain our famous Author's Cometography, so far as concerns our prefent Purpose.

Now they are propounded fo clearly and fully by our Author himfelf, that they in no wife need our Explication. Wherefore what follows, we fhall take Word for Word out of him.

Novemb. 15. 1708.

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## LECT, XXXVII.

T HAT Connets are higher than the Moon, and fhew themfelves to us in the Region of the Planets.

As the Want of a Diarnal Parallax hath raifed Comets above the Sublunary Regidns: fo their Descent into the Planetary Regions, is argued from their Annual Parallax. For those Comers which go forwards according to the Order of the Signs, are all of them, about the Time of their Difappearance, either Slower than ufual, or Restograde, if the Earth be betwixt them and the Sun; but fwifter than Ordinary, if the Earth, tends towards the Opposition. And on the Contrary, they which go contrary to the Order of the Signs, are fwifter than usual at the Time of their Disappearance, if the Eatth be betwixt them and . the Sun ; and flower than ordinary, or Retrograde, if the Earth be placed on the Contrary fide. This happens especially from the Motion of the Earth in its various Situation ; and it is here, as it is in the Planets, which according as the Motion of the Earth confpires with them, or is contrary to them, are fometimes Retrograde, fometimes feem to be mov'd more flowly, fometimes more fwiftly. If the Earth goes to the fame Part with the Comet, and is carried about the Sun with an Angula Motion more fwiftly, the Comet as feen from the Earth, by reason of its flower Motion, appears Retrograde; but if the Earth be carried more

more flowly, the Motion of the Comet ( that of the Earth being substracted ) becomes at least nower. And if the Earth becarried to the contrary Part, the Comet from thence appears fwifter. Now from this Acceleration, or Retardation, or Retrograde Motion, the Diftance of the Comet is thus Collected.

In Fig. 2. Plate 9. let YQA, YQB, YQC be Three observ'd Longitudes of the Comet about the Time of the Beginning of its Motion; and let  $\Upsilon Q F$  be the Longitude last observ'd, when the Comet begins to difappear. Let the Right Line A B C be drawn, the Parts whereof A B, and BC, which lie betwixt the Right Lines QA. and Q B, Q B, and QC, are one to the other, as the Times betwixt the Three first Observations. Let AC be produc'd to G, that AG may be to A B, as the Time betwixt the first Observation of ATUAP and the last, is to the Time betwixt the sirft Obfervation and the fecond; and let QG be join'd. Now if the Comet were mov'd uniformly in a Right Line, and the Earth either refted or went forward uniformly in a Right Line, the Angle YQG would be the Longitude of the Comet at the Time of the last Observation; the Angle FQG therefore, which is the Difference of the, Longitudes, arifes from the Inequality of the Motions of the Comet and the Earth. But this Angle, if the Earth and the Comet be moved to contrary Parts, is added to the Angle A Q G, and, fo renders the apparent Motion of the Comet fwifter; but if the Comet goes to the fame Part, with the Earth, it is substracted from the fame, and renders the Motion of the Comet either Slower, or perhaps Retrograde, as was faid a-This ; bove. • . .

This Angle therefore arifes chiefly from the Motion of the Earth, and is juftly to be reckoned for the Parallax of the Comet 1 fome Increase or Decrease of it, to wit, which may atile from the Comet's uneven Motion in its own Orb, being here neglected. But the Distance of the Comet is thus Collected from the Parallax.

In Fig. 4. Plate 9. let S represent the Sun, acT the Orbis Magnus, a the Place of the Earth in the first Observation, e the Place of the Earth in the Gecond Observation, T the Place thereof in the aft Obfervation, and Tr a Right Line drawn towards the Beginning of Aries. Let the Angle **TV** be taken equal to the Angle YQF; that is, equal to the Longitude of the Comet when the Earth is in T. Let at be join'd, and drawn out to g, that ag may be to ac, as AG is to AC, and g will be the Place which the Earth would reach unto at the Time of the last Observation, its Motion being continued uniformly in the Right Line ac; and therefore if  $g \gamma$  be drawn Parallel to T  $\gamma$ , and the Angle  $\gamma g V$  be taken equal to  $\Upsilon QG$ , this Angle  $\Upsilon gV$  will be equal to the Longitude of the Comet seen from the Place g, and the Angle TVg will be the Parallax which arifeth from the Transferring of the Earth out of the Place g into the Place  $\mathbf{T}$ ; and confequently V will be the Place of the Comet in the Plane of the Ecliptic. Now this Place V is wont to be below the Orb of Jupiter.

The fame Thing is Collected from the Curvature of the Way of Comets. These Bodies go forward almost in great Circles, fo long as they are mov'd more swiftly; but in the End of their Course, when that Part of the apparent Motion which ariseth from the Parallax bears a greater Proportion to the whole apparent Motion, they

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they are wont to decline from these Circles; and as oft as the Earth is mov'd to one Part, to be carried to the contrary. This Deflexion arises from the Parallax, because that it answers to the Motion of the Earth; and the notable Quantity of it hath by my Computation placed Comets when they disappear far enough below the Orb of Jupiter. From whence it follows, that in their Perigees and Perihelia, at which Times they are nearer, they descend oftentimes below the Orbs of Mars and the inferior Planets.

The Nearness of Comets is also confirm'd from the Light of their Heads. For the Splendor of a Celeftial Body which is illuminated by the Sun. and goes off into far diftant Regions, is diminish'd in the Quadruplicate Proportion of the Diffance : i.e. in one Duplicate Proportion, by reason of the increase of the Distance from the Sun; and alfo in another Duplicate Proportion, by reason of the Diminution of the apparent Diameter. From whence, if both the Quantity of Light, and the apparent Diameter of the Comet be given, the Diftance also will be given, by faying that the Distance is to the Distance of a Planet in the entire Proportion of Diameter to Diameter directly, and in the subduplicate Proportion of Light to Light inverfly." Thus the leaft Diameter of the Capillitium of the Comet of the Year 1682, being observ'd by the Famous Mr. Flamsteed thro' a Telescope of 16 Feet, and measur'd by a Micrometer, was 2'. o". But the Nucleas, or Star it felf, had scarce the roth Part of this Breadth, as being only 11 or 12" over. But in the Light and Clearnels of the Head, it exceeded the Head of the Comet of 1680, and even came near to Stars of the first and second Magnitude. Let us suppose Saturn with his Ring to be about 4 Times brighter Cc than

than that Star; and because the Light of the Ring doth almost equal the Light of the Intermediate Globe, and the apparent Diamerer of the Globe is about 21"; and confequently the Light of the Ring and Globe together doth almost equal the Light of a Globe, the Diameter whereof is 20"; the Distance of the Comet will be to the Diftance of Saturn, as 1 to V 4 inversity, and 12" to 20" directly; i.e. as 24 to 20, or 4 to 5. Again, the Comet of the Year 1665, in the Month of April, as Hevelius writes, did in Clearnels exceed almost all the Fixed Stars; yea, and Saturn it felf in regard of the Colour, which was far more lively. For the Comet was more lucid than that other which had appear'd in the End of the foregoing Year, and was to be compared with Stars of the first Magnitude. The Breadth of the Capillitium was about 6'; but the Nucleus, as compared with the Planets, by means of a Telescope, was plainly lefs than Jupiter; and fometimes was judged to be lefs than the intermediate Body of Saturn, fometimes equal thereto. Moreover, feeing the Diameter of the Capillitium of Comets doth feldom exceed 8' or 12", and the Diameter of the Nucleus or Central Star is about a 10th, or perhaps a 15th Part of the Diameter of the Capillitium ; It is manifest that these Stars are for the most part of the same apparent Magnitude with the Planets. From whence, fince their Light may oftentimes be compar'd with that of Saturn, and sometimes doth exceed it; It is manifest all the Comets in their Perihelia are placed either beneath Saturn, or not far above it. They are widely mistaken therefore, who Remove these Stars into the Region of the Fixed Stars; where certainly they could no more be Illuminated by our Sun, than the Planets which are here, are Illuminated by the Fixed Stars. We

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We have, whilft we have been Reafoning hitherto, not confidered the Obscuration of Comets by that very copious and groß Fume, wherewith the Head is encompass'd, which always shines as through a Cloud. For the more obscure the Body is rendred by this Fume, fo much the nearer it must needs approach to the Sun, that it may equal the Planets in the Plenty of Light reflected from it. From thence it feems probable. that Comets descend far below the Sphere of Saturn, as we have prov'd from their Parallax. But the fame Thing is especially confirm'd from their Tails. These arise either from the Reflexion of the Fume dispers'd through the Ether, or from the Light of the Head. In the former Cafe the Diftance of the Comets is to be diminish'd, left the Fume which always ariseth from the Head. should be propagated with an incredible Velocity and Expansion through too large Spaces. In the latter all the Light, as well of the Tail as of the Capillitium, is to be referr'd to the Nucleus of the Head. Therefore if we imagine all this Light gathered together within the Dilque of the Nucleus, the Nucleus would now certainly, as often as it fends forth a very great and fhining Tail, much exceed Jupiter it felf in Splendor. Since therefore it emits more Light, notwithstanding that it hath a much lefs Diameter, it must be much more Illuminated by the Sun, and confequently much nearer to the Sun. Moreover, the Heads which lie hid under the Sun, and do ar that Time put forth very great and resplendent Tails, like Beams of Fire, as sometimes hath been feen, ought by the fame Argument to be placed beneath the Sphere of Venus. For all that Light, if it be supposed to be gathered together into the Star, would sometimes exceed Venus it Cc2 felf,

felf, that I may not fay many Venus's in Splendor.

Laftly, The fame Thing is gathered from the Increase of the Light of the Head in the Recess of Comets from the Earth towards the Sun, and Decrease of the same in their Recess from the Sun towards the Earth. For the latter Comet of 1665, (as Hevelius observ'd) from what Time it begun to be feen; remitted always of its Motion, and confequently had pass'd the Perigee : but the Splendor of the Head daily increas'd, until the Comet, cover'd with the Solar Rays, ceased to appear. The Comet of Year 1682, as the fame Hevelius observ'd, in the End of the Month July, when it was first feen, mov'd very slowly, making about 40' or 45' every Day in its Orb. From that Time its Motion was continually increas'd until Septemb. 4. at which Time it arofe to about 5 Degrees. In this whole Time therefore it approach'd to the Earth. Which is also collected from the Diameter of the Head, meafur'd by a Micrometer : which Hevelius found Aug. 6. to be only 6'. 5", the Hair included ; whereas Sept. 2. it was 9'. 7". The Head therefore, in the Begining of its Motion, appear'd far lefs than in the End ; but nevertheless in the Beginning, in the Neighbourhood of the Sun, it was much brighter than about the End, as Hevelius relates. Therefore in all this Time, by reason of its Departure from the Sun, it decreas'd as to Light, notwithflanding its Access towards the Earth.

The Comet of the Year 1618, about the Middle of December; and that of 1680, about the End of the fame Month, mov'd most fwiftly; and confequently they were then in their Perigees: But the greatest Splendor of the Heads happened almost a Fortnight before, when they had just got

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out of the Rays of the Sun; and the greatest Splendor of the Tails was a little before, when they were nearer to the Sun. The Head of the former Comet, according to the Observation of Cylatus, Decemb. 1. seem'd greater than a Star of the first Magnitude; and Decemb. 16. ( it being now in its Perigee, ) it had fallen off a little in its Magnitude, but very much in its Splendor or Clearnels of Light. Jan. 7. Kepler being uncertain of the Head, made an End of Observing. December 12. the Head of the latter Comet was feen and observ'd by Mr. Flamsteed, in the Difance of 9 Degrees from the Sun; which a Star of the 2d Magnitude fcarce could have been. Decemb. 15. and 17. the fame appear'd as a Star of the 2d Magnitude, being at the fame time rendred lefs confpicuous by the Brightnefs of the Clouds which were about the Sun-fetting. Decemb. 26. It ing mov'd very swiftly, and being then almost in its Perigee, it was lefs than the Star call'd Os Pegasi, one of the third Magnitude. Jan. 2. It appear'd as a Star of the fourth Magnitude. Jan. 9. as one of the 5th. Fan. 12. by reason of the Splendor of the Moon increasing, it disappear'd. Jan. 25. It scarce equall'd a Star of the seventh Magnitude. According to this, if we take equal Times from the Perigee on this Side and on that, the Head, which being placed in Regions very remote, ought, by reason of equal Distances from the Earth, to have fhone equally; did on the Part betwixt the Perigee and the Sun, fhine most of all, on the other Part vanish'd out of fight. Therefore from the great Difference of · Light in these two Situations, the great Vicinity of the Comet to the Sun, in the former Situation, is rightly concluded. For the Light of Comets is wont to be regular, and to appear the greatest of all, C c 2

all, when the Heads are moved most fwiftly, and confequently are in their Perigees; only fo far as it becomes greater in the Neighbourhood of the Sun.

Coroll. (1.) Comets therefore shine by the Reflexion of the Light of the Sun.

Coroll. (2.) We may gather from what hath been faid, why Comets do so much frequent the Region of the Sun. If they were feen in Regions far beyond Saturn, they ought always to appear in the opposite Part to the Sun. For those which were in this Part would be nearer to the Earth, and the Sun by its Interpolition would obfcure the reft. But in running over the Hiftories of Comets, I have found that four or five Times more have been discover'd in the Hemisphere that is towards the Sun, than in the Opposite, befides others, no doubt not a few, which the Light of the Sun hath wholly hidden. The Thing is this; In their Descent to our Regions, they neither fend forth Tails, nor are fo illustrated by the Sun as to be feen by the naked Eye, until they have descended beneath Jupiter. Now the far greater Part of the Space described in so small an Interval about the Sun, is on that fide of the Earth which looks to the Sun; and in that greater Part thefe Stars, as being then, for the most Part, nearer to the Sun, are wont to be enlightned.

Coroll. (3.) From hence it is manifest that the Heavens are destitute of Resistance. For Comets taking oblique Ways, and sometimes contrary to the Course of the Planets, are mov'd every Way most freely, and hold their Motions for a long time, contrary to the Course of the Planets. I am mistaken, if they be not a kind of Planets; and which being in perpetual Motion, return in a Round. For whereas some Writers will have them

to be Meteors, taking their Argument from the perperual Changes of the Heads of them, This feems to want all Foundation. Their Heads are encompass'd with huge Atmospheres, and the Atmospheres ought to be more dense beneath. From whence it comes to pass, that it is not the Bodies themselves of the Comets, but Clouds about them. which those Mutations are seen in. Thus, if the Earth were feen from the Planets, it would fhine without doubt by the Light of its Clouds, and the firm Body would almost lie hid under those Clouds. Thus the Girdles of Fupiter, which are form'd in the Clouds of that Planet, change their Situation amongst themselves; so that the firm Body of Jupiter is difficultly difcern'd thro' those Clouds. And much more ought the Bodies of Comets to be hid under their Atmospheres, which are both more deep and more craffe.

Now to him that Revolves in his Mind the Orb of the Comet of 1680, and 168<sup>°</sup>, and the reft of the Phanomena it will be easily manifest, that the Bodies of Comets are Solid, Compact, Firm and Durable, like the Bodies of the Planets. For if they were nothing else than Vapours or Exhalations of the Sun, Earth and Planets, this Comet ought in its Transit through the Neighbourhood of the Sun to have been immediately diffipated. For the Heat of the Sun is as the Density of the Rays; that is, reciprocally, as the Square of the Diftance from the Sun. And therefore, fince the Diftance of the Comet from the Sun, Decemb. 8. at which Time it was in its Perihelion, was to the Diftance of the Earth from the Sun, as 6 to 1000, or thereabouts; the Heat of the Sun upon the Comet at that time was to the Heat of the Summer-Sun with us, as 1,000,000 is to 26: or as 28,000 is to 1. But the Heat of boiling C C 4 Water

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Water is about three times greater than that Heat which the dry Earth conceives from the Summer-Sun, as I have try'd my felf; and the Heat of Redhot Iron (if I guess right) is about three or four times greater than the Heat of boiling Water; and confequently the Heat which the dry Earth in the Comet contracted from the Rays of the Sun when it was in its Perihelion, might be about 2000 times greater than that of Red-hot Iron. Now by fo great an Heat as this, Vapours and Exhalations, and all Volatile Matter must have been prefently confum'd and diffipated.

The Comet therefore, in its Perihelion, contracted an immense Heat from the Sun, and would hold that Heat for a very long time. For a Globe of Red-hot Iron of one Inch Diameter, would scarce lose all its Heat in one Hour's Space, if it were expos'd to the open Air. But a greater Globe would keep its Heat longer, and this in the proportion of its Diameter; because the Surface (according to which it is cool'd by the Contact of the Ambient Air) is in that Proportion lefs, if compar'd with the quantity of the hot Matter included. And therefore a Globe of Red-hot Iron, of the Bigness of this Earth, i.e. of 40,000,000 Feet Diameter, would fcarce be wholly cool'd in fo many Davs, or 50,000 Years. I suspect nevertheless, that the Continuance of Heat, by reason of latent Causes, is increas'd in less Proportion than that, of the Diameter; and wish that the true Proportion were fearched out by Experiments.

Furthermore, it is to be noted, that the Comet in the Month of December, when it was fo heated by the Sun, fent forth a Tail far greater, and more refplendent than it had done before in November, when it had not yet reached to its Peri-

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Perihelion. And in general, all the Tails which exceed in Magnitude and Brightnefs are then feen, when the Star hath lately pafs'd through the Region of the Sun. The Heating therefore of the Comet conduceth to the Magnitude of the Tail. And from thence I am apt to conclude, that the Tail is nothing elfe but a most thin Vapour, which the Head or Nucleus of the Comet emits through its Heat.

Now there is a threefold Opinion concerning the Tails of Comets; that they are either the Beams of the Sun propagated thro' the Translucid Bodies of those Stars; or arise from the Refraction of the Light in the Progress thereof, from the Head of the Comet to the Earth; or, laftly, are a Cloud or Vapour arising continually from the Head of the Comet, and which is turn'd off to the Part opposite to the Sun. The first is the Opinion of those who are not yet instructed in the Knowledge of Optics. For the Beams of the Sun, let into a dark Chamber, are not feen there, any further than the Light is reflected from the Particles of Duft and Fumes floating in the Air; and confequently are much more bright in the Air when - ftuffed with groß Fumes, and ftrike the Senfe more forcibly; in a thinner Air are less perceiv'd, and in the Heavens, where there is no reflecting Matter, are not to be perceiv'd at all. Light is not seen as it is in the Beam, but as it is from thence reflected to our Eyes. For Sight is only by Rays which fall upon the Eyes. Some reflecting Matter therefore is requir'd in that Part of Heaven which the Tail takes up; Otherwife the whole Heaven it felf, which is illuminated by the Sun, would thine uniformly. The Second Opinion is urged with many Difficulties. The Tails are never varied with Colours, which yet are infeparable

feparable Concomitants of Refractions. The Light of the Fixed Stars and Planets, which is diffinctly transmitted to us, shews that the Celestial Medium is without any refractive Force. For as for what is faid, that the Fixed Stars have sometimes been seen with bright Streams by the Egyptians; this, because it is a thing which happens very feldom, is to be afcribed to the accidental Refraction of the Clouds. The Radiation also, and twinkling of the Fixed Stars, is to be referr'd to Refractions both in the Eyes and the tremulous Air; because, when they are seen through a Telescope, it vanisheth away.

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By the Tremor of the Air, and alcending Vapours it comes to pass, that the Rays are easily turn'd off by Turns from the narrow Space of the Pupil of the Eye; but from the wider Aperture of the Object Glass never. From hence it is, that in the former Cafe a twinkling is produc'd, but in the latter none at all; and the Absence of it in the latter Cafe demonstrates the regular Transmission of Light through the Heavens, without any fensible Refraction. And if any one should fay in this Place, that Tails are not wont to be feen in the Fixed Stars, only because their Light is weak and feeble; fo that their Secondary Rays have not Force enough to move the Eyes, that Tails fhould appear about them; He may take notice, that the Light of Fixed Stars may be increas'd by Telescopes above an Hundred Times, and yet no Tails are feen. The Light of Planets alfo is greater than that of Comets, but yet they have no Tails; yez, and Comets have the longest Tails, when the Light of their Heads is weak and very obtule. For the Comet of 1680, in the Month of December, at what time its Head did scarce equal a Star of the Second Magnitude, fent forth 2 Tail of notable Splendor unto 40, 50, 60 Degrees

Degrees of Length, and more. Afterwards, Jan. 27, and 28, the Head appear'd as a Star of the Seventh Magnitude only; but the Tail, which was of a dim Light indeed, but fensible enough, was feven or eight Degrees long, and with a very obfcure Light, which could fcarce be perceiv'd, it was ftretch'd forth 12 Degrees, and more, as was faid above. And Febr. 9, and 10, when the Head could no longer be feen with the naked Eye, I faw a Tail two Degrees long through a Telescope. Furthermore, if the Tail proceeded from the Refraction of Celestial Matter, and turn'd alide from the Opposition to the Sun, according to the Figure of the Heavens; that Deflection ought in the fame Regions of Heaven always to be to one Part. But the Comet of 1680, Decemb. the 28th. at 81 a Clock in the Evening at London, was in ¥ 8 deg. AI', with a Northern Latitude 28 deg. 6'; the Sun being in V3 18 deg. 26'. And the Comet of the Year 1557, Decemb. 29, was in X 8 deg. 41'. with Northern Latitude 28 deg. 40'; the Sun also being in v3 about 18°. 26. In both Cafes, the Earth was in the fame Place, and the Comet appear'd in the same Part of Heaven; yet notwithstanding in the former Cafe (by mine own and others Obfervations) the Tail of the Comet declin'd with an Angle of 4 Degrees from Opposition to the Sun towards the North; whereas in the latter (according to Tycho's Observation ) the Declination was 21 Degrees towards the South. The Refraction therefore of the Heavens being taken away, it remains that the Phanomena of the Tails are deriv'd from some reflecting Matter.

Now that the Tails do proceed from the Heads, and do alcend towards the Region oppolite to the Sun, is confirm'd from the Laws which they observe. As that lying in the Plains of the Orbs of

of the Comets paffing through the Sun, they decline from Opposition to the Sun, always unto that Part which the Heads going forward in those Orbs do leave behind them. That they appear to the Spectator when placed in these Plains, turn'd away directly from the Sun; but the Spe-Aator going aside from these Plains, the Deviation is perceiv'd by degrees, and grows every Day greater. That the Deviation, cateris paribus, is lefs when the Tail is more oblique to the Orb of the Comet, as also when the Head of the Comet approacheth nearer to the Sun; especially if the Angle of Deviation be taken at the Head of the Comet. Besides, that the Tails which deviate not, are ftraight; but those which do, are bowed. That the Curvature is greater, where the Deviation is greater, and more fensible when the Tail. cæteris paribus, is longer; for in fhort Ones the Curvature is hardly perceiv'd. That the Angle of Deviation is less at the Head of the Comet, greater at the other End of the Tail; and confequently, that the Tail on its Convex-fide looks to that Part from which the Declination is, and which is in a Right Line drawn from the Sun through the Head of the Comet in infinitum. And that the Tails which are longer and broader, and shine with a more lively Light, are upon their Convex-fides fomething more fplendid, and bounded on the Concave-fide with a Limit not very distinct.

The Phænomena therefore of the Tail depend upon the Motion of the Head, and not upon that Region of Heaven in which the Head is feen; and therefore are not made by a Refraction of the Heavens, but arife from the Head, which affords Reflecting Matter. For as in our Air the Smoke of a Body fet on Fire afcends upwards, and that either Perpendicularly if the Body refteth;

eth : or Obliquely, if it be mov'd to one Side : So in the Heavens, where Bodies gravitate to the Sun. the Fumes and Vapours ought to alcend from the Sun (as hath been already faid) and this in a Right Line if the Body refts, but Obliquely, if the Body be mov'd, and in going forwards always forfakes the Places from which the upper-Parts of the Vapour had afcended. And that Obliquity will be lefs, where the Afcent of the Va-pour is more quick; as in the Neighbourhood of the Sun, and near to the Surface of the fmoking Body. Now from the Diverfity of the Obliquity, the Column of the Vapour will be bowed. And because the Vapour on that fide of the Column which goes before is fomething more fresh, for this Reason it will be something more Denfe in the fame Place, and will therefore reflect Light more copioufly, and be bounded with more diffinct Limit. As to the fudden and uncertain Agitations of their Tails, and of the irregular Figures of the fame, which fome defcribe, I add nothing here; because they either arise from the Mutations that are in our Air, and the Motions of the Clouds which do in fome part obscure the Tail; or perchance from Parts of the Via Lastea, which may feem united with the Tails that pafs by them, and taken to be Parts of them.

Now that Vapours which fuffice to fill fuch vaft Spaces may arife from the Atmospheres of Comets, will be understood from the Rarity of our Air. For the Air, near to the Surface of the Air, possessing the Air, near to the Surface of the Air, possessing the fame weight; and therefore a Cylindrical Column of Air 850 Feet high, is of the fame Weight with a Column of Water of the fame Breadth, which is only one Foot high. But a Column

a Column of Air arising to the Top of the Atmosphere, doth equal in its Weight a Column of Water, which is about 22 Foot high ; and therefore if the inferior Part of the whole airy Column be taken away, which is of the Heighth of 8co Feet, the upper remaining Part will be equal in its Weight to a Column of Water of 22 Feet high. But from thence (according to the Hypothefis which hath been confirm'd by many Experiments, namely, that the Compression of the Air is as the Weight of the Atmosphere which lies upon it, and that Gravity is reciprocally, as the Square of the Diftance from the Center of the Earth) by making Computation according to Coroll. Prop. XXII. Book II. I found that Air, at the Heighth of a Semi-diameter of the Earth. from the Surface of the fame, is rarer than it is with us in a Proportion far greater than is that of all the Space beneath the Orb of Saturn to a Globe of one Inch Diameter. And therefore a Globe of our Air of one Inch Diameter in that Rarity, which it would have in the Heighth of a Semidiameter of the Earth, would fill all the Regions of the Planets unto the Sphere of Saturn, and much farther. Therefore fince Air, which is yet higher, grows more rare infinitely; and the Armosphere of a Comet, in alcending from its Center, is about ten times higher than the Surface of the Nucleus, and when the Tail doth afcend yet higher, the Tail ought to be rare in the highest degree. And altho' by reason of the more gross Atmosphere of Comets, and the great Gravitation of Bodies towards the Sun; and the Gravitation of the Particles of the Air and Vapours towards one another; it may come to pass, that the Air doth not grow Rare in fo great a Degree in the Heavenly

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venly Spaces and in the Tails of Comets; yet is it manifest from this Computation, that a very fmall Quantity of Air and Vapours fufficeth abundantly to all those Phænomena of the Tails. For the great Rarity of the Tails is also collected from the Stars which are visible through them. The Atmosphere of the Earth shining with the Light of the Sun, doth by its Craffitude, which is but for a few Miles, [frequently] both obscure, and even hide all the Stars and the Moon it felf; but the least Stars are observ'd to be visible without any detriment of their Clearness through the Tails of Comets, which are of an exceeding great Depth, and are at the fame time illuminated with the Light of the Sun. Nor is the Splendor of many of their Tails wont to be greater, than is that of our Air. when in a dark Chamber it reflects the Light of the Sun in the Form of a Beam, for the Breadth of an Inch or two.

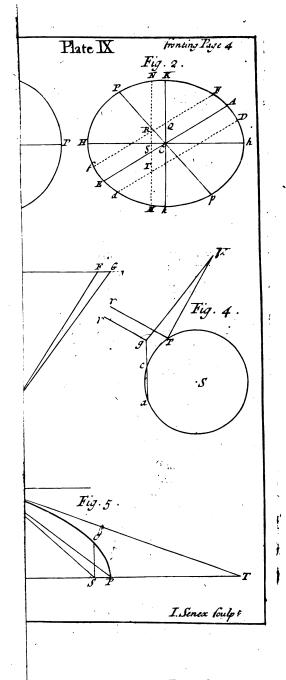
In what Time the Vapour afcends from the Head to the End of the Tail, may almost be known by drawing a Right Line from the End of the Tail to the Sun, and noting the Place where that Line cuts the Trajectory. For a Vapour in the End of the Tail, if it alcends straight from the Sun, begun to afcend from the Head, at what Time the Head was in the Place of Intersection. And tho' it doth not afcend ftraight from the Sun; yet by retaining that Motion which the Comet had before its Alcenfion, and Compounding the lame with the Motion of its Afcenfion, it afcends Obliquely. From whence the truer Solution of the Problem will be, that the Right Line which cuts the Orb, be Parallel to the Length of the Tail; or rather (by reafon of the Curvi-linear Motion of the Comet) that the fame deflect a little from the Line of the Tail.

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By this means I found, that the Vapour which was in the End of the Tail, Jan. 25. began to a-fcend from the Head, Decemb. the 11th, and confequently had spent above 45 Days in its Ascenfion. But all that Tail which appear'd Dec. 10. had afcended in the Space of those two Days which had pass'd from the Time of the Perihelion. Therefore the Vapour ascended most fwiftly in the Neighbourhood of the Sun, and afterwards went on afcending with a Motion still retarded by its own Gravity; and by its ascending increas'd the Length of the Tail; but the Tail, fo long as it appear'd, confifted almost all of the Vapour which had ascended from the Time of the Perihelion; and the Vapour which ascended first, and composed the End of the Tail, vanished not out of fight before that it ceafed to appear, by reafon of its great Distance both from the Sun which illustrated it, and from our Eyes. From whence alfo the Tails of other Comets which are short. do not afcend with a fwift and perpetual Motion from the Heads, and fo prefently vanish away, but are permanent Columns of Vapours, propagated from the Heads with a very flow Motion of many Days; which by partaking of that Motion of the Heads which they had at the Beginning, go on to be mov'd through the Heavens, together with the Heads. And from hence it is again collected, that the Heavenly Spaces are defitute of all Force of Refifting; fince not only the Solid Bodies of Planets and Comets, but also the exceeding thin and rare Vapours of the Tails of Comets do perform their Motions in them most freely and fwiftly; and hold the fame for a very long Time.

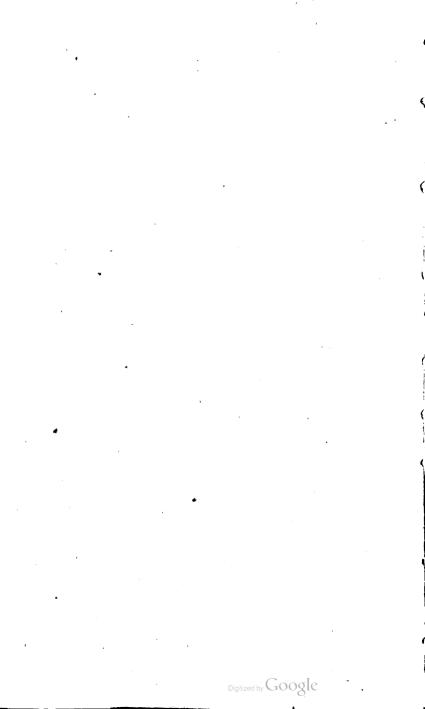
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Kepler



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Kepler ascribes the Ascent of the Tails from the Atmospheres of the Head, and their Progress to the Part opposite to the Sun, to the Action of the Rays of Light, which carries away with them that Matter of which the Tail confifts. And it is not altogether unreasonable to think, that a very thin Air may give way to the Rays of Light in free Spaces; notwithstanding that groß Substances cannot in our Regions be fenfibly impell'd or mov'd by the Solar Rays. Another is of Opinion, that there may be as well Light as heavy Particles of Matter; and that the Matter of the Tails are light, and by their Lightness ascend from the But fince the Gravity in Terrestrial Bodies Sun. is as the Matter in those Bodies, and confequently cannot be increas'd or diminish'd, the same Quantity of Matter remaining; I am prone to think, that that Afcent doth rather arife from the Rarefaction of the Matter of the Tails. The Smoke in a Chimney ascends by the Impulse of the Air in which it floats. That Air being rarified by the Heat, alcends by reason of the Diminution of its Specific Gravity, and carries away the Smoke entangled in it, together with it. What should then hinder, but that the Tail of a Comet should ascend in the fame manner from the Sun? For the Rays of the Sun do not agitate the Mediums through which they pass, but in their Reflection and Refraction. The Reflecting Particles being heated by that Action, will heat the Ethereal Air which is about them.

This, by the Heat communicated to it, will grow rare; and by reason of the Diminution of its Specific Gravity, by that Rareness will ascend, and D d carry

carry away the Reflecting Particles, of which the Tail confifts. It conduces alfo: to the Afcent of the Vapours, that they are turn'd about the Sunand by that Action endeavour to Recede from the Sun; whilft the Atmosphere of the Sun, and the Matter of the Heavens, do either wholly reft, or are sturn'd round flowly, only with a Motion, which they have receiv'd from the Rotation of the Sun. These are the Causes of the Ascent of the Tails in the Neighbourhood of the Sun where the Orbs are more Curve, and the Comets are within the more Denfe, and by that means the more heavy Atmosphere of the Sun. For the Tails which arife then, will, by keeping their Motion, and Gravitating in the mean while towards the Sun, be mov'd about the Sun in Ellipses after the manner of the Heads, and by that Motion always accompany the Heads, and flick to them. For the Gravity of Vapours towards the Sun, will no more caufe that the Tails should fall from the Heads afterwards towards the Sun, than the Gravity of the Heads can make, that they fhould fall from the Tails. Therefore by their common Gravity they will either fall together towards the Sun, or be retarded together in their Afcent ; and confequently that Gravity hinders not, but that the Tails and Heads should most easily receive, and afterwards most freely keep any Polition to one another, whatfoever it is, which they may receive from the Caufes which have been mention'd, or any other whatever.

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The Tails therefore which arife in the Perihelia of Comets, will go away with their Heads into far diftant Regions; and after a long Series of Years, return with the fame to us; or rather being

being there ratified, will by little and little vanifh away. For afterwards, in the Descent of the Heads to the Sun, new little flort Tails ought to be propagated from their Heads by a flow Motion. and afterwards to be Immenfely increas'd in the Perihelia of those Comets which defcend into the Atmosphere of the Sun. For the Vapour in those most free Spaces is perperually rarified and dilated. By which means it comes to pals, that the Tail is broader at the upper End. than near the Head of the Comet. Now it feems feafonable to think, that the Vapour being by that Rarefaction perpetually dilated, is dispersed at length through the whole Heaven, and is by little and little drawn unto the Planets, and mingled with their Atmospheres. For like as the Seas are altogether requird unto the Conflication of this our Earth; and this, that Vapours may be rais'd fufficiently out of them by the Heat of the Sun; which being gather'd together into Clouds. may fall down in Rains, and water and nourish all the Earth, for the Procreation of Vegetables; or being Condens'd in the cold Tops of Mountains, (as fome do reasonably enough suppose) runs down unto the Heads of Springs and Rivers: So to the Prefervation of Seas and Rivers in Planets, Comets feem necessary.; from the Condens'd Vapours and Exhalations whereof, what Liquor is spent by Vegetation and Corruption, and turn'd into dry Earth, is continually supplied and renew'd. For all Vegetables grow from Liquors or Juices; and then in great Part they pais by Putrefaction into dry Earth, and Mud arileth perpetually from the putrefied Liquors. Hence the Bulk of dry Earth is continually increas'd; and what is Humid, unlefs it be increas'd from elfewhere, ought per-Dd 2 petually

petually to decrease, and at length to fail. Yez, I sufpect that that Spirit which is the least indeed, but the most subtile and the best Portion of our Air, and which is required to the Life of all things, doth come from Comets especially.

The Atmospheres of Comets are in their Defcent to the Sun diminish'd by running out into Tails, and ( on that Part certainly which looks to the Sun) become more narrow: And on the other hand in their Recess from the Sun, at which time they run less out into Tails, they are enlarg'd; if so be Hevelins hath rightly noted the Phanomenons thereof. But they appear the leaft, when their Heads having been just now heated at the Sun, iffue forth in great and refulgent Tails; and their Nucleus's are perhaps furrounded with a more craffe and black Fume in the loweft Parts of their Atmosphere. For all Smoke which is ftirred up by a very great Heat, is wont to be fo much the more craffe and black. Thus the Head of the Comet of which we have been treating, did at equal Diftances from the Sun and the Earth, appear more obscure after the Perihelion than before. For in the Month of December, it was compar'd with Stars of the third Magnitude; whereas in November it equall'd Stars of the first and fecond. And they who had feen both, defcribe the former Comet as the greater. For November the 19th, this Comet, how obtuse soever the Light was where with it shined, did then equal Spice Virginz, as it appear'd to a Young man of Cambridge, and shin'd more clearly than it did afterwards. And Mr. Storer, in Letters which fell into my Hands, wrote, that the Head of it in December, at which time it fent forth the greatest and most refulgent 

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fulgent Tail of all, was fmall, and in its visible Magnitude, did fall far short of what it appear'd before the Rifing of the Sun in November. The Reason whereof he guessed to be this, That the Matter of the Comet at the beginning was more copious, and was wafted by degrees.

We may here also firly Note, that the Heads of other Comets, which have fent forth most great, and refulgent Tails, are describ'd as being somewhat dim and small. For in the Year 1668, March 5. New-Style, at 9 a-Clock in the Evening, the Reverend Father Valentine Estance, being then in Brafile, faw a Comet near to the Horizon at the South Weft, of a very little Head, and scarce to be feen, but with a Tail exceedingly Refulgent : fo that those who flood on the Shore might easily fee the Image of it reflected from the Sea. It. had the Appearance of a fhining Beam, of the Length of 22 Degrees, inclining from the West to the South, and almost Parallel to the Horizon. But this great Splendor endur'd only for two Days, and from that time notably decreas'd; and in the mean while that the Splendor decreas'd. the Tail did increase in Magnitude. From whence also it is said to have possess'd in Portugal almost a 4th Part of Heaven, ( or 45 Degrees ) ftretching from the West to the East, with a very great Splendor; nor did it all appear notwithstanding, the Head still in these Regions lying hid below the Horizon. From the Increase of the Tail, and Decrease of the Splendor it is manifest, that the Head was receding from the Sun; and was next to it about the beginning, as it was in the Comet of 1680. And we read of a like Comet in the Years 1101, or 1106; "The Star whereof was small " and Dd 2

ç. and obscure (as was that of 1680,) but the 66 Splendor which went forth from thence was æ very clear, and as it were a great Beam, reach, ec. ing to the East and North, as Hevelius hath it, " out of Simeon, the Monk of Durbam. It appear? in the Beginning of February about Evening, at the South-WVeft." But from thence, and the " Situation of the Tail, it is gathered, that the Head was near to the Sun. "It was diftant, faith Matthew of Paris, about one Cubit from the " Sun, from the third, (more truly the fixth "Hour) until the ninth, fending forth a long Ray from it felf." Such also was that fiery Comet describ'd by Aristotle, Book I. Me-teor 6. "The Head whereof was not seen the " first Day, because that it had Set before the " Sun, or at least under the Rays of the Sun; " but the following Day it was seen as far as its Situation allow'd. For it had left the Sun but Ç¢ . with a Diftance as fmall as might be; and then " Set By reason of the very great Ardor [of the "Tail,] the dispers'd Fire of the Head did not " yet appear; but afterwards, when the Tail did burn lefs, the Head was reftor'd to its former Ap-" pearance. And it extended its Splendor unto " a third Part of Heav'n, (i. e. unto 60 Degrees.) " Now its Appearance was in Winter-time, and "Afcending unto the Girdle of Orion, it there " Vanish'd.

That Comet of the Year 1618, which came with a very long Tail out of the Rays of the Sun, feem'd to equal, if not exceed a Star of the first Magnitude; but there have not a few greater Comers appear'd, which have had fhorter Tails. Some of them are related to have been equal to Jupiter Others to Venus, or even to the Moon.

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I faid before, that Comets are a kind of Planets, Revolving in very eccentric Orbits about the Sun. And like as amongft the Planets, which have no Tails, those are wont to be less which are revolv'd in lesser Orbits, and nearer to the Sun; fo likewife is it reasonable to think, that the Comets, which in their Perihetia approach nearer to the Sun, are for the most Part lesser, and are Revolv'd in lesser Orbits. But as for the Transverse Diameters of their Orbits, and their Periodical Times, we leave these Things to be determin'd from the Comparison of Comets, returning in the same Orbits after long Intervals of Time.

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#### LECT. XXXIX.

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# SYNOPSIS

OF THE

Aftronomy of COMETS.

By EDMUND HALLEY, L.L.D. Savilian Professor of Geometry at Oxford.

H E Ancient Egyptians and Chaldeans (if we may credit Diodorm Siculus) by a long Courfe of Obfervations, were faid to be able to predict the Apparitions of Comets. But fince they are also faid, by the help of the fame Arts, to have prognofticated Earthquakes and Tempests is 'tis pass all doubt, that their Knowledge in these Matters, was the Refult rather of meer Astrological Calculation, than of any Astronomical Theories of the Celeftial Motions. And the Greeks, who were the Conquerors of both those People, scarce found any other Sort of Learning amongst them, than 410

than this. So that 'tis to the Greeks themselves as the Inventors (and especially the great Hipparchus) that we owe the Aftronomy we have, and which is now improv'd to fuch a heighth. But yet, amongst the Greeks, the Opinion of Aristotle ( who wou'd have Comets to be nothing elfe, but Sublunary Vapours, or Airy Meteors) prevail'd fo far, that this most difficult part of the Afronomical Science lay altogether neglected ; for no Body thought it worth while to take notice of, or write about, the wandring uncertain Motions of what they effeemed Vapours floating in the Æther; whence it came to pais, that nothing certain, concerning the Motion of Comets, ican be found transmitted from them to us.

But Senece the Philosopher, having confider'd the Phænomena of two remarkable Comets of his Time, made no fcruple to place them amongst the Celestial Bodies; believing them to be Stars of equal Duration with the World, tho' he owns their Motions to be govern'd by Laws not as then known or found out. And at laft (which was no untrue or vain Prediction) he foretells, that there fhould be Ages sometime hereafter, to whom Time and Diligence should unfold all these Mysteries, and who, shou'd wonder how itwas possible the Ancients coa'd be ignorane of them, after fome lucky Interpreter of Naturo had thewis, in what parts of the Heaviens the Comets wander'd, subat fort of Beings and how great they were. Yet almost all the A. Aronomers differ'd from this Opinion of Seneca ; neither did Senera himfelf think fit to fet down those Phanomena of the Motion, by which he was enabled to maintain his Opinion; nor the Times of those Appearances, which might be of

#### Aftronomy of COMETS.

of use to Posterity, in order to the determining these things. And indeed, upon the turning over very many Histories of Comets. I find nothing at all that can be of service in this affair, before A. D. 1227. at which time Nicepborns Gregoras, a Constantinopolitan Historian and Astronomer, did pretty accurately describe the Path of a Comet amongst the Fix'd Stars, but was too lax as to the Account of the Time; so that this most doubtful and uncertain Comet, only deferves to be inferted in our Catalogue, for the fake of its appearing near Four-hundred Years ago.

The next of our Comets was in the Year 1472; which being the fwifteft of all, and neareft to the Earth, was observed by Regiomontanus. This Comer (Io fearful upon the account both of the Magnitude, of its Body, and the Tail) moved forty Degrees of a great Circle in the Heavens, in the Space of one Day; and was the first, of which any proper Observations are, come down to us. But all those that consider d Comets, until the Time of Txcha Brabe (that great Reftorer of Astronomy) believed, them to be below the Moon, and to took but Little Notice of them, reckoning them no other than Vapours.

But in the Year 1577, (Tyche letioully purluing the Study of the Stars, and having gotten large Inftruments for the performing Celeftial Menfurations, with far greater Care and Certainty than the Ancients could ever hope for) there appear d a very remarkable Comet; to the Observation of which Tyche vigorously applied himself; and found by many just and faithful Trials, that it had no Diurnal Parallax that was perceptible: And consequently was not only no Aerial Vapour, but also much higher than the

#### A SYNOPSIS of the

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the Moon; nay, might be plac'd amongft the Orbs of the Planets, for any thing that appear'd to the contrary; the cavilling Oppofition made by fome of the School-men in the mean time, being to no purpole.

Tycho was fucceeded by the most Sagacious Kepler. He having the Advantage of Tycho's Labours and Observations, found out the true Physical System of the World, and vastly improv'd the Science of Astronomy.

For he demonstrated that all the Planets perform their Revolutions in Elliptic Orbits, whofe Plains paß thro' the Center of the Sun, observing this Law, that the Area's of the Elliptic Sectors, taken at the Center of the Sun (which he proved to be in the common Focus of thele Elliples) are always proportional, to the Times in which the correspondent. Elliptical Arcs are describid. He discoverid also that the Diffances of the Planets from the Sun are in the Sesquialtera ratio of the Periodical Times, or (which is all one) that the Cubes of the Distances are as the Squares of the Times. This great Aftronomer had the opportunity of observing two Comets, one of which was a very remarkable one. And from the Observations of these (which afforded sufficient Indications of an Annual P2rallax) he concluded, That the Comets mov'd freely thro' the Planetary Orbs, with a Motion not much dif-ferent from a Rectilinear one; but of what kind be cou'd not precifely determine. Next, Hevelius ( a noble Emulator of Tycho Brabe ) following in Kepler's Steps, embraced the fame Hypothefis of the Rectilinear Motion of the Comets, himfelf accurately observing many of them. Yet he complain'd that his Calculations did not perfectly agree to what he observed in the Heayens: And was aware, that the Path of a Comet W AS

## Astronomy of COMETS.

was bent into a Curve Line concave towards the Sun. At length came that prodigious Comet of the Year 1680; which defcending (as it were from an infinite Diftance) perpendicularly towards the Sun, arofe from him again with as great a Velocity.

This Comet, (which was seen for four Months continually) by the very remarkable and peculiar Curvity of its Orb (above all others) gave the fittest Occasion for investigating the Theory of its Motion. And the Royal Observatories at Para and Greenwich having been for some time founded, and committed to the Care of most excellent Astronomers, the apparent Motion of this Comet was most accurately (perhaps as far as humane Skill cou'd go) observ'd by Mrs. Cassin and Flamsfreed.

Not long after, that great Geometrician the Illustrious Newton, writing his Mathematical Principles of Natural Philosophy, demonstrated not only that what Kepler had found, did necessarily obtain in the Planetary System; but also, that all the Phænomena of Comets wou'd naturally follow from the fame Principles ; which he abundantly illustrated by the Example of the aforefaid Comet of the Year 1680; fhewing at the fame time, a method of delineating the Orbits of Comets Geometrically; therein folving (not without meriting the highest Admiration of all Men) a Problem, whole Intricacy render'd it scarce Acceffible to any but himself. This Comet he prov'd to move round the Sun in a Parabolical Orb, and to describe Area's ( taken at the Center of the Sun ) proportional to the Times.



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A SYNOTSIS of the

Wherefore (following the Steps of Io Great 2 Man) I have attempted to bring the fame Method to Arithmetical Calculation; and that with all the Succefs I cou'd with. For, having collected all the Observations of Comets I cou'd; I have fram'd this following Table, the result of a prodigious deal of Calculation; which, tho' but small in Bulk, will be no unacceptable Prefent to Astronomers. For these Numbers are capable of representing all that has been yet obferved about the Motion of Comets, by the help only of the annex'd General Table; in the making of which I spar'd no Labour, that it might come forth perfect, as a Thing confectated to Posterity, and to last as long as Astronomy it felf.

36.

The Aftronomical Elements of the Motions in a Parabolic Orb of all the Comets that have been bitherto duely observ'd.

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8691	0801	1084	1083		1000	1077	1072	1005	1064	1001	1652	1018	1607	1596	1590	2585	1580	1577	1330	1532	1531	472	1337		An.	Com.
1698 27.44.15 11.46.	¥20.344	~ 28.15.	<b>W</b> 23.23.	021.10.3	1000V3 2. 2. 000.50.	11 20.49.I	<b>VS</b> 27-30-3	1118.02.	H31.14.	11.22.30.3	П28-10.	Ш16. 1.	8 20.21. 017. 2.	WI 2. I 2.3	11 1 5.30.4	0 7.42.3	J 18.57.2	P 2 5 . 5 2.	11/25-42-	П 20.27.	019.25. 017.50.	472 VS1 1.46.20	337, 1124-21.	8r. , "	Ascend.	Nodus
	_	005.48.40	1083 11223.23. 083.11. 0	10421 0 21.10.301 7.50. 0	0,00,50, 0	079.03.15	083.22.10	070.05. 0	021.18.36	032-35-50	079.28. 0	037.34. 0	017. 2. 0	30 <b>55.12.</b> 0	1 590 TK 1 5.30.4029.40.40 IT	0 7.42.30 0. 4. 0 1	00440. C	577 1 25.52. 0 74.32.45 8	1556 TX 25.42. 032. 6.30 VS	032.30. 0	017.50. 0	20 5.20. O	032.11. 0	109		Inclin.
0 1300.51.15	П 17.00.30	M 28.52. 0	1125.29.30	2.52.45	1 22.39.30	1077 11 20.49.1079.03.15 K 17.37. 5	1072 1327.30.3083.22.10 0 16.59.30	1005 IL 18.02. 070.05. 0 III 1.54.30	S(10.41.2)	1661 112230303235.50 025.58.40	079.28. 0 Y 28.18.40	T 2.14 0	2.16. 0	m 18.16. 0	11 6.54.30	8.51.		EL 9.22. 0	WS 8.50. 0	532 II 20.27. 032.30. 09021. 7. 0	1.39. 0	α	α 7.59.	gr. , ,		Peribelion.
5 69129	0 32500	51096	56020	1		20059	69739	10049	1025755	9 44851	84750		08 <i>9</i> 85	51293	57661	85 5601 O	59628	_		01605 0			0 40666	a Sole.	Periheli.	Diftan.
9.83966008.	9.511883 Sept.	9.982339 May	9.748343 July	9.7058771Sept	CV012 7.787186 Dec.	9-448 0	9.843476	9.027309	0-011044	9.651772	9-928140	9.579498	9.76849008.	9.710058	9.760882	0.038850	9.775450	9.263447	9.666424	9.706803	9.753583	9.734584 Feb.	9.609226 Fune	a Sole.	Peribelia.	Loe. Dift.
04. 8.16.57	•	-		Sept. 4.07.39	Dec. 8.00 6	Apr. 26.00.37=	9.843476 Feb. 20. 8.37	9.027309 Apr. 14 5.15=	1664 III 21.14. 021.18.30 10.41.25 102575 0.011044 Wou.24.11.52	9.651772 Jan. 16.23.41	9.928140 Nov. 2.15.40	9.579498 08. 29.12.23	08. 16. 3.50	9.710058 July 21.19.55	9.760882 Fan. 29. 2.45	0.0388 50 Sept. 27.19.20	9.775450 Nov. 28.15.00	9.263447 08. 26.18.45	9.666424 Apr. 21.20. 3	9.706803 08. 19.22.12	9.753583 Aug. 24.21.18	13	Fune 2. 6.24	d. h.		Temb. equat.
÷			87.53.30 Retrog	10	9-22-30	2 99 12. 5	109.29. o Direct		-	-	59-51-20 Direct		line -	83.56.20 Retrop			·	103.20. 0	103.8.0	30.40. 0	107-46-10	123.47.	3	<u>e</u> g	-	Perihelian
7. ofRetrog.	Direct	Direct	Retrog.	Retrog.	Direct	Retrog,	Direct	Retrog.	Retrog.	Direct	Direct	Direct	Betrog.	Retrog	Retrog	Direct	Direct	Retrog.	Direct	oDirett	Retrog	Retrog	Rerrog			

This Table needs little Explication, fince 'tis plain enough from the Titles, what the Numbers mean. Only it may be observ'd, that the Perihelium Diftances, are estimated in fuch Parts, as the Middle Diftance of the Earth from the Sun, contains 700000.

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A General Table for Calculating the Motions of Comets in a Parabolical Orbit.

1	<del></del>					
Med		Logar.		Mea		
mot.			1	mot.		pro dift.
0	gr. ' "	Sole.	1	0	gr. ' "	
1 1	1.31.40	0.000077		41	53.29.4	0.098300
2	3. 3.15	0.000309		42	54.27.32	0.102019
.3	4.34.43	0.000694		43	55.24.21	0.105752
4	6. 6. 0	0.001231	1	44	56.20.1	2 0.109490
5	7.37. 1	0.001921		45	57.15. 6	0.113240
	9. 7.43			46	58. 9. 3	
· 78	10.38. 2	0.003745	1 1	47	59. 2. 4	0.120756
	12. 7.54	0.004876		48	59-54-11	0.124518
9 10	13.37.17			49	60.45.25	0.128278
·	15. 6. 7			50	61.35.45	
11	16.34.20			51	62.25.14	0.135792
12	18. 1.54	0.010798	.	52	63.13.52	
13	19.28.47			53	64. 1.40	0.143291
14	20.54.54		1 1	54	64.48.38	0.147029
15	22.20.14			55	65.34.50	Concession of the local division of the loca
16	23.44.44	0.018783	1	56	66.20.13	0.1 54482
17	25. 8.22	0.021072		57	67.04.50	
18	26.31. 8	0.023470	·	58	67.48 42	0.161890
19	27.52.55	C.025969		<b>5</b> 9 60	68.31.50	0.165578
I -	29.13.47	0.028570	·		69.14.16	0.169254
21	30.33.40	0.031263		61	69.55.58	0.172914
22	31.52.32	0.034045		62	70.36.56	0.176557
23	33.10.23	0.036916	·	63 64	71.17.16	0.180188
24 25	34.27.12	0.039864		65	71.56 56	0.183803
· !	35.42.59	0.042892	· 1		72.35.57	0.187404
26	36.57.41	0.045989		66	73.14.15	0.190978
27 28	38.11.20	0.049154		67 68	73.51.59	0.194540
	39.23.54	0.052382		1.1	74.29.6	0.198085
	40.35.23	0.055668			75.05.38	0.201614
	41.45.47	0.059009	·		75.41.35	0.205122
	42.55.06	0.062400			76.16.56	0.200612
	44. 3.20	0.065838			76.51.43	0.212080
	45.10.29	0.069319			77-25.57	0.215529
	46.16.35	0.072839			77.59.41	0.218963 0.222378
	47.21.36	0.076396			78.32.54	
	48.25.33	0.079984		76	7. 5.35	0.225769
		0.083600	12		37.45	0.229142
		0.087244		8 8	30. 9.23	0.232488
		0.090910	. 1 6	9 8	80.40.34 1.11.16	0.235809
40 1	7 50. 30	0.094596	<u>_</u> °	~	4.11.10	0.239127

The General Table continued.

Med.	Angul. à	Logar.	1	Med.	Angul. à	Logar.
mot.	Perihelio.	pro dist.		mot.	Perihelio.	pro dift.
		à Sole.		0,	gr. ' ''	à Sole.
<u> </u>	gr.					
81	81.41.31	0.242416		142	102.32.41	0.407380
82	82.11.19	0.245684		144	103.00.31	0.411784
83	82.40.40	0.248933		146		0.416132
84	83. 9.34	0.252159 0.255366		148	103.54.31	0.420430
85	83.38. 4			150	104.20.43	0.424676
86	84. 6. 8	0.258552		152	104.46.22	0.428866
87	84.33.49	0.261720		154	105.11.33	0.433012
88	85. 1. 5	0.264865		156	105.36.16	0.437110
89	85.27.58	C.267989		158	106.00.32	0.441164
- 90	85 54.27	0.271092	- 1		106.24.23	0.445178
91	86.20.34	0+274176		162	106.47.47	° 449144
92	86.46.20	0.277239		164	107.10.44	0.453060
93 -	87.11.43	0.280284		166	107.33.17	0.456936
94	87.36.45	0.283306		168	107.55.27	0.460772
95	88.01.27	0.286308		170	108.17.14	0.464208
96	88.25.49	0.289293		172	108.38.37	0.468318
97	88.49.48	0.292252		174	108.59.39	0•472030
98	89.13.32	0.295201		176	109.20.20	°•475705
. 99	89.36.54	0.298122		- 178	109.40.40	0.479340
100	90.00.00	0.301030		180	110.00.40	0.482937
102	90.45.14	0.306782		182	110.20.20	0.486498
104	91.29.18	0.312469		184	110.39.41	0.490023
106	92.12.14	0.31,8060		186	110.58.44	0.493512
108	92.54.4	0.323587		188	111.17.28	0.496965
110	93.34.52	0.329042		190	111.35.55	0.500384
112	94.14.40	0.334424		192	111.54.05	0.503769
114	94.53.30	0.339736		194	112.11.58	0.507121
116	95.31.22	0.344979		196	112.29.34	0.510441
118	96. 8.22	0.350153		198	112.46.55	0.513729
120	96.44.30	0.355262		200	113. 4.00	0.516984
122	97.19.48	0.360306		204	113.37.25	0.523406
124	97.54.17	0.365284		208	114. 952	C.529705
126	98.28.00	0.370200		212	114.41.23	0.535886
128	99.00.57	0.375052		216	115.12.02	0.531958
130	99.33.11	0.379842		220	115.41.51	0.537922
132	100. 4.43	0.384576		224	116.10.52	0.553782
134	100.35.45	0.389252		228	116.39. 7	0.559538
136	101. 5.48	0.393868		232	117.06.38	0.565199
138	101.33.22	0.398428		236	117.33.27	0.570762
140	102. 4.19	0.402930		240		0 22

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The General Table continued.

Med.	Angul. d	Logar.		Med.	Ang. d	Logar.
mot.	perihelio.	pro dift.		mot.	peribelio	pro dift.
0	gr. / "	Sole.		0	gr. ' "	à Sole.
244	And the owner of the owner, where the ow	0.581616	÷.	620		0.882575
248				640	138. 3.58	0.892649
252		0.59212:	l I	660	138.33.21	0.902401
256	119.37.56	0.597252		680	139. 1.29	0.911866
260		0.602301		700	139.28.25	0.921012
264		0.507274	1	720	139.54.16	0.929907
261		0.612174		740	140.19. 5	0.938549
272		0.616998		760		0.946951
270	121.28.39	0.621750	l	780		0.955124
280		0.626438		800	141.28. 3	
284	122. 9.38	0.631056	ł	820	141.49.24 142.10.00	0.970836
288		0.635608	l	840 860	142.10.00	0.978397
292		0.640098		880	142.29.56	
296 300		0.644525 C.648893	1	500	143. 7.48	0.992970
1						
310		0.659559		920	143.25.51	1.000871
320		ು.669880 ೧.679876		940 960	144 00.10	1.013586 1.020155
340		c.689568		980	144.16.46	1.026 eRal
3.50		0.698970		1000	144.32.46	
360	the second se			1500	149.26. 8	
370				2000	1 52.26.15	
380				2500	1 54.32.20	1.313702
390				3000	156.7. 27	1.368678
400				3500	157.22.49	
410				4000	158.24.36	1.454950
420	130.36. 2	<b>0.75</b> 7930		4500	159.16.36	1 4901 25
430	131. 3.30	0.765516		<b>5</b> 000	160. 1.12	1.521521
440				5500	160.40. 5	
450		0.780148		6000	161.14.24	1.575718
460				6500	161.45.00	1.599460
470				7000	162.12.34	1.621417
480	133. 7.50			7500	162.37.34	1,641838
490	133.30.25			8000	163.00.23	.000922
500	133.52.20	the second se	-	8500	163.21.20	
520	134.34.18	0.826522		9000	163.40.42	1.695708
540	135.14. 0			5900	163.58.38	
500 580	135.51.28				164.15.201	
600	136.27. 60			0000	170.52. 02 172.45 <b>.</b> 442	200666
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## The Construction and Use of the General Table.

A S the Planets move in Elliptic Orbs, fo 66 A do the Comets in Parabolic ones, having the Sun in their common Focus, and describe 66 equal Areas in equal Times. Now fince all 33 Parabola's are fimilar to one another, there-" fore" if any determinate Part of the Area of a " given Parabola, be divided into any number of 65 Parts at liberty, 'there will be a like division " made in all Parabola's under the fame Angles, " and the Diftances will be proportional': Con-" fequently this one Table of ours will ferve for 66 all Comets." Thus far Dr. Halley.

But it is to be noted, that our famous Author doth not affert in this Place the Trajectories of Comets to be compleatly Parabolical; but only means, "that whereas they are indeed Elliptical, they are withal fo Eccentrical, that that Part of the Orbits of Planets which respects the Planetary World, and which we the Inhabitants of this Earth can fee, doth to little differ from the curvest Part of a Parabolic Line, that it may fafely, and without any fenfible Error, be assumed to be a Parabola. For it was before noted, that there may be Ellipfes of all Species, and that the Concentrical do at tength degenerate into Circles, 'the infinitely Ec-centrical into Parabola's." Nor is it therefore to be wonder'd at, if instead of an Ellipsis, a Figure of more difficult Contemplation, and generally of an unknown Species, we chule to ule a Parabola, a Figure more easy to be Contemplated, and of one Species only; in that Place, espe-Ee 2 cially

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cially where the Phænomena of Comets mark out to us Trajectories fcarce other than Parabolical." We have before fhew'd "that the Proportionality of the Area to the Time is common to Comets as well as Planets"; and fhall not go over the fame Thing again. It is alfo manifeft, "that like Figures, as Circles and Parabola's do admit and require, that the like Divifions of them, or their Proportional or Correspondent Parts should be express'd by the fame Numbers." But let our Author proceed.

"Now the Manner of the Calculation of this "Table is thus: In Fig. 5. Plate 9. let S be "the Sun, POC the Orbit of a Comet, P the Perihelion, O the Place where the Comet is 90 gr. diftant from the Perihelion, C any o-"ther Place. Draw the Right Lines C P, C S, and make S T, S R, equal to C S; and having "drawn the Right Lines C R, C T, (whereof the one is a Tangent, and the other a Perpendicu-"lar to the Curve) let fall C Q Perpendicular to the Axis P S R.

It is here as it is in the Planetary Aftronomy, (where we first enquire the Place of the Planet, or the Angular Diftance from the Axis of the Ellipsi, which we call the True Anomaly of the fame, together with the absolute Diftance from the Sun; Even fo here, we must in the first Place find out the like Angle and Diftance. But it is to be noted, that according to the Nature of all Parabola's, the Line SO is half the Latus Rectum. S P is a 4th Part of the fame Latus Rectum, or half of SO; and that a Tangent CT being drawn unto any Point C, and there being crected Perpendicular to the fame the Line CR, cutting the Axis; and there being let down from the fame Point,

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#### Aftronomy of COMETS.

Point C, to the Axis, the Perpendicular CQ cutting the Axis in Q; SC, SR and ST are equal amongft themfelves; and the Line QR is equal to SO, or half the *Latus Refum*. All which Things are well known from the Conics. But our Author proceeds.

"Now any Area, as COPS being giv'n, it is requir'd to find the Angle CSP, and the Difrance CS. From the Nature of the Parabola RQ is ever  $=\frac{1}{2}$  the Parameter or Latus Rectum of the Axis, and confequently if the Parameter be put =2, then RQ =1. Let CQ =2; PQ fhall  $=\frac{1}{2}zz$ , and the Parabolic Segment COP  $=\frac{1}{12}zzz$ ; But the Triangle CSP will  $=\frac{1}{2}z$ , and fo the Mixtilineal Area COPS =if  $z^{1} \times \frac{1}{4}z = 2$ , whence  $z^{1} \times 2z = 122$ . Wherefore refolving this Cubical Equation, z, or the Ordinate CQ will be known.

Thus far our Author. But it is to be well obferv'd, that we have here the Analytical way for finding the Coequate Anomaly in a Parabola from the Mean Anomaly given, that is, from the Area describ'd, which is every-where Proportional to the Time of the Description. Nor can the Angle CST, or the Coequate Anomaly be found directly from the Given Area or Mean Anomaly,... without Analysis. But that upon the Hypothesis, that the Line CQ, which is first to be fought, (for when that is found, the Angle CST will eafily be found, as will be manifest prefently) be called z, the Line PQ will be equal to  $\frac{1}{2}$  z z, is eafy to be demonstrated: For as RQ = I, is to 2CQ = z; fo the fame CQ = z is to QT, or zz, the half whereof confequently QP will be equal to zz. But that the Parabolic Segment COP, according to the fame Hypothesis, will be Ee 2 rightly

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rightly express'd by 13 z z z, easily follows from the Conics. For the Area COPSQ is to the Tri-: angle CPQ, or CPT equal to the fame)as 4 to 3; and confequently the Parabolic Area COP is \* to CPQ as 1 to 2; and fince the Triangle CPQ made of the Perpendicular CQ or z, drawn into half the Base izz, becomes izz, the 3d Part of it will be 12 z z z, equal to the Parabolic Area The Triangle CSP also is made of COP. the Perpendicular z, drawn into half the Bale i, equal to iz; and confequently the Sum of the Area's COP and CSP, or the whole Area COPS, Proportional to the Time, will be equal to the Sum of these Quantities, which is called a: or there will arife this Equation  $\frac{1}{12}z^3 \neq \frac{1}{2}z \neq = a_1$ and by multiplying on both Sides by 12,  $z^2 + 2z$ = 120; which is a Cubic Equation, the 2d and Ath Terms whereof are wanting. The Root therefore of this Equation being found, or the Value of z being found in Numbers by Dr. Halley's Method, or otherwife, the Length of the Line CQ will be known. Q. E. I. And now let us hear our Author himfelf.

Dr. Halley. " Now let the Area OPS be pro-" pos'd to be divided into One Hundred Parts ; " this Area is is of the Square of the Parameter, and " confequently 12 a is equal to that Square = 4. " If therefore the Roots of these Equations  $z^{i}$  + " z = 0, 04: 0, 08: 0, 12: 0, 16, or be fuc-" ceffively extracted, there will be obtain'd fo "many z or Ordinates CQ respectively, and " the Area SO P will be divided into One Hun-" dred equal Parts. And in like manner is the Cal-" culation to be continued beyond the Place O. "Now the Root of this Equation (fince R Q is " = I) is the Tabular Tangent of the Angle "CRQ, or of the Angle CSP, wherefore the Angle

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" Angle, CSP is given. And RC, the Secant " of the fame Angle C R Q, is a mean Proportional between R Q, or Unity and RT, which 66 " is the double of SC, as is plain from the Co-52 nics. But if SP be put = 1, and fo the Latus " Rettum = 4 (as in our Table) then R T will " be the Distance fought, viz. the double of SC in the former Parabola. After this manner " therefore, I compos'd the foregoing Table, " which ferves to reprefent the Motions of all " our Comets; of which hicherto there has been " none observed, but those that come within " the Laws of the Parabola,

Now that the Area OPS is a twelfth Part of the Square of the Latus Rectum, it is manifelt: Because, according to the Conics, the Area OPS is i of the Rectangle of OS, multiplied by SP; that is, of the Rectangle of half the Latus Rectum, multiplied by a 4th Part of the fame. For  $\frac{1}{2} \times \frac{1}{2} = \frac{1}{12}$  But any Numbers, as 4.8. 12. 16. If they be put in the 2d Place of Decimals, as here is done, will rightly express 100th Parts. And we are therefore content with a Right Angle as the principal Guide of Computation, because we want an entire Period in Parabola's. But because of the equal Angles SC, SR, the external Angle of the Isofccles Triangle CRS will be equal to the double Angle CRS. And there being giv'n confequently by the Tables of Tangents the Angle C R Q, the Double thereof. or the Angle CST; that is, the Coequate Anomaly of the Comet is found. In like manner, the Angle CST being now giv'n, if you make by the Golden Rule : As RQ = 1 is to the Secant of that Angle to be taken out of the fame Ta-E e A bles ; **-** .-

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bles; fo that Secant is to the 3d Proportional R T; The half hereof R S is equal to S C, or to the Diftance of the Comet from the Sun. Q. E. I.

Nov. 29. 1708.

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#### LECT. XL.

Halley. "It now remains, that we give the Rules for the Calculation, Dr. . . to fhew the way of determining the "Visible Place of a Comet, by these "Numbers. The Velocity of a Comet " moving in a Parabola, is every-where to the Velo-66 city of a Planet describing a Circle about the Sun, at " the same Distance from the Sun, as V 2 to 1. as " appears from Cor. 7. Prop. 16. Lib. 1. of the " Princip. Phil. Nat. Math. If therefore a Comet " in its Perihelion were suppos'd to be as far di-" ftant from the Sun as the Earth is, then the " Diurnal Area which the Comet wou'd describe, " woud be to the Diurnal Area of the Earth, as " d > to T And confequently, the Time of the  $\sqrt{2}$  to 1. And confequently, the Time of the " Annual Revolution, is to the Time in which " fuch a Comet wou'd describe the Quadrant of " its Orbit from the Perihelium, as 3. 14159, &c. " (that is the Area of the Circle) to  $\sqrt{\frac{5}{5}}$ 

That the Velocity in a Parabola is to the Velocity at the fame Diftance in a Circle, as √2 is to 1, or as 10 to 7 almost, was demonstrated in Prop. XXII. foregoing; or rather deduced

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deduced as it were a Corollary, from the Nature of Circular and Parabolic Curvity, and the Proportion of the Subtenfes of the Angle of Contact. But the Annual Time in an Elliptic Circle: or the Time of an Entire Revolution, represented by the whole Area of a Circle, which is to be eftimated from the Multiplication of half the Circumference by the Radius ; will be to the Time of the Description of a Quadrantal Arch in a Parabola, which is to be represented by a Quadrantal Area of the Parabola, to be estimated from the Multiplication of 3 of balf the Latus Rectum, by a quarter of the same Latus or Radius; as the Area's themfelves; or as the Heighths of the Rectangles to the Common Base; only so far as the Velocity of Description in a Parabola doth difturb and diminish that Proportion of the Times, in the Proportion of i to  $\sqrt{2}$ ; and therefore inflead of  $\frac{1}{2}$ . let v t be taken : and let the Numerator be doubled, becaufe of the Square Number two, the double of Unity; that is, for the Circle, let the Area of it 2, 141, 59, be taken, for the Parabola V. And thus the Truth of our Author's Reafoning will eafily be underftood.

Dr. Halley. "Therefore the Comet wou'd describe " that Quadrant in 109 Days, 14 Hours, 46 Minutes; and fo the Parabolic Area ( analogous " to the Area POS) being divided into One "Hundred Parts, to each Day there wou'd be al-" lotted 0.912.280 of those Parts, the Log. of " which, viz. 9.960128, is to be kept for con-" tinual use. But then the Times in which Co-" mets, at a greater or less Distance, wou'd de-" scribe fimilar Quadrants, are as the Times of " the Revolutions in Circles, that is, in the Selguiplicate Ratio of the Diftances: Whence the " Diur-

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"Diurnal Areas estimated in Centefinal Parts of the Quadrant (which Parts we put for Meafures of the mean Motion, like Degrees) are in each, in the Subsessment proportion of the Distance from the Sun in the Perihelion.

Mr. Whifton. The Mean Diurnal Motion, to wit, 0, 912, 280, to be express'd by a Negative Logarithm after the Old Manner- 0,029,872, is in this Place express'd in a New Way by a Positive One 9,960, 128, to avoid the Difficulty about the Negative Characteristic; but is prefently made equivalent to the wonted Form, by caffing away Ten in the Addition when occasion shall require. But our Author observes here rightly, that in divers Parabola's, a Quadrant is always reckon'd of the fame Number of Parts, I mean an Hundred; in fuch Sort nevertheless, that those Parts be indeed unequal, and according to the Magnitude of the Parabola greater or lefs, but not in that Proportion greater or lefs, in which the Diffances increase or decrease from the Sun, but in the Subfesquialteral Proportion of the fame: So that the Squares of the Diftances be betwixt themselves. as the Cubes of these Parts reciprocally.

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Dr. Halley. "These necessary Things premis'd, "let it be propos'd to compute the apparent Place "of any one of the foremention'd Comets for a-"ny given Time. Therefore,

[" 1. Let the Sun's Place be had, and the Log. of its Distance from the Earth.

2. Let the difference between the Time of the Peribelion and the Time given he gotten, in Days and Decimal Parts of Days. To the Log. of the Number, let there

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there be added the constant Log. 9.960.128, and the Gomplement Arithmetical of three balwes of the Log: of the Perihelion Distance of the Comet from the Sun : The Sum will be the Log. of the mean Motion, to be sought in the first Column of the general Table,

3. With the mean Motion let there be taken the correspondent Angle from the Perihelion in the Table, and the Log. for the Distance from the Sun: Then in Comets that are Direct; add, and in Retrograde ones subtract; if the Time be after the Perihelion, the Angle that found, to or from the Perihelion : or in Direct Comets, substract; and in Retrograde ones add; if the, Time be before the Perihelion; the foressaid Angle to or from the Place of the Perihelion; and so we shall have the Place of the Comet in its Orbit. And to the Log. for the Distance found, let there be added the Log. of the Distances to Perihelion; and the Sum will be the Log. of the sum Distance of the Comet from the Sum.

A. The Place of it is Node; together with the Place of the Comer in its Orbit, being given, let the Diffance of the Comer in its Orbit, being given, let the Diffance of the Comer from the Node be found; then the Inclination of the Plane being given; there will be given also (from the common Rules of Trigonometry) the Comet's Place reduced to the Reliptic, the Inclination or Heliocentric Latitude, and the Log. of the Curt Diffance.

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5. From these things given (by the very same Rules that we find the Planet's Places, from the Sam's Place and Distance given) we may obtain the Apparent or Geocentric Place of the Comet, together with the Apparent Latitude. And this it may be worth while to illustrate by an Example or two."] 428

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As to the Place of the Sun, and the Diffance thereof from the Earth, we have elfewhere taught how to find both by Aftronomical Calculation. But the Logarithms of the Diftances, we through fome neglect, omitted in that Place; and therefore shall add them in the End of this Work. But the Logarithm of Days is therefore added to the Given Logarithm of one Day, that the Motion of one Day may be understood to be multiplied by the Number of Days: For it is known, that the Addition of Logarithms doth infer the Multiplication of Numbers correfponding to those Logarithms.

And these Things may fuffice, if to be the Comet be suppos'd to pals in its Perihelion at a Diftance equal to a Radius of the great Orb. But if. which commonly is the Cafe, the Comer doth not pals at that Diftance, but at a greater was it is fometimes; or at a lefs, as oftner happens; that Area, proportional to the Time, is to be increased or diminish'd ; and this in the Sub-seferuiakeral Propertion of that least Diftance from the Sun ; to that at length that Area may rightly reprefent the Mean Anomaly. From whence the Logarithm of that Selqui-plicate Diftance is to be added to the former Sum of the Logarithms, and the Radius to be fubftracted according to the Exigence of the Golden Rule, to be practised in Logarithms; or which is the fame, the Arithmetical Complement only of that Selqui-alteral Logarithm is to be added. Neither ought it to feem ftrange, that in leffer Diftances we, by adding the Logarithm, obtain the true Proportion increas'd, and the same in greater Distances diminish'd : For Multiplication by a Fraction, or Decimal

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cimal Parts, doth no less diminish the Sum, than Multiplication by whole Numbers doth increase it. And the Thing is the same in Logarithmetical Addition. But we are to observe, that the Logarithms marked in the 3d Column of the General Table, are not the Logarithms of the Numbers of the Distances from the Sun, to be added over and above the Radius to the Mean Distance; but of Numbers, by the Multiplication of which, that true Distance were to be obtained. From whence the Logarithms of the same being superadded to one another, will easily give us the Logarithm of that whole Distance from the Sun. These Things being well understood, we shall be able to undertake and perform the Calculation.

#### EXAMPLE - L

Let is be requir'd to find the Place of the Comet of the Year 1664, March 1<sup>d</sup>, 7<sup>h</sup>, 00', P. M. London. That is, 96<sup>d</sup>, 19<sup>h</sup>, 8', after the Perihelion, which happen'd Novemb. 24°, 11<sup>h</sup>, 52'.

	•
Log. Dift. Perihel.	<b>0. 011044</b>
Log. Selquialt.	0. 016566
Comp. Arith.	9. 983434
	9. 960128
Log. Temp.	1. 985862
Log. Med. Mot.	1. 929424
Medius Motus	85.001
Perihel. R	10. 41. 25
Ang. Corresp.	82. 28. 05-
Comet in Orb. V	17. 3. 20
Ascend. Nod. II	21.14.00
Com. à Nodo	<b>2</b> 4. 10. 40
Red. ad Eclip.	22. 19. 05
Com. Helioc. &	18. 54. 55
Incl. Bor.	11. 46. 50
Log. pro dift.	0. 255369
Log. Perihel.	0. 011044
Co-fin. Incl.	9. 990754
Log. dift. Cur	0. 257167
Log. dift. O	9. 997918
Ο×	21. 44. 45
Com. Visus r	29. 18. 30
Lat. Vifa (Bor.)	8. 36. 15

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#### EXAMPLE II.

Let it be requir'd to find the Place of the Comet of the Tear 1683, July 23°, 12<sup>h</sup>, 25', P. M. London: Or, 13<sup>h</sup>, 40' Equat. Time. That w, 21<sup>d</sup>, 10<sup>h</sup>, 50' after the Peribelion.

Log. Dift. Perihel. Log. Sefquialt. Comp. Arith. Log. Temp. Log. Med. Mot. Medius Motus	9. 748343 9. 622514 0. 377486 9. 960128 1. 310723 1. 648337 44. 498
Perihel. II Ang. Correfp. Comet. in Orb. Y Nod. Defcend. X Com. à Nodo Red. ad Eclip. Com. Helioc. X Incl. Bor.	$\begin{array}{r} 25. \ 29. \ 30\\ 56. \ 47. \ 20-\\ 28. \ 42. \ 10\\ 23. \ 23. \ 00\\ 35. \ 19. \ 10\\ \underline{4. \ 48. \ 30}\\ 28. \ 11. \ 30\\ 35. \ 2. \ 00\end{array}$
Log. pro dift. Log. Perihel. Co-fin. Incl. Log. dift. Curt. Log. dift. O O Locus fl Com. Vifus S Lat. Bor.	0. 111336 9. 748343 9. 913187 9. 772866 0. 006104 10. 41. 25 5. 11. 50 28. 52. 00

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But now, that we may rightly perform this Calculation; It is to be noted,

(1.) That the Logarithm of the leaft, or perihelion Diftance, is only fet down here, that we may obtain the other Logarithm, which is Sefquialteral of the fame, or is thereto as 3 to 2.

(2.) That the Arithmetical Complement of this last Logarithm being added to the Constant Logarithm of one Day, doth make the Logarithm of the whole Time before or after the Perihelion. For working by Logarithms, the Numbers in the former of the Examples will be thus. The Logarithm of one Day is 9,960, 128; and the Logarithm of Days is 1, 985, 862. These alone being added together, would make the Logarithm of the Mean Motion, if the Perihelion Distance were equal to Unity, or the Radius of the great Orb : But when the Area of that Mean Motion is to be increas'd in the Proportion of that Sesquialteral Perihelion Distance to the Radius of the Annual Orbit, that Sefquialteral Logarithm 0,016,566, is to be added to the former Logarithm; and the Logarithm of the Number 10 is to be fubftracted; or, which comes to the fame. the Arithmetical Complement of the Sefqui-alteral Logarithm is only to be added : which is done in this Place. Now the Mean Motion will eafily be known, when the Logarithm of the fame is given.

(3.) The Mean Motion, or Mean Anomaly, being now given, the Angle Corresponding thereto in the General Table, is 83°. 28'. 5". (the intermediate proportional Parts being everywhere found, where there is Occasion, by the Golden Rule.) Which being deducted from the Place of the Perihelion in Leo 10°. 41'. 25", because of the Retrograde Motion of the Comet giveth

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us the Place of the Comet in its own Orb, 17°. 2'. 20", in Taurus,

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(4.) Substract this Place from the Place of the descending Node in Gemini; the Remainder will be the Diffance of the Comet from the Node; 34°. 10'. 40".

 $(\varsigma.)$  And now that we may reduce the Place of the Comet in its own Orb to the Ecliptic, we must refolve a Rectangular Spherical Triangle; and from the Given Angle and the Hypotenule, must find the other Sides. And for Reduction to the Ecliptic, for the Heliocentrical Longitude, the following Analogy will fuffice.

As Radius \_\_\_\_\_\_ 10, 000, 000 is to Cofin.of the Ang. 21°. 18'. 30".--9.969.248 So is Tangent, \_\_\_\_\_ 24. 10. 40. -9. 821. 890 To Tangent, \_\_\_\_\_ 9. 801. 128= =22°. 19'. 5"

Then for the Inclination, or Heliocentrical Latitude,

As Radius, \_\_\_\_\_\_ 10.000.000 To Sin. \_\_\_\_\_\_  $24^{\circ}$ . 10'. 40". \_\_\_\_\_ 9.749.552 So is Sine of 21. 18: 20 \_\_\_\_\_\_ 9.560.269 To Sine of 2 \_\_\_\_\_\_ 9.309.922= Ang. fought 5 \_\_\_\_\_\_ 9.309.922=  $11^{\circ}$ . 46'. 44,

(6.) For obtaining the Logarithm of the true Diftance of the Comet from the Sun, we must add the Logarithm for the Diftance from the Sun, which in the General Table belongs to the Mean Motion, to the Logarithm of the least or perihelium Diftance; that is, 0.255.269, to 0.011.044, which make 0.266,412. And then fay, As Radius — 10.000.000

To Dift. from Sun — 0, 266. 413 F f 434

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So is Co-fin of Inclin. — 9. 990. 754 To Curt. Dift. — — 0. 257. 167

Or, which comes to the fame; the three Logarithms are to be added, and the Logarithm of the Radius to be caft away; as is done in our Examples.

(7.) For obtaining the Geo-centrical Longitude of the Comet, or the visible Place in the Ecliptic, do thus. Substrate the Helio-centrical Longitude 1<sup>s</sup>. 18°. 54'-55'' out of the true Place of the Sun in the Ecliptic 11<sup>s</sup>. 21°. 44'. 45''; there will Remain the Angle of Commutation  $10^{5}$ .  $2^{\circ}$ . 49'. 50''; the Complement whereof unto a Circle is 1<sup>s</sup>.  $27^{\circ}$ . 10'. 10'', or,  $57^{\circ}$ . 10'. 10''. The half hereof is  $28^{\circ}$ . 35'. 5''. From whence fay,

As dift. of the Earth ---- 9.997.918

To Curt. Dift.of theCom.10.257.167

So is Radius \_\_\_\_\_ 10.000.000

To Tangent —  $10.259.249 = 61^{\circ}.10^{\circ}.2^{\circ}$ 

Now 45 Deg. being caft away, there refts-16.10.2. Therefore,

As Radius ——— 10.000.000 To Tang. 16°. 10'. 3".— 9.462.265 So is Tang. of Semi-Sum.-- 9.726.294 == 28°.25'.5". To Tang. of Semi-differ.— 9.198.559 == 8.58.26.

Which Half-difference being taken away out of the Half-Sum, there remains  $19^\circ$ : 26'. 29''; that is, the Parallax of the Orb. But the Parallax being in this Cafe fubftracted from the Heliocentrical Place of the Comet, the Geocentrical Place of the fame is  $\Upsilon$ . 29.  $18^\circ$ . 26. Something more exactly, as I fuppole, than our Author's Calculation hath it.

But if the Curtated Distance of the Comet from the Sun be less than the Distance of the Earth from the Sun, as it is in the other Example, we must work

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in the Calculation, as is done for the Inferior Planets, (like as we have Calculated here, as we do for the Superior.) And the Half difference of the Angles, which in that Cafe will reprefers the Elongation from the Sun, is to be added to the Longitude of the Sun in the Ecliptic, or fubfitracted from the fame, for obtaining the Geecentrical Place of the Comet.

(8.) For determining the Geocentrical Latitude of the Comet, we are to Work thus; (the Angle of Elongation being made up of the Aggregate of the Half-Sums.)

As Sin. of Ang. of Commut. 57.10.10. 9.924.423 Is to Sin. of Ang. of Elong. 37.33.41. 9.785 053 So is Tang. of Inclination-11.46.44. 9.319.161 To Tang. of Latitude ----(8.36.09.) 9.179.791

"At the Inftant of Time specified in the first " Example, 'twas observ'd (at London) that the " Comet applied to the second Star of Aries; fo " that it was found to be 9' more Northerly, 66 and 2' to the East, according to Dr. Hook's Ob-" fervation. But at that of the fecond Example, " I my felf ( near London, with the fame Instru-" ments whereby I formerly observ'd the Southern " Constellations) found the Place of the Comet 66 to be 5, 5°, 11'1, and 28°, 52' North Latitude, çc which agreed exactly with the Observation 66 made at Greenwich, almost at the very fame Mo-66 ment.

"As for the Comet of the Year 1680, which came almost to the very Sun it felf (being in its Perihelion, not above one third of the Semidiameter of the Sun distant from the Surface of it,) fince the Latur Restand of its Orb is fo very fmall, it could hardly be contained within the Limits Ff 2 <sup>66</sup> of the General Table, becaufe of the exceffive <sup>66</sup> Velocity of the Mean Motion. Wherefore in <sup>67</sup> this Comet, the beft way will be (after the <sup>67</sup> Mean Motion is found) to get from thence (by <sup>67</sup> the help of the foregoing Equation  $z^3 + 3z = \frac{4}{100}$ <sup>67</sup> of the Mean Motion) the Tangent of half the <sup>67</sup> Angle from the Perihelion, together with the <sup>67</sup> Log. for the Diftance from the Sun. Which be-<sup>67</sup> ing found, we are to proceed by the fame Rules, <sup>67</sup> as in the reft.

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"After this manner therefore, the Aftronomi-"cal Reader may examine thefe Numbers, which "I have calculated with all imaginable Care, from the Obfervations I could meet with. And I have not thought fit to make them publick before they have been by my felf duly examin'd, and made as accurate as 'twas poffible, not without the Labour of many Years. I have publifh'd "this Specimen of Cometical Aftronomy, as a *Prodromus* of a future Work I have in defign, left, happening to be prevented by Death, thefe Papers might chance to be loft, which every Man would not be capable to retrieve, by "reafon of the great Difficulty of the Calculation.

"Now it may not be amils to put the Reader in mind, that our five first Comets, (the third and fourth observed by Peter Apian, the fifth by *Paulus Fabricius*) as also the tenth, feen by Mestin, if I mistake not, in the Year 1596, are not for certain as the rest; for the Observations were made neither with sufficient Instruient disgreeing with themselves, and can by no means be reconciled with a regular Computus. The Comet which appeared in the Year 1684, was only taken notice of by Blanchinus, who observed it at Rome: And the last, which apient and the sufficient of the sufficient of the sufficient is peared

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pear'd in the Year 1698, was feen only by the' 66 Parifian Observers, who determin'd its Course " after a very uncommon manner. This Comet " was a very obscure one, and altho it may'd ~ fwift, and came near enough to our Earth ; yet 86 we, who are wont not to be incurious in these æ Matters, faw nothing of it. For want of Ob-" fervations, I have also left out of the foregoing " Catalogue, those two remarkable Comers which " have appear'd in this our Age, one in Novem-" ber in the Year 1689, the other in February, in ¢¢ the Year 1702. For they directing their Cour-"fes towards the Southern Parts of the World," and heing fcarce conficuous any where in and being fcarce confpicuous any where in " Europe, met with no Observers proper for the purpose. But if any one shall bring from India; or the Southern Parts, an accurate Series of Qb-"fervations, I will willingly fall to work again; "and undergo the Fatigue of representing their " Orbits in Numbers, as I have done the reft,

" By comparing together the Elements of the " Motions of these Comets, 'tis apparent, then " Orbits are dispos'd in no manner of Order; nor " can they, as the Planets are, be comprehended "within a Zodiac; moving indifferently every " way, as well recrograde as direct; from whence " it is clear, they are not carry'd about or mov'd " in a Vortical System. Moreover, the Distances " in their Perihelia are fometimes greater, fome-" times les; which makes me fuspect, there may " be a far greater Number of them, which may 66. move in Regions more remote from the Sur, " and being therefore very obfcure; and wanting " Tails, may pass by us unseen.

"Hitherto I have confider'd the Orbits of Co-" mers as exactly Parabolic; upon which Suppo-" fition it wou'd follow, that Comets being imf pell'd towards the Sun by a Centripetal Force, Ff 2 " would 438

would descend as from Spaces infinitely diftant, " and by their to falling acquire fuch a Velocity, 66 as that they may again fly off into the remotest a Parts of the Universe, moving upwards with a perpetual Tendency, fo as never to return again to the Sun. But fince they appear frequently 66 enough; and fince none of them can be found " to move with an Hyperbolic Motion, or a Mo-66 tion fwifter than what a Comet might acquire ¢C. by its Gravity to the Sun, 'tis highly probable 66 they rather move in very Eccentric Elliptic Or-" bits, and make their Returns after long Periods of Time: For fo their Number will be determinate, and, perhaps, not fo very great. Belides, " the Space between the Sun and the Fix'd Stars is " to immenfe, that there is room enough for a " Comet to revolve, tho' the Period of its Revo-<sup>16</sup> lution be vaftly long. Now, the Latus Rectum of 4 an Ellipsi, is to the Latus Rectum of a Parabola, CC. which has the fame Diftance in its Perihelium; ٤. as the Diftance in the Aphelium in the Ellipsi, 56 is to the whole Axis of the Ellipfis. And the Velocities are in a Subduplicate Ratio of the fame : Wherefore in very Excentric Orbits the Ratio 4 comes very near to a Ratio of Equality; and the **\$6** very small difference which happens on account 64 of the greater Velocity in the Parabola, is eafily CC. compensated in determining the Situation of the 66 Orbit. The principal Use therefore of this Table " of the Elements of their Motions, and that which 66 indeed induced me to conftruct it, is, that when-۶č ever a new Comet shall appear, we may be a-66 ble to know, by comparing together the Elećć ments, whether it be any of those which has 15 appear'd before, and confequently to deter-..... mine its Period, and the Axis of its Orbit, .... and to foretel its Return. And, indeed there EC. are many things which make me believe that "the

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the Comet which Apian observ'd in the Year " 1521, was the fame with that which Kepler and 33 Longomontanus more accurately describ'd in the cc Year 1607; and which I my felf have feen reçç turn, and observ'd in the Year 1682. All the " Elements agree, and nothing feems to con-66 tradict this my Opinion, besides the Inequality " of the Periodic Revolutions. Which Inequality •• is not fo great neither, as that it may not be \$6 owing to Physical Caufes. For the Motion of " Saturn is fo dilturbed by the reft of the Planets, " especially Jupiter, that the Periodic Time of 56 that Planet is uncertain for fome whole Days tor " gether. How much more therefore will a Co-66 met be subject to such like Errors, which rifes " almost four times higher than Saturn, and whole " Velocity, tho' increased but a very little, would " be fufficient to change its Orbit, from an Ellip-ČC. tical to a Parabolical one. And I am the more 4 confirmed in my Opinion of its being the fame; 10 for that in the Year 1456, in the Summer-time, " a Comet was feen passing Retrograde between " the Earth and the Sun, much after the fame cċ manner: Which tho' nobody made Observations ¢¢. upon it, yet from its Period and the manner of ćc its Transit, I cannot think different from those " I have just now mention'd. And fince looking ç c over the Histories of Comets I find, at an equal " Interval of Time, a Comet to have been feen a-66 bout Eafter in the Year 1205, which is another " double Period of 151 Years before the former : ČC . Hence I think I may venture to foretel, that it " will return again in the Year 1758. And, if it " fhould then fo return, we fhall have no reafon to " doubt but the reft may return alfo: Therefore " Aftronomers have a large Field wherein to exer-" cife themselves for many Ages, before they will be able to know the Number of these many and Ff4 great

#### A SYNOPSIS of the

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" great Bodies revolving about the common Cen-" ter of the Sun, and to reduce their Motions to " certain Rules. I thought indeed that the Comet 66 which appear'd in the Year 1522, might be the 66 fame with that observed by Hevelins in the Year " 1661. But Apian's Observations, which are the 66 only ones we have concerning the first of these " Comets, are too rude and inaccurate for any " thing of certainty to be drawn from them, in fo " nice a matter. But as far as probability from the equality of Periods, and fimilar appearance of " Comets, may be urged as an argument, the late " wondrous Comet of 168°, feems to have been the " fame, which was feen in the Time of our King " Henry I. Anno 1 106, which began to appear in the **6**< West about the middle of February, and continu-23 ed for many Days after, with fuch a Tail as was feen in that of 168?. And again in the Cor-" fulate of Lampadian and Oreftes, about the Year " of Chrift 521, fuch another Comet appeared, " in the West, of which Malela, perhaps an Eye-" witnefs, relates that it was winas is polieis, a great and fearful Star; that it appeared in the Weft, and " emitted upwards from it a long white Beam; " and was feen for 20 Days. It were to be wish'd the Hiftorian had told us what Time of the Year ¢C " it was feen ; but 'tis however plain, that the " Interval between this and that of 1106, is near-" ly equal to that between 1106 and 168°, viz. " about 575 Years. And if we reckon backward " fuch another Period, we shall come to the " 44th Year before Christ, in which Julius Casfar " was murder'd, and in which there appear'd a " very remarkable Comet, mentioned by almost " all the Hiftorians of those Times, and by Pliny " in his Natural Hiftory, lib. 11. c. 24. who recites " the Words of Augustus Cafar on this Occasion, " which lead us to the very Time of its Appearance,

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ance, and its Situation in the Heavens. Thefe' " Words being very much to our purpose, it may ..... not be amils to recite them. In ipfis Ludorum me-. .. orum diebus, (ydus crinitum per septem dies, in regione " Cali qua (ub Septentrionibus, est conspectum. Id oriebae c tur circa undecimam boram diei, clarumq; & omnibus " terris confpicuum fuit. Now these Ludi were de-66 dicated Veneri genetrici, (for from Venus the Ca-" fars would be thought to be descended,) and began with the Birth-day of Augustus, viz. Sept.22. " (as may be collected from a Fragment of an 66 Old Roman Calendar extant in Gruter, pag. 125.) " and continued for 7 Days, during which the Co-66 met appeared. Nor are we to fuppose that it was feen only those 7 Days, but possibly both before " and after. Nor are we to interpret the Words (ub Septentrionibus, as if the Comet had appear'd " in the North, but that it was feen under the Septem triones, or brighter Stars of Urla major. And as to its rifing Hora undecima diei, it can no ways رى be understood, unless the word diei be left out. " as it is by Suetonius; for it must have been very far from the Sun, either to rife at Five in the " Afternoon, or at Eleven at Night; in which "Cafes it must have appeared for a long time, and "its Tail have been to little remarkable, that it could by no means be call'd, Clarum & om-66 nibus Terris confpicuum Sydus. But supposing this 66 Comet to have traced the fame Path with that " of the Year 1680, the ascending part of the Orb will exactly represent, all that Augustus hath faid "concerning it; and is yet an additional Argu-"ment to that drawn from the Equality of the "Period. Thus 'tis not improbable but this Co-"met may have four times visited us at Intervals "of about 575 Years : Whence the Transverse "Diameter of its Elliptic Orb will be found 1'575×575 times greater than the annual Orb; 10

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" or 138 times greater than the mean Diffance of the Sun; which Diffance, tho' immenfely great, bears no proportion to that of the Fix'd Stars.

" I have lately found out a ready Method to 66 compute the Motion of Comets in these Ellip-66 tic Orbs, of which perhaps fhorthy we may ex-" hibit a Specimen, giving this Comet for an Ex-" ample. In the mean time, those that defire to know how to construct Geometrically the Orb " of a Comet, by three accurate Observations gi-\*\* ven, may find it at the End of the 2d Book of " Sir Ifaac Newton's Principles of Natural Philoso-66 phy, entituled De Syftemate Mundi, in the Words ~ of its renowned Inventor. Which have fince been •• more fully explain'd by my very worthy Col-" legue Dr. Gregery, in his learned Work of Aftro-66 nomia Physica & Geometrica.

"One thing more perhaps it may not be im-66 proper or unpleafant to advertife the Aftronoçç mical Reader; That fome of these Comets have 66. their Nodes fo very near the Annual Orb of the " Earth, that if it shall so happen, that the Earth ۶¢ be found in the Parts of her Orb next the Node **6**C of fuch a Comer, whilft the Comer passes by; **«** as the apparent Motion of the Comet will be " incredibly swift, so its Parallax will become " very fensible; and the proportion thereof to " that of the Sun will be given. Wherefore fuch " Transits of Comets do afford us the very best " means, tho' they feldom happen, to determine " the Diftance of the Sun and Earth: Which hi-" thereo has only been attempted by Mars in his " Opposition to the Sun; or else Venus in Peri-" gzo, whole Parallaxes, tho' triple to that of the Sun, are scarce any ways to be perceived by " our Instruments; whence we are still in great " Uncertainty in that Affair. This Use of Comets '' was

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Æ was the ingenious Thought of that excellent Ge-¢c ometrician Mr. Nicolas Fatio. Now the Comet Ġċ of 1472, had a Parallax above twenty Times 61 greater than the Sun's. And if the Comet of 60 1618.had come down, about the middle of March. 6c ro his descending Node; or if that of 1684, 66 had arriv'd a little fooner at its afcending Node. 66 they would have been yet much nearer the <sup>55</sup> Earth, and confequently have had more notable \* Parallaxes. But hitherto none has threaten'd the " Earth with a nearer Appulse, than that of 1680. " For by Calculation I find, that Novemb. 11°, 14, " 6', P. M. that Comet was not above the Semi-" diameter of the Sun to the Northwards of the Way of the Earth. At which time, had the " Earth been there, the Comet would have had 66 a Parallax equal to that of the Moon. as I . 66 take it. This is fooken to Aftronomers: But 62 what might be the Confequences of fo near 66 an Appulle; or of a Contact, or, laftly, of a " Shock of the Celestial Bodies, (which is by no ¢ç. means impossible to come to pass,) I leave to 65 be difcuss'd by the Studious of Physical Mat-66 ters.

FINIS.

#### ERRATA.

PAge 4. Line 15, dele, p. 5. l. ult. dele made less by the Diffance H L n 10 1 dele Made less by the Diftance H I. p. 10. 1. 5. dele X. J. 28. read m I. p. 11. 1. 22. r. Plate I. Fig. 2. p. 12. l. 12. dele or R Z. l. 12, r. Pl and KT. 1. 17. del. T. p 13. 1. 26. r. fince. 1. 29. r. Semi-ordinates. p. 15. l. 21. r. as long again or. p. 17. l. 8. r. LL. 1. 25. 1. of the. p. 18. 1. 4. 5. del. whether of the following Sections, or of the former. 1. 32. r. K.H. D.H. 1. 34. r. ib, bb. i a. b a. p. 19. l. 8, 9. r. double. p. 21. 1. 9. r. Fig. 6. p. 27. L. 20. del. smuld be laid down. 1. 32. del. fo. p. 28. l. 31. del. therefore. p. 29. l. 5. del. not. l. 6. del. to. p. 31. l. penuk. r. 21. l. ulc. r. therefore where the ambient Bodies are moved, these that relatively rest in those ambient Bodies are really moved. p. 38. l. 6.

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1. 6. r. a certain. p. 43, l. 17. del. the l. 33. r. of. p. 44. l. 18. t. the. p. 46 l. 13. r. it. p. 56. l. 18. r. coming with. P. 61. 1. 32. r. € or 1 €. 1. 37. r. 1 or 2€. p. 70. 1. 29. r. this. P. 74. 1. 0. r. Axels, Ropes, Strings. p. 77. 1. 27. r. C.E. p. 78. 1. 7. r. Corollary after. (p. 88. l. 14. r. CHB. p. 94. l. ult. r. if m. p. 97. l. 11. r. Plate 3. Fig. 5. p. 98. l. 19. r. db, DB. p. 99, l. 31. r. Eb. p. 105. l. 21. r. ax. l. 26. r. ed. l. 31. r. ga p. 106. 1. 25. r. a b1. 1. 29. r. 1b. 1. penult. r. Lines. p. 108 1. 20. del. bg. p. 112. 1. 1. r. are. 1.9. r. E H. 1. 29. r. mere the Delcent. : p. 116. l. 8. r. are .. 1 21. r. EA. p. 120. l. penult. r. Te. p. 121. . 9. r. I. Te. 1: 20, r. Fl. 1. 25, 26. r. l. and m. p. 122. 1. 3. T. Tb. p. 123. 1. 18. r. cr. p. 124. l. 5. del. D. l. 15. r. 5 F. sg will be leffer. p. 126. l. 23. r.  $\Delta i$ . l. penult. r.  $\Delta i$  quared, fo is the Quadruple of the. p. 127. l. 30. del. of. p. 128. l. 2. r. to deuble. 1. 28, 29. r. sg, the longest horizontal Range is the half. **B**: 129. penult. r. or 90 + 49. p. 131. b. 31. r. m. P. l. penult. r. in n. p. 134. l. 16. r. Fig. 6. p. 140. l. 13. r. 1. l. 20, 21. r. the Dimidiate or Subduplicate of 81 to 9. i.e. that of 81 to 27. p. 144. l. 21. r. along Bc. p. 150. l: 8. r. nearer. p. 152. l. 20. T. a. p. 152. l. 12. r. Point D. p. 161. l. 21. r. Pf. p. 166. 1. 3. r. Fig. 2. 1. 12. r. be taken as the Diftance, as. p. 168. 1. 10. T. Pq PF. 1. 20. r. as also do p q and x i. p. 169. l. 11, 12. r. at хi the fame Distance. 1. 15. r. the. p. 171. 1. 9 r. will. p. 175. 1. 6. r. a. l. penult. r. A B'd." p. 176. l. 14. r. 4. l. 17. r. AD. l. 24 T. 18. r. 50. l. 26. r. 1130. or. p. 177. l. penult. r. 1 × 10. rp. 178. l. 5. r. B.D. p. 180. l. 2. r. ARPB. p. 181. l. 16. r. ARPV. p. 185. l. 17. r. 2 19. l. 25. r. Fh. l. antepen. r. sq. p. 186. l. 3. r. | 57. p. 187. l. 1. r. h. l. 21 r. 4672 85. l. 22. 25. r. Fh. 1. 32 del. T. p. 188. l. 5. r. the Number 2. p. 192. 1. 29. r. Coroll. 1. p. 193. l. 2. del. and. p. 196. l. 34. del. which. p. 197. l. 25. r. a. p. 200,204, Orc. r. Syzygics. p. 210. I. 28, 29. r. nine. p. 219. l. 12. r. Attraction. p. 237.1. 22: r. pag. 1.34. r. 31. p. 241. l. 19. r. HI, KL. p. 256. 1. 31. r. + 1. 32. del. PS. p. 260. l. 21. r. Decemb. 3. 1705. 1. ult. r. on the fame Side of the Plane. p. 262. Sepe for + r. x. 1. 29. r. that of the. p. 265. l. penult. r. Angle of Incidence. p. 269 l. i 1. del, B. l. 13. Marg. del. Lat. p. 270. l. 11. del. O. l. 31. 82 p. 271. l. 7, 10, 11, 12. instead of Glass r. Spectrum. p. 272. I. ult. r. For. p. 279-1. 22. add and I X NF will be equal to 3 R X N P. or I: 3 R :: N P : N F. p. 280. 1. 4. r. as well as. 1. 8. r. the doubled arc Ff. p. 284. 1. 34. r. internal. p. 286. . I. 18. r. a. ) p. 303. l. amepenuk. r. the lame in. p. 305. l. ulr. r. p. 106. p. 321. l. 22. del. to. p. 345. r. 227500. 220. 94. ·p. 347. t. 123. 100. 25 . 19. 15. p. 365. l. 1. del. not. l. 31, r. two or three Hours. p. 372. l. 4. r. Now. p. 375. l. 4. r. Scarce 10. p. 383. 1. 24. T. Q C. ..





